



Electron maps: from photon to electron visibilities

Università di Genova
DIMA | Dipartimento di Matematica

STIX Co-Location

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From photon to electron visibilities

Photon visibilities:

$$V(u, v; \epsilon) = \mathcal{F}(I(x, y; \epsilon)) = \int \int I(x, y; \epsilon) e^{2\pi i(xu + yv)} dx dy \tag{1}$$

Bremsstrahlung equation:

$$I(x, y; \epsilon) = \frac{a}{4\pi R^2} \int_{\epsilon}^{\infty} N(x, y) \bar{F}(x, y, E) Q(\epsilon, E) dE \tag{2}$$

$$N(x, y) = \int_0^{\ell} n(x, y, z) dz$$

$n(x, y, z)$ is the local density of target particles along the line-of-sight depth $\ell(x, y)$

$$\bar{F}(x, y; E) = \frac{1}{N(x, y)} \int_0^{\ell(x, y)} n(x, y, z) F(x, y, z; E) dz$$

$F(x, y, z; E)$ is the differential electron flux spectrum at the point (x, y, z) in the source

From photon to electron visibilities

Photon visibilities:

$$V(u, v; \epsilon) = \mathcal{F}(I(x, y; \epsilon)) = \int \int I(x, y; \epsilon) e^{2\pi i(xu + yv)} dx dy \tag{1}$$

Bremsstrahlung equation:

$$I(x, y; \epsilon) = \frac{a}{4\pi R^2} \int_{\epsilon}^{\infty} N(x, y) \bar{F}(x, y, E) Q(\epsilon, E) dE \tag{2}$$

Electron visibilities:

$$W(u, v, E) = \frac{a}{4\pi R^2} \int \int N(x, y) \bar{F}(x, y; E) e^{2\pi i(xu + yv)} dx dy \tag{3}$$

Bremsstrahlung equation for visibilities:

$V(u, v; \epsilon)$

Measured photon visibilities

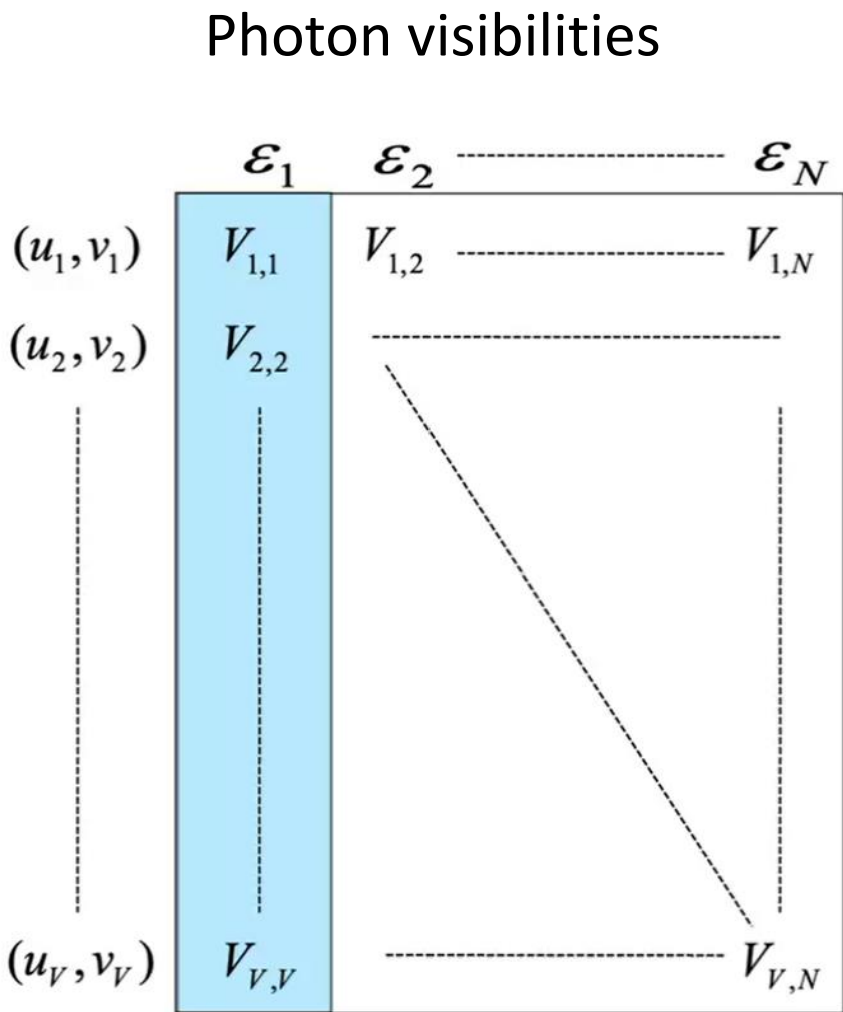
$$= \int_{\epsilon}^{\infty}$$

$W(u, v; E)$

Electron visibilities

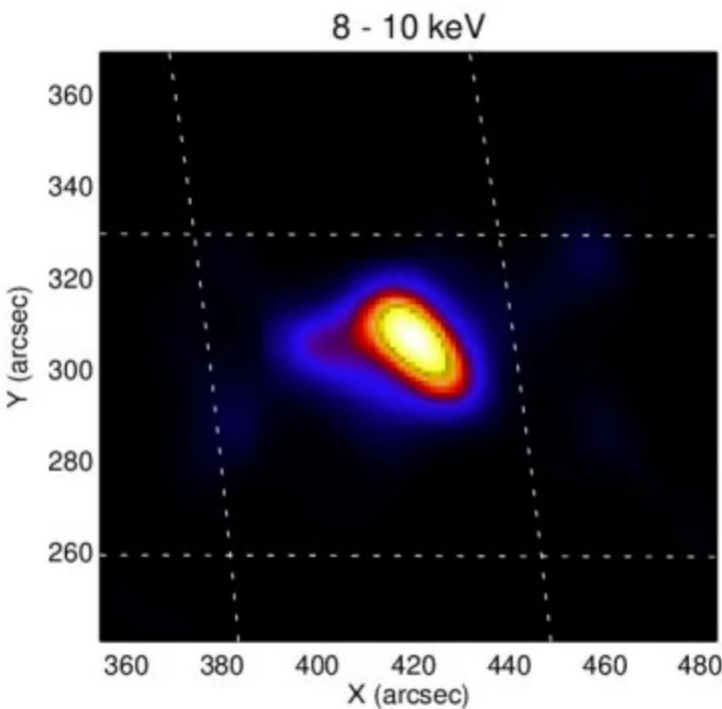
$$Q(\epsilon, E) dE \tag{4}$$

Photon visibilities



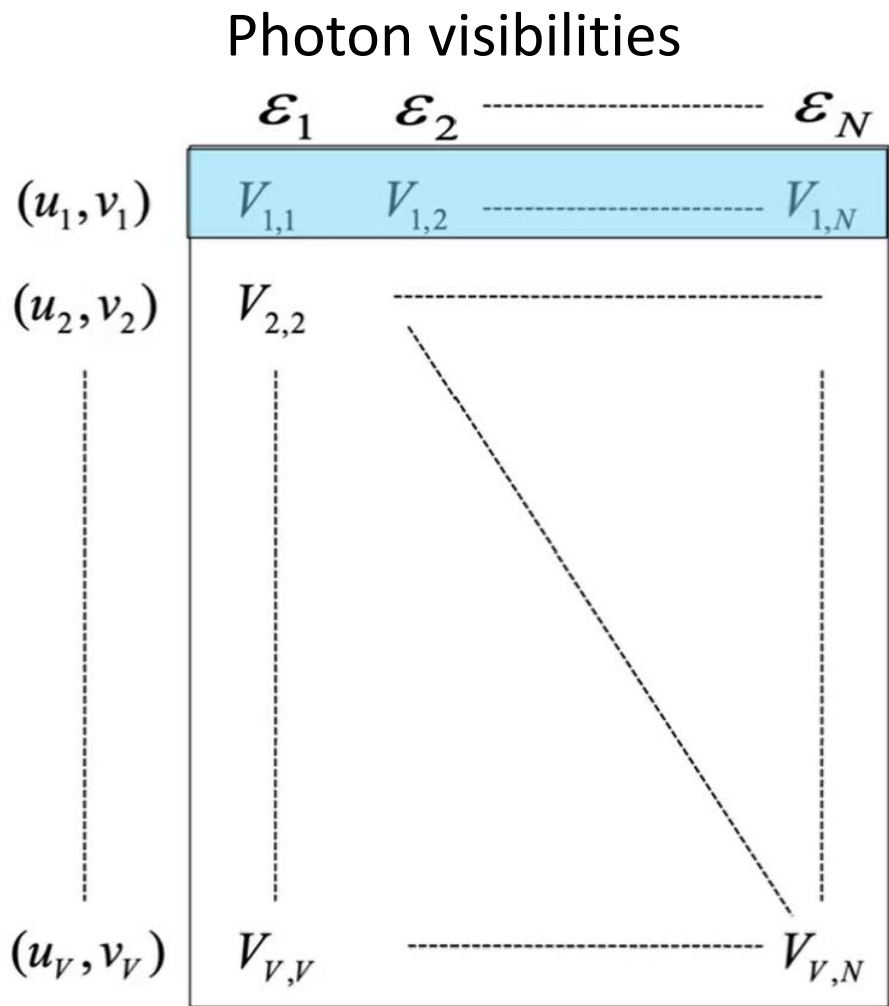
May 7, 2021

Time range = 18:51:00 – 18:53:40



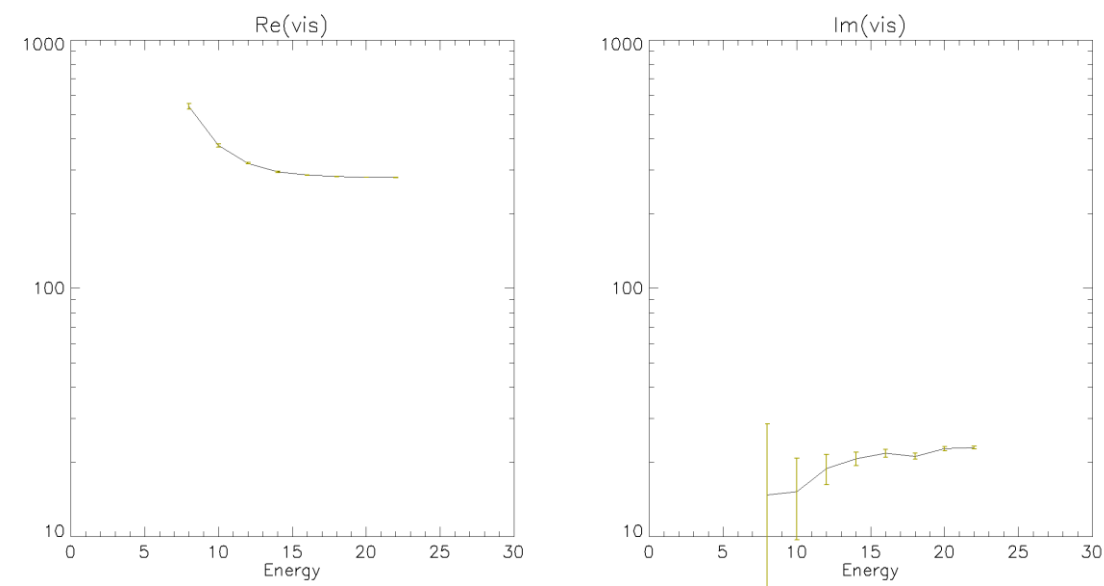
Reconstruction provided by MEM_GE, from photon visibilities.

Photon visibilities



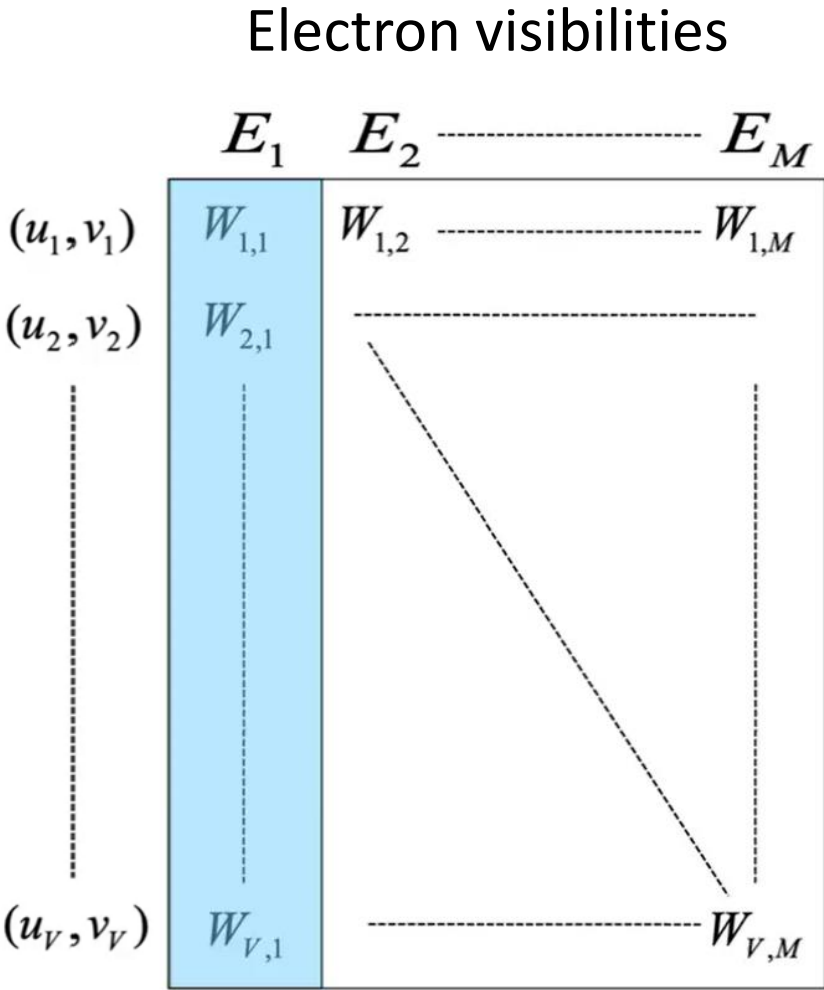
May 7, 2021

Time range = 18:51:00 – 18:53:40



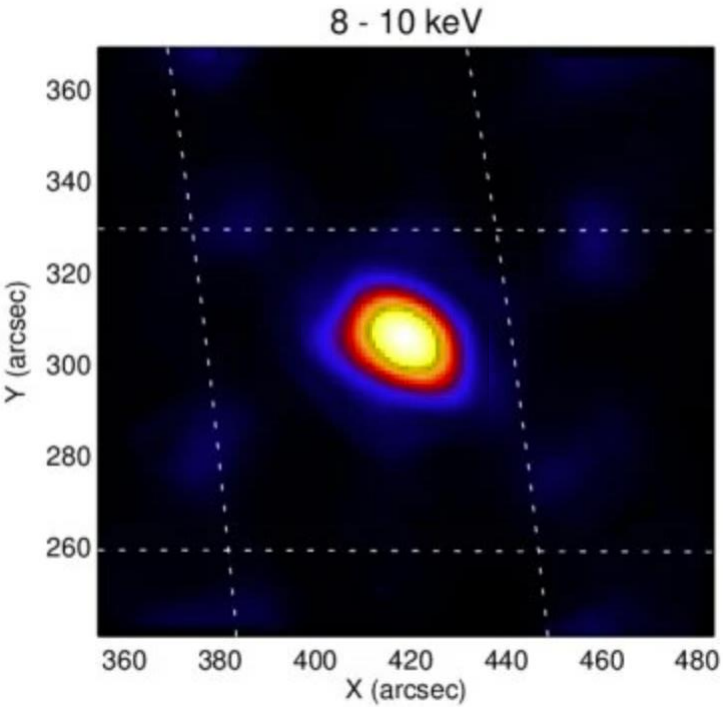
Real part (*on the left*) and imaginary part (*on the right*) of observed photon visibilities in $(u, v) = (0.002, -0.001)$ for eight energy bands.

Electron visibilities



May 7, 2021

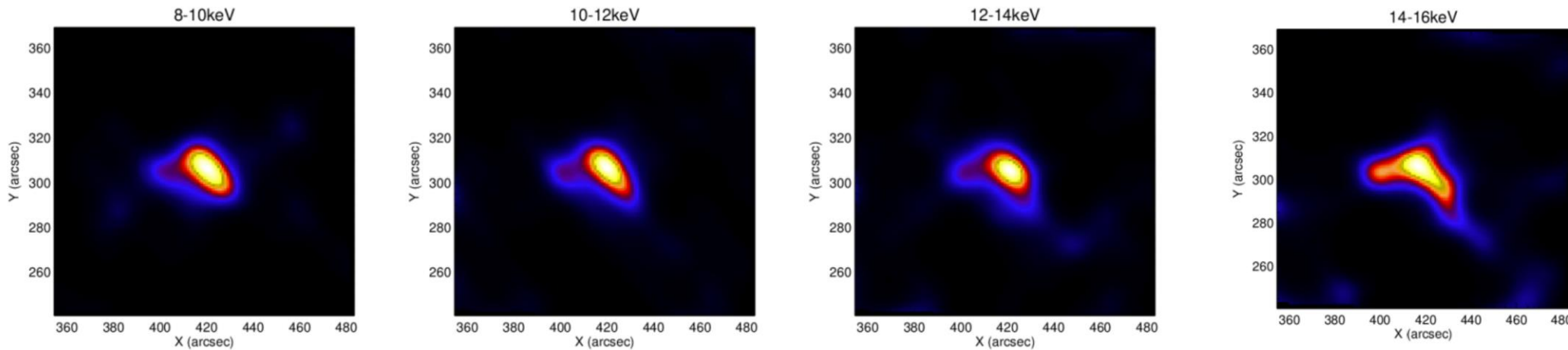
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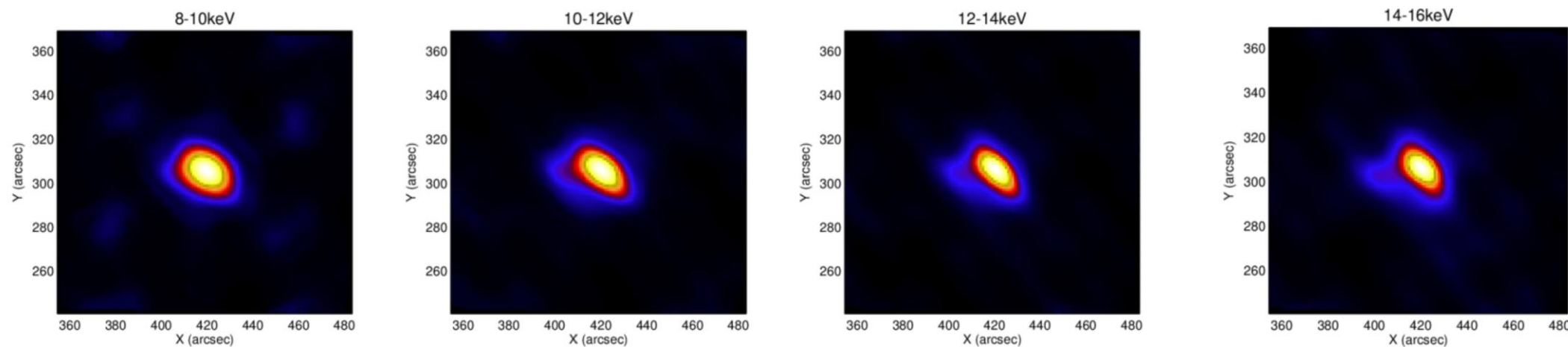
Reconstruction provided by MEM_GE, from electron visibilities.

Results – May 7, 2021

PHOTON VISIBILITIES

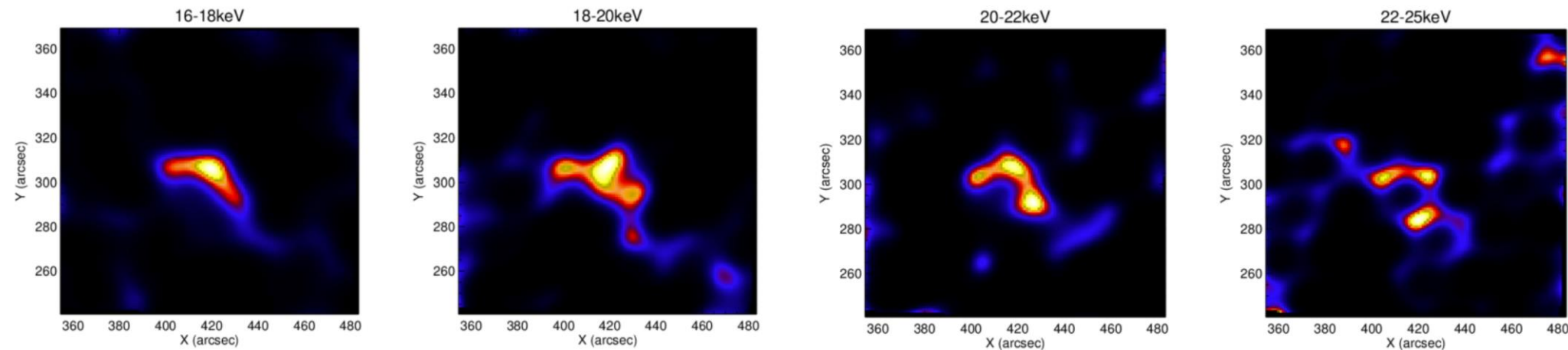


REGULARIZED ELECTRON VISIBILITIES

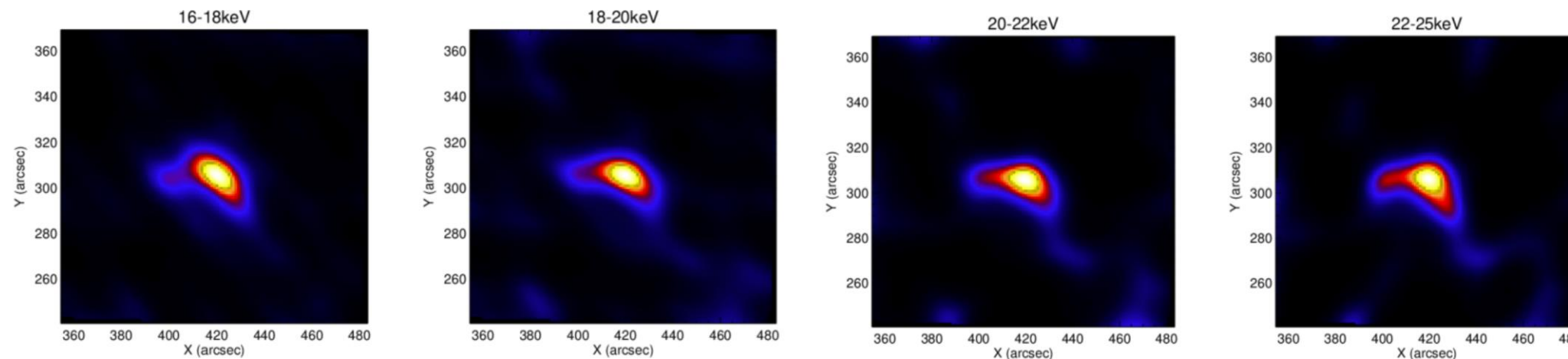


Results – May 7, 2021

PHOTON VISIBILITIES



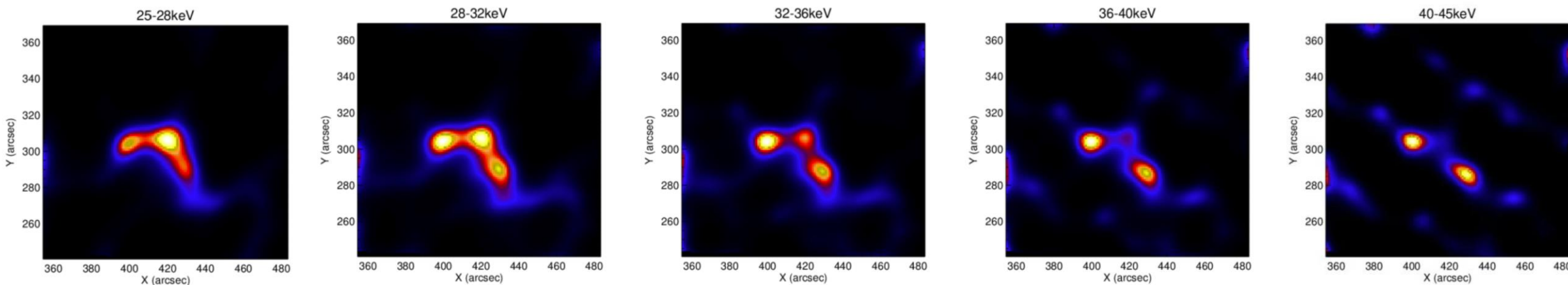
REGULARIZED ELECTRON VISIBILITIES



Results – May 7, 2021

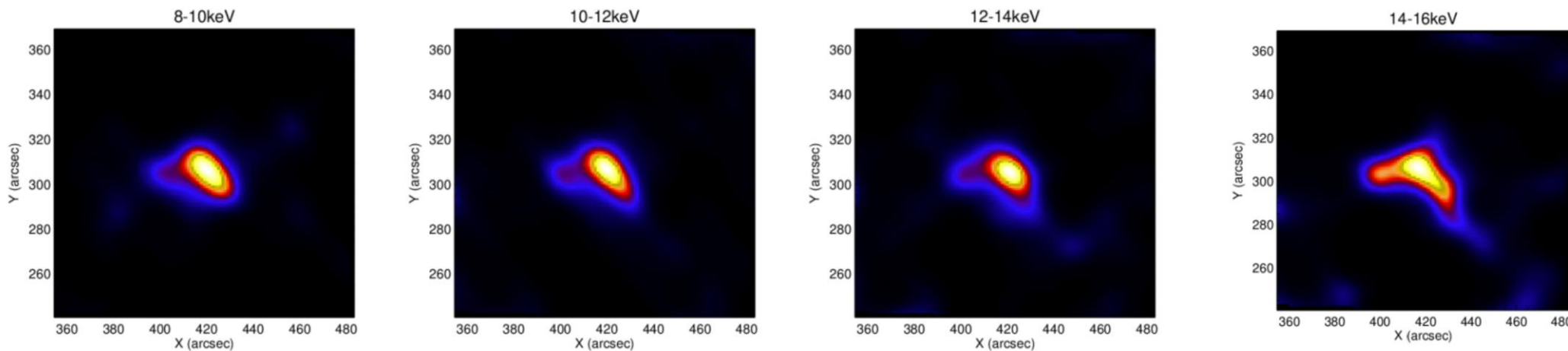
PHOTON VISIBILITIES

REGULARIZED ELECTRON VISIBILITIES

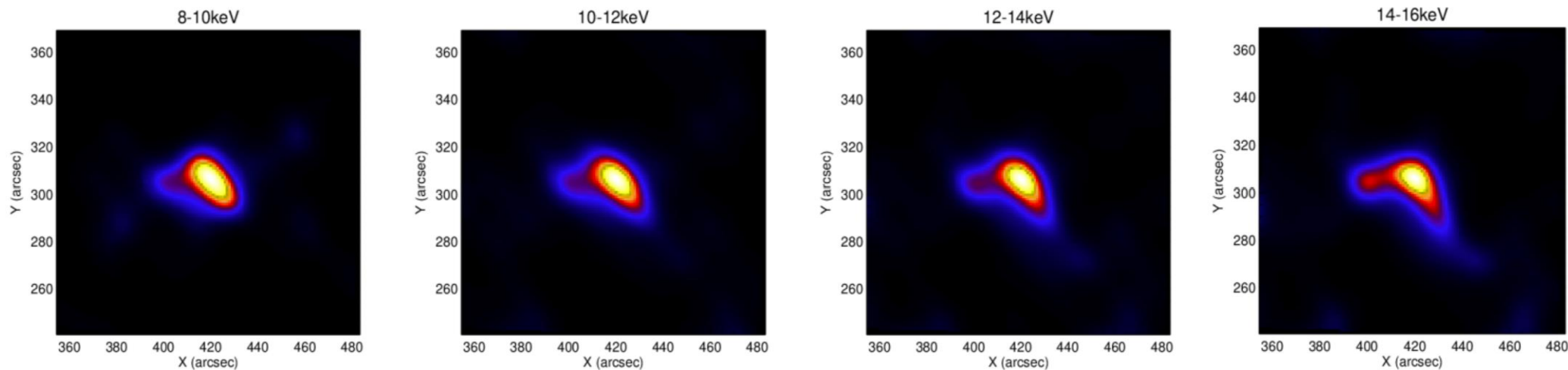


Results – May 7, 2021

PHOTON VISIBILITIES

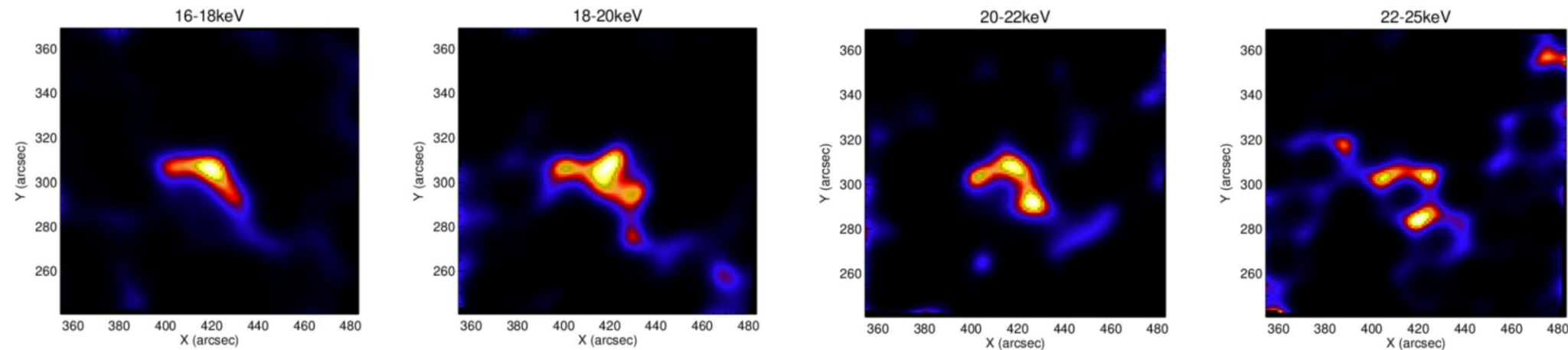


REGULARIZED PHOTON VISIBILITIES

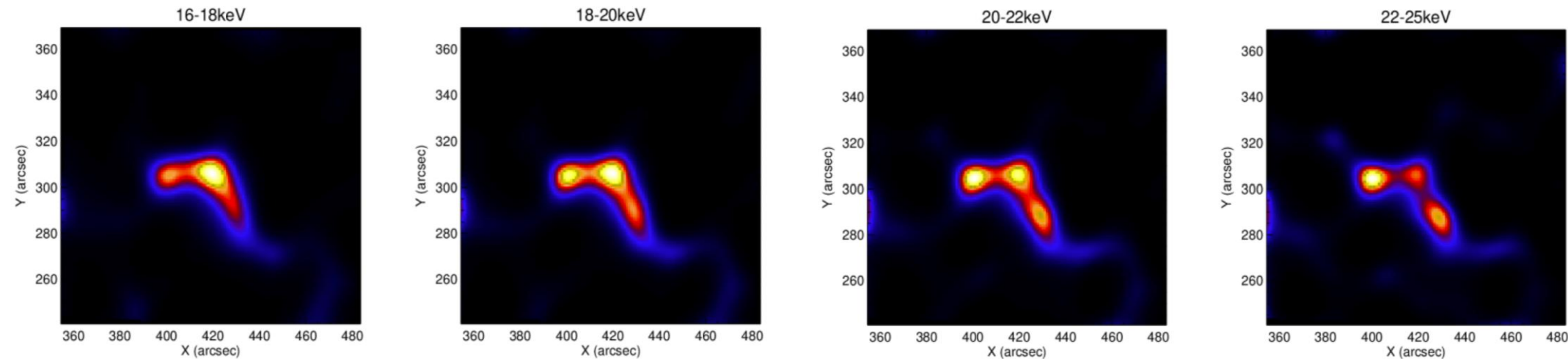


Results – May 7, 2021

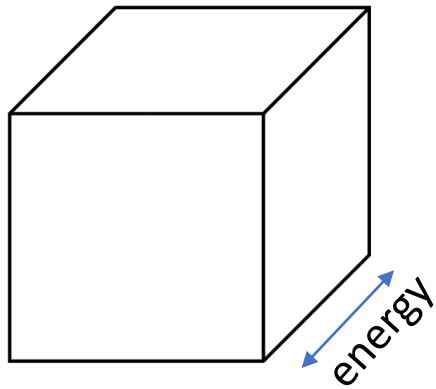
PHOTON VISIBILITIES



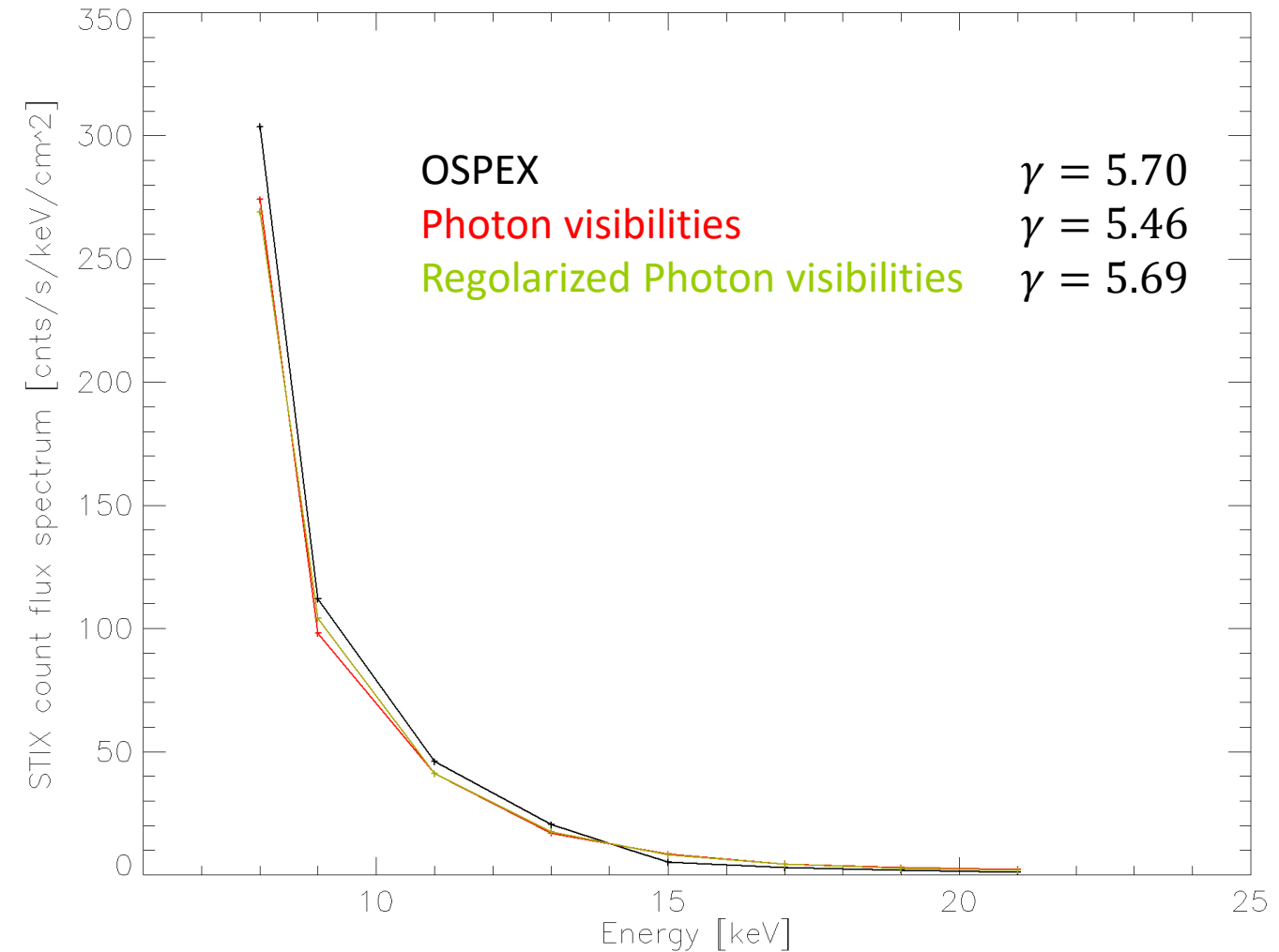
REGULARIZED PHOTON VISIBILITIES



Photon spectrum



1. For each energy bin consider the total flux in the recovered map;
2. Consider the total flux as a function of the energy and fit with a power law
 $A\epsilon^\gamma$
3. Consider the flux spectrum provided by OSPEX and fit it with a power law.





Multi-scale CLEAN

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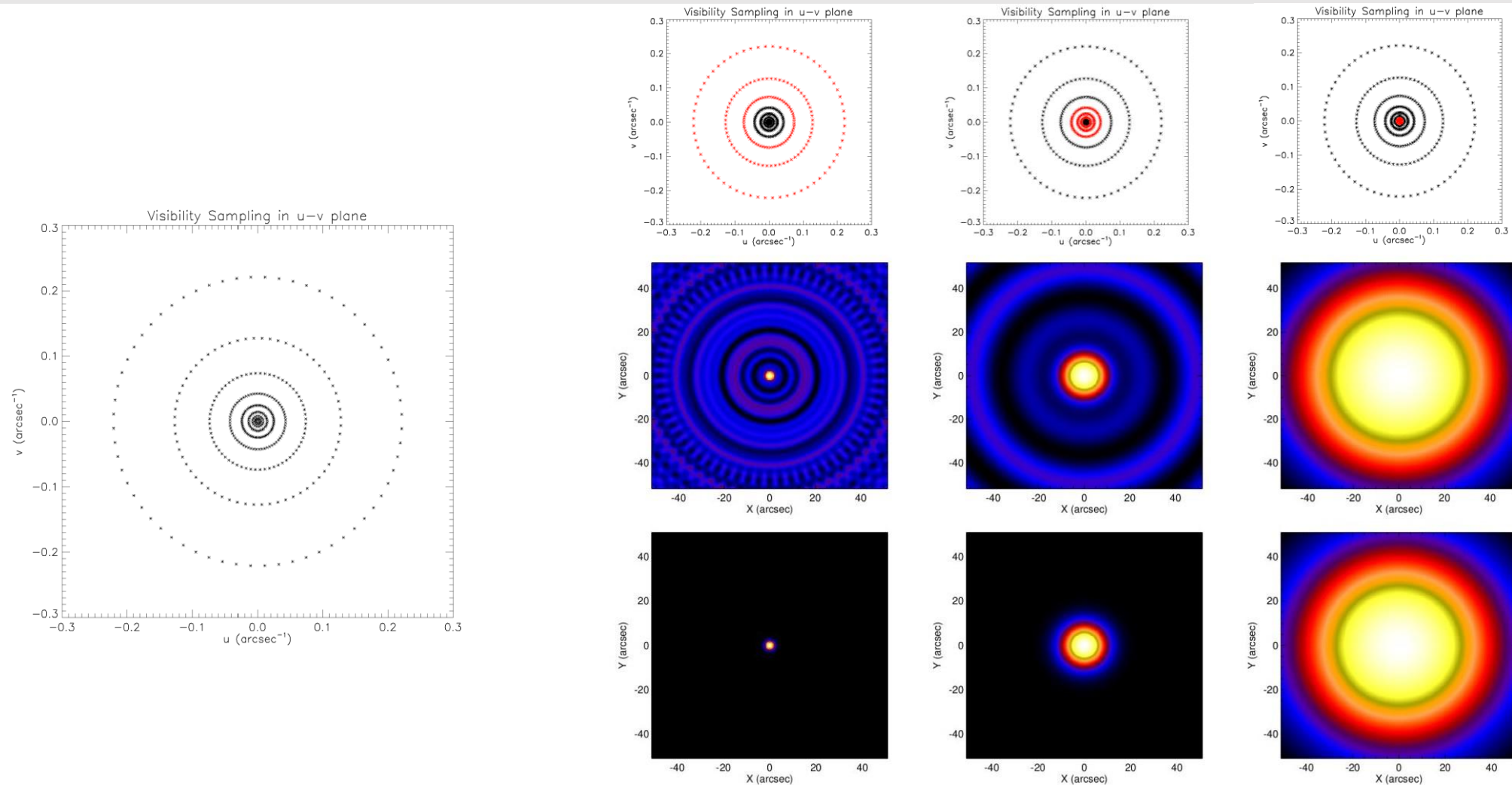
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Multi-scale CLEAN algorithm



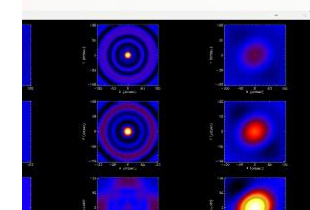
Top row: visibilities grouped in the same scale. Middle row: the three PSF components using 3 scales: det. 1-2-3, det. 4-5-6, det. 7-8-9. Bottom row: the three corresponding basis functions.

Multi-scale CLEAN algorithm

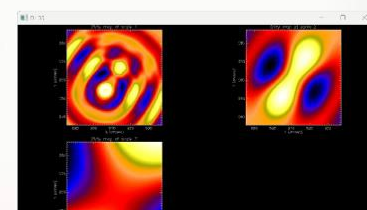
5025
5025
5641

```
scale = [1,1,1,2,2,2,3]
```

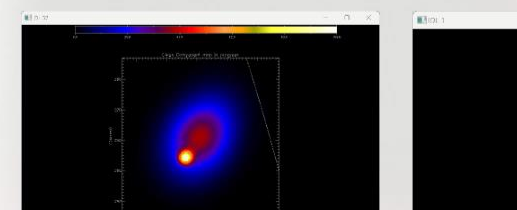
26



 IDL 35



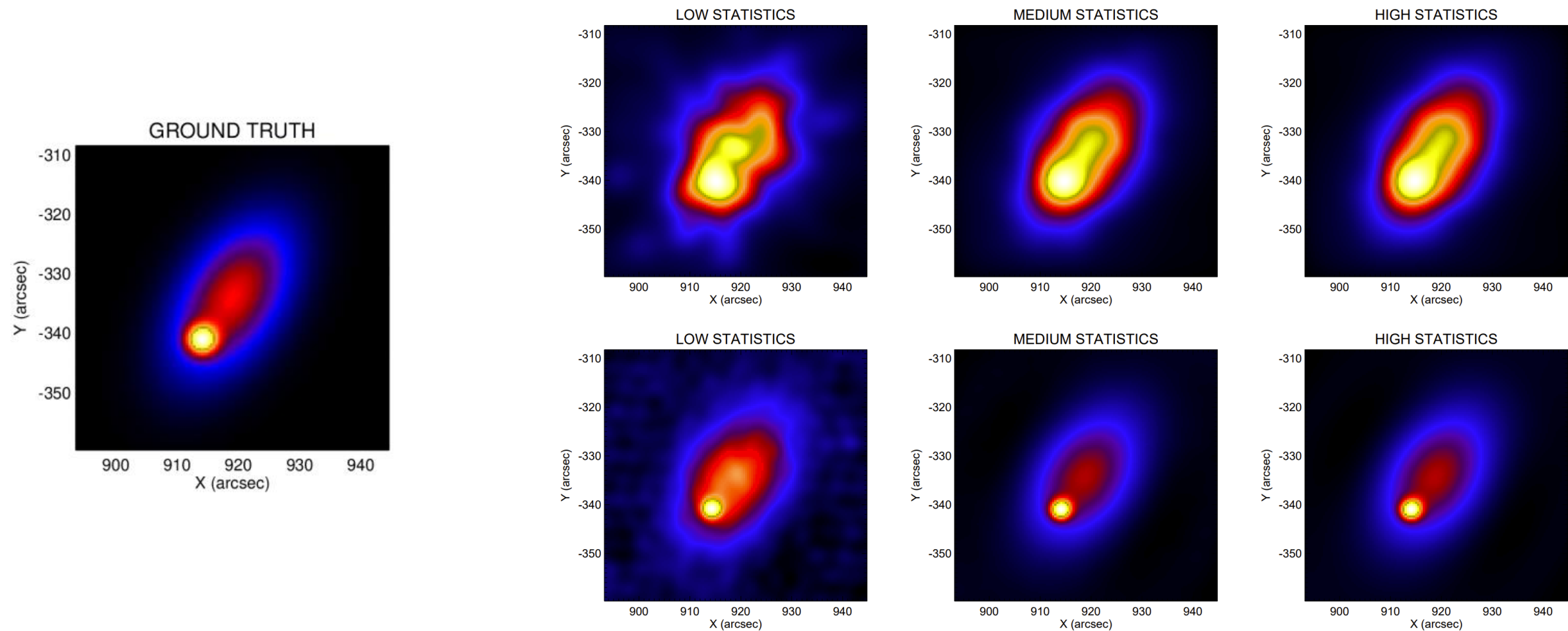
IDL 37 IDL 1



 IDL 1



Results – simulated visibilities

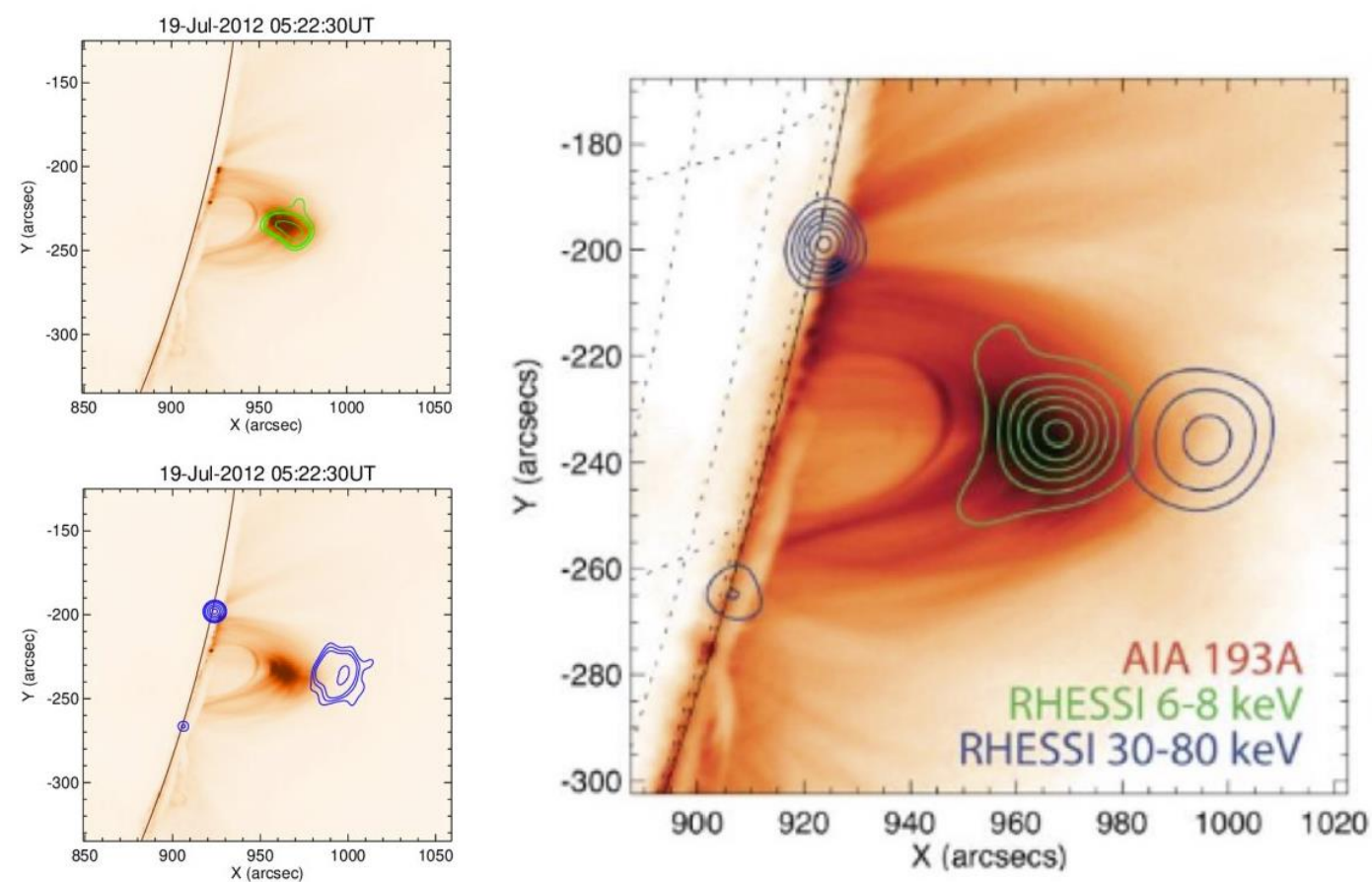


CLEAN vs multi-scale CLEAN for the reconstruction of the synthetic hard X-ray source. Reconstructions obtained by standard CLEAN (*top row*) and multi-scale CLEAN (*bottom row*) in the case of low, medium and high statistics.

Results – experimental visibilities

Energy range: 6 – 8 keV
3 scales: det. 3-4, det. 5-6, det. 7-8-9
Contour levels: 55, 65, 70, and 90%

Energy range: 30 – 80 keV
3 scales: det. 3-4, det. 5-6-7, det. 8-9
Contour levels:
Footpoints: 15, 20, 30, 50, 70, and 90%
Coronal source: 2, 2.5, 3, and 5%



Left panel: Multi-scale CLEAN reconstructions superimposed to the 193 Å extreme ultra-violet map recorded by SDO/AIA. *Right panel:* the EUV emission provided by AIA with superimposed the level curves provided by the two-step CLEAN algorithm.

Multi-scale CLEAN algorithm

We exploit the fact that hard X-ray telescopes are characterized by a PSF that can be written as the sum of a finite number of PSF components, each one filtering a specific portion of the (u, v) plane

We group the circles in the (u, v) plane in N disjoint sets, we can define a number N of experimental PSF components, i.e. of dirty beams $\{K_j(x, y)\}_{j=1, \dots, N}$ such that

$$K(x, y) = \sum_{j=1}^N K_j(x, y) \quad K_j(x, y) = \sum_{l=1}^{N_j} \exp[-2\pi i(xu_l^{(j)} + yv_l^{(j)})] \delta u \delta v \quad j = 1, \dots, N.$$

where N_j and $\{u_l^j, v_l^j\}_{l=1, \dots, N_j}$ are the number and the set of visibilities belonging to the j – *th* subset of sampled circles in the (u, v) plane.

Using the same approach, we can define a set of N dirty maps

$$I^D(x, y) = \sum_{j=1}^N I_j^D(x, y) \quad I_j^D(x, y) = \sum_{l=1}^{N_j} V(u_l^{(j)}, v_l^{(j)}) \exp[-2\pi i(xu_l^{(j)} + yv_l^{(j)})] \delta u \delta v \quad j = 1, \dots, N$$

Multi-scale CLEAN algorithm

The source image $I(x, y)$ can be modeled as the superposition of the basis functions

$$I(x, y) = \sum_{i=1}^N \sum_{q_i=1}^{Q_i} I_{q_i} m_i(x - x_{q_i}, y - y_{q_i}) + B(x, y) \quad \Longrightarrow$$

at scale i , for $i = 1, \dots, N$, there are Q_i sources, each one placed at (x_{q_i}, y_{q_i}) and with peak intensity I_{q_i} for $q_i = 1, \dots, Q_i$

$$I^D(x, y) = \sum_{j=1}^N I_j^D(x, y) = \sum_{j=1}^N \left(\sum_{i=1}^N \sum_{q_i=1}^{Q_i} I_{q_i} (m_i * K_j)(x - x_{q_i}, y - y_{q_i}) + (K_j * B)(x, y) \right) \quad (5)$$