

Through the Cosmic Lens: Deep Dive into Resolved Stellar Populations

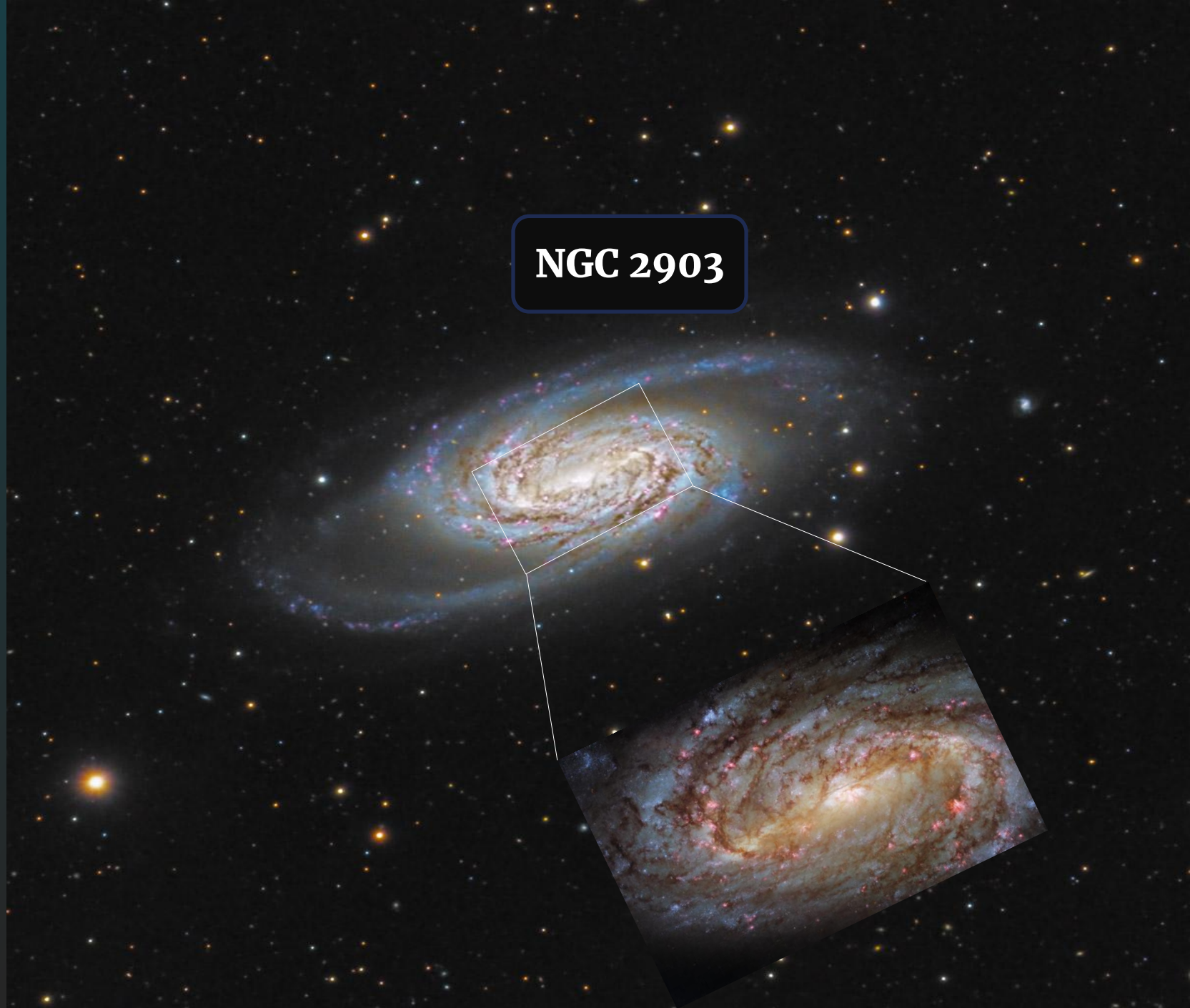


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NGC 2903

Project Goals and Scientific Motivation

Galaxy formation and evolution

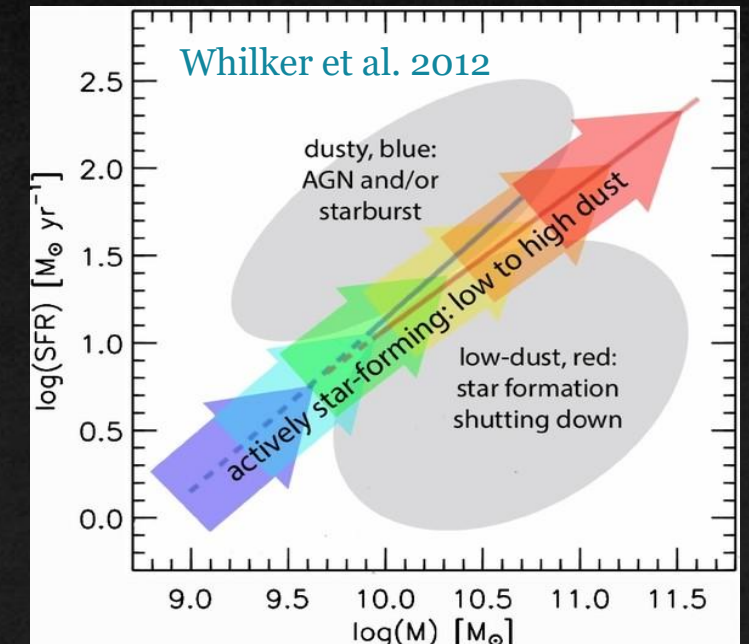
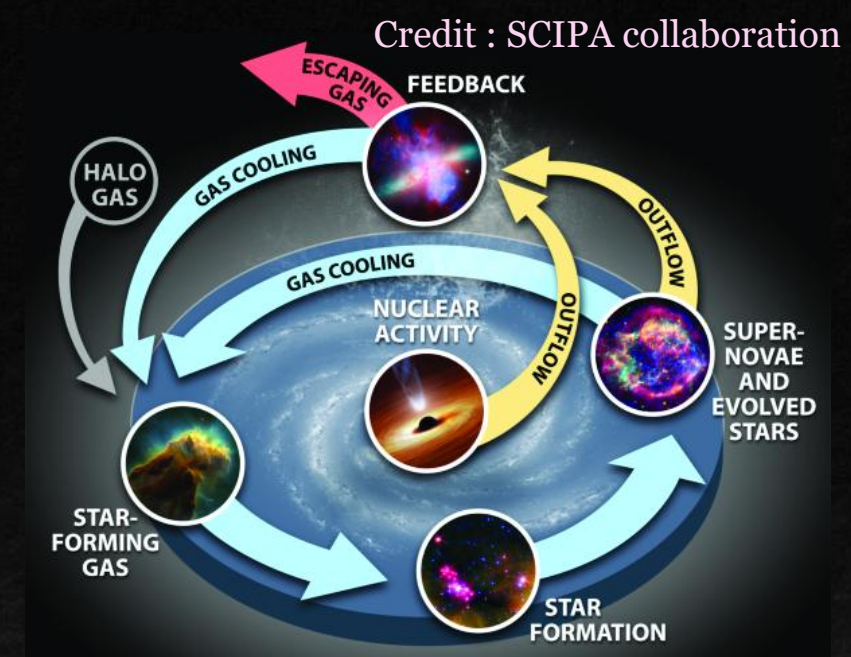
Morphology and demographics of SP

Baryon Cycle of stars – SFH

Metallicity variation – insitu SFH or radial migration

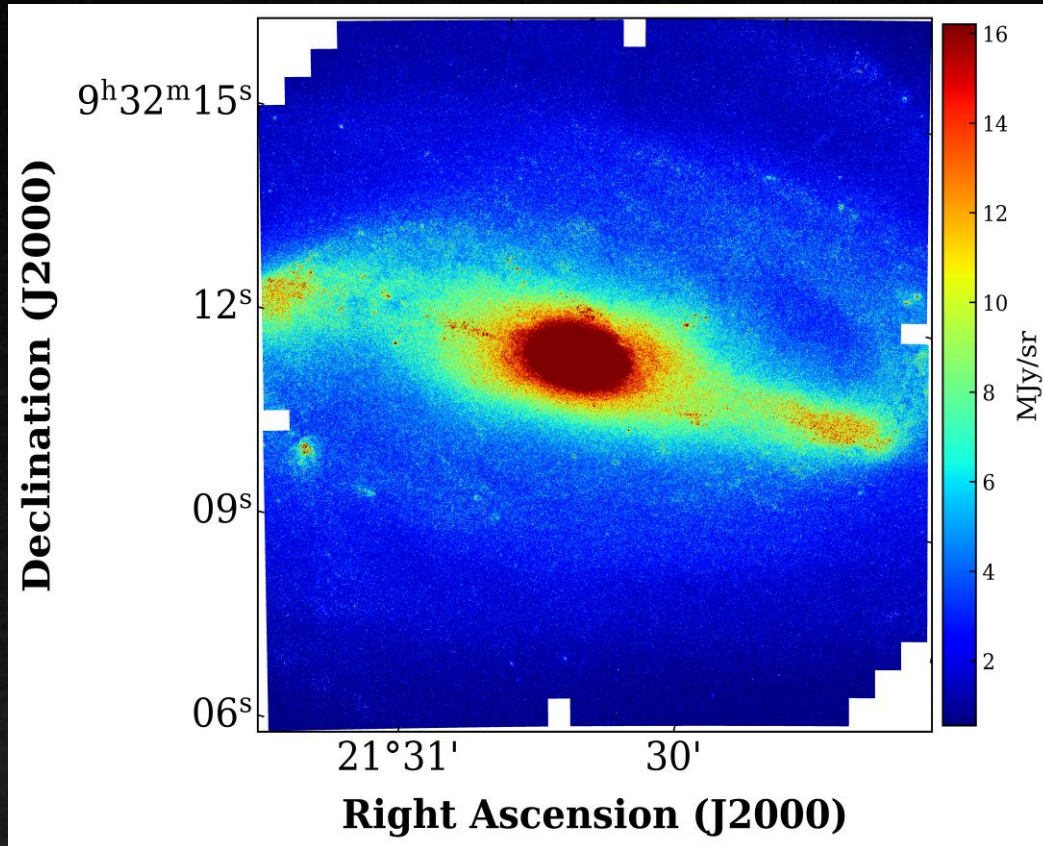
Age gradient – inside out formation scenario

Nearby Galaxies



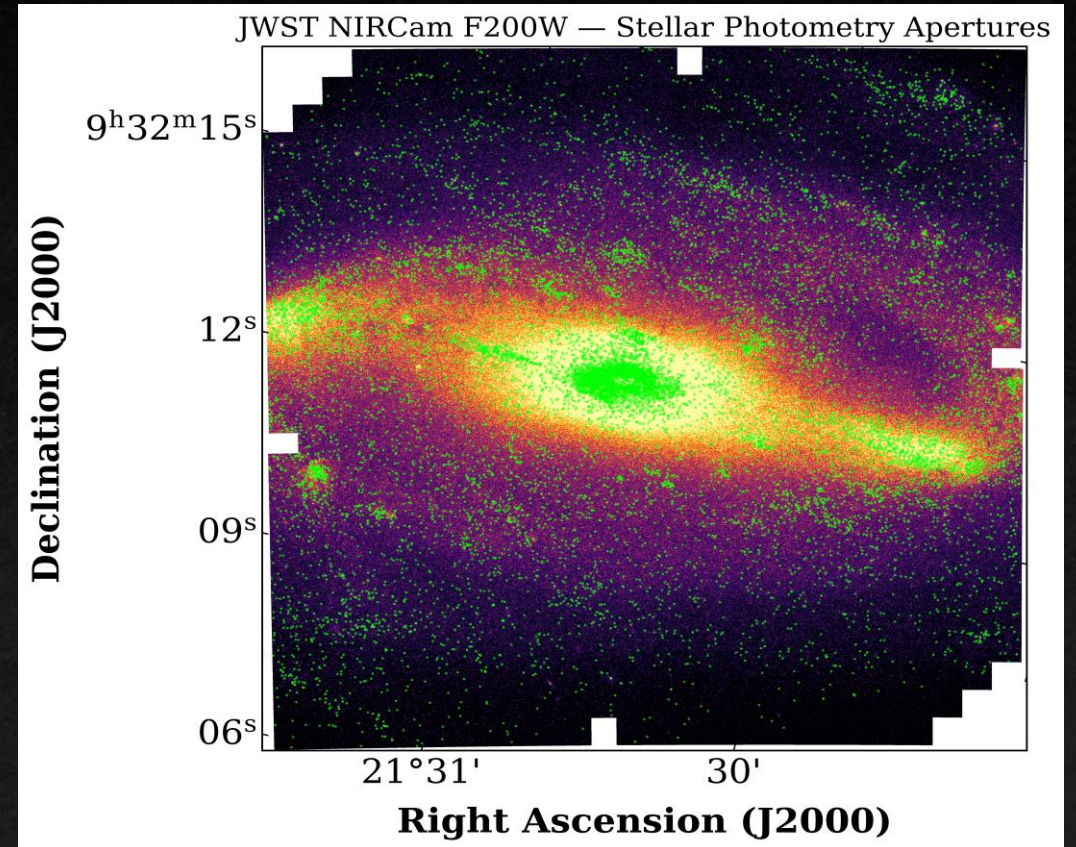
Unveiling the Star Formation Next Door

Sample data



JWST/NIRCam F200W image of NGC 2903

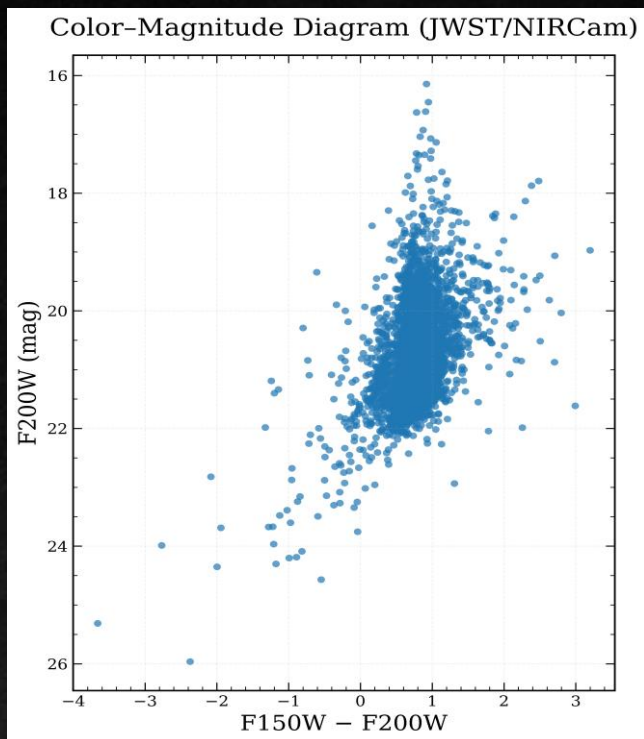
Source Detection



Identified stars plotted over galaxy image

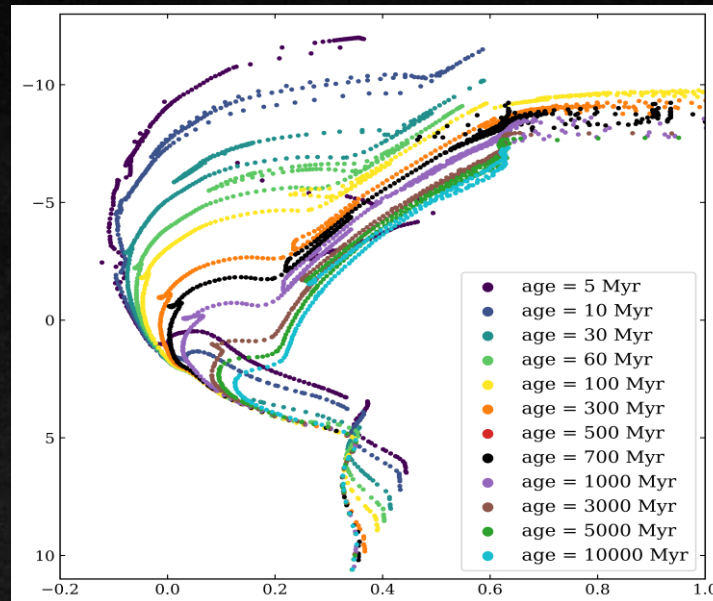
From Observation to Understanding

Observed CMD



Color Magnitude Diagram
for identified stars within
galaxy

Theoretical Isochrones

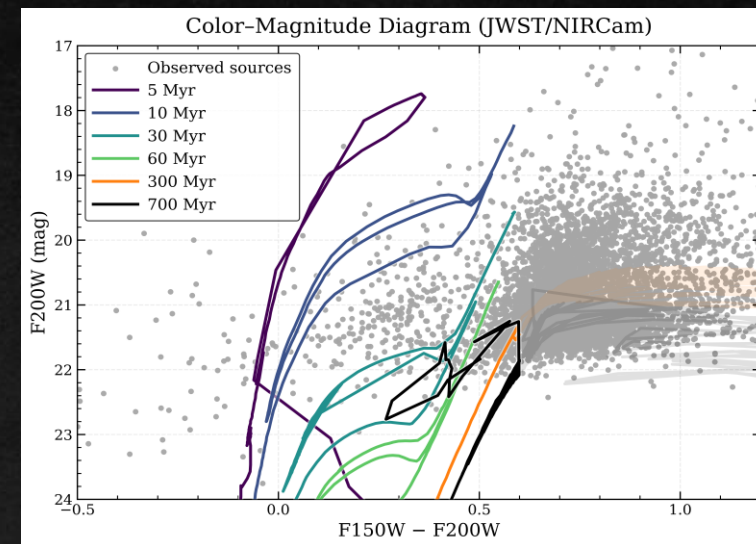
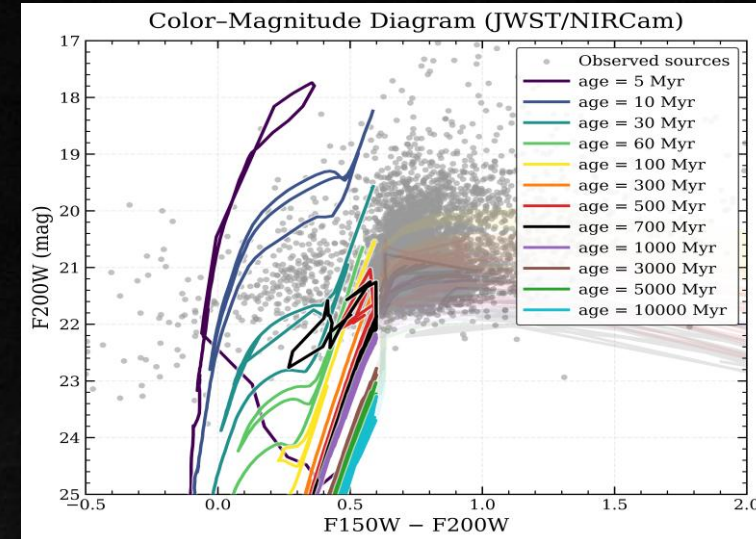


PARSEC-COLIBRI isochrones

Metallicity – 0.02 (Solar)

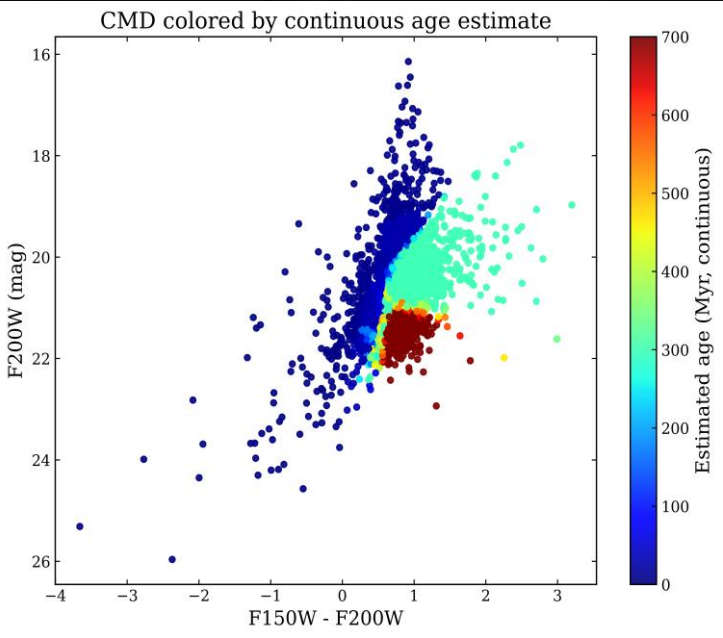
IMF – Kroupa (2001)

$E(B-V) = 0.03$, $m-M = 29.74$

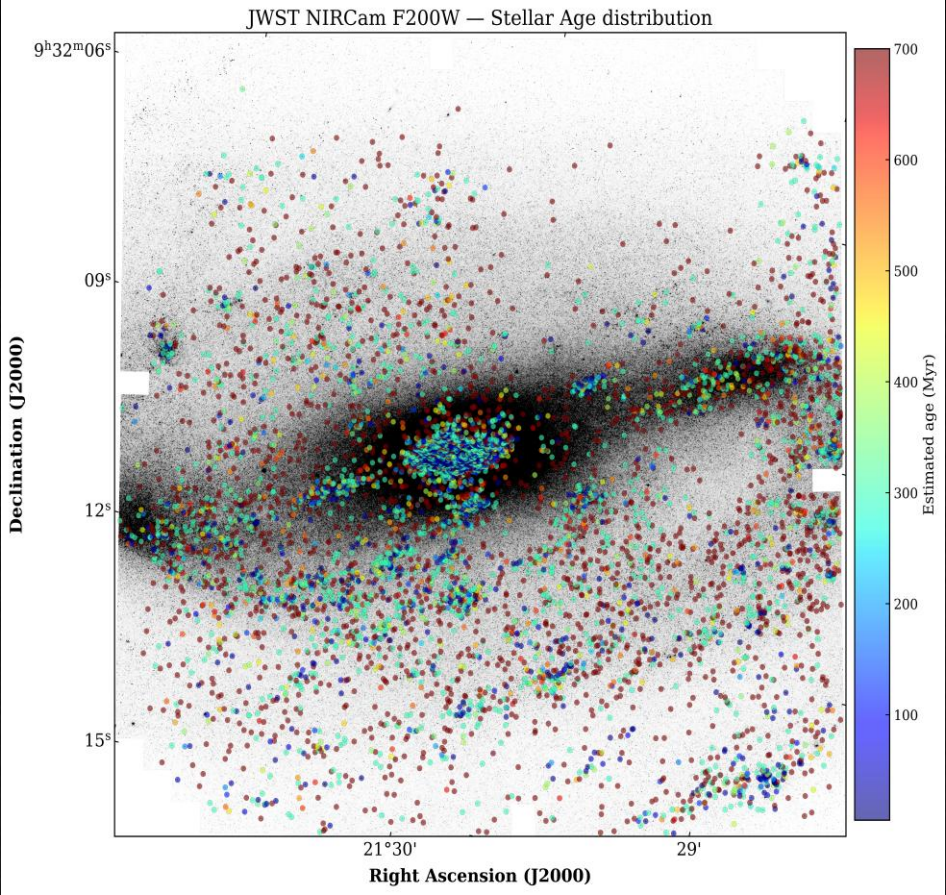


Observations + Theoretical
tracks analysis

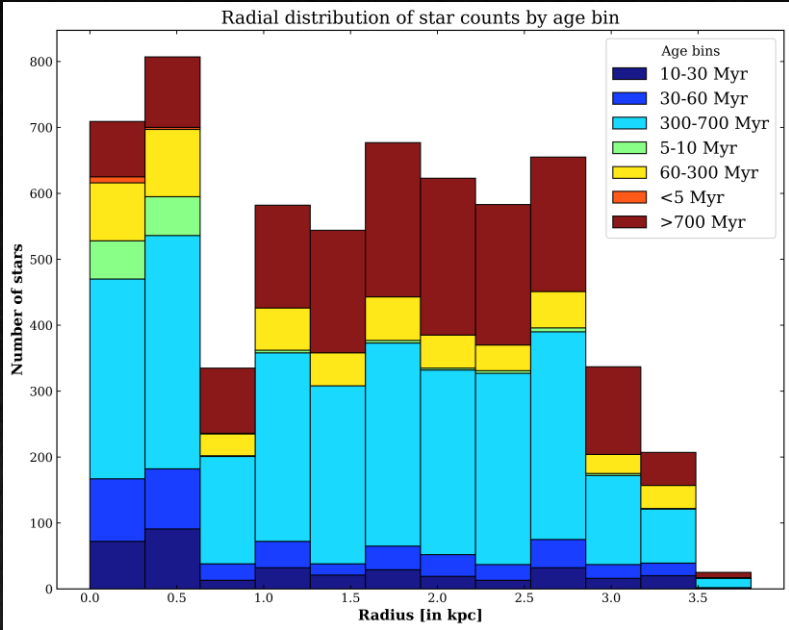
Result : Age distribution



Continuous stellar age distribution on CMD



Age distribution across the galaxy



Histogram of radial age distribution

Journey Beyond the Stars

Dust and PAH detection and distribution

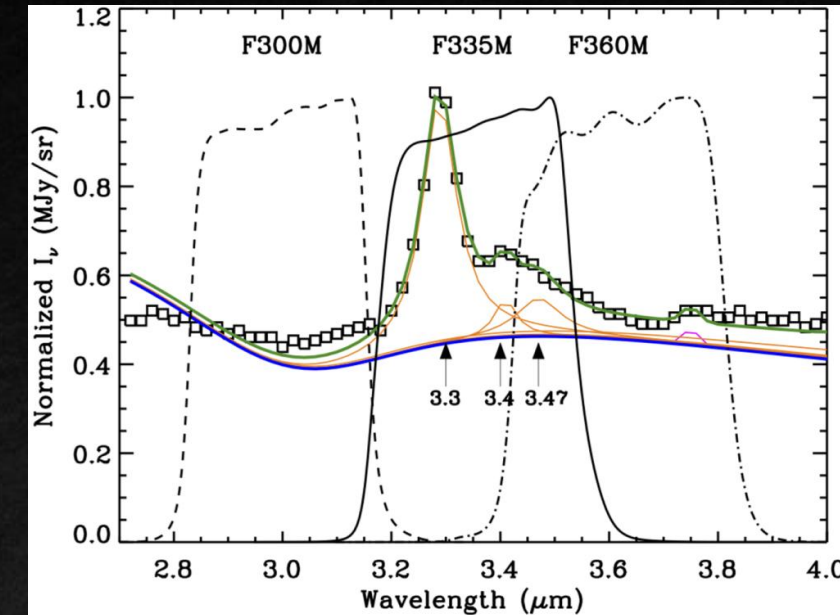
Formation mechanism of PAH – Chemical perspective

Type of dust that dominates the ISM
Fraction of dust in the form of PAHs,
 $R_{\text{PAH}} = (F_{770\text{W}} + F_{1130\text{W}})/F_{2100\text{W}}$

Balance b/w PAH formation and destruction in HII regions

and much more.....

Sandstrom et al. 2023



NIRCam F300M, F335M, and F360M
wavelength coverage, showing PAH,
aliphatic and Pfund γ emission