

Laboration 2: RGBD-cameras

Sensors and Sensing

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*All code for this exercise can be found at
<https://github.com/tgolsson/sensors-laboration2-xtion>*

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1 Theory and motivation

1.1 RGBD-cameras

RGBD-cameras, short for *Red-Green-Blue-Depth*-camera, is a type of low-cost camera commonly used for robot vision. The concept became widely popular with the release of the Microsoft Kinect in late 2010.

These cameras consist of two separate parts: one normal color-based camera, and one infrared sensor with accompanying projector. The sensing consists of projecting a deterministic pattern onto the scene using an infrared emitter, and then unprojecting by comparing the image to previously captured patterns at known depths. By interpolating through these patterns, a full depth-image is generated.

1.2 Noise

A common problem in any type of sensing is the introduction of noise into the system. This noise can come from many sources, and be predictable or unpredictable. Examples of noise sources could be frequency hum from electric circuits, flickering lights, air pollution or pure inaccuracy. This noise can skew the results of sensors that make algorithm much more error prone.

There are many approaches to reduce noise. Proper calibration and good testing environments is a good start, but this can only reduce external noise. Internal noise of the sensor needs to be analyzed and minimized on a much lower-level such as by using specially constructed algorithms. For sensors that generate some sort of sequence, one very naive (but nonetheless effective) approach is the use of smoothing.

2 Implementation

The purpose of this exercise is to calibrate an RGBD-camera and investigate its characteristics. Then, several smoothing algorithms shall be evaluated to reduce noise in the depth sensor.

2.1 Hardware and environment

This exercise was performed using an *ASUS Xtion Pro*. The camera was connected over *USB2* to a laptop running Linux kernel 4.2.5. The communication to the camera is done using the *Robot Operating System* [ROS] version *Indigo Igloo*. All ROS packages used are compiled directly from GitHub development branch for Indigo Igloo.

Other software used includes the OpenCV libraries, version 2.4.12.2-1.

2.2 Camera setup

2.3 ROS setup

Instead of using RVIZ to view the data as before, a custom ROS-node can be used. As there are three types of data - color image, depth image, and pointcloud, there are three listeners setup to receive this data. An example of the data on these topics can be found in the embedded files below¹. The `depth_image` and `color_image` are stored in the OpenCV `y[a]ml` format, and the point cloud is stored in the *Point Cloud Library* [PCL] `pcd` format.

 Pointcloud file  RGB-image file

2.4 Camera calibration

```
1 image_width: 640
2 image_height: 480
3 camera_name: rgb_PS1080_PrimeSense
4 camera_matrix:
5   rows: 3
```

¹If these embedded files do not work, such as if you use Adobe Acrobat, please download them from the `report` folder in the GitHub repository. Verified to work with Okular and Evince on Linux.

```

6   cols: 3
7   data: [546.589906, 0, 319.850785, 0, 545.054549, 240.484512, 0, 0, 1]
8   distortion_model: plumb_bob
9   distortion_coefficients:
10  rows: 1
11  cols: 5
12  data: [0.073113, -0.09862, 0.002178, 0.001978, 0]
13  rectification_matrix:
14  rows: 3
15  cols: 3
16  data: [1, 0, 0, 0, 1, 0, 0, 0, 1]
17  projection_matrix:
18  rows: 3
19  cols: 4
20  data: [546.70282, 0, 320.499416, 0, 0, 547.7890619999999, 240.692826,
        0, 0, 0, 1, 0]

```

2.5 Noise characterization

The sensors used in the camera suffer from noise, and the first requirement for improvement is to objectively quantify the error. First, 5 measurements of 10 seconds each were recorded with *rosvbag*. The measurements were taken *50, 100, 150, 200*, and *300* cm from a white wall. The camera was placed on a table, so that the camera normal was perpendicular to the wall. This ensures that all measurements are made against the same baseline.

First, the errors were quantified at varying distance with set window sizes. Fig. 1 shows the error and variances for a 20x20 window, and fig. 2 on the next page shows the same for a 40x40 window. The most obvious pattern can be seen in the left-most plot in the two figures: there is a very strong correlation between distance and both variance and error. This has been shown by other researchers, and is logical: the further the distance, the more things can interfere with measurements such as light conditions, dust particles, and so on.

Another interesting pattern is that the mean error barely changes between the two windows; however, a larger window reduces the variance significantly. This is not so much related to reduced errors, but a result of the definition of a variance. If the errors are assumed to adhere to some sort of distribution $X \in (-\infty, \infty)$, then $\sum_{X_1}^X = \bar{X}$, i.e., the static error. From this it follows that the variance will then also be the pure noise in the signal. However, if too few points are sampled the sum will only approximate the mean, and the variance will become unstable.

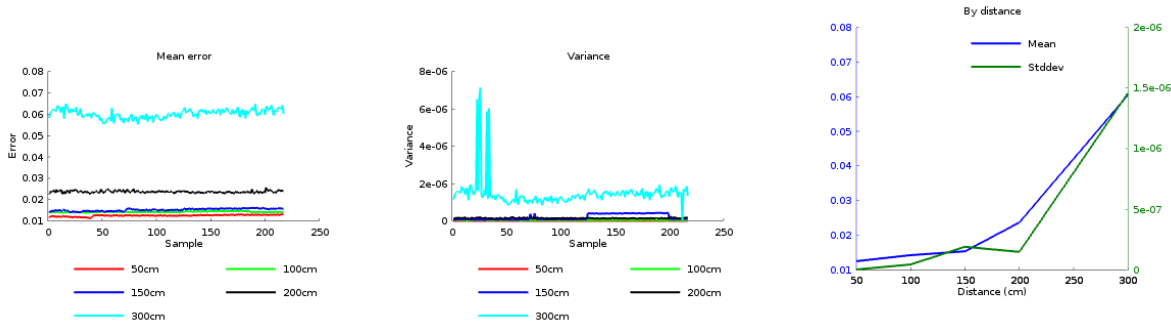


Figure 1: The error and variance for a 20x20 pixel window at the center.

2.6 Noise filtering

A large issue when implementing the filters was handling the large number of invalid values in the input. In some situations, a majority of the elements were NaN, which obviously makes it hard to draw any conclusions from the data. The *gaussian* and *median* filters from **OpenCV**

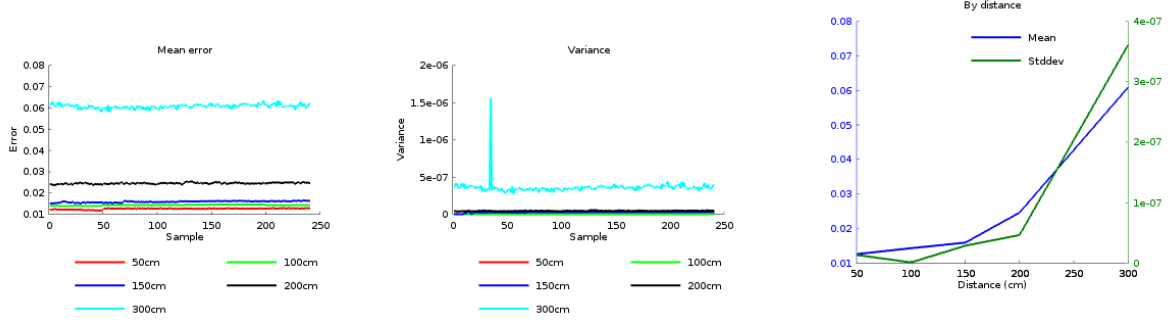


Figure 2: The error and variance for a 40x40 pixel window at the center.

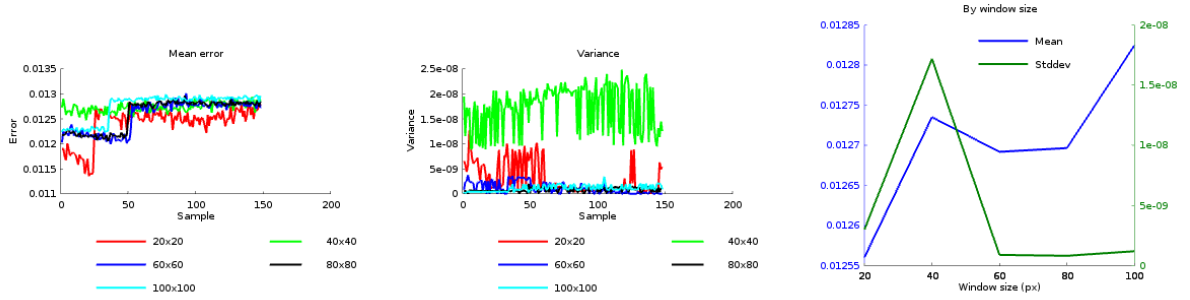


Figure 3: The error and variance at 50cm for various window sizes.

work fine without any modification. The bilateral filter however suffers a segmentation fault when there are NaN values in the input.

In order to solve this issue, all NaN-values in the input are set to 10000 before being passed to the bilateral filter. By the edge preserving property of bilateral filtering, this should reduce the impact they have on the actual filtering process. After the filtering is done, every cell which was set to 10 000 is set to 0; and the mean and variance is calculated on all non-zero elements.

Another solution we considered - but did not implement - was to use local median-filtering at each NaN. We deemed this to be a very computationally expensive process, but could possibly yield better results.

3 Results

References