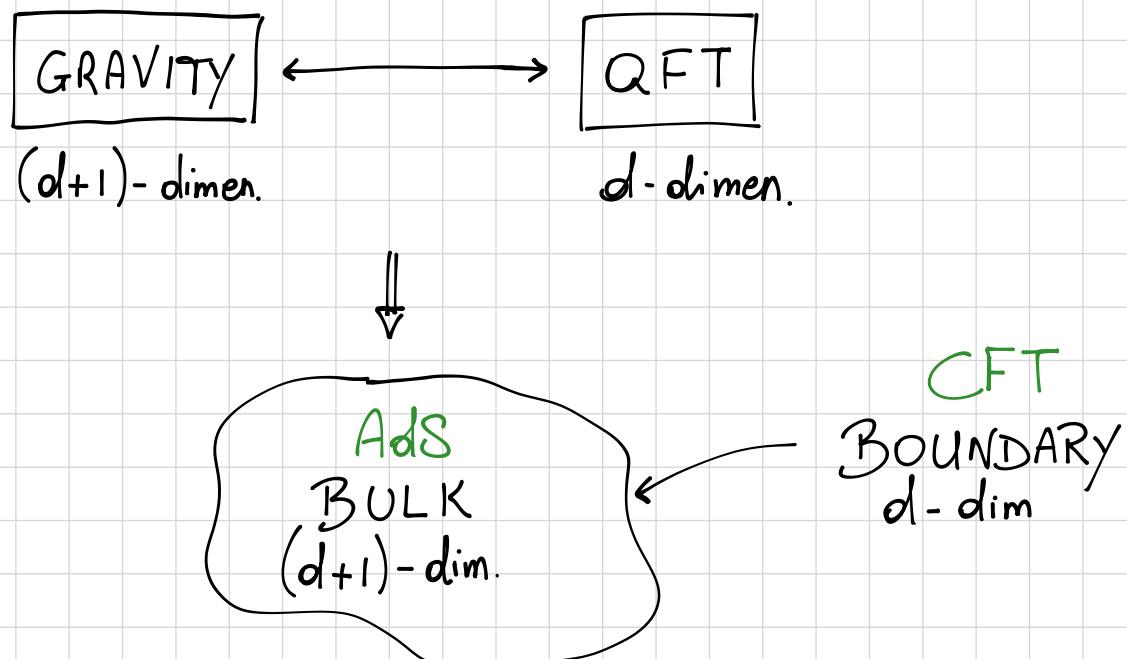


• INTRODUCTION To AdS / CFT CORRESPONDENCE [D. Martelli]

1997 Maldacena:



If SUSY is present \Rightarrow SCFT on boundary \leftrightarrow SUGRA in bulk

↓
as a limit of string theory

GRAVITY

$$S_{EH} = \frac{1}{2K_D^2} \int d^D x \underbrace{\sqrt{-g}}_e R \quad \text{where } R = g^{\mu\nu} R_{\mu\nu}, \quad R_{\mu\nu} = R_{\mu}^{\lambda}{}_{\nu\lambda}$$

$$2K_D^2 = 16\pi G_D$$

* FERMIONS \rightarrow Orthonormal frame:

$$g_{\mu\nu} = e_{\mu}^a e_{\nu}^b \eta_{ab} \quad [e_{\mu}^a := \text{"vielbein"}]$$

$$\Rightarrow \text{"spin connection": } \omega_{\mu}^{ab} = e_{\nu}^a \Gamma_{\sigma\mu}^{\nu} e^{\sigma b} + e_{\nu}^a \partial_{\mu} e^{\nu b}$$

$$\begin{aligned} \text{"Curvature 2-form": } R_{\mu\nu}^{ab} &= \partial_{\mu} \omega_{\nu}^{ab} - \partial_{\nu} \omega_{\mu}^{ab} + \omega_{\mu}^{ac} \omega_{\nu}^b{}_c - \omega_{\nu}^{ac} \omega_{\mu}^b{}_c \\ &\downarrow \\ R &= e_{a}^{\mu} e_{b}^{\nu} R_{\mu\nu}^{ab} \end{aligned}$$

$$\hookrightarrow \text{RARITA-SCHWINGER ACTION: } S_{RS} = \frac{1}{2K_D^2} \int d^D x e \bar{\psi}_{\mu} \gamma^{\mu\rho} D_{\rho} \psi_{\mu}$$

where ψ_{μ} : gravitino;

$$\gamma^{\mu\rho} = \gamma^{\mu}[\gamma^{\nu}\gamma^{\rho}]$$

$$D_{\rho} \psi_{\mu} = \partial_{\rho} \psi_{\mu} + \frac{1}{4} \omega_{\rho}^{ab} \gamma_{ab} \psi_{\mu}.$$

$\Rightarrow S = S_{EH} + S_{RS} \rightarrow$ invariant under local SUSY transf.

$$\downarrow$$

$$\varepsilon \rightarrow \varepsilon(x)$$

s.t.

$$\delta e_\mu^a = \frac{1}{2} \bar{\varepsilon} \gamma^a \psi_\mu$$

$$\delta \psi_\mu = D_\mu \varepsilon = \partial_\mu \varepsilon + \frac{1}{4} \omega_\mu^{ab} \gamma_{ab} \varepsilon$$

$$\Rightarrow \boxed{\delta S = 0}$$

$$\rightarrow S_{EH} = \frac{1}{2K_D^2} \int d^D x e e^\mu e^\nu R_{\mu\nu} \xrightarrow{\text{Ricci}} \delta S_{EH} = \frac{1}{2K_D^2} \int d^D x \left[- \bar{\varepsilon} \gamma^\mu \psi^a e^\nu \gamma^\rho + e \frac{1}{2} \bar{\varepsilon} \gamma^\mu \psi_\rho \right]$$

$$= \frac{1}{2K_D^2} \int d^D x e \left[\frac{1}{2} R g_{\mu\nu} - R_{\mu\nu} \right] \bar{\varepsilon} \gamma^\mu \psi^\nu$$

$$\rightarrow \delta S_{RS} = \frac{1}{2K_D^2} \cdot \frac{1}{2} \int d^D x e \bar{\psi}_\mu \gamma^{\mu\rho} R_{\nu\rho}^{ab} \gamma_{ab} \varepsilon = \dots = - \delta S_{EH}$$

$$\delta(D_\nu \psi_\rho) = D_\nu \delta \psi_\rho + \text{cubic in } \psi_\mu$$

\Rightarrow if we keep higher orders we have to study case by case!
 \hookrightarrow add more fields!

$$\rightarrow EOM: \begin{cases} R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = T_{\mu\nu}(\psi_\mu) \\ \gamma^{\mu\rho} D_\nu \psi_\rho = 0 \end{cases}$$

\Rightarrow set fermions = 0: $\psi_\mu = 0 \Rightarrow R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 0 \Leftrightarrow R_{\mu\nu} = 0$ "Ricci-FLAT"

These sol. are connected to SUSY solut:

$$\begin{aligned} \delta e_\mu^a &= 0 \quad \checkmark \quad \xrightarrow{\text{Killing Spinors / Kill spin. eqn.}} \\ \delta \psi_\mu &= \boxed{D_\mu \varepsilon = 0} \quad \checkmark \quad \left(D_\mu = \partial_\mu + \frac{1}{4} \omega_\mu^{ab} \gamma_{ab} \right) \\ &\downarrow \\ &\text{non trivial eqn for: 1) } g_{\mu\nu} \quad 2) \varepsilon \end{aligned}$$

Solutions: $\mathcal{E} = \text{const}$ (s.t. $\partial_\mu \mathcal{E} = 0$)

$$g_{\mu\nu} = \eta_{\mu\nu} \Rightarrow \omega_\mu^{\alpha\beta} = 0 \Rightarrow D_\mu \mathcal{E} = 0$$

NB $R_{\mu\nu} = 0$; $D_\mu \mathcal{E} = 0 \Rightarrow \text{"SPECIAL HOLONOMY MANIFOLDS"}$

$\hookrightarrow \text{SUGRA/SUSY} \longleftrightarrow \text{GEOMETRY}$

Now add a COSMOLOGICAL CONST.:

$$S = \frac{1}{2K^2} \int d^D x e \left[R - \bar{\psi}_\mu \gamma^{\mu\nu\rho} D_\nu \psi_\rho + \frac{(D-1)(D-2)}{L^2} - \frac{(D-2)}{2L} \bar{\psi}_\mu \gamma^{\mu\nu} \psi_\nu \right]$$

$\hookrightarrow \delta e_\mu^\alpha = \text{UNCHANGED}$

$\delta \psi_\mu = \hat{D}_\mu \mathcal{E} = D_\mu \mathcal{E} - \frac{1}{2L} \gamma_\mu \mathcal{E}$

cannot be absorbed in ω_μ (only one γ_μ)

$$\Rightarrow \text{e.o.m.: } R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = \frac{1}{2L^2} (D-1)(D-2) g_{\mu\nu}$$

where we took a background s.t. $\psi_\mu \equiv 0$.

Now take the trace and simplify: $R_{\mu\nu} = -\frac{D-1}{L^2} g_{\mu\nu}$
 $\Rightarrow \text{"EINSTEIN SPACES"}$

Further trace: $R = -\frac{D(D-1)}{L^2} < 0 \Rightarrow \text{NEGATIVE CURVATURE!}$

We now have $D_\mu \mathcal{E} = \frac{1}{2L} \gamma_\mu \mathcal{E}$ (which $\Rightarrow R_{\mu\nu} = -\frac{(D-1)}{L^2} \phi_{\mu\nu}$)

ANTI de SITTER (AdS_D)

AdS_D is a solution to (maximally (super)symm.)

$$\left\{ \begin{array}{l} R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = \frac{1}{2L^2} (D-1)(D-2) g_{\mu\nu} \\ D_\mu \epsilon = \frac{1}{2L} \gamma_\mu \epsilon \end{array} \right.$$

$$\Rightarrow R_{\mu\nu\rho\sigma} = -\frac{1}{L^2} (g_{\mu\rho}g_{\nu\sigma} - g_{\mu\sigma}g_{\nu\rho}) \quad (\text{MAX SYMM})$$

	$\Lambda < 0$	$\Lambda > 0$
Lorentzian	AdS_D	dS_D
Euclidean	H_D ↳ "Hyperbolic Spaces"	S^D

Take $\mathbb{R}^{2, D-1} \rightarrow X_0, X_D, X_i \quad i = 1, \dots, D-1$

$$\Rightarrow X_0^2 + X_D^2 - \sum_{i=1}^{D-1} X_i^2 = L^2 \quad \begin{matrix} \nearrow \text{RADIUS OF AdS} \\ \text{(even though it's not compact)} \end{matrix}$$

$$\hookrightarrow ds^2(\mathbb{R}^{2, D-1}) = -dX_0^2 - dX_D^2 + \sum_{i=1}^{D-1} dX_i^2$$

Isometry group $O(2, D-1)$.

Introduce

"Global AdS": $X_0 = L \cosh \rho \cos \tau$ $\rho \in \mathbb{R}^+$
 $X_D = L \cosh \rho \sin \tau$ $\tau \in [0, 2\pi]$
 $X_i = L \sinh \rho \omega_i$ s.t. $\sum_i \omega_i^2 = 1$

↓
coord on
 S^{D-2}

$$\Rightarrow ds^2(\text{AdS}_D) = L^2 \left(-\cosh^2 \rho \frac{d\tau^2}{\downarrow} + d\rho^2 + \sinh^2 \rho d\Omega_{D-2}^2 \right)$$

"universal cover": $\tau \in [0, 2\pi] \rightarrow \mathbb{R}$.

$$NB: \rho \rightarrow 0 \Rightarrow \mathbb{R}_\tau \times \mathbb{R}^{D-1}$$

$\rho \rightarrow \infty \Rightarrow \text{"CONFORMAL BOUNDARY"}$

Now introduce Poincaré coord:

$$X_0 = \frac{1}{2r} (1 + r^2(L^2 + \vec{x}^2 - t^2))$$

$$X_D = L r t$$

$$X_i = L r x_i \quad i = 1, \dots, D-2$$

$$X_{D-1} = \frac{1}{2r} (1 - r^2(L^2 - \vec{x}^2 + t^2))$$

$$\hookrightarrow ds^2 = L^2 \left(\frac{dr^2}{r^2} + r^2 \underbrace{(-dt^2 + d\vec{x}^2)}_{dx^\mu dx_\mu} \right)$$

Minkowski

$\Rightarrow r \rightarrow \infty : \text{"CONFORMAL BOUNDARY"} \quad (ds^2 \sim L^2 r^2 dx^\mu dx_\mu)$

$r \rightarrow 0 : \text{HORIZON}$

Other coords: $(r = \frac{1}{z})$

$$ds^2 = L^2 \frac{dz^2 + dx^\mu dx_\mu}{z^2} \Rightarrow z \rightarrow 0 : \text{"CONF. BOUND"}$$

and $(r = e^y)$

$$ds^2 = L^2 (dy^2 + e^{2y} dx^\mu dx_\mu)$$

MINIMAL GAUGED SUGRA IN $D=4$

\Rightarrow FIELD CONTENT:

$$\text{Majorana} \Rightarrow \psi_\mu = \psi_\mu^1 + i\psi_\mu^2 \quad (\text{Dirac})$$

$$e_\mu^a \text{ (graviton)} + \psi_\mu^1 \text{ (gravitino)} + A_\mu, \psi_\mu^2$$

$\Rightarrow A_\mu$ gauging on $SO(2) \sim U(1)$ R-Symmetry

Change notation: $\mathcal{D}_\mu = \nabla_\mu - \frac{i}{L} A_\mu = \partial_\mu + \frac{1}{4} \omega_\mu^{ab} \gamma_{ab} - \frac{i}{L} A_\mu$

$$\mathcal{S} = \frac{1}{2K^2} \int d^4x e \left[R - 2 \bar{\psi}_\mu \gamma^{\mu\nu\rho} \mathcal{D}_\nu \psi_\rho + \frac{2}{L} \bar{\psi}_\mu \gamma^{\mu\nu} \psi_\nu + \frac{6}{L^2} - F_{\mu\nu} F^{\mu\nu} + \frac{i}{4} F^{\mu\nu} \bar{\psi}_\mu \gamma_{[\mu} \gamma^{\rho\sigma} \gamma_{\nu]} \psi_\sigma \right]$$

"Einstein-Maxwell" (+ Λ)

$$\Rightarrow \delta e_\mu^a = \text{Re} [\bar{\epsilon} \gamma^\mu \psi_\mu]$$

$$\delta \psi_\mu = \hat{\mathcal{D}}_\mu \epsilon \quad \left(\hat{\mathcal{D}}_\mu = \nabla_\mu + \frac{1}{2L} \gamma_\mu + \frac{i}{4} F_{\nu\rho} \gamma^{\nu\rho} \gamma_\mu \right)$$

$$\delta A_\mu = \text{Im} [\bar{\epsilon} \psi_\mu]$$

\Rightarrow KSE: $\hat{\mathcal{D}}_\mu \epsilon = 0 \Rightarrow$ eqn for $\phi_{\mu\nu}$, A_μ and ϵ

e.o.m.: Maxwell: $d * F = 0$

$$\text{Einstein: } R_{\mu\nu} + \frac{3}{L^2} g_{\mu\nu} = 2 \left(F_\mu^\rho F_{\nu\rho} - \frac{1}{4} F^2 g_{\mu\nu} \right)$$

$(F_{\mu\nu} = 0 \rightarrow \text{AdS}_4 \text{ is the solution})$

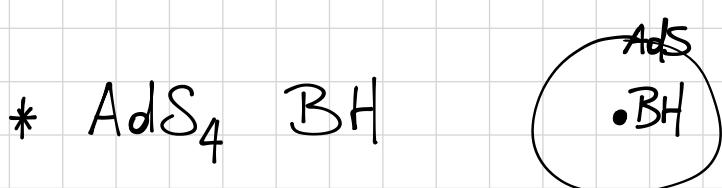
We are interested in solutions that at least approach AdS ASYMPTOTICALLY.

$$\hat{D}_\mu \mathcal{E} = 0 \quad \xrightarrow{\text{Killing vec.}} \quad \text{e.g.: } V_\mu = i \bar{\mathcal{E}} \gamma_\mu \mathcal{E} \rightarrow \text{real vector} \quad \text{s.t.} \quad \begin{aligned} 1) \quad & \nabla_{(\mu} V_{\nu)} = 0 = \mathcal{L}_V g_{\mu\nu} \\ 2) \quad & \mathcal{L}_V F_{\mu\nu} = 0 \end{aligned}$$

Simple ex of SUSY sol:

* "SELF-DUAL" sol.: $*F = -F \Rightarrow R_{\mu\nu} = -\frac{3}{L^2} g_{\mu\nu}$
 $(d*F = 0)$

→ plug into KSE: $C_{\mu\nu\rho\sigma} = \frac{1}{2} \epsilon_{\mu\nu\tau\lambda} C^{\tau\lambda}_{\rho\sigma}$
(Weyl tensor)
(QUATERNIONIC KÄHLER)



$$ds^2 = -f(r) dt^2 + \frac{1}{f(r)} dr^2 + r^2 ds^2(H^2)$$

metric on hypers. (as $ds^2(S^2)$)

$$f(r) = \left(\frac{r}{L} - \frac{L}{2r} \right)^2 \Rightarrow \text{SUSY PRESERVED}$$

Compare REISSNER-NORDSTRÖM BH: $f_{RN}(r) = 1 - \frac{2MG}{r} + \frac{q^2 G}{4\pi r^2}$

• NO SUSY (RN) vs SUSY (AdS_4)

• $r \rightarrow \infty$: $f(r) \sim \frac{r^2}{L^2}$: asymptotically AdS_4

$$ds^2 \underset{r \rightarrow \infty}{\sim} \frac{L^2}{r^2} dr^2 + r^2 \left(-\frac{dt^2}{L^2} + ds^2(H^2) \right) \quad f_{RN}(r) \sim 0 : \text{asymptotically FLAT}$$

not exactly S^2

for $\frac{q^2}{4\pi G} = M^2$
we have BPS BH
(SUSY)

The fact that $f(r)$ is SUSY \Rightarrow sth similar to BPS!

NB: $D=4$ gauged SUGRA can be embedded in $D=11$ SUGRA

$D=11$ SUGRA

Field content: e_μ^a ψ_μ (graviton + gravitino)
 Majorana

$$C_{\mu\nu\rho} \rightarrow G = dC \text{ ("flux")}$$

$$\Rightarrow S_{11} = \frac{1}{2K_{11}^2} \int d''x \left[\sqrt{-g} \left(R - \frac{1}{2} G^2 \right) - \frac{1}{8} C \wedge G \wedge G \right] + \text{[Seimions]}$$

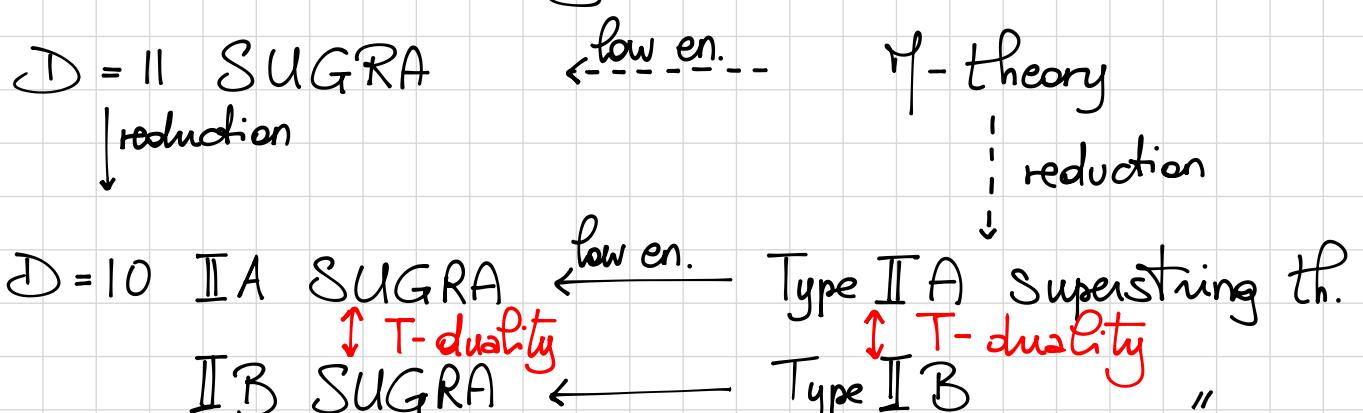
$$\hookrightarrow \text{KSE: } 0 = \delta \psi_\mu = \hat{D}_\mu \varepsilon \quad \text{where} \quad \hat{D}_\mu = \nabla_\mu + \frac{1}{288} \times \\ \times (\gamma_\mu^{\nu\rho\sigma} - 8\delta_\mu^\nu \gamma^{\rho\sigma}) G_{\nu\rho\sigma}$$

$$\text{e.o.m.: } R_{\mu\nu} = \dots$$

$$8\text{-form} \leftarrow [d * G + \frac{1}{2} \underbrace{G \wedge G}_{4+4=8\text{form}} = 0 \quad \text{"Maxwell eqn."} \right. \\ \left. \underbrace{11-4=7\text{form}}_{7+1=8\text{form}} \right]$$

$D=11$ SUGRA is usually referred to as low-energy limit of M-theory

it contains all the others through dim. reduction (KK)



Two well known solutions:

- * $\text{AdS}_4 \times S^7$ with $G \propto \text{vol}(\text{AdS}_4)$
- * $\text{AdS}_7 \times S^4$ " $G \propto \text{vol}(S^4)$

They both preserve 32 SUSY charges → maximally SUSY

⇒ Type IIA [& Type IIB] SUGRA

Reduction: $\mu = 0, \dots, 10$

$$\psi_\mu \begin{cases} \psi_i^+, \psi_i^- \text{ (Majorana Weyl GRAVITINOS)} \\ \lambda^+, \lambda^- \text{ (Majorana-Weyl DILATINOS)} \end{cases}$$

$$\begin{aligned} C_{\mu\nu\rho} &\xrightarrow{\quad} C_{ijk} \longrightarrow A_{ijk}^{(3)} \quad \text{RR field} \Rightarrow D2 \text{ (el)}, D4 \text{ (mag)} \\ C_{\mu\nu\rho} &\xrightarrow{\quad} C_{ij} \longrightarrow \beta_{ij} \quad \text{NSNS field} \longrightarrow F1 \text{ (el)}, NS5 \text{ (mag)} \\ q_{\mu\nu} &\xrightarrow{\quad} \begin{cases} q_{ij} \\ q_{\mu\mu} \end{cases} \longrightarrow e^\phi \quad \text{NSNS dilaton} \\ q_{\mu\nu} &\xrightarrow{\quad} q_{ii} \longrightarrow A_i^{(1)} \quad \text{RR field} \Rightarrow \text{coupling to } D0 \text{ (electr.)}, \\ &\qquad\qquad\qquad D6 \text{ (magn.)} \end{aligned}$$

↓
Type IIA SUGRA (= massless Type IIA Superstring)

$$NB: C^{(p+1)} \rightarrow *dC^{(p+1)} = d\tilde{C}^{(\bar{p}+1)} \rightarrow \tilde{p} = 6 - p$$

In SUGRA D-branes are sol. to e.o.m.

e.g. D2-brane solution → $A_{ijk}^{(3)}$ q_{ij} e^ϕ

What about Type IIB spectrum?

NSNS spectrum is universal ⇒ β_{ij} q_{ij} e^ϕ ✓ → F1, NS5

$$\begin{aligned} RR &\rightarrow A^{(0)} \quad A^{(2)} \quad A^{(4)} \\ &\downarrow \quad \downarrow \quad \downarrow \\ D(-1) \quad D1 \quad D3 \quad D7 \quad D5 &\longrightarrow F^{(5)} = *F^{(5)} \quad (SELF-DUAL) \\ &\qquad\qquad\qquad F^2 = 0 \end{aligned}$$

Now focus on D-brane Solutions.

Set to 0 all unnecessary fields. From string th:

IIA, IIB

$$\underbrace{S_{sf}}_{\text{string frame}} = \frac{1}{(2\pi)^7 l_s^8} \int d^{10}x \sqrt{-g} \left\{ e^{-2\phi} (R + 4 \partial_\mu \phi \partial^\mu \phi) - \frac{2}{(8-p)!} \overline{F}_{p+2}^2 \right\}$$

\downarrow to Einstein frame

$$g_{\mu\nu}^{\text{EF}} = e^{-\phi/2} g_{\mu\nu}^{\text{sf}}$$

$$l_s^2 = \alpha'$$

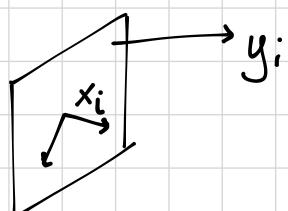
$$\text{coupling const: } q_s = \langle e^\phi \rangle$$

depends on ϕ

$$\hookrightarrow S_{\text{EF}} = \frac{1}{(2\pi)^7 l_s^8} \int d^{10}x \sqrt{-g^E} \left(R_E - \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{2}{(8-p)!} e^{ap\phi} \overline{F}_{p+2}^2 \right)$$

Derive e.o.m. and consider the Ansatz:

$$ds^2 = H_p^{-1/2}(r) (-dt^2 + dx_1^2 + \dots + dx_p^2) + H_p^{1/2}(r) (dy_1^2 + \dots + dy_{9-p}^2)$$



x_i : on the brane

y_i : orth. dir.

$$\Rightarrow r^2 = y_1^2 + \dots + y_{9-p}^2$$

$$\Rightarrow \text{Potential: } A_{t, x_1, \dots, x_p}^{(p+1)} = -\frac{1}{2} (H_p(r))^{-1} \quad \text{and} \quad e^\phi = q_s (H_p(r))^{\frac{3-p}{4}}$$

s.t. $r \rightarrow \infty \rightarrow$ recover Poincaré

$$\rightarrow \text{Taking into e.o.m.} \Rightarrow H_p(r) = \underbrace{1}_{\text{constant}} + \underbrace{\frac{L^{7-p}}{r^{7-p}}}_{\text{quant. cond.}}$$

Harmonic in transverse space

constant + quant. cond.:

$$\frac{1}{(2\pi l_s)^{p+1} q_s} \underbrace{\int d^{8-p}x * F_{p+2}}_{\text{cycle wrapping the brane at } \infty} = N \in \mathbb{N} \Rightarrow L^{7-p} = (2\pi)^{p-2} d_p l_s^{7-p} q_s N$$

no. of p-branes

... other factors of 2π ...

Remarks:

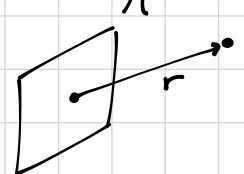
- * these are SUSY solutions (they solve the KSE) $\nearrow \frac{1}{2}$ SUSY
- * \exists non-SUSY versions called "black ϕ -branes"
- * for general ϕ they are singular as $r \rightarrow 0$
(except for $\phi = 3$)

Choose $\phi = 3$:

$$e^\phi = g_s$$

$$H_3(r) = 1 + \frac{L^4}{r^4}, \quad L^4 = 4\pi g_s (\alpha')^2 N = (2\pi)^4 \underbrace{\frac{g_s (\alpha')^2 N}{4 V_{\text{op}}(S^5)}}_{\pi^3}$$

i.e. "near brane"



Consider $\phi = 3$. The limit $r \rightarrow 0$ ("near horizon" limit)

Let

$\alpha' \rightarrow 0$ with $u = \frac{r}{\alpha'}$ FIXED (energy scale)

$\hookrightarrow G_{10} = 2^3 \pi^6 (\alpha')^4 g_s^2 \rightarrow 0$ keep g_s fixed



brane theory DECOUPLES FROM THE BULK

and α' corrections in SUGRA are suppressed.

Then we find:

$$H_3 = H \xrightarrow{\alpha' \rightarrow 0} \frac{4\pi g_s N}{u^4 (\alpha')^2} \Rightarrow ds^2 = \alpha' \left[(4\pi g_s N)^{1/2} \left(\frac{du^2}{u^2} + ds^2(S^5) \right) + \frac{u^2}{(4\pi g_s N)^{1/2}} dx^\mu dx_\mu \right]$$

NB: the N branes are in spher. coord.

all on top of each other

\Rightarrow Near brane geometry looks like $AdS_5 \times S^5$

$L = \text{radius of } AdS_5 =$
" " S^5

D3 preserves
SUSY (all 32
actually...)

The "curvature" \sim set by L

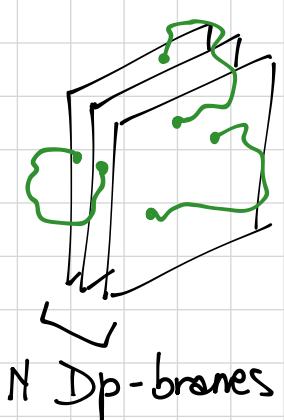
$$\hookrightarrow \text{Small curvature} \Rightarrow \frac{L^4}{\ell_s^4} \gg 1 \Rightarrow g_s N \gg 1$$

Moreover to string loop corrections $N \gg 1$ (fixed g_s)

view SUGRA from ST low en.- limit $\Rightarrow d' \rightarrow 0$ suppresses stringy corr.
 $\Rightarrow N \gg 1$ " the loop contrib.

D-BRANE SYM THEORY

1) D-brane \hookrightarrow SUGRA SOL. of CLOSED STRING
 \rightarrow EXTENDED OBJ. with (open) STRINGS ATTACHED



Minkowski $R^{1,p}$

\rightarrow low energy effective th. on N coincident Dp-branes

\downarrow
 $(p+1)$ -dim. SYM

$$G = U(N)$$

- FIELDS:

* non Abelian A_μ

* 9-p scalars in ADJ. REP of $U(N)$

* gauginos λ



SUSY fixes uniquely the action

$$\mathcal{W} = 1 \quad \text{Dp-brane} : S_{DBI} = -T_p \int d^{p+1}x e^{-\phi} \sqrt{-\det(g + 2\pi\alpha' F)}$$

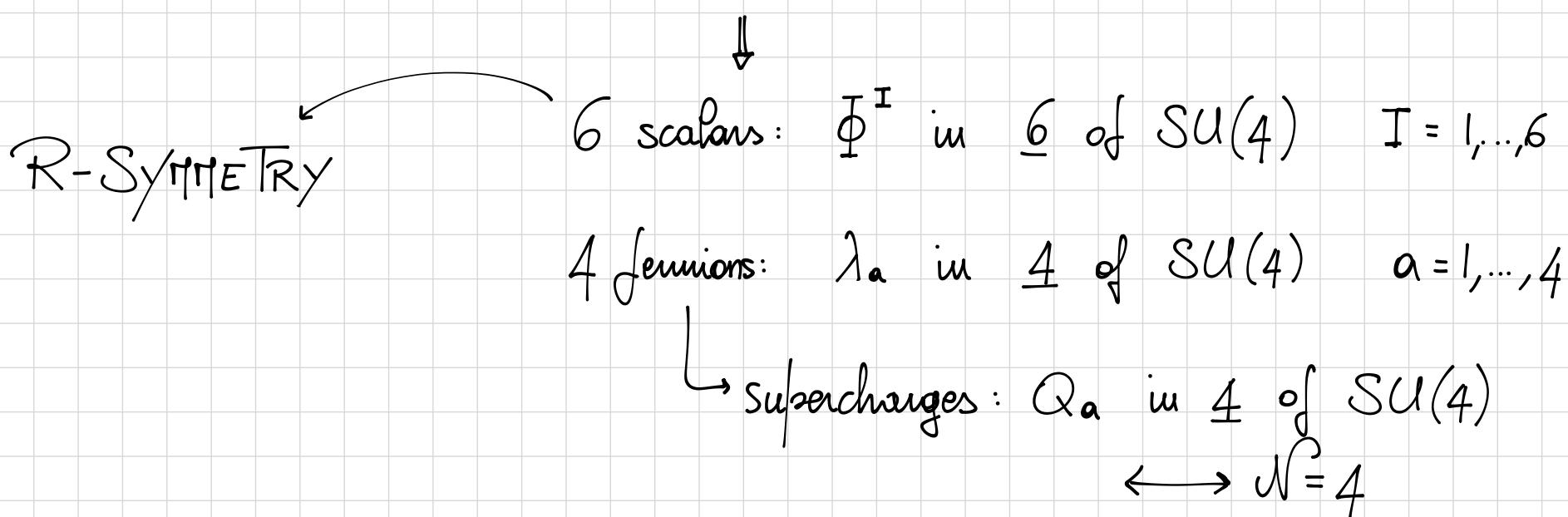
$$T_p = \frac{1}{(2\pi)^p (\alpha')^{p+1} g_s}$$

$$\text{Now take } \alpha' \rightarrow 0 \text{ with } g_s \propto (\alpha')^{\frac{p-3}{2}} \text{ fixed} \Rightarrow S \sim \frac{1}{g_{YM}^2} \int d^{p+1}x e^{-\phi} F^2 \Rightarrow g_{YM}^2 = g_s (2\pi)^{\frac{p-2}{2}} (\alpha')^{\frac{p-3}{2}}$$

For $\phi = 3 \rightarrow 4\pi g_s = g_{YM}^2$:

- A_μ in 1+3 dim
- 6 scalar
- λ
- 16 SUSY $\longrightarrow \boxed{N=4 \text{ SYM}}$

$\Rightarrow \mathcal{L}$ has manifest SU(4) GLOBAL SYMM: $so(4) \sim so(6)$



\Rightarrow FULL SYMM. \Rightarrow Poincaré \times SO(6)
of the \mathcal{L} in $p+1$ dimensions

$$\begin{array}{c} \downarrow \\ 4 P_\mu \\ 6 J_{\mu\nu} \end{array}$$

$N=4$ SYM is a $\left\{ \begin{array}{ll} 1 & \mathcal{D} \\ 4 & K_\mu \end{array} \right.$
conformal theory

15 generators $\Rightarrow SO(2,4)$

~~~ "actual" full symmetry:  $SO(2,4) \times SO(6)$  (bosonic)

$\Rightarrow$  Maldacena 1997:

The following are equivalent:

① Type IIB ST in background of  $AdS_5 \times S^5$  w/  $N$  units of RR flux and  $=$  radius of  $AdS_5$  and  $S^5$

$$L = (4\pi g_s (\alpha')^2 N)^{1/4}$$

②  $W=4$  SYM in  $d=1+3$  w/  $G = SU(N)$  s.t.

$$g_{YM}^2 = 4\pi g_s$$

This referred to as "STRONG FORM" of the correspondence.

In ST side we have

$$\alpha' \rightarrow 0$$

$$g_s N \gg 1 \quad \text{for the validity of SUGRA.}$$

$$N \gg 1$$

Define 't Hooft coupling:

$$\lambda := g_{YM}^2 N$$

where  $N \rightarrow \infty$  at fixed  $\lambda \Rightarrow$  't Hooft / "large  $N$ " limit

- curvature  $\sim \frac{1}{L^2} \sim \frac{1}{\lambda^{1/2} \alpha'}$

$$\hookrightarrow \mathcal{L}_{eff} = \frac{\alpha'}{L^2} \mathcal{L}_1 + \left(\frac{\alpha'}{L^2}\right)^2 \mathcal{L}_2 + \dots =$$

$$= \lambda^{-1/2} \mathcal{L}_1 + \lambda^{-1} \mathcal{L}_2 + \dots \Rightarrow \alpha' \rightarrow 0 \sim \lambda \rightarrow \infty.$$

## LARGE N LIMIT

Gravity  $\Rightarrow$  large no. of branes

QFT side  $\Rightarrow$  consider th. containing ADJ (of  $U(N)$  in front.)

(

$$\ln Z = \sum_{g=0}^{\infty} N^{2-2g} f_g(\lambda)$$

$\Rightarrow$  for large  $N \rightarrow g=0$  dominates:  $\ln Z \xrightarrow{N \rightarrow \infty} N^2 f_0(\lambda) =$

$$= N^2 (c_0 + c_1 \lambda + c_2 \lambda^2 + \dots)$$

$\rightarrow$  OK if  $\lambda \ll 1$  but this is not the case.

Use the "DOUBLE LINE NOTATION":

$$i \sim j := \begin{array}{c} \xleftarrow{i} \\ \xrightarrow{j} \end{array}$$

$$\sim := \begin{array}{c} \xleftarrow{i} \\ \xrightarrow{j} \end{array} \quad \begin{array}{c} \xleftarrow{i} \\ \xrightarrow{j} \end{array}$$

$$\sim := \begin{array}{c} \xleftarrow{i} \\ \xrightarrow{j} \end{array} \quad \begin{array}{c} \xleftarrow{i} \\ \xrightarrow{j} \end{array}$$

$$\Rightarrow \mathcal{L}_{\text{YM}} \sim (\text{d}A)^2 \sim (dA + g_{\text{YM}} A^2)^2 = dA^2 + g_{\text{YM}} A^2 dA + g_{\text{YM}}^2 A^4$$

$$\text{Then introduce } A \rightarrow \frac{1}{g_{\text{YM}}} A \Rightarrow \mathcal{L}_{\text{YM}} = \frac{N}{\lambda} (dA^2 + A^3 + A^4)$$

$\hookrightarrow$  in double line notation, for any loop  $\Rightarrow$  factor  $N$

Consider for example:

$$\begin{array}{c}
 \text{blob} = \text{circle with 3 loops} \\
 \rightarrow 3 \text{ loops} \rightarrow N^3 \\
 \rightarrow 3 \text{ propagators} \rightarrow \left(\frac{\lambda}{N}\right)^3 \\
 \rightarrow 2 \text{ wavy lines} \rightarrow \left(\frac{N}{\lambda}\right)^2 \\
 \Rightarrow A \sim N^2 \lambda = N^3 g_{YM}^2
 \end{array}$$

$$\begin{array}{c}
 \text{blob with hole} = \text{circle with 2 loops} \\
 \rightarrow 2 \text{ loops} \rightarrow N^2 \\
 \rightarrow 6 \text{ propagators} \rightarrow \left(\frac{\lambda}{N}\right)^6 \\
 \rightarrow 4 \text{ vertices} \rightarrow \left(\frac{N}{\lambda}\right)^4 \\
 \Rightarrow A \sim N^0 \lambda^2
 \end{array}$$

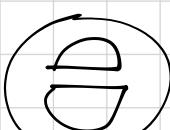
Therefore:

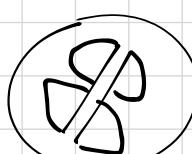
$$N^L \left(\frac{\lambda}{N}\right)^E \left(\frac{N}{\lambda}\right)^V = N^{L+V-E} \lambda^{-V+E} = N^X \lambda^{E-V}$$

Euler charact.

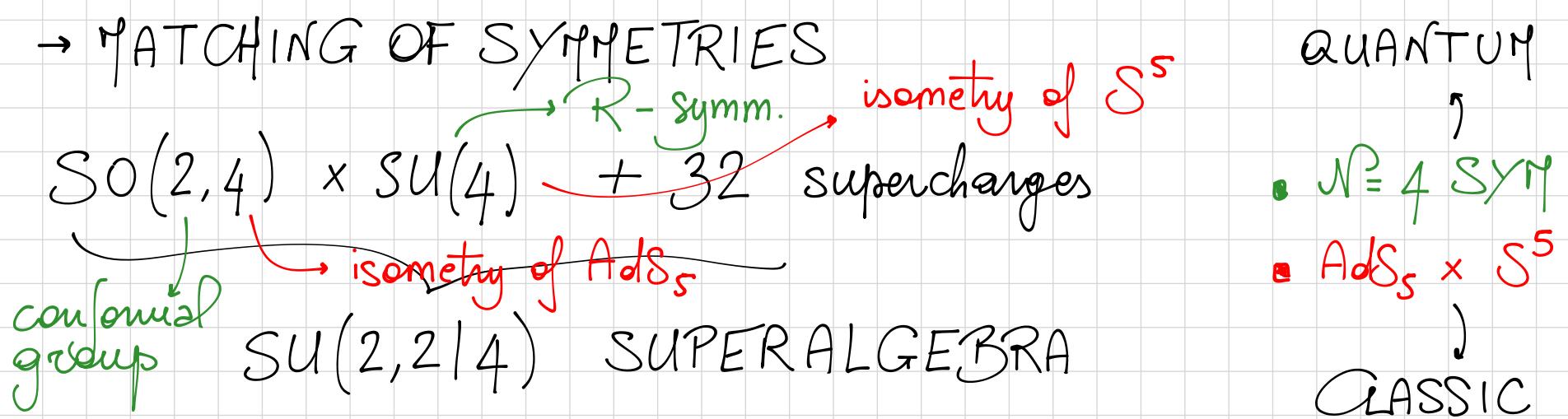
$$\text{CLOSED SURF. : } X = 2 - 2g$$

$\sum$  genus

NB:  :  $g=0 \rightarrow$  we can draw it on a sphere  
(PLANAR DIAGRAM)

  $\rightarrow g=1 \Rightarrow$  TORUS ( $X=0$ )

$\Rightarrow$  LARGE  $N \longrightarrow$  PLANAR LIMIT (strong 't Hooft coupling)



REMARKS:

- \* keep  $AdS_5$
  - replace  $S^5 \rightarrow M_5$ : we can break (some) SUSY w/ diff. symm. on QFT side
- manifold (e.g.: Sasaki-Einstein)

⇒ "PHYSICS IN THE BULK" = "PHYSICS ON THE BOUNDARY"

Therefore:

$AdS_5$ :

$g_{\mu\nu}$   
 scalars  
 etc.

I CAN'T SIMPLY  
 COMPARE THEM!  
 (on  $AdS_5$  there are no  
 gauge indices)

CFT (e.g.:  $N=4$  SYM)

$A_\mu, \phi^i, \lambda^a$  in ADJ of  $SU(N)$

↑  
 NOT gauge inv.

(not physical obs.)

We must assemble fields in gauge inv. operators. E.g.:

$$\text{Tr}(F_{\mu\nu} F^{\mu\nu})$$

$$\text{Tr}(\phi^I \phi^J \phi^K \dots)$$

$$\text{Tr}(F_{\mu\nu} F^{\mu\nu} \phi^I)$$

We need to derive from ST a SD theory  $\Rightarrow$  KK REDUCTION

$\Rightarrow$  KK SPECTRUM of TYPE IIB on  $\text{AdS}_5 \times S^5$

$$\begin{cases} x = \text{coordinates on } \text{AdS}_5 \\ y = " \quad S^5 \end{cases}$$

Generic field:

$$\phi(x, y) = \sum_n \phi_n(x) Y_n(y) \quad \text{spherical harmonics on } S^5$$

The KK spectrum (in 5 steps):

1) Write every field of Type IIB SUGRA:

$$\phi = \dot{\phi} + \delta\phi \quad \begin{array}{l} \xrightarrow{\text{FLUCTUATIONS}} \\ \downarrow \quad \text{BACKGROUND (it solves e.o.m.)} \\ \text{AdS}_5 \times S^5 \end{array}$$

2) Plug e.o.m. and linearize around  $\dot{\phi} \rightarrow$  linear eq. for  $\delta\phi$

3)  $\delta\phi = \sum (\text{harmonics})$

4) Coupled eq.'s  $\rightarrow$  DIAGONALIZE the linear system

5) COMPUTE EIGENVALUES

The symm. is  $SU(2,2|4) \rightarrow$  eigenv. will be labelled by "QUANTUM NO.s"  
 ↳ these modes are the "fields in  $AdS_5$ "

Back in  $\mathcal{N}=4$  SYM:

- focus on  $SO(2,4) \supset SO(1,1) \times SO(1,3)$

$$\begin{array}{ccc} & \downarrow & \downarrow \\ & \text{DILATATION} & \text{LORENTZ} \\ (\mathcal{D}) & & (J_{\mu\nu}) \end{array}$$

Under  $\mathcal{D}$ :

$$\begin{aligned} x^\mu &\rightarrow \omega x^\mu \\ \Theta(x) &\rightarrow \omega^{\Delta} \Theta(\omega x) \end{aligned}$$

scaling dim.

$$\hookrightarrow \left. \begin{array}{l} [\mathcal{D}, P_\mu] = -i P_\mu \\ [\mathcal{D}, K_\mu] = i K_\mu \end{array} \right\} \text{classical} \rightarrow \hat{\mathcal{D}} \Theta = -i \Delta \Theta$$

$$\mathcal{D}(P_\mu \Theta) = -i (\Delta + 1) P_\mu \Theta$$

$$\mathcal{D}(K_\mu \Theta) = -i (\Delta - 1) K_\mu \Theta$$

Unitarity requires a finite spectrum s.t.  $K_\mu \Theta_{\text{low}} = 0$ .



$$K_\mu \Theta_{\text{lowest}} = 0 \quad (\text{analogous to } \alpha_n |0\rangle, n > 0)$$



"PRIMARY OP."

(acting w/  $P_\mu$  I find "descendants")

Use SUSY  $\Rightarrow$  we have Q's:

$$[\mathcal{D}, Q] = -\frac{i}{2} Q \rightarrow \Delta(Q\Theta) = \Delta + \frac{1}{2}$$

in  $\mathcal{N}=4$  there are 16 Q's

$\Rightarrow$  "LONG REP" of superalg. ↳ they all form the same rep

→ in this case "Δ is not protected" ( $\Delta = \Delta(N, \lambda)$ )

↳ NB: SUGRA is a specific limit for  $\lambda \Rightarrow$  op. in CFT must not depend on  $\lambda$

⇒ if (comb. of Q's)  $\Theta = 0 \Rightarrow \Theta$  is called CHIRAL PRIMARY

↳ "SHORT REP."

↳ for this rep,  $\Delta$  is UNIQUELY FIXED IN TERMS OF THE R-SYMMETRY (SU(4))

⇒ R-charges  $\sim$  integers

(for  $\mathcal{N}=4$ :  $(q_1, q_2, q_3)$  Casimirs of  $SU(2)^3 \times SO(6)$ )

TOTALLY SYMM., TRACELESS OF  $SO(6)$

e.g.:

$$\Theta_n = \text{Tr} \left( \phi^{(I_1} \dots \phi^{I_n)} \right)$$

( ↳  $\Delta = n, n \geq 2$ )

Then  $Q^4 \Theta_{n+2} = \text{Tr} \left( F_{\mu\nu} F^{\mu\nu} \phi^{(I_1} \dots \phi^{I_n)} \right) \rightarrow \Delta = n+4 \quad n > 0$

↳ on the  $AdS_5$  part those correspond to scalars:

$$\left( \square_{AdS_5} - m^2 \right) \phi = 0$$

comes from KK-red  $\Rightarrow$  Casimirs of  $SO(6)$

$$\Rightarrow \Delta = 2 + \sqrt{4 + m^2 L^2} \text{ on } AdS_5$$

$$\left[ \Delta = \frac{1}{2} \left( d + \sqrt{d^2 + 4m^2 L^2} \right) \text{ on } AdS_{d+1} \right]$$

$$d^2 + 4m^2 L^2 \geq 0 \Leftrightarrow m^2 \geq -\frac{d^2}{4L^2}$$

$\Rightarrow$  BF BOUND

even the tachyon has bounds

# THE AdS/CFT "MASTER FORMULA"

⇒ must answer two questions:

- 1) What is the physical motivation to associate a field  $\phi$  in the bulk to an operator  $\mathcal{O}$  on the boundary?
  - 2) In a CFT the observables are

$$\langle \theta_1 \dots \theta_n \rangle$$

How do I relate them to the gravity side?

$$\mathcal{Z}_{\text{QFT}}[J] = \mathcal{Z}_{\text{string}}[\phi_0]$$



$$J = \phi_0$$

where

$$\mathcal{Z}_{\text{QFT}}[J] = \int \mathcal{O}[\text{dynam. fields}] e^{-S_{\text{QFT}} + \int d^4x \Theta(x) J(x)}$$

↓ ↑  
 N=4 Sym:  $A_\mu, \phi^I, \lambda^a$       composite op.s      sources

$$\Rightarrow \langle \Theta(x_1) \dots \Theta(x_L) \rangle = \frac{\delta^L}{\delta J(x_1) \dots \delta J(x_L)} Z_{QFT}[J] \Big|_{J=0}$$

On the gravity side:

$$\mathcal{Z}_{\text{string}}[\phi_0] = \int \mathcal{O}_0[\text{dyn. fields}] e^{-S_{\text{string}}}$$

$\downarrow$   
 $\phi_0 = \phi \Big|_{\partial A \cup S_S}$

Then  $Z_{QFT}[J] = Z_{\text{string}}[\phi_0] \rightarrow J = \phi_0$

We use the "WEAK FORM" (computable) :

$$\begin{aligned} N &\rightarrow \infty \\ \lambda &\sim \frac{1}{(\alpha')^2} \rightarrow \infty \end{aligned} \Rightarrow Z_{\text{string}}[\phi_0] = e^{-N^2 S_{\text{SUGRA}}[\phi_0]} + O(\alpha') \times \underset{\substack{\Downarrow \\ \text{stringy} \\ \text{correct.}}}{\text{ }} \quad \boxed{Z_{QFT}[J] = e^{-N^2 S_{\text{SUGRA}}[\phi]}}$$

We then need to compute  $S_{\text{SUGRA}}[\phi_0]$ :

- 1) solve e.o.m. for  $\phi$  w/ prescribed  $\phi|_{\partial \text{AdS}} = \phi_0$
- 2) plug  $\phi$  back into  $S_{\text{SUGRA}} \rightarrow S[\phi_0]$

Remarks:

- a)  $S$  is typically divergent  $\rightarrow$  HOLOGRAPHIC RENORMALIZATION
- b) in practice  $S_{\text{SUGRA}}$  is an effective action ( $d=5$ ) from KK reduction from Type IIB ST ( $d=10$ ).

$$\text{In } d=5 : g_{\mu\nu} = g_{\mu\nu}(\text{AdS}_5) + \delta g_{\mu\nu}$$

$$e^{-S_{\text{SUGRA}}[\delta g_{\mu\nu}]} = \langle e^{\int d^4x \sqrt{g} g_{\mu\nu} T^{\mu\nu}} \rangle$$

energy-mom. tensor  
 of the DUAL FIELD  
 THEORY

↓  
 QUANTUM OPERATOR

$$e^{-S_{\text{SUGRA}}[SA_i]} = \langle e^{\int d^4x \sqrt{g} A_i J^i} \rangle$$

$A_\mu$ : gauge field in the bulk with  $A_\mu|_{\partial \text{AdS}} = A_i$

$J_i$ : conserved current

That is  $A^\mu J_\mu \rightarrow A_\mu \rightarrow A_\mu + \partial_\mu \Lambda \Rightarrow \partial^i J_i = 0$

Consider now:

- fixed  $g_{\mu\nu}$  ( $\text{AdS}_5$ )
- fluctuations of  $\phi$  on top of  $g_{\mu\nu}$

$$\Rightarrow (\square_{\text{AdS}} - m^2) \phi = 0$$

$\xrightarrow{z=\varepsilon \rightarrow 0}$  boundary

$$\text{Choose } ds^2(\text{AdS}_5) = \frac{L^2}{z^2} (dz^2 + dx^i dx_i) \quad (\text{coord. } x^i, z, i=1, \dots, 4)$$

$\hookrightarrow$  use  $t \rightarrow -it$  to have  $dx^2$

$$\square = \frac{1}{\sqrt{g}} \partial_\nu (\sqrt{g} g^{\mu\nu} \partial_\mu) = \frac{1}{L^2} z^5 \partial_z (z^{-3} \partial_z) + \frac{z^2}{L^2} \square_{4d}$$

$\hookrightarrow \sqrt{g} = \left(\frac{L}{z}\right)^5$

$$\text{Consider: } \phi(z, x^i) = \int \frac{d^4 p}{(2\pi)^4} \phi_p(z) e^{ip \cdot x}$$

$$\Rightarrow z^5 \partial_z (z^{-3} \partial_z) \phi_p(z) - p^2 z^2 \phi_p(z) - m^2 L^2 \phi_p(z) = 0$$

$$\hookrightarrow \phi(z) = \phi_0(x) z^\alpha \text{ ansatz}$$

$$z^\alpha \phi_0 [\alpha(\alpha-4) - m^2 L^2] - z^{\alpha+2} p^2 \phi_0 = 0$$

as  $z \rightarrow 0$ , I must impose

$$\alpha(\alpha-4) - m^2 L^2 = 0 \Rightarrow \alpha_\pm = 2 \pm \sqrt{4 + m^2 L^2}$$

Then for  $z \rightarrow 0$ :

$$\phi(x, z) \sim \phi_0(x) z^{\alpha_-} + \phi_1(x) z^{\alpha_+} + \dots$$

$$\text{NB } \alpha_- - \alpha_+ < 0$$

Then  $z^{\alpha_-}$  is dominant.  $\Rightarrow \underbrace{\phi_0}_{\text{is the b.c. of } \phi(x,z)}$  is the b.c. of  $\phi(x,z)$ .

$\phi_1, \phi_2, \dots$  are specified from  $\phi_0$ .

$\rightarrow \alpha_-$  is related to  $\Delta$ :

$$\begin{aligned} x^\mu &\rightarrow \omega x^\mu \\ z &\rightarrow \omega z \end{aligned} \Rightarrow \phi \text{ is inv.}$$

$$\phi_0 \rightarrow \phi_0 \omega^{-\alpha_-} \Rightarrow \int d^4x \phi_0 \Theta_\Delta \Rightarrow \Theta_\Delta \sim \omega^{4-\alpha_-} \Theta_\Delta$$

$$\boxed{\Delta = 4 - \alpha_- = \alpha_+}$$

$$\Rightarrow \Delta = 2 + \sqrt{4 + m^2 L^2}$$

Going back to the KK spectrum  $AdS_5 \times S^5$ :

$$m^2 L^2 = k(k-4) \rightarrow k \geq 2$$

(comes from mixing modes  $\pi, b$ )

$$\Rightarrow \Delta = 2 + \sqrt{4 + k^2 - 4k} = k \Rightarrow \text{the modes have increasing } \Delta !$$

## HOLOGRAPHIC (SCALAR) TWO-POINT FUNCTION

$$\langle \Theta_\Delta(x) \Theta_\Delta(y) \rangle = ?$$

$\Rightarrow Z \sim e^{-S[\phi_0]} \rightarrow$  I need  $S$  up to  $\phi^2$  since I'm interested in two point functions.

Consider :

$$\left( \square_{AdS} - m^2 \right) \phi = 0$$

$$\text{Take } u = p\bar{z} \rightarrow u^5 \frac{d}{du} \left( u^{-3} \frac{d}{du} f(u) \right) f(u) - u^2 f(u) - m^2 L^2 f(u) = 0$$

where  $f(u) \stackrel{u=p\bar{z}}{=} \phi(\bar{z})$ ; then  $f(u) = u^2 F(u)$ :

$$u^5 \frac{d}{du} \left( u^{-3} (2uF + u^2 F^2) \right) - u^4 F - m^2 L^2 u^2 F = 0$$

$$\Leftrightarrow \text{two solutions: } \begin{cases} I_{\Delta-2}(u) \\ K_{\Delta-2}(u) \end{cases} \quad F(u) = A I_{\Delta-2}(u) + \beta K_{\Delta-2}(u)$$

$$\hookrightarrow u \rightarrow \infty (\bar{z} \rightarrow \infty : \text{Poincaré}) : I_{\Delta-2}(u) \sim \frac{e^u}{\sqrt{u}}$$

$$\Downarrow \qquad \qquad \qquad K_{\Delta-2}(u) \sim \frac{e^{-u}}{\sqrt{u}}$$

$$A = 0$$

$$u \rightarrow 0 (\bar{z} \rightarrow 0 : \text{boundary}) : K_{\Delta-2}(u) \sim \frac{\Gamma(\Delta-2)}{2} \left( \frac{2}{u} \right)^{\Delta-2}$$

$$\Rightarrow f(u) \sim u^{4-\Delta} \quad \checkmark$$

$$\text{So far: } \phi(x, \bar{z}) = \beta \cdot (p\bar{z})^2 K_{\Delta-2}(p\bar{z}) e^{ip \cdot x}$$

$$\hookrightarrow \bar{z} = \epsilon \rightarrow \phi(x, \bar{z}) = e^{ip \cdot x}$$

$$\Rightarrow \phi(x, \bar{z}) = \frac{(p\bar{z})^2 K_{\Delta-2}(p\bar{z})}{(p\epsilon)^2 K_{\Delta-2}(p\bar{z})} e^{ip \cdot x}$$

NB: the final result has  $\epsilon \rightarrow 0 \Rightarrow$  it must not depend on  $\epsilon$  !!!

Plug back the phase

$$S = C \int d^5x \sqrt{g} ((\partial\phi)^2 + m^2\phi^2) \stackrel{\text{B.P.}}{=} 0$$

$$= C \underbrace{\int d^5x \sqrt{g} (-\phi \square \phi + m^2\phi^2)}_{= 0} + C \int_{\text{boundary}} d^4x \sqrt{h} \phi n^\mu \nabla_\mu \phi =$$

$$= 0$$

$$n^\mu = k \left( \frac{\partial}{\partial x} \right)^\mu$$

s.t.  $n^\mu n^\nu g_{\mu\nu} = 1$

$$k = \frac{x}{L}$$

$$= C L^3 \int_{z=\epsilon} d^4x \frac{1}{z^3} \phi \partial_z \phi =$$

$F(\epsilon, p)$  := "Flux FACTOR"

$$= CL^3 (2\pi)^4 \delta^4(p+q) \underbrace{\left[ \frac{1}{z^3} \phi(z) \partial_z \phi(z) \right]}_{z=\epsilon=0}$$

↑ Fourier expansion

⇒ Taylor expand  $F(\epsilon p)$  for  $\epsilon \rightarrow 0$ : there are divergent terms  $\frac{1}{\epsilon^\#}$

$$\underbrace{\frac{1}{\epsilon^\#} + \dots + \text{finite}}_{\text{}} + O(\epsilon)$$

can we remove these? (Renorm.)

Now focus on the finite part. We have

$$F(\epsilon, p) \Big|_{\text{finite}} \sim p^{2\Delta-4} \rightarrow \langle \hat{\theta}(p) \tilde{\theta}(q) \rangle \sim \delta^4(p+q) p^{2\Delta-4}$$

$$\Rightarrow \langle \theta(x) \theta(y) \rangle \sim \frac{1}{|x-y|^{2\Delta}} \quad (\text{anti-Fourier transf.})$$

# CORRELATION FUNCTIONS IN $x$ -SPACE $\Rightarrow$ WITTEN DIAGRAMS

- propagators: bulk-to-boundary propagator

$\hookrightarrow$  Witten's diagrams ( $\neq$  Feynman because NOT quantum)

$\Rightarrow$  Higher n-point functions, other than scalars, too!

$\hookrightarrow$  AIM: compute  $S = C \int d^4x \sqrt{g} ((\partial\phi)^2 + m^2\phi^2)$

$\hookrightarrow$  start from  $(\square_{AdS} - m^2)\phi^2 = 0$

Define

$$\rightarrow \partial_z K(z, x, x') = \delta^4(x - x')$$

$$\Rightarrow \phi(x, z) = \int d^4x' K(z, x, x') \phi(x', z)$$

$$\Rightarrow \text{SOLUTION: } K(z, x, x') = C \frac{z^\Delta}{(z^2 + |x-x'|^2)^\Delta} \Rightarrow C = \frac{\Gamma(\Delta)}{\pi^2 \Gamma(\Delta-2)}$$

$\hookrightarrow K \rightarrow 0$  if  $x - x' \neq 0$  as  $z \rightarrow 0$

$K \sim z^{-\Delta} \rightarrow \infty$  if  $x = x'$  as  $z \rightarrow 0$

After IBP:

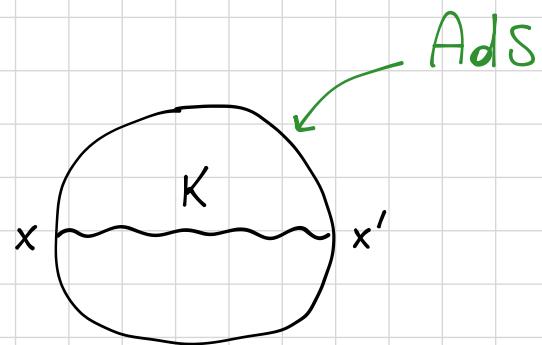
$$S = C \int_{z=\epsilon} d^4x L^3 \frac{1}{z^3} \phi(x, z) \partial_z \phi(x, z) \quad \text{NB: } \phi(x, z) \stackrel{z \rightarrow 0}{\sim} z^{4-\Delta} \phi_0(x)$$

Insert  $\phi = \int K \phi_0$ :

$$S \sim \int_{z=\epsilon} d^4x \frac{1}{z^3} z^{4-\Delta} \phi_0(x) z^{\Delta-1} \int d^4x' \frac{\phi_0(x')}{|x-x'|^{2\Delta}} =$$

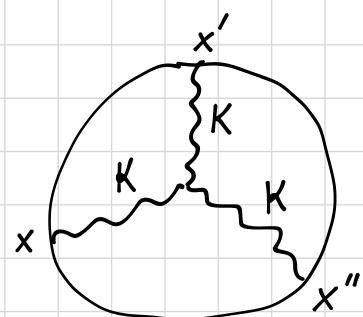
$$= \int d^4x \int d^4x' \frac{\phi_0(x) \phi_0(x')}{|x-x'|^{2\Delta}}$$

This is related to:



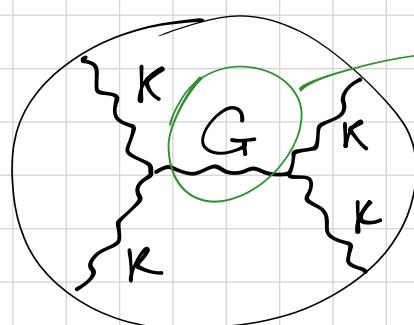
once we take the right no. of funct. derivatives.

For the 3 point function:



$\Rightarrow$  we need  $S \sim \int \phi^3$  to have it!

For 4 point functions:



bulk-to-bulk propagator

$$\Rightarrow S_{\text{bulk}} = S_{\text{skin}} + \int d^5x \sqrt{g} \frac{1}{3} \lambda \phi^3$$

$$\hookrightarrow (-\square + m^2) \phi = \lambda \phi^2$$

$\Rightarrow$  PERTURBATION THEORY

$$\phi = \phi^{[0]} + \phi^{[1]} + \phi^{[2]} + \dots$$

where  $\phi^{[0]}(x, z) = \int d^4x' K(z, x, x') \phi_0(x')$ . Then:

$$\phi^{[1]} = \lambda \int d^4x' d\bar{x}' G(z, z', x, x') (\phi^{[0]}(x', \bar{x}'))^2$$

$$(-\square + m^2) G = \frac{1}{g} \delta(z - \bar{z}') \delta^4(x - x')$$

# HOLOGRAPHIC RENORMALIZATION

$z = \epsilon \rightarrow 0 \Rightarrow \exists$  DIVERGENCES !!! e.g.: in  $\langle \Theta(p) \Theta(q) \rangle$

What are the symm. we want to preserve?

\* SYMMETRIES OF AdS!

$$\rightarrow S_{HE} \sim \int d^d x \sqrt{g} \left( R + \frac{(D-1)(D-2)}{L^2} \right)$$

$$\text{for AdS: } R = -\frac{D(D-1)}{L^2}$$

$$S_{HE} \Big|_{\text{on-shell}} \sim \int d^{D-1} x \int_0^{+\infty} \frac{dz}{z^D} \rightarrow \infty$$

THIS IS WHERE THE PROBLEM LIES

$\Rightarrow$  This is a general feature:  $g_{\mu\nu} \neq \text{AdS}_{d+1}$



$g_{\mu\nu}$ : asymptotically (locally) AdS

I can choose

$$ds^2 = \frac{1}{z^2} \left( dz^2 + g_{ij} dx^i dx^j \right)$$

ONLY IF  $d$  EVEN

they dep. on  $g_{(0)}$

$$\text{where } g_{ij}(x, z) = g_{(0)ij}(x) + \dots + z^d \left( g_{(d)ij}(x) + h_{(d)ij}(x) \ln z^2 \right) + \dots$$

↳ FEFFERMAN-GRAHAM

$\Rightarrow$  corresponds to the expansion of  $\phi$  w/  $\Delta = d \Rightarrow$  compatible with  $m^2 = 0$ .

Then we can plug it into the action:

$$S_{\text{reg}} \Big|_{\text{on-shell}} = \int d^d x \sqrt{g} \left[ \epsilon^{-2\Delta+d} a_{(0)} + \epsilon^{-2\Delta+d+2} a_{(2)} + \dots + \ln \epsilon a_{(d)} + O(\epsilon^0) \right]$$

$\Rightarrow$  Define counterterms:

$$S_{\text{ct}} = -(\text{divergent part of } S_{\text{reg}} \Big|_{\text{on-shell}})$$

$$\hookrightarrow S_{\text{sub}} = S_{\text{reg}} + S_{\text{ct}} = \int d^d x \sqrt{g} \quad (\text{FINITE})$$

$\rightarrow$  compute, then  $\epsilon \rightarrow 0$

$$\Rightarrow S_{\text{ren}}[\phi_0] = \lim_{\epsilon \rightarrow 0} S_{\text{sub}}[\phi_0, \epsilon]$$

e.g.:

Go back to KG eqn:

\* SCALAR in fixed metric on  $\text{AdS}_{d+1}$ :

$$(\square_{\text{AdS}} - m^2) \Phi(x, z) = 0$$

$$\Rightarrow \Phi(x, z) = z^{d-\Delta} \phi(x, z)$$

$$\hookrightarrow (\Delta(\Delta-d) - m^2 L^2) \phi(x, z) + (1+d-2\Delta) z \partial_z \phi(x, z) +$$

$$+ z^2 (\partial_z^2 \phi(x, z) + \square_d \phi(x, z)) = 0$$

$$\Rightarrow \text{Ansatz: } \phi(x, z) = \phi_0(x) + z^2 \phi_{(2)}(x) + \dots \quad (\text{odd powers are set to 0})$$

$\downarrow$

$$\text{at 1st NON TRIV. ord: } (1+d-2\Delta) \phi_{(2)}(x) + \phi_{(2)}(x) + \frac{1}{2} \square \phi_{(0)}(x) = 0$$

$$\Rightarrow \phi_{(2)}(x) = \frac{-\square \phi_{(0)}}{2(d+2-2\Delta)}$$

$$\text{Then } \phi_{(4)} = \frac{1}{4(2\Delta - d - 4)} \square \phi_{(2)}$$

:

$$\phi_{(2n)} = \underbrace{\frac{1}{2n(2\Delta - d - 2n)}}_{\square} \phi_{(2n-2)}$$

AT SOME THIS CAN VANISH

if at some point  $\exists k \mid 2\Delta - d - 2k = 0$ , then add

$$\phi(x, z) = \phi_{(0)}(x) + \dots + z^{2\Delta-d} (\phi_{(2k)}(x) + \rho_m z^2 \tilde{\phi}_{(2k)}(x)) + \dots$$

↓ w/ this

$$\tilde{\phi}_{(2k)} = -\frac{1}{2} \square \phi_{(2k-2)}$$

$$\Rightarrow S = C \int d^d x \sqrt{g} \left( (\partial \Phi)^2 + m^2 \Phi^2 \right) =$$

$$= C \int d^d x \frac{1}{\epsilon^{2\Delta-d}} \left( (d-\Delta) \phi^2 + \phi \in \partial_\epsilon \phi \right)$$

$$\Rightarrow \epsilon \partial_\epsilon \phi = 2\epsilon^2 \phi_{(2)} + 4\epsilon^4 \phi_{(4)} + \dots$$

$$\hookrightarrow S = -CL^{d-1} \int_{z=\epsilon}^d x \epsilon^{d-2\Delta} \left[ (d-\Delta) \phi_{(0)}^2 + (2d-2\Delta+2) \phi_{(0)} \phi_{(2)} \epsilon^2 + \dots \right]$$

Referring to the previous expansion:

$$\alpha_{(0)} = -CL^{d-1} (d-\Delta)$$

$$\alpha_{(2)} = CL^{d-1} \frac{d-\Delta+1}{d-2\Delta+2} \phi_{(0)} \square \phi_{(0)}$$

Now determine

$\phi_0(x)$  as a function of  $\bar{\Phi}(x, z)$ :

$$\phi_{(0)} = \epsilon^{-d+\Delta} \left( \bar{\Phi}(x, \epsilon) - \frac{1}{2(2\Delta-d-2)} \square_y \bar{\Phi}(x, \epsilon) \right)$$

$\gamma_{ij} = \frac{1}{\epsilon^2} \delta_{ij}$

$$\phi_{(2)} = \epsilon^{-d+\Delta-2} \frac{1}{2(2\Delta-d-2)} \square_y \bar{\Phi}(x, \epsilon)$$

Therefore:

there are  $\epsilon$  terms in this but we hide them to show the full AdS symmetry

$$S_{ct}[\bar{\Phi}, \epsilon] = G_L^{d-1} \int_{z=\epsilon}^d d^d x \sqrt{g} \left[ (d-\Delta) \bar{\Phi}^2 + \frac{1}{2\Delta-d-2} \bar{\Phi} \square_y \bar{\Phi} \right] + \dots$$

Sufficient to renorm. up to  $\Delta < \frac{d}{2} + 2$

NB:  $\Delta = \frac{d}{2} + 1 \rightarrow$  add  $\ln$  term:

$$\frac{1}{2\Delta-d+2} \bar{\Phi} \square_y \bar{\Phi} \xrightarrow{\text{ANOMALY}} -\frac{1}{2} \ln \epsilon \bar{\Phi} \square_y \bar{\Phi}$$

↓ explicitly insert  $\epsilon$ !

## HOLOGRAPHIC WEYL ANOMALY

→ focus on metric using FG expansion: det.  $g_{(2n)ij}(x)$ .

$$\rightarrow S = \frac{1}{16\pi G} \int d^{d+1} x \sqrt{g} \left( R + \frac{d(d-1)}{L^2} \right) - \frac{1}{8\pi G} \int \partial^d x \sqrt{g} K$$

induced metric on boundary

$K = \gamma^\mu \nabla_\mu n_\nu$

The additional term will cancel some of the div  $\Rightarrow$  NOT A CT!

(NB the boundary term DOES NOT affect the e.o.m.)

$$\Rightarrow ds^2 = \frac{L^2}{z^2} (dz^2 + g_{ij} dx^i dx^j) \text{ where } g_{ij} = g_{(0)ij} + z^2 g_{(2)ij} + \dots$$

$$(NB: g_{(2)ij} = \frac{L^2}{d-2} (R^{(0)}_{ij} - \frac{1}{2(d-1)} R^{(0)} g_{(0)ij})$$

Up to  $g_{(d-2)ij}$  everything is det., but we know few things on  $g_{(d-2)ij}$ .

e.g.:

$$\text{Tr } g_{(4)} = \frac{1}{4} \text{Tr } g_{(2)}^2 \quad \xrightarrow{\text{w/ } g_{(0)ij}}$$

$$\rightarrow S_{ct} = \frac{1}{8\pi G} \int_{z=\epsilon}^d \sqrt{\gamma} \left[ \frac{d-1}{L} + \frac{L}{2(d-2)} R + \frac{L^3}{2(d-4)(d-2)^2} \underbrace{\left( R_{ij} R^{ij} - \frac{d}{4(d-1)} R^2 \right)}_{\text{computed w/ fall } \gamma_{ij}} + \dots \right]$$

$$e.g.: d=4 \Rightarrow -\frac{L^3}{2(d-2)^2} \ln \in \left( R_{ij} R^{ij} - \frac{1}{12} R^2 \right)$$

In CFT  $\rightarrow D$  invariance leads to

$$\langle g^{ij} T_{ij} \rangle = A \neq 0 \rightarrow \text{ANOMALY}$$

at quantum level.

$$S_{ct} = \frac{1}{8\pi G} \int d^4x \sqrt{g} \left[ \frac{3}{L} + \frac{L}{4} R - \frac{L^3}{8} (R^{ij} R_{ij} - \frac{1}{3} R^2) \ln \epsilon \right]$$

↪  $\langle T \rangle = A \neq 0 \Rightarrow$  WEYL ANOMALY

$$ds^2 = L^2 \left( \frac{d\rho^2}{4\rho^2} + \frac{1}{\rho} g_{ij}(\rho, x) dx^i dx^j \right) \quad \rho = z^2$$

$\Rightarrow P_{BH}$  diffeom in the bulk (= Weyl var. at  $\partial$ ) :

$$\begin{cases} \rho = \rho'(1 - 2\sigma(x')) \\ x^i = x^{i'} + a^i(x^i, \rho') \end{cases}$$

$$\text{where } a^i(x, \rho) = \frac{L^2}{2} \int_0^\rho d\hat{\rho} \hat{g}^{ij}(x, \hat{\rho}) \partial_j \sigma(x)$$

$$\Rightarrow g_{ij} \rightarrow g_{ij} + 2\sigma(1 - \rho \frac{\partial}{\partial \rho}) g_{ij} + 2 \nabla_{(i} a_{j)}$$

↪ for  $\rho \rightarrow 0 \rightarrow a_i \rightarrow 0$

$$g_{ij} \xrightarrow[\text{at boundary}]{} g_{ij}(1 + 2\sigma) \simeq e^{2\sigma} g_{ij}$$

$$\text{Then } S_{PBH}(S + S_{ct}) = S_{PBH} S_{ct} = \int_{\rho=\epsilon} d^4x \left( \delta \rho \frac{\delta}{\delta \rho} + \delta g_{ij} \frac{\delta}{\delta g_{ij}} \right) S_{ct} =$$

↓

$-2\sigma \rho \quad 2\sigma g_{ij}$

picks up the  $\ln \epsilon$  term in  $S_{ct}$

$$\Rightarrow A = \frac{1}{16\pi G} 2L^3 \left( \frac{1}{8} R^{ij} R_{ij} - \frac{1}{24} R^2 \right)$$

$\Rightarrow$  Field theory (general CFT) :

$$\langle T^i_{\ i} \rangle = \frac{c}{16\pi^2} I_4 - \frac{a}{16\pi^2} E_4$$

$$\Rightarrow I_4 = R_{ijk\rho} R^{ijkl} - 2R_{ij} R^{ij} + \frac{1}{3} R^2$$

$$E_4 = R^{ijkl} R_{ijkl} - 4R^{ij} R_{ij} + R^2$$

$\Rightarrow$  Then

$$1) \quad a = c$$

$$2) \quad a = \frac{\pi L^3}{8G_s}$$

$$\Rightarrow a = \frac{\pi^3 N^2}{4 \text{vol}(S^5)} = \frac{N^2}{4}$$

$$G_5 = \frac{G_{10}}{L^5 \text{vol}(S^5)}$$

$$16\pi G_{10} = (2\pi)^7 \ell_s^8 g_s^2$$

$$L^4 = \frac{(2\pi)^4 g_s \ell_s^4 N}{4 \text{vol}(S^5)}$$

Now dfr. with  $\sqrt{= 4}$  SYM :

$$a = c = \frac{N^2}{4} - \frac{1}{4}$$

$\rightarrow$  The Weyl anomaly DOES NOT depend on the coupling.

NB: the comp. goes on even w/o  $S^5$ :  $a = \frac{\pi^3 N^2}{4 \text{vol}(Y_5)}$ .

$\downarrow$

5D manifold

DEFORATIONS of  $\text{AdS}_5$ :

$$\text{AdS}_5 \times S^5$$

deform  $\text{AdS}_5$       deform  $S^5$

keep  $S^5$       keep  $\text{AdS}_5$

a linear comb.

# SUPERSYMMETRY BREAKING

→ Introduce finite  $T$ :

1) go to Euclidean

2) compactify time  $t_E \sim t_E + \beta$

PARTITION F.

$$Z_E[\beta] = \int D\phi e^{-S_E[\phi]} = \text{Tr } e^{-\beta H}$$

$$\phi(\vec{x}, t) = \phi(\vec{x}, t_E + \beta)$$

$$\psi(\vec{x}, t_E) = -\psi(\vec{x}, t_E + \beta)$$

SUSY break.

What about GRAVITY side?

⇒ Do the same!

↳ e.g. SCHWARZSCHILD BH (not AdS, but good to start with)

$$d=4 \Rightarrow ds^2 = + \left(1 - \frac{2MG}{r}\right) dt^2 + \frac{dr^2}{1 - \frac{2MG}{r}} + r^2 ds^2(S^2)$$

EUCL.

⇒ Horizon at  $r_* = 2MG \Rightarrow$  IN EUCL. THERE IS NO HORIZON

↳ is it regular? YES but we check for possible sing. which we can pick up somewhere else

NB:  $\tau \sim \tau + \beta \rightarrow$  ~polar coord → in  $(r, \tau)$  the space looks like  $\mathbb{R}^2$

↳ change coord:  $\rho^2 = r - 2MG$

Near  $\rho \rightarrow 0$ :

$$ds^2 \simeq \frac{1}{r_*} \rho^2 d\tau^2 + 4 r_* d\rho^2 + r_*^2 ds^2(S^2) = \\ = 4 r_* \left( \rho^2 \left( \frac{d\tau}{2r_*} \right)^2 + d\rho^2 \right) + r_*^2 ds^2(S^2)$$

$$\Rightarrow \frac{\mathcal{I}}{2r_*} \in [0, 2\pi] \Rightarrow \frac{1}{T} = 8\pi N G \quad (\text{temperature})$$

$$\hookrightarrow dU = T dS \Rightarrow S = \int dM \frac{1}{T(M)} = 4\pi G M^2 = \frac{A_{\text{Hd}}}{4G}$$

$\downarrow$   
 $U \sim M$

Now consider BLACK HOLES in AdS:

- 1) Poincaré patch
- 2) Global coord.

$\Rightarrow$  Poincaré Coords.

Go back to D3-brane solution.

Deform:

$$ds^2 = (H(r))^{-\frac{1}{2}} \left( -f(r) dt^2 + d\vec{x}^2 \right) + (H(r))^{\frac{1}{2}} \left( \frac{dr^2}{f(r)} + r^2 ds^2(S^5) \right)$$

$$* H(r) = 1 - \frac{L^4}{r^4}$$

$$* f(r) = 1 - \underbrace{\left(\frac{r_*}{r}\right)^4}_{\text{"BLACKENING FACTOR"}}$$

deform.  $\Rightarrow$  not SUSY!

Now take

1) near-horizon

2)  $t \rightarrow -i\tau$

$$3) z = \frac{L^2}{r}$$



$$ds^2 = \frac{L^2}{z^2} \left( \left( 1 - \frac{z^4}{z_*^4} \right) d\tau^2 + d\vec{x}^2 + \frac{dz^2}{1 - \frac{z^4}{z_*^4}} \right) + L^2 ds^2(S^5)$$

with  $\tau \sim \tau + \beta$ .

THIS IS ALSO SUSY SOL.

↓ AdS<sub>5</sub> in Poincaré patch.

Now define  $\rho : \frac{z}{z_*} = 1 - \frac{\rho^2}{L^2}$

$$\Rightarrow z \rightarrow z_* \leftrightarrow \rho \rightarrow 0 \longrightarrow ds^2 \xrightarrow{\rho \rightarrow 0} d\rho^2 + \frac{4}{z_*^2} \rho^2 d\tau^2 + \frac{L^2}{z_*^2} d\vec{x}^2$$

$$\Rightarrow \frac{1}{\beta} = T = \frac{1}{\pi z_*}$$

$$\hookrightarrow S = \frac{A}{4G_S} \Rightarrow A = \left( \frac{L}{z_*} \right)^3 \text{vol}(R^3)$$

$$\begin{aligned} \frac{S}{\text{vol}(R^3)} &= \frac{L^3}{z_*^3} \cdot \frac{L^5 \text{vol}(S^5)}{G_{10} \cdot 4} = \frac{N^2}{\text{vol}(S^5)} \cdot \frac{\pi^5}{2} T^3 = \\ &= \frac{1}{2} \pi^2 T^3 N^2 \end{aligned}$$

$\Rightarrow$  this is hol. dual of  $\mathcal{N}=4$  SYM @ temp =  $T$  ( $N \rightarrow \infty, \lambda \gg 1$ )

↪ SUSY broken + strong coupl.  $\Rightarrow$  plasma?

The comparison can be made in free theory:  $S = \frac{S}{\text{vol}(\mathbb{R}^3)}$  in free  $\text{sym} @ T$

$$S_{\text{free}} = \frac{2}{3} \pi^2 N^2 T^3$$

while  $S = \frac{1}{2} \pi^2 N^2 T^3 \Rightarrow \frac{4}{3}$  mismatch.

$$\Rightarrow \frac{S_{\lambda=\infty}}{S_{\lambda=0}} = \frac{3}{4} \simeq 75\% \quad \text{while in } \ell@CD \quad \frac{S_{\lambda=\infty}}{S_{\lambda=0}} = 80\% !$$

Now let's look at:

\* finite charge density  $\Rightarrow$  chemical potential  
 $J^0$   $\alpha_0$   
 $(\rightarrow \partial_i J^i = 0)$

$$\Rightarrow \int d^d x J^i a_i \quad \xrightarrow{\text{bulk: } A_\mu \text{ (gauge field)}}$$

$$\Rightarrow AdS_5 \times S^5 \longrightarrow \text{symm.} = SO(6) \supset U(1)^3 \text{ (Cartan subalg)}$$

Charge dens. at bound. corresponds to ELECTRICALLY CHARGED BULK SOL./BH:

$$A \rightarrow a_0 dt \text{ for } z \rightarrow 0$$

→ EFFECTIVE SUGRA (Einstein-Maxwell +  $\Lambda$ )

$$ds^2 = \frac{L^2}{z^2} \left( -f(z) dz^2 + d\vec{x}^2 + \frac{dz^2}{f(z)} \right)$$

$$f(z) = 1 - \left(\frac{z}{z_*}\right)^4 (1 + c\mu^2) + \\ + c\mu^2 \left(\frac{z}{z_*}\right)^6$$

$$\Rightarrow 2 \text{ param: } z_* \sim \frac{1}{T}, \mu$$

$$A = \mu \left( 1 - \left(\frac{z}{z_*}\right)^2 \right) dt \xrightarrow{z \rightarrow 0} \mu dt \quad (a_0 = \mu)$$

$$\rightarrow T = \frac{1}{\pi z_*} \left( 1 - \frac{c\mu^2}{2} \right)$$

$$A \text{ for } z \rightarrow \hat{z} \mid f(\hat{z}) = 0 \text{ (near-horizon)} \Rightarrow \boxed{A \rightarrow 0}$$

## THERMODYNAMICS

↳ Grand-canonical ensemble:

$$\Omega(\mu, T, V) = U - TS - \mu N$$

We have

$$Z_{QFT} = e^{-\beta \Omega} = e^{-S_{\text{SUGRA}}}$$

$$\downarrow \quad \Omega = T S_{\text{SUGRA}}(T, \mu, V)$$

Now consider  $\text{AdS}_5$  in global coord. ( $\mathcal{W} = 4 \text{ Sy} \mathbb{R} \times S^3$ )  
+ finite  $T$

$\Rightarrow$  Euclidean  $S_\beta^1 \times S^3$

$$\Rightarrow Z_{\text{QFT}} [S_\beta^1 \times S^3] = \sum_i Z_{\text{SUGRA}} [X_i]$$

inner  
 possible manifolds (all of them)  
 $(\text{SUSY} \rightarrow \text{unique})$   
 $(\text{SUSY} \rightarrow \text{not unique})$

$$\Rightarrow X_1 : \text{"thermal AdS}_5\text{"} \rightarrow ds^2 = \left(1 + \frac{r^2}{L^2}\right) d\tau^2 + \frac{dr^2}{1 + \frac{r^2}{L^2}} + r^2 ds^2(S^3)$$

$\hookrightarrow r \rightarrow 0 : S^1 \times \mathbb{R}^4$

$$\Rightarrow X_2 : ds^2 = f(r) d\tau^2 + \frac{1}{f(r)} dr^2 + r^2 ds^2(S^3)$$

$$f(r) = 1 - \frac{\mu}{r} + \frac{r^2}{L^2}$$

$$( \text{horizon: } f(r_h) = 0 )$$

$$\Rightarrow T = \frac{1}{2\pi} \frac{|f'(r_h)|}{2} = \frac{1}{4\pi r_h L^2} (2L^2 + 4r_h^2) \quad f(r_h) = 0$$

$$\hookrightarrow T_{\min} = \frac{1}{2\pi L} \sqrt{d(d-2)} \quad (d=4)$$

$\Rightarrow$  Sum over two sol:

$$\begin{aligned} Z_{\text{QFT}} &= e^{-S[X_1]} + e^{-S[X_2]} = \\ &= e^{-S[X_1]} \left( 1 + e^{-(S[X_2] - S[X_1])} \right) \end{aligned}$$

on-shell  
 act.

$$\Rightarrow S[X_2] - S[X_1] = \frac{1}{T} \frac{r_h^{d-2}}{2K^2} \text{vol}(S^{d-1}) \left( 1 - \frac{r_h^2}{L^2} \right)$$

$r_h = L$  is a phase transition at charact.  $T \Rightarrow \text{CONF.} / \text{DECONF.}$   
 PHASE TRANS.

$$\Rightarrow (\Theta(x))^+ = x^{-2\Delta_\theta} \Theta\left(\frac{x^\mu}{x^2}\right)$$

$$(\Theta(x)|_0)^+ = x^{-2\Delta_\theta} \langle 0 | \Theta\left(\frac{x^\mu}{x^2}\right)$$

Now take the usual limit:  $(y^\mu = \frac{x^\mu}{x^2})$ :

$$\langle \Theta | = \lim_{y \rightarrow \infty} y^{2\Delta_\theta} \langle 0 | \Theta(y)$$

$$\hookrightarrow \langle \Theta(x_1) \Theta(x_2) \rangle = \langle 0 | \left( x_1^{-2\Delta_\theta} \Theta(y_1) \right)^+ \Theta(x_2) | 0 \rangle$$

$$= x_1^{-2\Delta_\theta} \langle 0 | \left( e^{y_1 P} \Theta(0) e^{-y_1 P} \right)^+ e^{x_2 P} \Theta(0) e^{-x_2 P} | 0 \rangle$$

$$= x_1^{-2\Delta_\theta} \langle 0 | \Theta(0) e^{y_1 K} e^{x_2 P} \Theta(0) | 0 \rangle =$$

$$= x_1^{-2\Delta_\theta} \langle \Theta | e^{y_1 K} e^{x_2 P} | \Theta \rangle =$$

$$|\Theta\rangle \text{ is scalar} \quad \sim x_1^{-2\Delta_\theta} \left( \langle \Theta | \Theta \rangle + \frac{x_1^\mu x_2^\nu}{|x_1|^2} \langle \Theta | K_\mu P_\nu | \Theta \rangle \right)$$

$$= x_1^{-2\Delta_\theta} \left( \langle \Theta | \Theta \rangle + \frac{x_1^\mu x_2^\nu}{|x_1|^2} \langle \Theta | [K_\mu, P_\nu] | \Theta \rangle \right) =$$

$$= x_1^{-2\Delta_\theta} \left( \langle \Theta | \Theta \rangle + \frac{x_1^\mu x_2^\nu}{|x_1|^2} \langle \Theta | (2\Delta_\theta \eta_{\mu\nu} - 2P_\mu^\nu) | \Theta \rangle \right)$$

$$= x_1^{-2\Delta_\theta} \underbrace{\left( 1 + 2\Delta_\theta \frac{x_1 \cdot x_2}{|x_1|^2} \right)}_{=1} \langle \Theta | \Theta \rangle.$$

NB we know  $\langle \Theta(x_1) \Theta(x_2) \rangle = \frac{1}{|x_{12}|^{2\Delta_\theta}}$  Expansion at  $x_2 \rightarrow 0$

NB • Reflection positivity in EQFT:

$$\langle \psi | \psi \rangle \geq 0$$

Define  $|\Theta\rangle$  s.t.:

$$\langle \Theta | \Theta \rangle = 1.$$

Do the descendants have positive norm?

$$|P_0|\Theta\rangle|^2 = \langle \Theta | K_0 P_0 |\Theta \rangle = 2\Delta_\Theta \langle \Theta | \Theta \rangle = 2\Delta_\Theta$$

$\Rightarrow$  it must be  $\Delta_\Theta \geq 0$  (for SCALARS!)

In general

$$\langle \Theta^I | \Theta_J \rangle = \delta_J^I \rightarrow (P_\mu | \Theta^I \rangle)^\dagger (P_\nu | \Theta_J \rangle) = \langle \Theta^I | K_\mu P_\nu | \Theta_J \rangle =$$

$$= \langle \Theta^I | [K_\mu, P_\nu] | \Theta_J \rangle = \xrightarrow{\text{int rep. of } \eta_{\mu\nu}}$$

$$= \langle \Theta^I | 2\Delta_\Theta \eta_{\mu\nu} - 2R(\eta_{\mu\nu})_I^J | \Theta_J \rangle =$$

$$= 2\Delta_\Theta \eta_{\mu\nu} \delta_J^I - 2 \langle \Theta^I | R(\eta_{\mu\nu})_I^J | \Theta_J \rangle =$$

$$= 2\Delta_\Theta \eta_{\mu\nu} \delta_J^I - 2 R(\eta_{\mu\nu})_J^I$$

$\Rightarrow \Delta_\Theta \geq \text{MAX EIGENVALUE OF } R(\eta_{\mu\nu})_J^I$

$$\Rightarrow R(M_{\mu\nu})^I_J = \frac{1}{2} \left( L^{ab} \right)_{\mu\nu} \cdot \left( S_{ab} \right)^I_J$$

$$(L^{ab})_{\mu\nu} = \delta_\mu^a \delta_\nu^b - \delta_\nu^a \delta_\mu^b$$

$$\Rightarrow L^{ab} S_{ab} = \frac{1}{2} \left[ (L + S)^2 - \underbrace{L^2 - S^2} \right] =$$

Casimir oper.

$$= \frac{1}{2} [ \text{Cas}(V_1) + \text{Cas}(r) - \text{Cas}(r \otimes V_1) ]$$

Suppose  $r$  is traceless-symm. w/  $\ell$  indices:

$$\text{Cas}(\ell) = \ell(\ell+d-2)$$

$$\Rightarrow \Delta \geq \frac{1}{2} \left[ \text{Cas}(\ell) + \underbrace{\text{Cas}(\ell=1)}_{\text{Cas}(V_\ell \otimes V_1)} - \text{Cas}(V_\ell \otimes V_1) \right]$$

$$V_1 \otimes V_\ell = \underline{\underline{V_{\ell-1}}} \oplus \dots$$

Therefore

$$\Delta \geq \ell + d - 2$$

↪ the larger the spin the larger the dim.

$$\Rightarrow |P_\mu P^\mu |\phi\rangle \geq 0 \Rightarrow \Delta_0 \geq \frac{d-2}{2} \quad (\ell=0)$$

NB:  $\mathbb{I} \rightarrow \Delta = 0, \ell = 0$

These are called UNITARITY BOUNDS.

When one of them is saturated, then one of the desc. has zero norm. E.g.:

$$J_\mu \Rightarrow \Delta_J \geq d-1$$

When  $\Delta_J = d - 1 \Rightarrow |\mathcal{P}_\mu |J^\mu\rangle|^2 = 0$   
 ↪ " $\partial_\mu J^\mu = 0$ " as a quantum op.  
 ("SHORT MULTIPLET")

If  $\ell = 2$ :

$$T^{\mu\nu} \rightarrow \Delta_T = d \Rightarrow |\mathcal{P}_\mu |T^{\mu\nu}\rangle|^2 = 0 \Rightarrow \partial_\mu T^{\mu\nu} = 0 \quad \checkmark$$

## OPERATOR PRODUCT EXPANSION

$$\underbrace{\bullet O_1(x_1) \bullet O_2(x_2)}_{|4\rangle} = \sum_{\Delta, \ell} C_{\Delta, \ell} |O_{\Delta, \ell}\rangle$$

$$\Rightarrow \langle O_1 O_2 \dots O_n \rangle$$

take two of them and replace  $|4\rangle$   by the state containing the irreps of them.

$$\Rightarrow O_1(x_1) O_2(x_2) = \sum_k C_{12k}(x_{12}, \partial_2) O_k(x_2)$$

$$\quad \quad \quad = \sum_k \tilde{C}_{12k}(x_{12}, x_{23}, \partial_3) O_k(x_3)$$

↗ valid INSIDE CORR. FUNCT.

"OP E"

$$\text{e.g.: } \underbrace{\langle O_1(x_1) O_2(x_2) O_3(x_3) \rangle}_{\int_{123}} = \sum_k C_{12k}(x_{12}, \partial_2) \langle O_k(x_2) O_3(x_3) \rangle$$

$$\int_{123}^{''}$$

$$x_{12}^{\Delta_1 + \Delta - \Delta_3} x_{13}^{\Delta_1 - \Delta_2 + \Delta_3} x_{23}^{-\Delta_1 + \Delta_2 + \Delta_3}$$

$$\Delta_1 = \Delta_2 = \Delta, \Delta_3 \text{ arb.}$$

$$= \sum_k C_{12k}(x_{12}, \partial_2) \frac{\delta_{k,3}}{x_{23}^{2\Delta_0}} =$$

$$= \underbrace{C_{123}(x_{12}, \partial_2)}_{\text{DETERMINE } C_{123}} \frac{1}{x_{23}^{2\Delta_3}}$$

$$\Rightarrow C_{123}(x, \partial) = \int_{123} x^{\Delta_3 - 2\Delta} \left( 1 + \frac{1}{2} x \cdot \partial + \alpha x^\mu x^\nu \partial_\mu \partial_\nu + \dots \right)$$

where  $\alpha = \frac{\Delta_3 + 2}{8(\Delta_3 + 1)}$ .

## SASAKI-EINSTEIN / $\mathcal{N}=1$ SCFT DUALITY

$$\frac{1}{4} \text{SUSY} \leftarrow \text{AdS}_5 \times S^5 \rightarrow \text{replace w/ } Y_5$$

Gravity side  $\rightarrow$  Type IIB:

$$\begin{aligned} \text{Ansatz} \quad & \left\{ \begin{array}{l} ds_{10}^2 = ds^2(\text{AdS}_5) + ds^2(Y_5) \\ F_5 = \frac{L}{L} (\text{vol}(\text{AdS}_5) + \text{vol}(Y_5)) \end{array} \right. \end{aligned}$$

$$\Rightarrow R_{ij} = -\frac{4}{L^2} g_{ij} \quad \text{"EINSTEIN METRIC"}$$

$$\rightarrow 10d: \epsilon = \chi \otimes \gamma \quad \begin{matrix} \downarrow \\ \text{AdS}_5 \end{matrix} \quad \begin{matrix} \downarrow \\ Y_5 \end{matrix}$$

$$\text{KSE} \Rightarrow \nabla_\mu \chi = \frac{1}{2L} \gamma_\mu \chi \Rightarrow \nabla_i \gamma + \frac{i}{2L} \gamma_i \gamma = 0$$

(check that  $Y_5$  has to be Einstein)

REMARKS:

- 1) generically  $\exists \geq 3$  solution with  $3$  Dirac
- 2) any SE manifold have CANONICAL KILLING VECTOR

$$\bar{\gamma} \gamma^\mu \gamma = V^\mu = \left( \frac{\partial}{\partial x} \right)^\mu$$

$$\hookrightarrow \frac{1}{L^2} ds^2(Y_5) = \underbrace{ds^2(M_4)} + (d\psi + a)^2$$

$$R_{ij}^{(4)} = 6 g_{ij}^{(4)} \quad (\text{K\"ahler})$$

SYMMETRY IS  $SO(2,4) \times U(1)$  :

\*  $\mathcal{W}=1$  (bosonic) Superconformal alg.

$\hookrightarrow$  dual 4d is SCFT ( $\mathcal{W} \geq 1$ )

Strategy:

- i) look at  $\mathcal{W}=1$  Lagrangians
- ii) require conf. invariance (CFT)

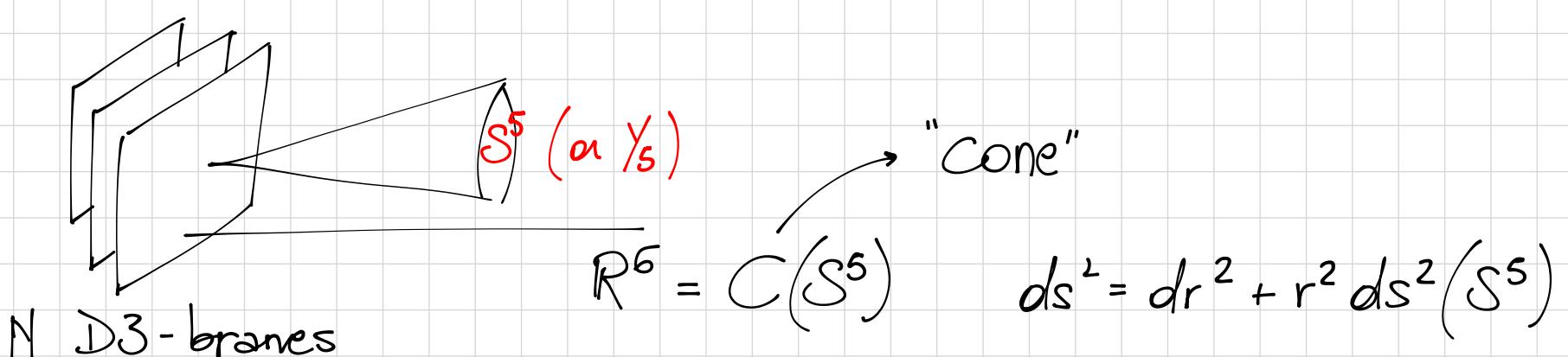
Once we have a candidate for  $\mathcal{W}=1$  SCFT :

$\Rightarrow$  CHECKLIST:

- 1) Symm.
- 2) KK spectrum
- 3) anomalies  $\rightarrow$  Weyl

Take D3-brane picture:

$\hookrightarrow F_5$  must come from this!



$$\Rightarrow ds_{10}^2 = (H(r))^{-\frac{1}{2}} ds^2(\mathbb{R}^{1,3}) + (H(r))^{\frac{1}{2}} \underbrace{(dr^2 + r^2 ds^2(S^5))}_{\text{Non Compact variety}}$$

Calabi-Yau man.

$$\nabla_i E = 0$$

(Special Holonomy)

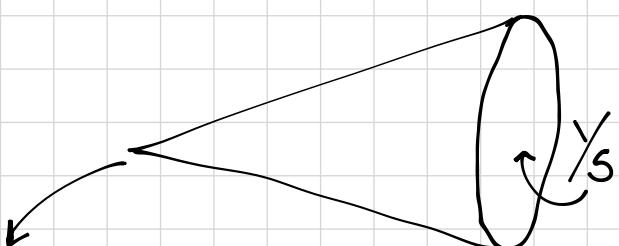
→ instead of  $C(S^5) \rightarrow C(Y_5)$



CALABI-YAU CONE  
from SE manifold

$$\Rightarrow H(r) = 1 + \frac{L^4}{r^4} \text{ is unchanged.}$$

Take near horizon limit:



SINGULAR

(but not visible to curv. invariants) → we cannot cut it out as usual

↳ CALABI-YAU SINGULARITIES

↳ ALGEBRAIC GEOMETRY  
(study of manifolds in  $\mathbb{C}^N$ )



the field th. comes from  
this

KEY CONSTRAINTS on our  $\mathcal{N}=1$  SCFT:

\* it comes from  $N$  D3-branes

$\Rightarrow$  SYM theory  $\rightarrow$  it cannot be pure SYM (not conf.)  
nor  $\mathcal{N}=4$  (it's  $\mathcal{N}=1$ )



we must add MATTER in rep. of  $G$

\*  $\mathcal{N}=1$  SUSY  $\Rightarrow$  multiplets and Lagrangians have det. structure

\* it has  $AdS_5$  dual  $\Rightarrow$  CFT

The mult:

- vector: 
$$\begin{cases} A_\mu^a \\ \lambda^a \\ \bar{\lambda}^a \end{cases}$$
 in ADJ of  $G$

- chiral: 
$$\begin{cases} \phi_i \\ \bar{\psi}_i \\ F_i \end{cases}$$
 in R of  $G$

$$\begin{aligned} \Rightarrow \mathcal{L} \supset & -\frac{1}{4} \text{Tr}(F_{\mu\nu} F^{\mu\nu}) + (\bar{\lambda}^a \phi)^+ (\bar{\lambda}^\mu \phi) + |F|^2 + \\ & + \bar{\lambda}^a \bar{\lambda}_a + (\text{fermions}) \end{aligned}$$

Integrate out  $F_i$ ,  $D^a$ :

$$F_i = \frac{\partial \mathcal{W}}{\partial \phi^i} \quad "F\text{-terms"} \Rightarrow |F|^2 + D^a D_a \sim V(\phi)$$

$$D^a = \phi^i T_{ij}^a \phi^j \quad "D\text{-terms"}$$

$\Rightarrow$  Vacuum moduli space: CONST SOLUTION:

$$V = 0 \Rightarrow D^a = 0 \quad (\text{enforced by } \delta \mathcal{L} = 0, \delta \lambda = 0)$$

SUSY Vacuum mod. sp.  $\Rightarrow$  KÄHLER VARIETY

In some cases it can be CY (any dim. because it's not the phys. spacetime)

$\Rightarrow$  SE  $\rightarrow C(Y_5)$ : CY  $\rightarrow$  SUSY  $\rightarrow$  vacuum moduli space

$\Rightarrow G$  is gauge group  $G = SU(n)^X$  with different  $g_{YM}^2$

$$\Rightarrow \mathcal{W} = \sum_{k=1}^M h_k^a \bar{\phi}_j^a \phi_j^k \quad \Rightarrow \text{flow } u \xrightarrow{v} c \xrightarrow{IR} \text{must flow to conf. point}$$

additional couplings

$\Rightarrow R$  is usually  $(N, \underline{N})$ : "bi-fundamental"

We use quartic Superpot.

Impose Conf. inv.  $\Rightarrow \beta_i = 0$

$\hookrightarrow \lambda \gg 1 \quad N \rightarrow \infty \Rightarrow$  How can we compute  $\beta$ ?

For  $\omega = 1$  th. the  $\beta$  is exact ("Russian  $\beta$ -function")  
NSVZ

We have:

$$g_{YM}, h_k \Rightarrow \beta(g_{YM}) = 0, \beta(h_k) = 0$$



$$\beta(g_{YM}) = 0 \Leftrightarrow 0 = 1 - \sum_I (1 - R_I)$$

chiral fields charged  
under gauge fields

$$\beta(h_k) = 0 \Leftrightarrow \sum_I R_I - 2 = 0 \quad \xrightarrow{\text{sum over the chirals}}$$

R-charges

$$\Rightarrow R = \frac{2}{3} \Delta \quad \text{for chiral fields}$$

R-charge of  $\phi$

→ we need to fix the R-charges  $\Rightarrow \alpha$ -MAXIMIZATION

# KLEBANOV-WITTEN DUALITY

$\text{AdS}_5 \times SE_5 \longleftrightarrow \text{W}^1 \text{ SYM} + \text{matter flowing to SCFT}$

$$\hookrightarrow SO(4) \simeq SU_1(2) \times SU_2(2)$$

$$\Rightarrow \underline{\text{metric}}: ds^2(T'') = \frac{1}{6} (d\epsilon_1^2 + \sin^2 \epsilon_1 d\phi_1^2 + 1 \leftrightarrow 2) +$$

$$+ \frac{1}{9} (d\psi - \cos \epsilon_1 d\phi_1 - \cos \epsilon_2 d\phi_2)^2$$

$$\phi_1, \phi_2 \in [0, 2\pi]$$

$$\psi \in [0, 4\pi]$$

$$\epsilon_1, \epsilon_2 \in [0, \pi]$$

$$\Rightarrow \text{vol}(T'') = \frac{1}{6^2} \cdot \frac{1}{3} \left( \int_0^{2\pi} d\phi_1 \right)^2 \left( \int_0^\pi d\phi_1 \sin \epsilon_1 \right) \int_0^{4\pi} d\psi = \frac{16}{27} \pi^3 = \frac{16}{27} \text{vol}(S^5)$$

$\hookleftarrow C(T'') := \text{"CONIFOLD"}$

$$\Rightarrow 1 \text{ eq. in } \mathbb{C}^4: z_1^2 + z_2^2 + z_3^2 + z_4^2 = 0$$

$\Rightarrow SO(4) \text{ SYMM. } \nearrow$

We can also use:  $\tilde{z}_1 \tilde{z}_2 = \tilde{z}_3 \tilde{z}_4 \quad \tilde{z}_i \in \mathbb{C}^4$

$$\tilde{z}_1 = A_1 B_1 \quad \tilde{z}_3 = A_1 B_2$$

$$\tilde{z}_2 = A_2 B_2 \quad \tilde{z}_4 = A_2 B_1$$

$$\Rightarrow A_i \sim \lambda A_i \quad B_i \sim \lambda^{-1} B_i \quad \lambda \in \mathbb{C}^*$$

$$\Rightarrow |A_1|^2 + |A_2|^2 - |B_1|^2 - |B_2|^2 = 0 \quad \lambda = \delta e^{i\alpha}$$

Impose  $A_i \sim e^{i\alpha} B_i \sim e^{-i\alpha} B_i$  (complex ident.)

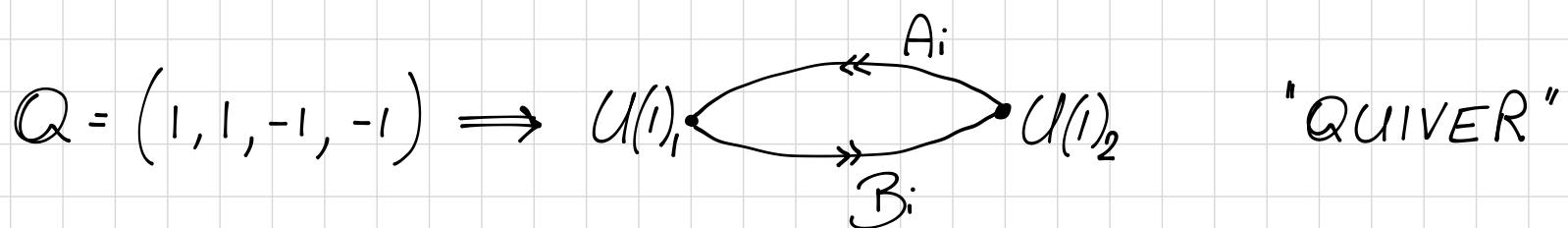
$\Rightarrow$  Kähler quotient:  $\mathbb{C}^4/\mathrm{U}(1)$

$\hookrightarrow$  charges:  $Q = (1, 1, -1, -1)$

$\hookrightarrow$  we have then  $A_i \in \mathcal{L}$  of  $SU(2)_1$   
 $B_i \in \mathcal{L}$  of  $SU(2)_2$

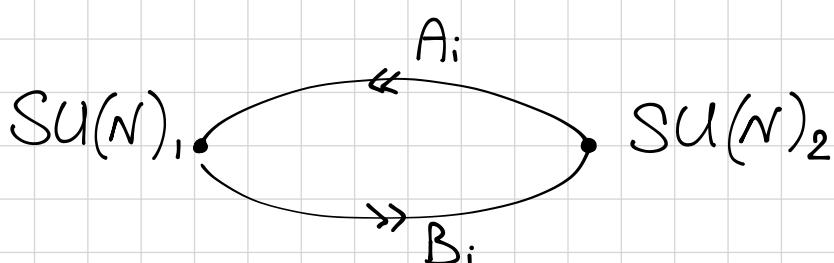
We can look at this as vac. mod. sp. of  $W=1$  SCFT:

$\Rightarrow$  gauge th. is  $U(1)_1 \times U(1)_2$



add  $W(\phi) = \lambda \epsilon^{ij} \epsilon^{kl} \text{Tr}[A_i B_k A_j B_l]$  to promote to  $SU(2)_1 \times SU(2)_2$

Effectively makes  $A_i, B_i$  abelian on the VMS = Sym N (conifold):



$$+ W = \lambda \epsilon^{ij} \epsilon^{kl} \text{Tr}(A_i B_k A_j B_l)$$

is then our QFT  $\Rightarrow$  CHECKLIST:

\* couplings:  $g_1, g_2, \lambda$

$$\beta_1 = 0 \Leftrightarrow 1 - \frac{1}{2} (1 - R[A_1] + 1 - R[A_2] + 1 - R[B_1] + 1 - R[B_2]) = 0$$

assume  $R[A_1] = R[A_2] = R[A]$  (same for  $B$ ):

$$R[A] + R[B] = 1$$

Now assume  $\mathbb{Z}_2$  symm. ( $A \leftrightarrow B$ ):

$$R[A] = R[B] = \frac{1}{2} \quad (\text{below unitarity bound})$$

$\Rightarrow$  KK spec. works

$\Rightarrow$  Weyl anomaly:

$$a = \frac{3}{32} (3\text{Tr } R^3 - \text{Tr } R)$$

$$c = \frac{1}{32} (9\text{Tr } R^3 - 5\text{Tr } R)$$

$$\text{NB: } \text{Tr } R^\alpha = \sum_{\text{ferm}} \underbrace{\left( \underbrace{R_i - 1}_{\text{scal}} \right)^\alpha}_{\text{ferm}}$$

$$\rightarrow \text{Tr}(R) = N^2 \left[ \left(\frac{1}{2} - 1\right) \cdot 4 + 2 \right] = 0$$

$\uparrow \quad \uparrow$

4                   $\lambda \quad (R[\lambda]=1)$

$$\text{Tr}(R^3) = N^2 \left( \left(\frac{1}{2} - 1\right)^3 \cdot 4 + 2 \right) = N^2 \cdot \frac{3}{2}$$

$$\Rightarrow a = c = \frac{27}{64} N^2 = \frac{\pi^3 N^2}{4 \text{vol}(T^{1,1})}$$

We can forget  $\mathbb{Z}_2$  symm.  $\Rightarrow$  1 eq for 3 coupl  $\Rightarrow$   $\exists$  2d manifold

$$\rightarrow \underbrace{R[A] + R[B]}_x = 1$$

$$\hookrightarrow \text{Tr } R = 0 \quad \forall x; \quad \text{Tr } R^3 = N^2 \left( 2(x-1)^3 + 2(-x)^3 + 2 \right)$$

$$\Rightarrow \alpha = \frac{27}{16} N^2 \times (1-x) \Rightarrow$$

$\hookrightarrow \exists \alpha_{\max} ? \Rightarrow \alpha(x_{\max} = \frac{1}{2}) = \frac{27}{64} N^2$

$\alpha$ -MAXIM.: CFT  $\Rightarrow \alpha$  is max in  
R-symm. space

$\Rightarrow$  the geom. dual is:

SASAKI-EINSTEIN /  $W=1$  SCFT

(in  $d=2$  it's C-maximization).