# CSE305 Concurrent Programming: N-Body simulation project

Martin Lau, Oscar Peyron, Ziyue Qiu

May 31, 2025

## 1 Introduction

This document outlines the N-Body simulation project for CSE305, which includes how we approached the problem, our project structure, code breakdown and strategies, and encountered difficulties.

To get started quickly, try typing make nbody in the project directory, and generate a basic gif file. For information on how to execute the code, check our README file.

For this project, our aim is to be able to simulate systems of bodies with forces interacting with one another in 2D (such as orbiting planets around the solar system with gravitational forces, or with particles interacting with each other with Coulomb forces). This therefore includes two sections:

- 1. Simulation of the N bodies and their evolving states.
- 2. Visualization of recorded telemetries into an animated output.

Below is a general view of our project repository file structure.

The work has been split as follows. The arrangement is not defining, as we like to help each other out in different parts. Our files have still been mostly separated to properly separate who worked on what functionalities.

- 1. Martin handles general project structure, core classes, visualization, and the naive and its optimized algorithm implementation.
- 2. Zivue handles the Barnes Hutt algorithm.
- 3. Oscar handles the Particle Mesh algorithm implementation (simple, thread based and cuda implementation)

# 2 Core components

#### 2.1 Main elements

There are 3 primary classes defined used throughout our project: Vector, Body, System. These are defined in core.cpp/hpp.

- 1. Vector holds information on a pair of numbers, and also allows for operations with other vectors/scalars (as opposed to using std::pair).
- Body holds information on a given particle or body, including its mass, coordinates, velocity, and acceleration. It also contains an update method which updates its position and velocity based on acceleration.
- 3. System stores a collection of bodies and its recorded telemetry from the simulations we are going to do. It also contains the visualization function, which takes its recorded telemetry and outputs an animated file.

Note: This organization is heavily inspired from one assignment from CSE306 Computer Graphics

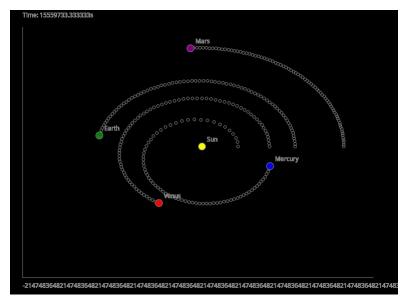
Elaborating more on the telemetry stored within the System class, this is stored as a vector of vector of Vector. What a mouthful! Our simulations creates different steps/frames. Each step/frame contains the positions of all bodies inside the system (this is a vector of Vector). To have the entire telemetry, we have a vector of frames, or the aforementioned data structure for the telemetry.

## 2.2 Visualization code

To visualize the code, we opt to create a gif animation after the telemetry is recorded, using the ImageMagick/Magick++ libraries. The bulk of the visualization code is located in core.cpp as a method for the System class. As of present, the visualization function works in two steps. First, it creates the frames for the animation, then writes the frames to a gif file. The frame creation is relatively quick, and most of the visualization time is actually spent in one line (writeImages(frames.begin(), frames.end(), name)).

We will look into further solutions to try and decrease the time taken to create the visualization gifs, like reducing image quality, or the number of frames.

For testing purposes, there is also a visualizer in visualizer.py which creates a animation using Python's matplotlib library, working significantly faster, allowing us to test the correctness of our telemetries.



# 3 Naive algorithm

So far, we have a basic implementation of the naive algorithm without multi threading (most of my time was taken bugfixing core functionalities and getting visualization to work).

#### Algorithm 1 Naive simulation outline

```
Require: System of bodies with masses m_i, initial positions \vec{r}_i, velocities \vec{v}_i
Require: Time step \Delta t, number of steps N
   telemetry \leftarrow \emptyset
   telemetry.append(initial positions)
   for step \leftarrow 0 to N-1 do
        for each body i do
            \vec{a}_i \leftarrow \vec{0}
                                                                                                              ▶ Reset accelerations
       end for
       for i \leftarrow 0 to n-1 do
            for j \leftarrow i+1 to n-1 do
                 F_{ij} \leftarrow \text{ComputeGravitationalForce}(body_i, body_j)
                 \vec{a}_i \leftarrow \vec{a}_i + \vec{F}_{ij}/m_i
                 \vec{a}_j \leftarrow \vec{a}_j - \vec{F}_{ij}/m_j
            end for
        end for
        for each body i do
            \vec{v}_i \leftarrow \vec{v}_i + \vec{a}_i \Delta t
                                                                                                                  ▶ Update velocity
            \vec{r}_i \leftarrow \vec{r}_i + \vec{v}_i \Delta t
                                                                                                                  ▶ Update position
        end for
        telemetry.append(current positions)
   end for
Ensure: Position history for all bodies stored in telemetry
```

Other aspects (parallelizing the update steps, parallelizing forces computations, avoiding race conditions) will be implemented later on.

# 4 Particle Mesh based algorithm

I have implemented a sequential Particle mesh algorithm, as well as a thread based.

Here below is the pseudo-code for the sequential implementation as well as for the thread-based implementation

## Algorithm 2 Particle-Mesh Simulation, using Nearest-Grid-Point (NGP)

```
Require: System universe, time step \Delta t, grid size N, spatial extent R
 1: telemetry \leftarrow \emptyset
 2: telemetry.append(initial positions)
 3: boundaries \leftarrow [-R, R]
 4: Compute cell size h \leftarrow \frac{2R}{N}
 5: Initialize mass density grid M[G][G] \leftarrow 0
 6: Initialize potential grid \Phi[G][G] \leftarrow 0
 7: Initialize FFTW input/output arrays and plans
    for each time step s = 1 to S do
         Clear mass density grid M
 9:
         for each body b in universe do
10:
             i \leftarrow \left\lfloor \frac{x_b - \min_x}{h} \right\rfloor
11:
             j \leftarrow \left| \frac{y_b - \min_y}{h} \right|
12:
13:
              if (i,j) in bounds then
                  M[i][j] \leftarrow M[i][j] + \text{body.mass}
14:
              end if
15:
         end for
16:
         Copy grid mass to FFTW input array
17:
         Compute FFT of mass density using forward FFT
18:
         for each (i, j) in frequency domain do
19:
              Compute wave numbers (k_x, k_y)
20:
              Compute k^2 \leftarrow k_x^2 + k_y^2
21:
              if k^2 > 0 then
22:
                  Multiply by -\frac{G}{k^2} in frequency domain
23:
24:
              else
                  Set value to zero
25:
              end if
26:
         end for
27:
         Compute inverse FFT to obtain gravitational potential
28:
         Normalize inverse FFT result and store in potential grid \Phi
29:
30:
         for each body b in universe do
             i \leftarrow \left\lfloor \frac{x_b - \min_x}{h} \right\rfloor
31:
             j \leftarrow \left[ \frac{y_b - \min_y}{h} \right]
32:
              if (i, j) is valid and not at boundary then
33:
                  Compute force via central difference of potential:
34:
                  \vec{a}_i \leftarrow -\left(\frac{\Phi[i+1][j] - \Phi[i-1][j]}{2h}, \frac{\Phi[i][j+1] - \Phi[i][j-1]}{2h}\right)
35:
              else
36:
                  \vec{a}_i \leftarrow (0,0)
37:
              end if
38:
         end for
39:
         for each body i do
40:
              \vec{v}_i \leftarrow \vec{v}_i + \vec{a}_i \Delta t
                                                                                                              ▶ Update velocity
41:
              \vec{r_i} \leftarrow \vec{r_i} + \vec{v_i} \Delta t
                                                                                                              ▶ Update position
42:
43:
         end for
         telemetry.append(current positions)
45: end for
46: Cleanup FFTW plans and memory
```

## Algorithm 3 Parallel Particle-Mesh Simulation using Nearest-Grid-Point (NGP)

```
Require: System universe, time step \Delta t, grid size N, spatial extent R, number of threads T
 1: telemetry \leftarrow \emptyset
 2: telemetry.append(initial positions)
 3: boundaries \leftarrow [-R, R]
 4: Compute cell size h \leftarrow \frac{2R}{N}
 5: Initialize mass density grid M[N][N] \leftarrow 0
 6: Initialize potential grid \Phi[N][N] \leftarrow 0
 7: Initialize FFTW input/output arrays and plans
    for each time step s = 1 to S do
         Clear mass density grid M
 9:
         Parallel for each thread t = 1 to T
10:
            Assign a chunk of bodies to thread t
11:
         for each body b assigned to thread t do
12:
             i \leftarrow \left\lfloor \frac{x_b - \min_x}{h} \right\rfloor
13:
             j \leftarrow \left| \frac{y_b - \min_y}{h} \right|
14:
             if (i, j) in bounds then
15:
16:
                  Lock M[i][j]
                  M[i][j] \leftarrow M[i][j] + \text{body.mass}
17:
                  Unlock M[i][j]
18:
             end if
19:
         end for
20:
         End parallel for
21:
22:
         Copy mass grid to FFTW input array
         Compute FFT of mass density using forward FFT
23:
24:
         for each (i, j) in frequency domain do
             Compute wave numbers (k_x, k_y)
25:
             k^2 \leftarrow k_x^2 + k_y^2
26:
             if k^2 > 0 then
27:
                 Multiply by -\frac{G}{k^2} in frequency domain
28:
             else
29:
30:
                  Set value to zero
             end if
31:
         end for
32:
         Compute inverse FFT to obtain gravitational potential
33:
         Normalize result and store in potential grid \Phi
34:
         Parallel for each thread t = 1 to T
35:
36:
            Assign a chunk of bodies to thread t
         for each body b assigned to thread t do
37:
             i \leftarrow \left\lfloor \frac{x_b - \min_x}{h} \right\rfloor
38:
             j \leftarrow \left| \frac{y_b - \min_y}{\iota} \right|
39:
40:
             if (i,j) valid and not at boundary then
                 Compute force via central difference:
41:
                 \vec{a}_b \leftarrow -\left(\frac{\Phi[i+1][j] - \Phi[i-1][j]}{2h}, \frac{\Phi[i][j+1] - \Phi[i][j-1]}{2h}\right)
42:
             else
43:
                  \vec{a}_b \leftarrow (0,0)
44:
             end if
45:
         end for
46:
         End parallel for
47:
         Parallel for each thread t = 1 to T
48:
            Assign a chunk of bodies t = 1 to T
49:
         for each body b assigned to thread t do
50:
             \vec{v}_b \leftarrow \vec{v}_b + \vec{a}_b \Delta t
51:
             \vec{r}_b \leftarrow \vec{r}_b + \vec{v}_b \Delta t
52:
53:
         end for
         End parallel for
54:
         telemetry.append(current positions)
                                                               5
56: end for
57: Cleanup FFTW plans and memory
```

# 4.1 Time performance of the particle-mesh simulation (sequential and parallel

The simulation where done on MacBook air with M2 chip containing 8 cores. All simulations where done with grid size = 10, a spatial extent R of 10000 and time increment dt = 0.2. Additionally, each simulation includes asteroids with high radius for their orbit (between 5 and 3000 as radius) and the rest being small asteroids with smaller radius for their orbit (between 0.5 and 5 as radius). The reason why I included big differences in radius is to allow between uniformity in the space domain.

Parallel computation was done through parallelizing via the bodies. Each body's mass was assigned in the grid via multiple threads and the computation of their acceleration was also done through parallelization.

The fast fourier transforms as well as the computation of the potential was not done in parallel since, assigning a thread would have been very inefficient if multiple bodies where assigned on the same grid. Indeed, locking the cells of each grid to avoid race conditions would have given a performance similar to a sequential one.

However, one can still implement fast-fourier transform in parallel within the library fftw3. Unfortunately, I was not able to make it work on my machine.

#### Time performance of the sequential particle-mesh simulation

Number of Bodies	Execution Time (ms)		
4	52		
5	54		
50	109		
100	157		
1,000	941		
2,000	3,013		
5,000	7331		
10,000	14 053		

Table 1: Execution time of the particle-mesh simulation for varying numbers of bodies (grid size = 10)

#### Time performance of the parallel particle-mesh simulation

Number of Bodies	Threads $= 5$	Threads $= 7$	Threads $= 10$
4	1661	_	_
5	2209	2848	4058
50	2295	2906	3263
100	2298	3034	3165
1000	3436	3800	5006
2000	5075	5865	6595
5000	7576	6244	6757
10000	18230	12306	13881

Table 2: Execution time (in milliseconds) for varying numbers of bodies and thread counts (grid size = 10).

Looking at performance we see that until 2000 bodies, sequential time-performance is better than for the parallel implementation. However, for higher number of bodies, the parallel implementation with 7 threads displays better performance. 7 threads is particularly optimal as it is very close to the number of cores in the machine on which we simulate (8 cores), allowing optimal use of the CPU.

# 5 Barnes–Hut Algorithm

The Barnes–Hut algorithm reduces the complexity of the classical  $O(N^2)$  N-body force computation to approximately  $O(N \log N)$  by hierarchically clustering distant bodies. Our implementation follows these main stages:

#### Algorithm 4 Barnes–Hut Simulation Outline

```
Require: System of bodies with masses m_i, positions \vec{r}_i, velocities \vec{v}_i
Require: Time step \Delta t, number of steps N, opening angle \theta
 1: telemetry \leftarrow \emptyset
 2: telemetry.append(initial positions)
 3: for step \leftarrow 0 to N-1 do
        [minB, maxB] \leftarrow computeBounds(universe)
 4:
        root \leftarrow createRootNode([minB, maxB])
 5:
        for each body b_i do
 6:
 7:
            insertBody(root, b_i)
        end for
 8:
        computeMassDistribution(root)
 9:
        for each body b_i do
10:
            \vec{a}_i \leftarrow \mathtt{forceOnBody}(b_i,\mathtt{root},\theta)/m_i
11:
12:
        end for
        updateBodies(universe, \Delta t)
13:
        freeQuadTree(root)
14:
        telemetry.append(current positions)
15.
Ensure: Position history for all bodies stored in telemetry
```

- 1. Compute Simulation Bounds. We first call
  - computeBounds(const System&) to find an axis-aligned square enclosing all bodies (with a small padding).
- 2. Quadtree Construction. We represent space by a pointer-based quadtree of QuadNode objects:
  - createRootNode(const Bounds&) allocates the root covering the full region.
  - insertBody(QuadNode \*, Body&) recursively subdivides nodes so that each leaf contains at most one body.
- 3. Mass-Center Distribution. A post-order traversal aggregates mass and center-of-mass at every internal node:
  - computeMassDistribution(QuadNode \*) computes totalMass and centerOfMass.
- 4. Force Computation. For each body  $b_i$ , we traverse the tree and apply the opening-angle criterion  $\theta$ :
  - forceOnBody(const Body&, QuadNode, double theta) approximates distant clusters as single masses, or recurses into children when closer.
- 5. Parallelization (Shared-Memory). To exploit multicore CPUs, we compute forces in parallel:
  - computeForcesParallel(System&, QuadNode, double theta) spawns a pool of std::threads, partitions the bodies into chunks, and each thread calls forceOnBody on its subset.
  - Accelerations are stored per-thread and then written back, avoiding fine-grained locks.
- 6. **Time Integration.** Once all accelerations are known, we update velocities and positions in one pass:
  - updateBodies(System&, double dt) applies

$$\mathbf{v} \leftarrow \mathbf{v} + \mathbf{a} \, \Delta t, \quad \mathbf{x} \leftarrow \mathbf{x} + \mathbf{v} \, \Delta t.$$

- 7. Cleanup. The quadtree is deallocated to avoid memory leaks:
  - freeQuadTree(QuadNode \*) recursively deletes all nodes.

## Concurrency Aspects

- Multithreading in C++: we illustrate basic shared-memory concurrency by partitioning the force computation across threads. This uses standard std::thread and per-thread buffers, avoiding complex locking.
- Concurrent Data Structures: insertion into the quadtree could be parallelized by feeding bodies into a thread-safe queue; our current version remains serial for clarity but is structured to allow a concurrent insertBody using a mutex per node or a lock-free pointer array.
- GPU / PRAM Illustration (Optional): under USE\_CUDA, we provide simulateBruteForceGPU(System&,double) which launches an  $O(N^2)$  CUDA kernel. Each GPU thread computes one body's net force, demonstrating the PRAM model in practice.
- Correctness & Testing: we compare parallel results to a single-threaded reference on small N, and use tools like ThreadSanitizer to detect data races. Telemetry outputs (positions over time) are also verified for physical invariants (e.g. center-of-mass motion).