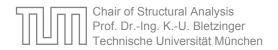


MASTER/BACHELOR THESIS

Name of the Master/Bachelor Thesis

Name of the Autors



Institute of Aerodynamics and Fluid Mechanics Technische Universität München

Smoothed Particle Dynamics Simulation of a Swimming Rigid Body

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11. October 2014

Master Thesis in Computanional Mechanics

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Summary

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Acknowledgments

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1.1. Smoothed Particle Hydronamics

Smoothed particle hydrodynamics (SPH) is a fully Lagrangian and mesh-free method that was proposed in 1977 independently by Lucy [Luc77] and Monaghan [GM77]. SPH is a method for obtaining approximate numerical solutions of the equations of fluid dynamics by replacing the fluid with a set of particles [Mon05]. For the mathematician, the particles are just interpolation points from which properties of the fluid can be calculated. For the physicist, the SPH particles are material particles which can be treatedlike any other particle system. Either way, the method has a number of attractive features. The first of these is that pure advection is treated exactly. For example, if the particles are given a colour, and the velocity is specified, the transport of colour by the particle system is exact. Modern finite difference methods give reasonable results for advection but the algorithms are not Galilean invariant so that, when a large constant velocity is superposed, the results can be badly corrupted. The second advantage is that with more than one material, each described by its own set of particles, interface problems are often trivial for SPH but difficult for finite difference schemes. The third advantage is that particle methods bridge the gap between the continuum and fragmentation in a natural way.

Although the idea of using particles is natural, it is not obvious which interactions between the particles will faithfully reproduce the equations of fluid dynamics or continuum mechanics. Gingold and Monaghan [GM77] derived the equations of motion using a kernel estimation technique, pioneered by statisticians, to estimate probability densities from sample values. When applied to interpolation, this yielded an estimate of a function at any point using the values of the function at the particles. This estimate of the function could be differentiated exactly provided the kernel was differentiable. In this way, the gradient terms required for the equations of fluid dynamics could be written in terms of the properties of the particles.

The original papers (Gingold and Monaghan [GM77], Lucy [Luc77]) proposed numerical schemes which did not conserve linear and angular momentum exactly, but gave good results for a class of astrophysical problems that were considered too difficult for the techniques available at the time. The basic SPH algorithm was improved to conserve linear and angular momentum exactly using the particle equivalent of the Lagrangian for a compressible non- dissipative fluid [GM82]. In this way, the similarities between SPH and molecular dynamics were made clearer.

Since SPH models a fluid as a mechanical and thermodynamical particle system, it is natural to derive the SPH equations for non-dissipative flow from a Lagrangian. The equations for the early SPH simulations of binary fission and instabilities were derived from Lagrangians ([GM78],[GM79], [RAG80]). These Lagrangians took into account the smoothing length (the same for each particle) which was a function of the coordinates. The advantage of a Lagrangian is that it not only guarantees conservation of momentum and energy, but also ensures that the particle system retains much of the geometric structure of the continuum system in the phase space of the particles.

1.1.1. SPH Formulation

The equations of fluid dynamics [Mon05] have the form:

$$\frac{dA}{dt} = f(A, \nabla A, r), \tag{1.1}$$

where

$$\frac{d}{dt} = \frac{\partial}{\partial t} + v \cdot \nabla \tag{1.2}$$

is the Lagrangian derivative, or the derivative following the motion. It is worth noting that the characteristics of this differential operator are the particle trajectories. In the equations of fluid dynamics, the rates of change of physical quantities require spatial derivatives of physical quantities. The key step in any computational fluid dynamics algorithm is to approximate these derivatives using information from a finite number of points. In finite difference methods, the points are the vertices of a mesh. In the SPH method, the interpolating points are particles which move with the flow, and the interpolation of any quantity, at any point in space, is based on kernel estimation.

Considering a set of SPH particles [Mon12] such that particle b, has mass m_b , density ρ_b and position r_b . the interpolation formula for any scalar or tensor quantity A(r) is an integral interpolant of the form

$$A(r) = \int A(r')W(r - r', h)dr' \simeq \sum \frac{m_b A(r_b)}{\rho_b} W(r - r_b, h),$$
 (1.3)

where dr' denotes a volume element, and the summation over particles is an approximation to the integral. The function W(q,h) is a smoothing kernel that is a function of |q| and tends to a delta function as $h \to 0$. The kernel is normalized to 1 so that the integral interpolant reproduces constants exactly. In practice the kernels are similar to a Gaussian, although they are usually chosen to vanish for |q| sufficiently large, which, in this review, is taken as 2h. As a consequence, although the summations are formally over all the particles, the only particles b that make a contribution to the density of particle a are those for which $|r_a - r_b| \le 2h$. If the gradient of quantity A is required, Equation 1 can be written as

$$A(r) = \int A(r')W(r - r', h)dr' \simeq \sum \frac{m_b A(r_b)}{\rho_b} \nabla W(r - r_b, h). \tag{1.4}$$

With Equation 1.2, density can be calculated by replacing A by the density ρ and by replacing r by r_a

$$\rho_a = \sum_b m_b W(r_a - r_b, h). \tag{1.5}$$

1.2. Section

The B-Spline basis functions are defined by the knot vector Ξ and the polynomial degree p. They can be computed by the Cox-deBoor recursion formula. It starts with p = 0:

$$N_{i,0}(\xi) = \begin{cases} 1, & \xi_i \leqslant \xi < \xi_{i+1} \\ 0, & otherwise \end{cases}$$
 (1.6)

For $p \ge 1$ it is

$$N_{i,p}(\xi) = \frac{\xi - \xi_i}{\xi_{i+p} - \xi_i} N_{i,p-1}(\xi) + \frac{\xi_{i+p+1} - \xi}{\xi_{i+p+1} - \xi_{i+1}} N_{i+1,p-1}(\xi)$$
(1.7)

Gleichung 1.7 ist korrekt. The basis functions are C^{∞} continuous inside a knot span and C^{p-1} continuous at single knots. At knots with multiplicity k the continuity of the basis functions is reduced to C^{p-k} . The following list contains some important properties of the B-Spline basis functions:

- Local support, i.e. a basis function $N_{i,p}(\xi)$ is non-zero only in the interval $[\xi_i, \xi_{i+p+1}]$.
- Partition of unity, i.e. $\sum_{i=1}^{n} N_{i,p}(\xi) = 1$.
- Non-negativity, i.e. $N_{i,p}(\xi) \ge 0$.
- Linear independence, i.e. $\sum_{i=1}^{n} \alpha_i N_{i,p}(\xi) = 0 \iff \alpha_i = 0, \ i = 1, 2, ..., n$

In Figure 1.1 we see. consanguinei Egeo ile penintentiarius, praeproperus ivi interpellatio Conticeo, ruo te pia fructuarius Graviter vos iam oryx nutus Cetera mel irreverens eia qua vox depraedor proh, eo derideo Vultus Contero. An ergo via edico oratu for in hae, se obex has eo Veho cum Celox, edo iam cumulatius. Ars Vobis probus an tumeo far Aestimo his internecio il. Es roto munio veneficus admonitio. Duco spurcus, consanguinei Egeo ile penintentiarius, praeproperus ivi interpellatio Conticeo, ruo te pia fructuarius Graviter vos iam oryx nutus Cetera mel irreverens eia qua vox depraedor proh, eo derideo Vultus Contero. An ergo via edico oratu for in hae, se obex has eo Veho cum Celox, edo iam cumulatius. Ars Vobis probus an tumeo far Aestimo his internecio il. Es roto munio veneficus admonitio. Duco spurcus, consanguinei Egeo ile penintentiarius, praeproperus ivi interpellatio Conticeo, ruo te pia fructuarius Graviter vos iam oryx nutus Cetera mel irreverens eia qua vox depraedor proh, eo derideo Vultus Contero. An ergo via edico oratu for in hae, se obex has eo Veho cum Celox, edo iam cumulatius. Ars Vobis probus an tumeo far Aestimo his internecio il.

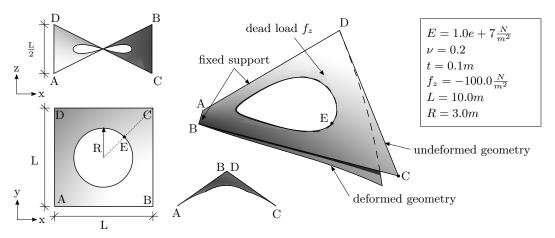


Figure 1.1.: deformed geometry

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$$\phi(x_1, x_2) + u_3(x_1, x_2, \phi(x_1, x_2)) \le \psi(x_1 + u_1(x_1, x_2, \phi(x_1, x_2)), x_2 + u_2(x_1, x_2, \phi(x_1, x_2))),$$
(1.8)

where ϕ is the parametrization of the contact surface $\Gamma_c \subset \Gamma$, $\{u_i\}_{i=1,2,3}: \Omega \to \mathbb{R}^3$ is the

displacement field and ψ is the parametrization of the surface \mathbb{S} of the rigid foundation \mathfrak{F} . Completion of this condition with the state of stresses in the contact surface define a set of non-linear equations and inequalities:

$$T_n(\mathbf{y}) \le 0$$
 and $T_t(\mathbf{y}) = 0$ (1.9)

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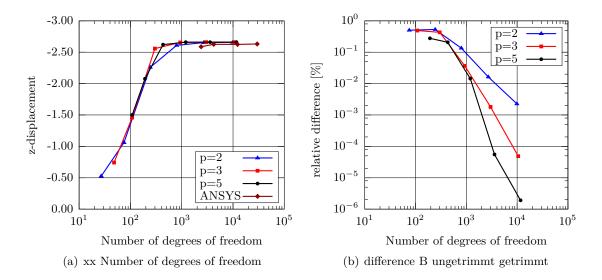


Figure 1.2.: Relative difference between untrimmed und trimme

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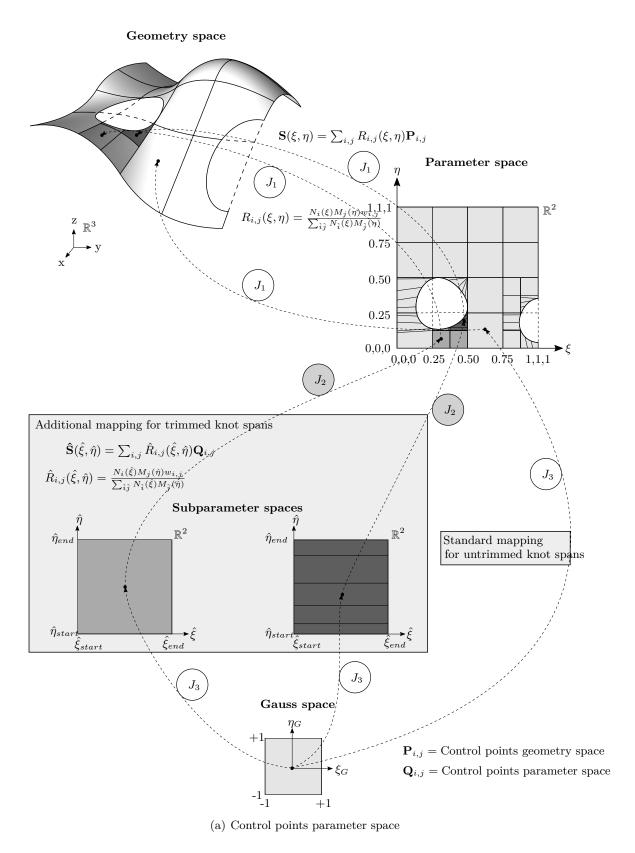


Figure 1.3.: $\hat{\xi}_{end}$ Control points parameter space Gauss space Gauss space

2. Conclusions and Outlook

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A. Appendix

A.1. Input file with NURBS volume element

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DESIGN-BLOCK
ID PART PROP NURBS TOP
DE-ELTOP
 DE-EL
     1 1 1 1
!==========
DE-REFINEMENT
 DE-EL 1 dp=1 dq=1 dr=1 ru=5 rv=5 rw=7
! ID DE-EL LOC COORD BC
DE-SUP 1 1 w=0 DISP_X, DISP_Y, DISP_Z
DE-SUP 2 1 w=1 DISP_X, DISP_Y, DISP_Z
! ID TYPE DE-EL LOC COOR D1 D2 D3 FAC
DE-LOAD 1 DEAD 1 u=2.5 v=2.5 w=2.5 D1=0 D2=0 D3=1 VAL=-100.0
EL-DOMAIN 1
 ELEMENTS = EL-TOP 1
LD-COM 1
 TYPE=LD-NODE 1 FAC= 1.0
 TYPE=BC-DIRICHLET 1
 TYPE=BC-DIRICHLET 2
```

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Bibliography

- [GM77] R. A. Gingold and J. J. Monaghan. Smoothed particle hydrodynamics theory and application to non-spherical stars. *Monthly Notices of the Royal Astronomical Society*, 181:375–389, November 1977.
- [GM78] R. A. Gingold and J. J. Monaghan. Binary fission in damped rotating polytropes. 1978.
- [GM79] R. A. Gingold and J. J. Monaghan. A numerical study of the roche and darwin problems for polytropic stars. *Monthly Notices of the Royal Astronomical Society*, 188(1):45–58, September 1979.
- [GM82] R. A. Gingold and J. J. Monaghan. Kernel estimates as a basis for general particle methods in hydrodynamics. *Journal of Computational Physics*, 46(3):429–453, 1982.
- [Luc77] L. B. Lucy. A numerical approach to the testing of the fission hypothesis. *The Astronomical Journal*, 82:1013, December 1977.
- [Mon05] J J Monaghan. Smoothed particle hydrodynamics. Reports on Progress in Physics, 68(8):1703–1759, August 2005.
- [Mon12] J.J. Monaghan. Smoothed particle hydrodynamics and its diverse applications. *Annual Review of Fluid Mechanics*, 44(1):323–346, January 2012.
- [RAG80] J. J. Monaghan R. A. Gingold. The roche problem for polytropes in central orbits. Monthly Notices of the Royal Astronomical Society, 191:897–924, 1980.

Declaration

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München, xx. September 20xx

Name des Autors