

EFFECTS OF WAVE PACKET PROFILES ON NEUTRINO OSCILLATIONS

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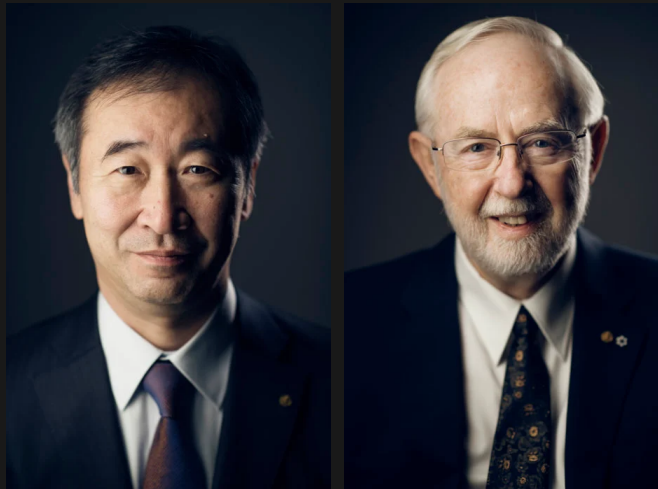
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Takaaki Kajita & Arthur B. McDonald, Nobel Prize in Physics 2015
“for the discovery of neutrino oscillations, which shows that neutrinos have mass”

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where $T = t_d - t_p$ and $L = x_d - x_p$.

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Transition probability found by

$$P_{\nu_\alpha \rightarrow \nu_\beta}(L) = \int dT \left| \mathcal{A}_{\nu_\alpha \rightarrow \nu_\beta}(T, L) \right|^2 .$$

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Kayser [4] first to study oscillations with wave packets, and Giunti [5] first to obtain explicit results with Gaussians

[3] S. Nussinov, Phys. Lett. B **63**, 201 (1976).

[4] B. Kayser, Phys. Rev. D **24**, 110 (1981).

[5] C. Giunti, C. W. Kim, and U. W. Lee, Phys. Rev. D **44**, 3635 (1991).

Why a Gaussian?

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“A gaussian momentum distribution is the most convenient one for the calculation of several integrations ...

Other distributions which are sharply peaked around an average momentum lead to the same results after their approximation with a gaussian ...

Therefore, the gaussian momentum distributions can be taken as approximations of the real momentum distributions from the beginning.” [6]

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Expanding about minimum $p = P$, we require

$$\frac{1}{4!} \left| g^{(iv)}(P) \right| \ll \frac{1}{2} \left| g''(P) \right|^2$$

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Lorentzian wave packets

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Neutrinos in the context of the Mössbauer effect described by a Lorentzian wave packet [8, 9]

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$$\tilde{\psi}(p; \bar{p}, \gamma) = \mathcal{N} \left[\frac{\gamma}{(p - \bar{p})^2 + \gamma^2} \right]$$

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Moments undefined. Cannot be approximated by a Gaussian!

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$$\tilde{\psi}(p_\mu; a_\mu) = \mathcal{N} \exp \left[- a_\mu p^\mu \right],$$

where $a_\mu = (\alpha, -\beta) \in \mathbb{C}^2$ and transforms as a vector.

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Gaussian minimises Heisenberg-Robertson uncertainty relation

$$\sigma_x \sigma_p = \left| \frac{1}{2i} \langle [\hat{x}, \hat{p}] \rangle \right| = \frac{\hbar}{2} .$$

Assuming that position and momentum are independent

Robertson-Schrödinger uncertainty relation

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$$\psi(p; \bar{x}, \bar{p}, \sigma_p) = \mathcal{N} \exp \left[-\frac{(p - \bar{p})^2}{4\sigma_p^2} + i\bar{x}p \right]$$

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$$\tilde{\psi}(E_p, p; \alpha, \beta) = \mathcal{N} \exp[-\alpha E_p + \beta p] ,$$

where α, β determined by the moments of velocity and space(time)

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RMU:

$$\mathcal{A}_{\nu_\alpha \rightarrow \nu_\beta}(T, L) \sim \sum_i U_{\alpha i}^* \frac{(T - i\alpha)m_i K_1 \left(-m_i \sqrt{(T - i\alpha)^2 + (L - i\beta)^2} \right)}{\sqrt{(T - i\alpha)^2 + (L - i\beta)^2}} U_{\beta i}$$

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2. RMU wave packets follow semi-classical trajectories, so different mass eigenstates do not separate!

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Controversy in early 2000s whether group velocities can be equal [12-14]

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3. $E_1/E_2 \simeq 1$ in ultra-relativistic regime, not generally true for m_1/m_2

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Only agrees in semi-classical regime, for $\sigma_x \gg \lambda_c \equiv 1/m$

Summary

- Neutrino wave packets could have a non-Gaussian profile
- If described by RMU wave packets, then neutrinos are highly localised, and decoherence is heavily suppressed
- Propagation at equal velocities should be taken seriously, and could be experimentally tested

Evan Gale,
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