

New applications of effective field theories

Rothstein

Effective Field Theory:

$\vec{F} = m\vec{a} \Leftarrow S = T - V \rightsquigarrow$ ok, but how do I know V for every system?

→ Is there a way to find e.v.m.? \Rightarrow extremization procedure [action \rightarrow extreme \rightarrow classical con.]

↳ CLAIM: The action can be determined by knowing:

- 1) symm of S
- 2) symm of ground state (α_S)

↳ OK if only interested in gapless dofs (lowest energy)

The mean is GOLDSTONE THEOREM \Rightarrow Goldstone bosons are those gapless dofs
([counter examples])

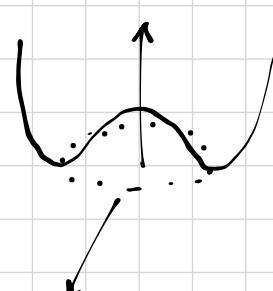
Build the action: ALGORITHM

- 1) Determine relevant d.o.f.
- 2) Action consistent with symmetries

→ Goldstone Theorem $G \xrightarrow{\text{broken}} H$:

T : unbroken generators

X : broken "



VACUUM MANIFOLD: $T|0\rangle = |b\rangle$

$\frac{G}{H}$ coset space \rightarrow equivalence class

ex: show that cosets only form a group if H is a normal subgroup $(ghg^{-1} \in H)$

$$(g h) \cdot (g' h') = (g h g^{-1})(g g') h' \in \frac{G}{H}$$

$\underbrace{\in H}_{\in G/H}$

$$(g h)^{-1} = h^{-1} g^{-1} = g^{-1} g h^{-1} g^{-1} = g^{-1} (g h g^{-1})^{-1} \in \frac{G}{H}$$

NB $V \rightarrow \infty$ limit \Rightarrow superselection rule: $\langle \phi_1 | \phi_2 \rangle = 0$ [you cannot pass between two points in the vacuum manifold]

locally is consistent \rightarrow Goldstone bosons

\rightarrow Proof (relativistic): \rightarrow every broken current couples to a massless particle

$$\begin{aligned} \lim_{q \rightarrow 0} \int d^d x q_\mu e^{iq \cdot x} \langle 0 | T[j^\mu(x) \phi(0)] | 0 \rangle &= \lim_{q \rightarrow 0} \int d^d x i \epsilon_{\mu\nu\rho} e^{iq \cdot x} \partial_\nu \langle 0 | T(j^\mu(x) \phi(0)) | 0 \rangle = \\ &= \lim_{q \rightarrow 0} q_\mu V^\mu(q^2) \quad \text{acting on } \Theta \text{ functions} \\ &= \lim_{q \rightarrow 0} q_\mu (q^\mu F(q^2)) = q^2 F(q^2) \\ &= \delta_{\mu,0} \langle 0 | [Q, \phi] | 0 \rangle \neq 0 \sim \langle 0 | \delta \phi | 0 \rangle \end{aligned}$$

$\xrightarrow{\text{then}} \lim_{q \rightarrow 0} q^2 F(q^2) \neq 0 \Rightarrow F(q^2) \xrightarrow{q \rightarrow 0} \frac{1}{q^2}$ PARTICLE!

\Rightarrow every broken current produces one Goldstone boson

\Rightarrow NON RELATIVISTIC:

$$\langle 0 | [Q, \phi] | 0 \rangle \neq 0 = \langle 0 | Q \phi(0) - \phi(0) Q | 0 \rangle = \int d^3 x \left[\langle 0 | j^\mu(x, t) \sum_n | n \rangle \langle n | \phi(0) | 0 \rangle - \dots \right]$$

\Rightarrow Assuming translational invariance:

$$\begin{aligned} &= \int d^3 x \left[e^{i \vec{p}_n \cdot \vec{x} - i E_n t} \langle 0 | j^\mu(0) \sum_n | n \rangle \langle n | \phi(0) | 0 \rangle - \dots \right] \neq 0 \\ &= (2\pi)^3 \left[\delta^3(\vec{p}_n) e^{-i E_n t} \langle 0 | j^\mu(0) \sum_n | n \rangle \langle n | \phi(0) | 0 \rangle - \dots \right] \neq 0 \quad \text{width} \ll \text{energy when } E \rightarrow 0 \end{aligned}$$

\Rightarrow state s.t. $E \rightarrow 0$ when $p \rightarrow 0 \Rightarrow$ can be anything (not necessarily a particle)

\hookrightarrow e.g.: solid lattice:

3 boosts
3 rotat.
3 translat.

broken \Rightarrow only 3 Goldstones (phonons)

\Rightarrow how do we define an action if we don't know how many dofs there are?

broken generators

$$U = e^{i \vec{X} \cdot \vec{\pi}(x)} \in G \quad (\text{local fluctuation on vacuum manifold})$$

↓
fictitious Goldstone bosons

$$\hookrightarrow g U(\pi) = U(\pi') \underbrace{h(\pi, g)}$$

H leaves everything invariant

$\Rightarrow H \rightarrow \pi$ transform linearly $H: \pi \rightarrow L\pi$

$G \rightarrow \pi$ " nat-lin (eg $\pi \rightarrow \pi + a$)

ex show that this is true:

$$\begin{aligned} [T, T] &\sim T \\ [T, X] &\sim X \end{aligned}$$

show that this is true first.

$S - TUx$

$$[S_i, S_j] = i \epsilon_{ijk} S_k$$

$$[T_a, X_p] = i f_{ap} S_i \Rightarrow f_{ap} = 0 \text{ if } i=a,b,c$$

(for compact lie group you can always choose a basis where struct const are totally antisym.)

\Rightarrow Maurer-Cartan 1-forms: $U^{-1} \partial_\mu U = T_a \omega_a + X_a \partial_\mu \pi^a$

under $\in G$ transf.: $* X_a \omega_a \rightarrow h(\pi, g) \omega_a h^{-1} h'(\pi, g) \Rightarrow$ transf. os if under H .

ex prove these

$$**: T^a A_a \rightarrow h T_a A_a h^{-1} + i (h h')$$

How do I build an action for π ?

\Rightarrow LOCALITY: (ways to define)

* micro-causality: $(x-y)^2 < 0 \quad [O(x), O'(y)] = 0$

* action of finite no. of derivative coupling (∞ no. \rightarrow couples everything \rightarrow against micro-causality)

* GR equivalence: well defined derivative expansion (analytic for $q \rightarrow 0$)

\hookrightarrow NO LOCALITY NO SCIENCE

$$L = (\partial_\mu \phi)^2 + c_1 \phi^2 + c_2 \phi^4 + \dots + c_3 (\partial_\mu \phi)^n + \dots \Rightarrow \text{No predictive power}$$

$(\infty \text{ no. of terms})$

\Rightarrow every EFT MUST have some power counting scheme allows to truncate with well defined of "theoretical error"

$$\text{eg } d=4 \rightarrow [\phi] = 1 \quad [S] = 0 \quad [O] = 1^d \quad \left[\int C_\alpha O \, d^d x \right] = 0 \rightarrow C_\alpha = 4 - d$$

Then $L = \sum_i \frac{C_i}{(A)^{d+4}} \partial_i \rightarrow \Lambda$: symm. breaking scale

$$\text{e.g.: } g \ll \Lambda \rightarrow [\phi(x), \dot{\phi}(x')] \sim \delta^3(x-x') \sim g^3 \Rightarrow [\phi] = \Lambda$$

① Show that:

$$\cdot e^{-A} \frac{d}{dx} e^A = \int_0^1 dy e^{-yA} \frac{dA}{dx} e^{yA}$$

$$\cdot e^A B e^{-A} = B + [A, B] + \frac{1}{2} [A, [A, B]] + \dots \text{ s.t. } e^{-A} \frac{d}{dx} e^A = A' + \frac{1}{2} [A', A] + \frac{1}{3} [[A', A], A]$$

② Construct COSET for $SO(n) \rightarrow SO(n-1)$ up to \mathcal{J}^4

$$\Rightarrow \text{from } U^{-1} D_u U = X^a D_u^a(\pi) + T^a A_a^a \rightarrow L = (D_u(\pi))^2 + \dots$$

\rightarrow suppose $\exists q$ transforms linearly in G : $q \rightarrow gq$

$$q := U(\pi) \tilde{q} \rightarrow g U(\pi) \tilde{q} = U(\pi') h(g, \pi) \tilde{q} = U(\pi') \tilde{q}'$$

$$\Rightarrow \tilde{q} \xrightarrow[G]{h} \underbrace{\tilde{q}}_{H}$$

$$\Rightarrow D_u \tilde{q} = (D_u \tilde{q} + A_u^a T^a \tilde{q}) \longrightarrow I \text{ can build invariants from this!}$$

\rightarrow SPACETIME SYMM.: non compact group!

\hookrightarrow in this case $[T, T] \sim T$, $[T, X] \sim T + X$

$$\Rightarrow U = e^{i \vec{\pi} \cdot \vec{x}}$$

\rightarrow Look at non relativistic cases \rightarrow Galilean symmetries

\hookrightarrow sth in GS $\Rightarrow \exists$ a defect in spacetime:

point particle: $x^i(t) \rightarrow$ breaks boost inv. + 3 transl.

string: $y^i(x, t) \rightarrow$ breaks boost + 2 transl + 1 rot...

membrane: $y^i(x^a, t)$

$$U \rightarrow e^{i H t - i \vec{p} \cdot \vec{x}} e^{i \vec{p} \cdot \vec{y}(x, t)} e^{i \vec{\eta} \cdot \vec{k}} U \Rightarrow \tilde{p}: \text{unbroken translations}$$

p : broken transl

η : broken Goldstone \rightarrow "feynons"

Galileian alg.

$$[K_i, P_j] = i \hbar \delta_{ij}$$

$$[L_i, L_j] = i \epsilon_{ijk} L_k$$

$$[L_i, K_j] = i \epsilon_{ijk} K_k$$

$$[H, K_i] = i P_i$$

\Rightarrow COSET for NR p.p.:

broken:

$$P_i$$

$$K_i \text{ (boost)}$$

unbroken:

$$L_i$$

$$M$$

$$\text{einheit (NR: 1; R: } \gamma)$$

$$\text{MC 1-form} \Rightarrow U^{-1} \partial_t U = i \tilde{E} \left(P_0 + B \pi^e P^e + D \eta^e K^e + M A_0 \right) =$$

$$= e^{-i\vec{\eta} \cdot \vec{K}} e^{i\vec{y}(t) \cdot \vec{p}} e^{-iHt} \partial_t \left[e^{iHt} e^{-i\vec{p} \cdot \vec{y}(t)} e^{i\vec{\eta} \cdot \vec{K}} \right] =$$

$$= e^{-i\vec{\eta} \cdot \vec{K}} e^{i\vec{y} \cdot \vec{p}} e^{-iHt} \left((iH)[...] + e^{iHt} (-i\dot{\vec{y}} \cdot \vec{p}) \dots + \dots (i\dot{\vec{\eta}} \cdot \vec{K}) e^{i\vec{\eta} \cdot \vec{K}} \right)$$

$$= \underbrace{e^{-i\vec{\eta} \cdot \vec{K}} (iH) e^{i\vec{\eta} \cdot \vec{K}}}_{\text{boost transf for } H} + e^{-i\vec{\eta} \cdot \vec{K}} (-i\dot{\vec{y}}(t) \cdot \vec{p}) e^{i\vec{\eta} \cdot \vec{K}} + i\vec{\eta} \cdot \vec{K} =$$

$$= i \left[H - \vec{p} \cdot \vec{\eta} + \frac{1}{2} M \vec{\eta}^2 \right] - i\vec{y}(t) \cdot (\vec{p} - M\vec{\eta}) + i\vec{\eta} \cdot \vec{K}$$

$$E=1 \Rightarrow P_0 = H$$

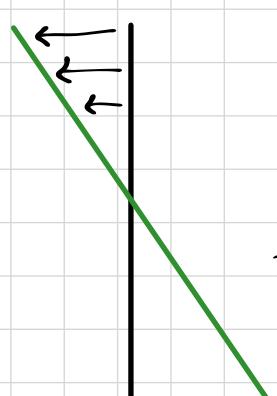
$$A_0 = \frac{1}{2} \vec{\eta}^2 + \dot{\vec{y}}(t) \cdot \vec{\eta}$$

$$D\pi = - \left[\dot{\vec{y}} + \vec{\eta} \right]$$

$$D\eta^e = \dot{\eta}^e$$

$\Rightarrow \eta$ is a redundant Goldstone boson! \Rightarrow shouldn't be there!

Suppose a membrane:



\rightarrow I can rotate or locally translate!

\hookrightarrow REVERSE HIGGS MECHANISM

$\Rightarrow [X, P] \sim X'$ (i.e.: [broken gen, unbroken] $\sim X'$)

Remove redundant Goldstone \rightarrow Impose constraint (compatible with symmetries)

$$\hookrightarrow \boxed{D\pi = 0} \Rightarrow \vec{i}\dot{\gamma} = -\vec{\eta} \Rightarrow \text{the boost is just the velocity of the particle}$$

\Rightarrow ACTION: we have $D\eta^a = -\ddot{i}\dot{\gamma}^a$

$$\text{then: } S = \int dt C_1 (D\eta^a)^2 = \int dt C_1 (\ddot{i}\dot{\gamma}^a)^2 \Rightarrow \text{which doesn't make sense!}$$

NB : locality \Rightarrow differential equation
stability \Rightarrow at most II order derivative } arXiv: 1506.0221

$\Rightarrow (\ddot{i}\dot{\gamma}^a)^2$ is fine as long as it is subleading \rightarrow OK for lengths given by the size of the particle

However I'm allowed to consider

$$\int Q A_0 dt \rightarrow \text{invariant up to a total derivative} \leftarrow \text{example of Wess-Zumino term.}$$

$$A_0 \propto \vec{\eta}^2 = \vec{i}\dot{\gamma}^2$$

$$\Rightarrow \int [C_2 \vec{i}\dot{\gamma}^2 + C_1 (\ddot{i}\dot{\gamma}^a)^2 + \dots] dt \rightarrow \text{"experimentally": } C_2 = \frac{1}{2} m$$

$$[S] = E \cdot T = L M \frac{L}{T^2} T = \frac{ML^2}{T} \quad \boxed{C_1 = ?}$$

$$[(i\dot{\gamma})^2] = \frac{L^2}{T^4} T = \frac{L^2}{T^3}$$

$$[C_1] = T^2 M \quad \begin{cases} \downarrow \\ C_1 \sim \frac{R^2}{V_s^2} M \end{cases} \quad \begin{matrix} \text{radius} \\ \text{of the particle} \\ \downarrow \\ \text{mass} \end{matrix}$$

\Rightarrow everything breaks down when the perturbation becomes of the same order as the LO.

2. interacting particles:



+ 3 boosts \Rightarrow inverse Higgs

Goplers modes: translation of the whole system + rotations [3 transl + 2 rot] (3D)

$$\text{Gap mode: } \omega = \sqrt{\frac{K}{m}} \quad (\text{contractions})$$

1 is unbroken

COSET CONSTRUCTION

Broken: $\vec{L}, \vec{T}, \vec{R}$

Unbroken: M, H

$$U = e^{iHt} e^{-i\vec{\eta} \cdot \vec{P}} e^{i\vec{\eta} \cdot \vec{K}} e^{i\vec{L} \cdot \vec{\theta}(t)}$$

$$\text{MC: } U^{-1} \partial_t U = (H + D\pi^e P^e + D\eta^e K^e + D\theta^i L^i + MA_0)$$

$$\Rightarrow D\pi^e = \eta^e + i j^e$$

$$A_0 = \vec{\eta} \cdot \dot{\vec{X}} + \frac{1}{2} \vec{\eta}^2$$

$$D\theta^i = -\dot{\theta}^i$$

$$D\eta^i = \dot{\eta}^i + \epsilon^{ijk} \dot{\eta}^j \theta^k$$

$$D\pi = 0 \Rightarrow j^e = -\eta^e$$

a) Using $g(U(\pi)) = U(\pi') h(\pi, g)$ determine how $\eta^i, \theta^i, j^e(t)$ transforms under boosts

b) Determine $D\eta^i$ to all orders in θ .

Then:

$$S = \int dt \left[\underbrace{c_1 A_0(t)}_{\text{WZ term}} + I^{ij} \dot{\phi}^i \dot{\phi}^j + \beta_i (D\eta^i = \dot{x}^i + \ddot{x}^j \partial^k \epsilon_{jk}^i + \dots) \right]$$

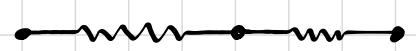


Use e.o.m. to eliminate \ddot{x} "order reduction":

$$S = \int dt \left(-\frac{1}{2} m \dot{x}^i \dot{x}^i + C \dot{x}^i D(x)^i \right) \quad x^i \rightarrow \frac{C}{m} D^i(x) + x^i \text{ cancels } O(c) \text{ but generates } O(c^2)$$

Write down first non vanishing term beyond canonical kinetic term

Multiple Oscillators

 modes: $\omega = 0, \sqrt{\frac{k}{m}}, \sqrt{\frac{3k}{m}} \dots$

What happens if the chain is ∞ ? "Large N limit"

$$\text{Diagram: } \cdots \xrightarrow[a]{\leftarrow} \xrightarrow[k]{\rightarrow} \cdots$$

$$W = 2\sqrt{\frac{k}{m}} \sin\left(\frac{qa}{2}\right), \quad q = \frac{2\pi}{L} n \quad L = Na \quad n = 1, \dots, N$$

$\Rightarrow q \rightarrow 0 \rightarrow \omega \sim q \Rightarrow$ waves! New gapless mode!

\downarrow
 ∞ chain \Rightarrow translational symm.

\downarrow
 spontaneously broken! \rightarrow PHONONS

Consider a inhomogeneous lattice:

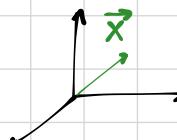
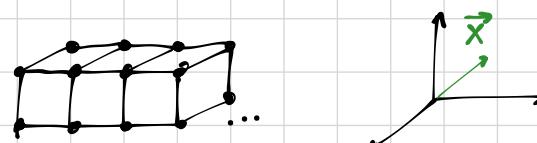
$$\cdots \xrightarrow[m_1]{\leftarrow} \xrightarrow[m_2]{\rightarrow} \cdots \Rightarrow \omega = \left(\frac{2ku}{m_1 m_2}\right)^{1/2} aq \quad u = \frac{m_1 m_2}{m_1 + m_2}$$

\hookrightarrow I must look at it from larger scales \Rightarrow coarse grain

Build coset for relativ. p.b. ($E \neq 1$)

\Rightarrow EFT of solids (1405.7384, 1712.07795)

\rightsquigarrow Enhanced Symmetry: rotational and transl. INTERNAL Symmetry



Internal Symmetry:

\vec{R} : int. rotations

\vec{Q} : int. transl.

\vec{L} : spacetime rotations

\vec{P} : " brausl."

$$\rightarrow \Phi^i(\vec{x}, t) \quad \text{s.t. } \langle 0 | \Phi^j(x, t) | 0 \rangle = x^j = \delta_a^i x_a$$

\Rightarrow Symm. breaking pattern: $SO(3)_{\text{int}} \otimes SO(3)_{\text{ST}} \rightarrow SO(3)_{\text{int+ST}}$

$$T_{\text{int}}^3 \otimes T_{\text{ST}}^3 \longrightarrow T_{\text{int+ST}}^3$$

Effective Field Theory:

Broken: $Q_i, \vec{P}_i, L_i, R_i, K_i$

Unbroken: $H, M, L_i + R_i, Q_i + P_i$

$$\rightarrow U = e^{i\bar{P} \cdot x} e^{i\bar{R} \cdot \vec{\eta}} e^{i\vec{Q} \cdot \vec{\pi}} e^{i\bar{R} \cdot \theta} \quad \text{where } \bar{P} = (H, Q^i + P^i)$$

$\gamma_i(t) \xrightarrow{\text{continuous}} \pi(x, t)$

$$\Rightarrow U^{-1} \partial_\mu U = E_\mu^A (\nabla_A \pi + \dots)$$

(, galileian boosts : $E_0^0 = 1, E_i^j = \delta_i^j, E_0^i = \eta^i, E_i^0 = 0$)

$$\rightarrow A_0 = \frac{1}{2} \vec{\eta} \quad \nabla_0 \pi^i = \dot{\pi}^i - \eta^a (\delta^a \pi^i + \partial^j \epsilon^{ija}) + \eta^i \quad \nabla_j \eta_i = \partial_j \eta_i$$

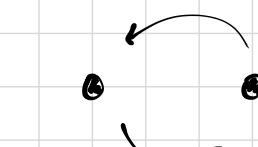
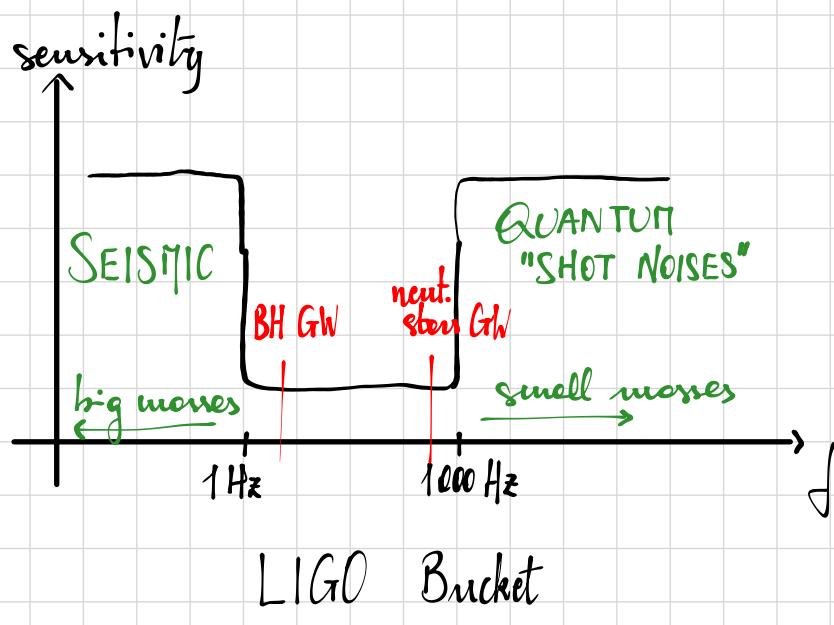
$$A_i = \eta_i \quad \nabla_j \pi_i = \partial_j \pi_i + \partial_k \epsilon_{ijk} \quad \nabla_0 \theta_i = \dot{\theta}_i$$

$$\nabla_i \theta_j = \partial_i \theta_j$$

\rightarrow inverse Higgs: $\nabla_0 \pi^i = 0 \rightarrow$ solve for η \rightarrow only π remains

$$\epsilon_{ij} \nabla_j \pi_i = 0 \rightarrow$$
 solve for θ

GRAVITY



\Rightarrow no radiation: I.S.C.O.: inner stable orbit
radiation: collapse

$$@ \text{ISCO} \quad f_{\text{plus}} \sim \frac{1}{M}$$

Signal: $\Rightarrow F(M_1, M_2, S_1, S_2, I)$ internal struct.

\rightarrow Use EFT:

SCALES:

R : radius of compact object

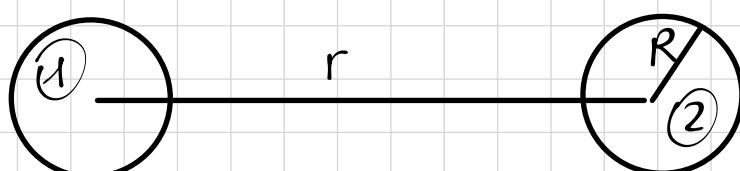
r : radius of orbit

$$R \ll r \ll \lambda$$

$$\lambda \sim \frac{r}{v}$$

as long as $v \ll 1$ we expand in v

EXPANSION PARAMETERS: $\frac{R}{r}, \frac{v}{c}, \frac{\hbar}{L}$ \rightarrow Feynman diag: so small. we ignore loops



① $r \gg R$: treat system as interacting p.p.

$$S = S_0 + S\left(\frac{R}{r}\right) \rightarrow \text{finite size effects}$$

$$(i=1,2) \quad S_0 = -\sum_i m_i \int d\lambda_i \sqrt{\frac{dx^{\mu}}{d\lambda} \frac{dx^{\nu}}{d\lambda}} \eta_{\mu\nu} + q_i \int v^{\mu} A_{\mu} d\lambda \quad x^{\mu}(\lambda): \text{World line}$$

FINITE SIZE effects (LO)

Syn: Lorentz, Reparam. invan. (RPI), Gauge invan.

$$S_{FS} = \sum_i \int d\lambda_i F_{\mu\nu} F^{\mu\nu} C_i \sqrt{v^2} + \sum_i \int d\lambda_i \left(V_i^{\mu} F_{\mu\nu} F^{\nu\rho} V_{i\rho} \right) \frac{C_i^2}{\sqrt{v^2}}$$

$$V^{\mu} = \frac{\partial x^{\mu}}{\partial \lambda}$$

NB: $F \rightarrow \bar{F} + SF \Rightarrow C^1, C^2 := \text{en. polarizability}$

$$\textcircled{1} \quad S = \int C_1 F^2 \sqrt{V^2} + C_2 (V \cdot F)^2 / \sqrt{V^2} \quad (\text{see notes})$$

↓ NR limit

$$\int_G E^2 + C_B B^2, \quad A^\mu = A_B^\mu + \delta A^\mu$$

show that $\delta A^\mu(z) = -\frac{C_B}{2\pi} \frac{(\vec{z} \times \vec{B})'}{|z|^3} \delta_\mu^\nu - \frac{C_E}{2\pi} \delta_\mu^\nu \frac{z^\lambda E_\lambda}{|z|^3}$

$$\vec{P} = -2C_E \vec{E} \quad C_E = -2\pi \alpha_E$$

$$\vec{M} = -2C_B \vec{B} \quad C_B = -2\pi \alpha_B$$

e.g.: perfect conductor: $C_E = \pi R^3, C_B = \frac{\pi}{2} R^3$

Consider charged part



1) solve exactly

2) solve through EFT

$$Z[J] = e^{-iW[J]} = \int d\phi e^{-i(s+J\phi)} = e^{iE[J]t}$$

$$L = -\frac{1}{4} F^2 + J \cdot A$$

$$J^\mu = \sum_i q_i \int d\lambda v_i^\mu \delta(x^\mu - x_i^\mu(\lambda))$$

$$\rightarrow W[J] = -\frac{i}{2} \int dx dy J^\mu(x) G_{\mu\nu}(x-y) J^\nu(y) = -iE[J]t$$

→ NR limit (Feynman gauge): $= \frac{d^4 k}{(2\pi)^4}$

$$W = -\frac{i}{2} \int \frac{[d^4 k]}{k^2 + i\epsilon} \int dt dt' [e^{ik \cdot (x_1(t) - x_2(t))} \cdot$$

$$v^\mu = (1, \vec{v})$$

$$\cdot q_1 q_2 (1 - \vec{v}_1(t) \cdot \vec{v}_2(t)) + q_1 q_2 (1 - \vec{v}_1(t) \cdot \vec{v}_2(t)) e^{+ik \cdot (x_1(t) - x_2(t))} + \text{self-energy}]$$

STATIC LIMIT: $V \rightarrow 0$ term:

$$W = -\frac{i}{2} \int \frac{[d^4 x]}{k^2 + i\epsilon} dt dt' (q_1 q_2) e^{i k_0 (t-t') - i \vec{k} \cdot (\vec{x}_1 - \vec{x}_2)}$$

indep of + in
static limit

$$= -i \int \frac{dk_0}{2\pi} \frac{d^3k}{(2\pi)^3} q_1 q_2 e^{-i\vec{k} \cdot (\vec{x}_1 - \vec{x}_2)} \frac{e^{-i\vec{k}^2}}{\vec{k}^2 - k_0^2} \cancel{\delta(\vec{k})}$$

$$= +it \int \frac{d^3k}{(2\pi)^3} q_1 q_2 e^{-i\vec{k} \cdot (\vec{x}_1 - \vec{x}_2)} = \frac{i q_1 q_2 t}{4\pi |\vec{x}_1 - \vec{x}_2|} \Rightarrow$$

$$\rightarrow E[J] = \frac{q_1 q_2}{4\pi r}$$

NB: $O(v^2)$: $v \rightarrow \overset{\circ}{k} \rightarrow 0$ $\sim v$ corrections

static limit: force is instantaneous!

→ moving particle: CORRECTIONS



$$\frac{1}{\vec{k}^2 - \vec{k}'^2} \sim -\frac{1}{\vec{k}^2} + \frac{\vec{k}'^2}{\vec{k}^4} + \dots$$

EX

$$V^{O(v^2)} = -\frac{q_1 q_2}{8\pi r} \left[\vec{v}_1 \cdot \vec{v}_2 + \vec{v}_1 \cdot \frac{(\vec{x}_1 - \vec{x}_2)}{|\vec{x}_1 - \vec{x}_2|} v_2 \cdot \frac{(\vec{x}_1 - \vec{x}_2)}{|\vec{x}_1 - \vec{x}_2|} \right]$$

↗
 ok only for trivial
 theory

Now reproduce from EFT:

$$E[J] = R + iI \rightarrow ? \rightarrow \text{DECAY/RADIATION}$$

$$\frac{1}{k^2 + i\epsilon} \rightarrow \frac{1}{k^2} + i\pi \delta(k^2)$$

$$\Rightarrow k_0 = |\vec{k}| \xrightarrow{\text{ON SHELL}} \text{RADIATES!}$$

(3)
EFT

Consider

$$A_\mu = \bar{A}_\mu + A_\mu^P$$

↓ radiation mode ↓ potential
 $\vec{k}^0 \sim |\vec{k}|$ $k^0 \ll |\vec{k}|$
 $\vec{k}_\text{pot}^0 = \left(k^0 \sim \frac{1}{r} \sqrt{n^{n+1}}, \vec{k} \sim \frac{1}{r} \right)$
 $\vec{k}_\text{rad}^0 = \left(\frac{\vec{v}}{r}, \frac{\vec{v}}{r} \right)$

$$\mathcal{L}[A \rightarrow \bar{A} + A^P] = -\frac{1}{2} [F[\bar{A} + A_P]^2] + J \cdot (\bar{A} + A^P)$$

→ Prime EFT Directive: every term in action must scale homogeneously in power containing param.

• How do fields scale?

• How do deriv of fields scale?

$$\rightarrow \langle A_P A_P \rangle = \int d^4 k \frac{e^{ik \cdot (x_1 - x_2)}}{k^2 - \vec{k}^2} \Rightarrow A_P \sim v^{1/2}$$

↓
 $d^4 k \sim \frac{d^3 k}{V^4}$

$$\langle \bar{A}(x_2) \bar{A}(x_1) \rangle \sim \int \frac{d^4 k}{V^4} e^{ik \cdot (x_1 - x_2)} \Rightarrow \bar{A} \sim v$$

$$\rightarrow [L^0_{[\bar{A} + A^P]}] = \int d^4 x \frac{1}{2} \bar{A}_\mu \square \bar{A}^\mu + \int d^4 x \frac{1}{2} \left(A_\mu^P \left(\partial_0^2 - \vec{\partial}^2 \right) A^{\mu P} \right) +$$

$$+ \sum_i q_i \int dt_i A_0^P(x_i(t)) + \sqrt{\bar{A}}$$

Now compute LO potential:

EFT

$$x_1(t) \times \rightarrow -\frac{i}{\hbar} \vec{p}_2 \text{ from } -\vec{p}^2$$

Coulomb potential!

Now consider $L^{O(v^2)} = \frac{1}{2} \int d^4x A_P^\mu \partial_\mu A_P^\nu + \sum_i q_i \int \vec{v} \cdot \vec{A}_P(x_i(t)) dt$

$$\Rightarrow \int \frac{d^4k}{k^4} (k_0)^2 e^{ik(r-x_1-x_2)} O(v^2)$$

We did everything for A_P , now?

$$\int \partial \bar{A} \partial A_P e^{iS[\bar{A}, \bar{A}_P]} + \int J(A + \bar{A}_P) =$$

$$= \int \partial \bar{A} e^{iV(x_1 - x_2)T} + [T] + \int T, \bar{A}]$$

From $\textcircled{1}$ $\textcircled{2} \rightarrow \xrightarrow{r} \bullet$ (Bound state: we eliminated short dist loops!)

How does radiation couple to the BS:

$$\int J \cdot A = \int dt (q_1 \bar{A}(x_1) \cdot v_1 + q_2 \bar{A}(x_2) \cdot v_2)$$

\rightarrow we removed $x_1 - x_2$ from the syst \rightarrow we want the theory of 1 single particle.

\rightarrow MULTIPLE EXPANSION
in CM frame

$$\int v^\mu \bar{A}_\mu = \int v^\mu \bar{A}_0 - \vec{v} \cdot \vec{\bar{A}} dt \approx \int v^\mu \bar{A}(0, t) - v^\mu \vec{x} \cdot \vec{\partial} \bar{A} - \vec{v} \cdot \vec{\bar{A}}$$

NR limit: $v^\mu = 1$:

$$L[\vec{A}] = \sum_i q_i \int dt A^i(0,t) - \sum_i q_i \int dt \vec{x}_i \cdot \vec{\partial} A^i(0,t) - \underset{\text{EFT}}{\underset{(5)}{\dots}}$$

$$\begin{aligned} \sum_i q_i &= Q \underbrace{\int A^i(0,t) dt}_{\text{MONOPOLE}} - \sum_i q_i \int dt [\vec{x}_i \cdot \vec{\partial} A^i(0,t) + \vec{x} \cdot \vec{\partial} A^i(0,t)] \\ &= Q \int A^i dt + \int dt (q_1 \vec{x}_1 \cdot \vec{E} + q_2 \vec{x}_2 \cdot \vec{E}) + \dots \\ &\quad \underbrace{\int dt \vec{p}(t) \cdot \vec{E}}_{\text{DIPOLE!}} \end{aligned}$$

Power Loss: $\langle 0 | \dot{k} \rangle_J = \dots$ $\partial_i A_0 - \partial_0 A_i$

$$\Rightarrow \sum_{p \in e} |\dot{M}|^2 = \sum_{p \in e} |\dots|^2 =$$

$$\Rightarrow = \sum_{p \in e} \left| \int dt \langle 0 | \vec{p}(t) \cdot \vec{E} | k \rangle \right|^2 = \sum_{p \in e} \int dt dt' p_i(t) \cdot p_j(t') \langle 0 | E_i(t) | k \rangle \cdot \langle 0 | E_j(t) | k \rangle$$

$$\sum \epsilon_\mu \epsilon_\nu^* = -g_{\mu\nu}$$

$$= \sum_{p \in e} \int dt dt' p_i(t) p_j(t') e^{ik_0 t} e^{-ik_0 t'} (k_0 \epsilon_i - k_i \epsilon_0) (k_0 \epsilon_j^* - k_j \epsilon_0)$$

$$= - \int dt dt' p_i(t) p_j(t') e^{ik_0(t-t')} (-k_0^2 \delta_{ij} + k_i k_j)$$

$$\Rightarrow \sum_{p \in e} \int \frac{d^3 k}{(2\pi)^3 2k} |\dot{M}|^2 = - \int dt dt' p_i(t) p_j(t') \int \frac{d^3 k}{(2\pi)^3 2k} \left(\frac{1}{2} \vec{k}^2 + \frac{1}{3} \delta_{ij} \vec{k}^2 \right)$$

$$= \frac{2}{3} \int dt dt' p_i(t) p_j(t') \frac{4\pi}{(2\pi)^3} \frac{1}{2} \int dk k^3$$

mult by γ energy

$$P = \frac{1}{T} \int_0^\infty dk k^4 \frac{1}{3} \frac{1}{2\pi^2} \int dt dt' e^{-i\vec{k}(t-t')} \vec{P}(t) \cdot \vec{p}(t')$$

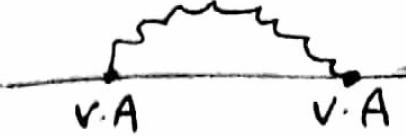
$$= \frac{1}{8} \frac{1}{2\pi^2} \int dt dt' \vec{\ddot{p}}(t) \cdot \vec{\ddot{p}}(t') \int_{-\infty}^{+\infty} dk e^{ik(t-t')}$$

$$= \frac{1}{6\pi} \int dt \vec{\ddot{p}}(t) \cdot \vec{\dot{p}}(t)$$

$$\Rightarrow \bar{P}_{\text{avg}} = \frac{1}{6\pi} \langle \vec{\dot{p}} \cdot \vec{\dot{p}} \rangle$$

Radiation reaction force

↪ IN-W formulation



1405.5532
YM as comet | ① EFT

→ causal propagator

nothing off shell (at LO: $G(x-y)$ instead of $G^F(x-y)$)

$$\rightarrow iM = \int dz dz' (-q^2) V_\mu(z) V_\nu(z') \underbrace{\langle A^\mu(x(z)) A^\nu(x(z')) \rangle}_{\downarrow}$$

$$\ddot{x}^\mu = \frac{e^2}{4\pi} \left[\frac{2}{3} \dot{x}^\mu - \dot{x}^\mu (\dot{x})^2 \right] \quad \text{A.D.L.} \quad \frac{1}{2\pi} \delta((x(z) - x(z'))^2) \cdot$$

(as small perturbation)

BH and binary inspiral

- EFT of binary inspiral (NRGR: "non relativ. GR")

1) p.p. approx: $\mathcal{S} = -m \int d\tau + \sqrt{\frac{dx^\mu}{d\lambda} \frac{dx^\nu}{d\lambda}} g_{\mu\nu} d\lambda$

$$+ \int C_1 R(g) d\tau + \int C_2 \frac{V_\mu V_\nu}{\sqrt{g}} R^{\mu\nu}(g) d\tau$$

everything with symme

they do not contribute:

$$R^{\mu\nu} = 0 \text{ in vacuum}$$

→ LO finite size effect:

$$C_E \int E^2 d\tau + C_B \int B^2 d\tau \xrightarrow{\text{on-shell}}$$

$C_{\mu\nu\sigma}^{\text{Weyl}}$: Weyl tensor

$$\text{where } E_{\mu\nu} = C_{\mu\nu\sigma} V^\sigma V^\nu \equiv R_{\mu\nu\sigma} V^\sigma V^\nu$$

$$B_{\mu\nu} = \frac{1}{2} * C_{\mu\nu\sigma} V^\sigma V^\nu \equiv \frac{1}{2} * R_{\mu\nu\sigma} V^\sigma V^\nu$$

→ order V^0 {

$$NB: * C_{\rho\sigma\mu\nu} = \epsilon_{\rho\sigma\alpha\beta} C^{\alpha\beta}_{\mu\nu}$$

$\Rightarrow v$ expands \rightarrow post Newtonian expansion (PN)

$$v^{2m} \Rightarrow v^0: 5\text{-PN}$$

Power counting

$$M = M_1 \sim M_2, \quad r, \quad G_N \sim \frac{1}{M_{pl}^2}$$

\rightarrow Expansion parameters: v, L^{-1}

$$\Rightarrow \text{Virial theorem: } \frac{1}{2} mv^2 - G \frac{m^2}{r} \Rightarrow v^2 \sim \frac{m}{r M_{pl}^2}$$

$$L \sim mvr \Rightarrow vL \sim \frac{M^2}{M_{pl}^2}$$

$$dr \sim \frac{r}{v}$$

$$Md\tau - \frac{Mr}{v} \sim \frac{L}{v^2}$$

Therefore

$$L_0 \text{ action} \sim L$$

$$\rightarrow L[g] = L[\bar{h} + H]$$

radiation potential \rightarrow to be integrated out

\rightarrow work in background field gauge:

$$S_{GF} = M_{pl}^2 \int d^4x \sqrt{h} T^\mu T_\mu$$

$$T_\mu = \partial_\alpha H^\alpha_\mu - \frac{1}{2} \partial_\mu H^\alpha_\alpha$$

$$\hookrightarrow \vec{\partial} H \sim \frac{1}{r}$$

$$\partial_\theta H \sim \frac{v}{r}$$

$$\Rightarrow \langle HH \rangle \sim \int \frac{d^4 k}{k^2} \sim v \rightarrow H \sim \frac{\sqrt{v}}{r}$$

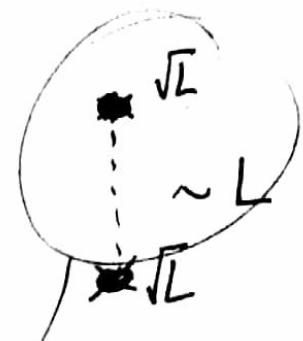
(3)
EFT

Coupling radiation - worldline: $g_{\mu\nu} = \eta_{\mu\nu} + \frac{H_{\mu\nu}}{M_{pl}}$

$$-H d\tau = H \left[1 - \vec{v}^2 + \frac{H_{00}}{M_{pl}} - 2 \frac{H_{0i} v^i}{M_{pl}} + \frac{v^i v^j H_{ij}}{M_{pl}} \right]^{1/2}$$

$$\rightarrow L = -\frac{H}{2M_{pl}} H_{00} - \frac{H}{M_{pl}} v^i H_{0i} - \frac{H}{4M_{pl}} v^2 H_{00} - \frac{H}{2M_{pl}} v^i v^j H_{ij} - \frac{H}{M_{pl}^2} H_{00}^2 + \dots$$

* $\hookrightarrow L^\circ = -\frac{H}{2M_{pl}} H_{00} \sim \sqrt{vL} \frac{\sqrt{v}}{r}$
 $\Rightarrow S^\circ \sim \int L^\circ d\tau \sim \sqrt{L}$



potential: $H = \left(-\frac{M_1}{2M_{pl}} \right) \left(-\frac{M_2}{2M_{pl}} \right) \cdot \int \frac{[d^4 k]}{(-k^2)} (-1) P_{\infty, \infty} e^{ik(r_i - r_f)}$
 $P_{\mu\nu; \alpha\beta} = \frac{1}{2} \left\{ \eta_{\mu\alpha} \eta_{\nu\beta} + \eta_{\mu\beta} \eta_{\nu\alpha} - \frac{2}{d-2} \eta_{\mu\nu} \eta_{\alpha\beta} \right\}$

\Rightarrow Newton's Law

* $\hookrightarrow L_1 = -\frac{H}{M_{pl}} v^i H_{0i}$ ⊗ ④ ○
* $\hookrightarrow L_2 = -\frac{H}{4M_{pl}} v^2 H_{00} - \frac{H}{2M_{pl}} v^i v^j H_{ij} - \frac{H}{8} \frac{H_{00}^2}{M_{pl}^2}$

\rightarrow 1PN Potential (E.I.H.)



propagator correct
 $\frac{1}{k_0^2 - k^2} \rightarrow \frac{1}{k^2} + \frac{k_0^2}{k^4}$

$\hookrightarrow S = S_M + S_{\text{BULK}} \rightarrow M_{pl}^2 \int \sqrt{g} R d^4 x$

$\Rightarrow 3\text{-gravitons: } \sim \int d^4 x \ H^3 \ j^2 \sim \frac{v^2}{\sqrt{L}} \rightarrow$ contributes at $\mathcal{O}(v^2)$

→ All in all the potential becomes:

$$V_{\text{EIH}}^{(\text{PN})} = - \frac{GM_1M_2}{2r} \left[3(v_1^2 + v_2^2) - 7v_1 \cdot v_2 - \frac{(v_1 \cdot r)(v_2 \cdot r)}{r^2} \right] + \\ + G_N^2 \frac{M_1M_2(M_1 + M_2)}{2r^2}$$

NB: matching calc lead to
 $C_E = C_B = 0 \rightarrow$ symmetries???

GRAVITATIONAL RADIATION

$$L = V(r_1, v_1, v_2) + \int \sqrt{g} R(h) + \\ + \int [Q_{ab} E^{ab} - \frac{1}{3} T^{ab} B_{ab} + \dots]$$

↑
mass quadrupole ↑
current quadrupole

$$\rightarrow \text{power law formula: } P = \frac{G_N}{8} \langle \ddot{\tilde{Q}} \dot{\tilde{Q}} \rangle$$

→ Tree level:

$$- \frac{1}{2M_{pl}} \int d^4x T_{\mu\nu} \left[h^{\mu\nu}(0) + x^\rho \partial_\rho h^{\mu\nu}(0) + \frac{1}{2} x^\rho x^\lambda \partial_\rho \partial_\lambda h^{\mu\nu}(0) + \dots \right]$$

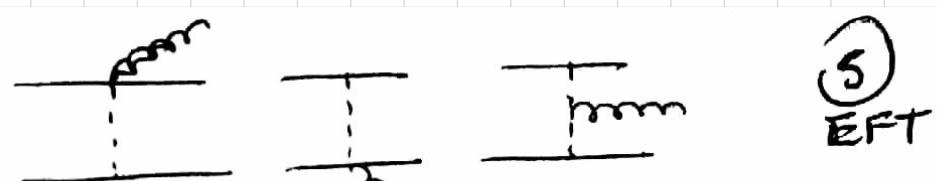
$T_{00} \sim V^0$ LO
 $T_{0i} \sim V$
 $T_{ij} \sim V^2$

$$\Rightarrow - \frac{1}{2M_{pl}} \int d^4x \left[\cancel{T_{00}} + \cancel{2T_{0i}} \left(\overset{M}{T}{}^0{}_0 h_{00} + 2 \overset{P_i}{T}{}^{0i} h_{0i} \right) - \right.$$

$P_i = 0$ in rest frame

$$\left. - \frac{1}{2M_{pl}} [\partial_i h_{0k} - \partial_k h_{0i}] \underbrace{\left(T^{0i} x^k - T^{0k} x^i \right)}_{L^{ik}} + \dots \right]$$

→ They do not radiate higher order terms

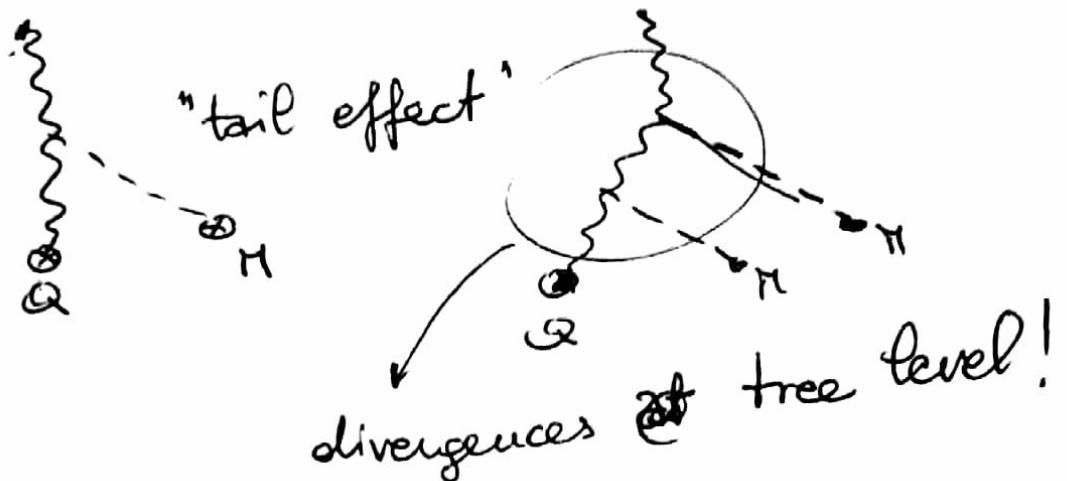


(5)
EFT

$$H \rightarrow (H_1 + H_2) + \frac{1}{2} H_1 v_1^2 + \frac{1}{2} H_2 v_2^2 - G \frac{H_1 H_2}{r}$$

→ LO power:

$$\otimes \text{~~~~~} Q \cdot E$$



Dimensional regularization: $\omega^2 G_N^2 m^2 \frac{107}{105^\epsilon} \left[\frac{1}{\epsilon} - \ln \frac{k^2}{\mu^2} \right]$

$$\Rightarrow \mu \frac{d}{\mu} Q_{ij}(\mu, \omega) = -2(\omega G_N m)^2 \frac{107}{105^\epsilon} \delta^{ij}$$

RGE function for quadrupole $Q_{ij}(\mu, \omega) = \left(\frac{\mu}{\mu_0} \right)^{-\frac{214}{105}} (\omega G_N m)^2 Q_{ij}(\mu_0, \omega)$

