

# Astrodynamics HW 3

2011135042 Tae Hyeun Kim

1.

$$\mathbf{r}_{in} = 7000km$$

## Kepler's 3rd law

$$P^2 = \frac{4\pi^2}{\mu} a^3$$

$$\mu = 398600.4418km^3/s^2$$

$$P_{fin} = 12hours \times 60min \times 60sec$$

In [17]: Rin = 7000;

In [16]: mu = 398600.4418;

In [13]: Pfin = 12\*3600;

In [14]: Rfin = cbtr(mu\*Pfin^2/4/pi^2);

$$\mathbf{r}_{fin} = 26610.2228km$$

In [15]: Atrans = (Rin + Rfin)/2;

$$a_{trans} = \frac{r_{in} + r_{fin}}{2}$$

$$= 16805.1114km$$

In [16]: e = 1-(Rin/Atrans);

## using vis-viva equation

$$\Delta V = \sqrt{\frac{2\mu}{r} - \frac{\mu}{a}}$$

by vis-viva equation

$$\mathbf{V}_{in} = \sqrt{\frac{2\mu}{r_{in}} - \frac{\mu}{a_{in}}}, \mathbf{V}_{fin} = \sqrt{\frac{2\mu}{r_{fin}} - \frac{\mu}{a_{fin}}}$$

$$\mathbf{V}_{trans,a} = \sqrt{\frac{2\mu}{r_{in}} - \frac{\mu}{a_{trans}}}, \mathbf{V}_{trans,b} = \sqrt{\frac{2\mu}{r_{fin}} - \frac{\mu}{a_{trans}}}$$

$$\Delta \mathbf{V}_a = \sqrt{\frac{\mu}{r_{in}}} - \sqrt{\frac{2\mu}{r_{in}} - \frac{\mu}{a_{trans}}}$$

$$\Delta \mathbf{V}_b = \sqrt{\frac{\mu}{r_{fin}}} - \sqrt{\frac{2\mu}{r_{fin}} - \frac{\mu}{a_{trans}}}$$

```
In [10]: Vin = sqrt((2mu/Rin)-(mu/Rin));
```

```
In [44]: Vfin = sqrt((2*mu/Rfin)-(mu/Rfin));
```

```
In [45]: Vtrans_a = sqrt((2mu/Rin)-(mu/Atrans));
```

```
In [46]: Vtrans_b = sqrt((2mu/Rfin)-(mu/Atrans));
```

```
In [50]: deltaVa = Vtrans_a - Vin;
```

```
In [52]: deltaVb = Vfin - Vtrans_b;
```

$$\begin{cases} \Delta V_a = 1.9496 km/s \\ \Delta V_b = 1.3724 km/s \end{cases}$$

```
In [53]: deltaV = deltaVa + deltaVb;
```

$$\begin{aligned} \Delta V &= \Delta V_a + \Delta V_b \\ &= 3.3220 km/s \end{aligned}$$

**2.**

$$Period = 318 min$$

```
In [14]: P = 318*60;
```

```
In [18]: mu = 398600.4418;
```

```
In [28]: a = cbrt(P^2*mu/4/pi^2);
```

```
In [56]: incl = 10*2pi/360;
```

In identical equatorial plane, we can do inclination-only maneuver on two nodes which is ascending and descending nodes.

$$P^2 = \frac{4\pi^2}{\mu} a^3$$

$$a = 15433km, e = 0.5$$

```
In [5]: e = 0.5;
```

```
In [7]: nu = 45/360*2pi;
```

```
In [36]: nu2 = (180+45)/360*2pi;
```

```
In [29]: r = a*(1-e^2)/(1+0.5*cos(nu));
```

```
In [41]: r2 = a*(1-e^2)/(1+0.5*cos(nu2));
```

```
In [30]: a*(1-e^2);
```

$$\begin{aligned} r &= \frac{a(1-e^2)}{1+e\cos\nu} (km) \\ &= \frac{15432.8071(1-0.5^2)}{1+0.5\cos\nu} (km) \\ &= \frac{11574.6053}{1+0.5\cos\nu} (km) \end{aligned}$$

**by vis-viva equation**

$$\begin{aligned} v &= \sqrt{\frac{2\mu}{r} - \frac{\mu}{a}} \\ &= \sqrt{\frac{\mu}{r} \left( 2 - \frac{1-e^2}{1+0.5\cos\nu} \right)} \end{aligned}$$

```
In [45]: Vasc = sqrt((mu/r)*(2-((1-e^2)/(1+0.5*cos(nu)))));
```

```
In [46]: Vdesc = sqrt((mu/r2)*(2-((1-e^2)/(1+0.5*cos(nu2)))));
```

at descending node,

$$V_{descending} = 4.3239km/s$$

at ascending node,

$$V_{ascending} = 8.2096 \text{ km/s}$$

$$\Delta \vec{V}_{i,only} = \vec{V}_{fin} - \vec{V}_{in}$$

```
In [62]: cosfpa1 = sqrt((1-e^2)*a^2/r/(2a-r));
```

```
In [63]: cosfpa2 = sqrt((1-e^2)*a^2/r2/(2a-r2));
```

```
In [61]: delV_a = 2cosfpa1*Vasc*sin(incl/2);
```

```
In [64]: delV_d = 2cosfpa2*Vdesc*sin(incl/2);
```

1) ascending node

$$\sin(5^\circ) = \frac{\Delta V_{i,only,ascending}/2}{2\cos(\phi_{fpa})V_{ascending}}$$

$$\Rightarrow \Delta V_{i,only,ascending} = 2\cos(\phi_{fpa})V_{ascending}\sin(5^\circ)$$

$$= 1.3846 \text{ km/s}$$

2) descending node

$$\sin(5^\circ) = \frac{\Delta V_{i,only,descending}/2}{2\cos(\phi_{fpa})V_{descending}}$$

$$\Rightarrow \Delta V_{i,only,descending} = 2\cos(\phi_{fpa})V_{descending}\sin(5^\circ)$$

$$= 0.6613 \text{ km/s}$$

**3.**

$$altitude_{sat1} = 191.344 \text{ km}$$

$$altitude_{GEO} = 35781.35 \text{ km}$$

Note. the initial and final orbits are circular, whereas the transfer orbits are elliptical.

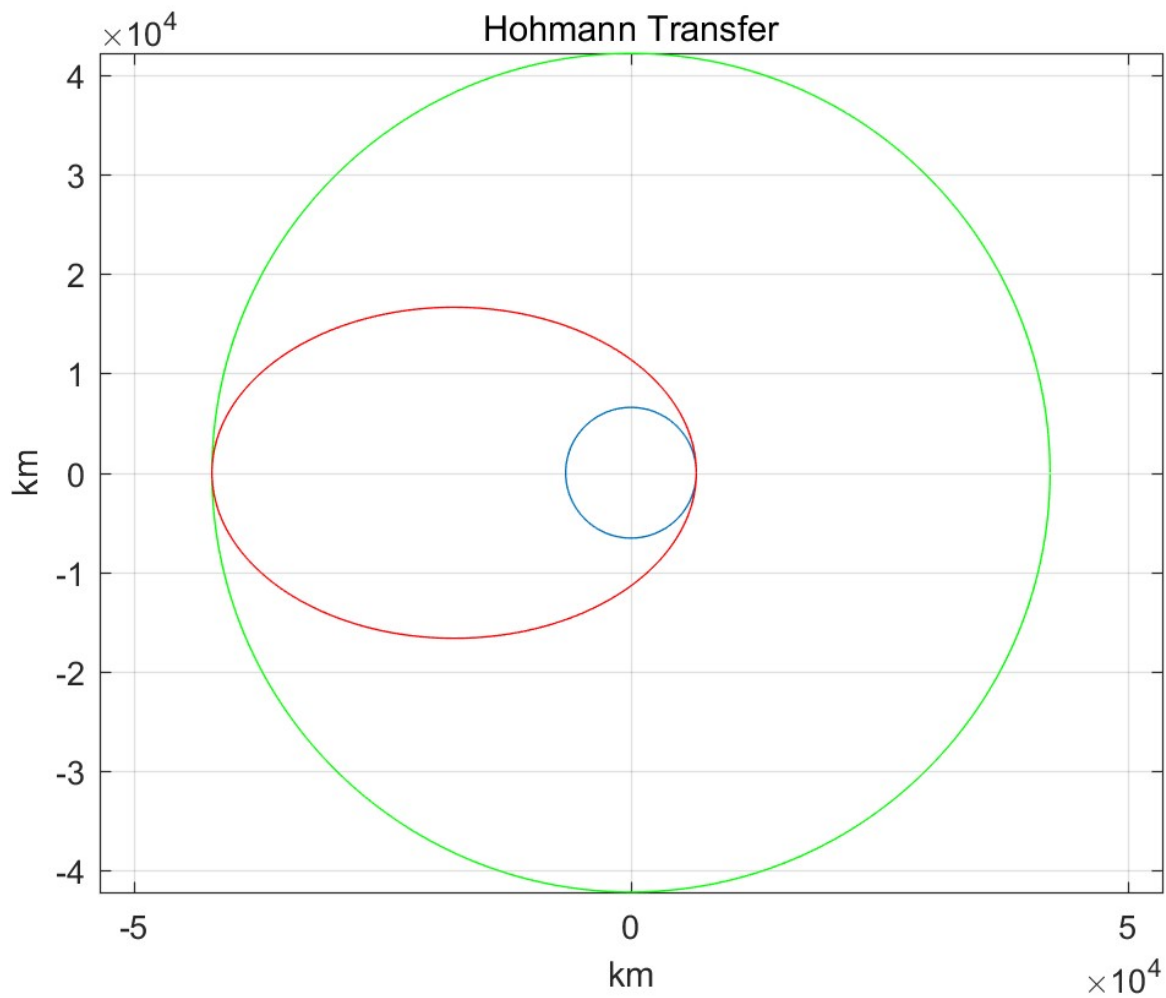
For the One-tangent burn transfer, it is given as follows.

$$\nu_{trans,b} = 160^\circ$$

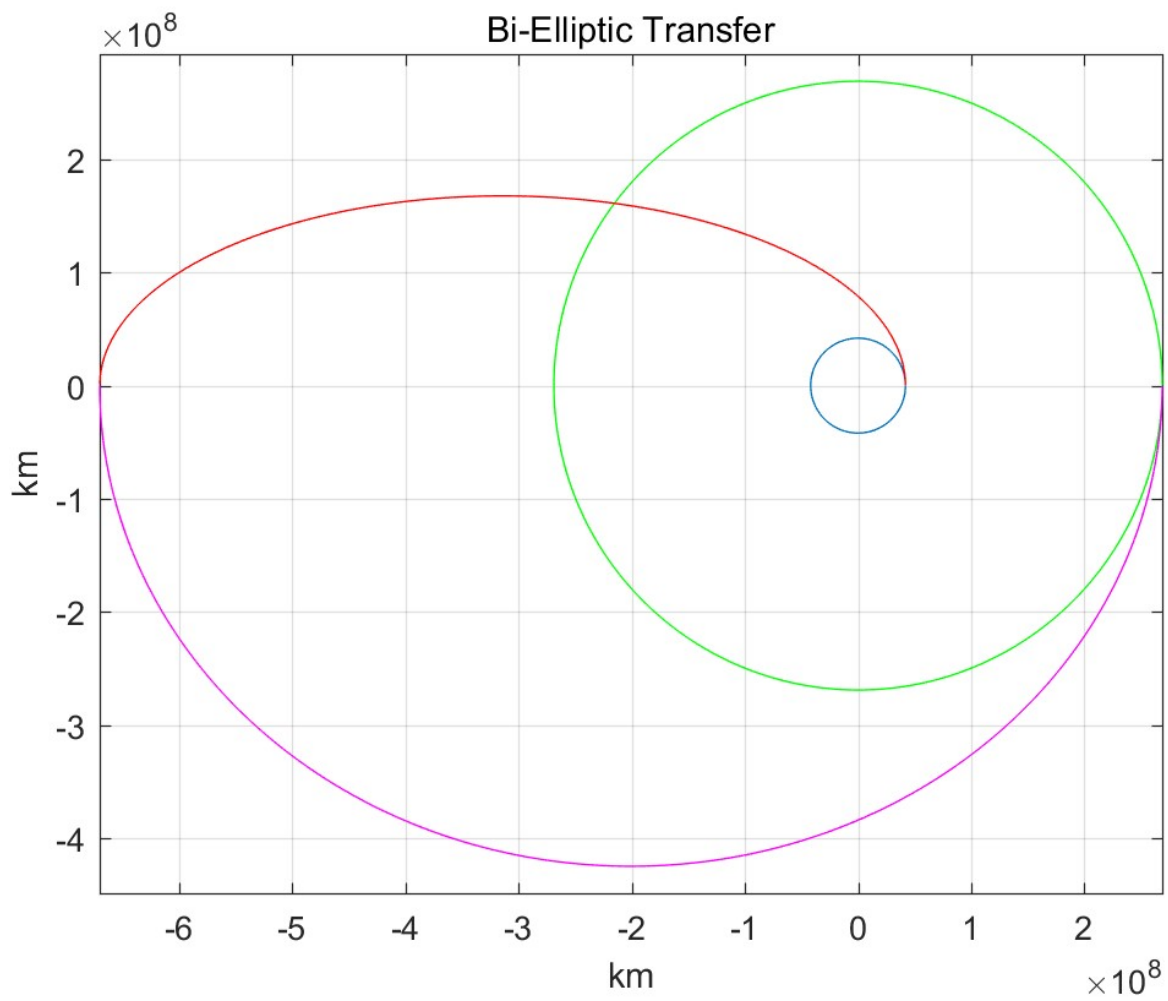
the total change in velocity, transfer time and comparing the results from each transfer.

hohmann(191.344+re, 35781.35+re, 0, 0, 0, pi)

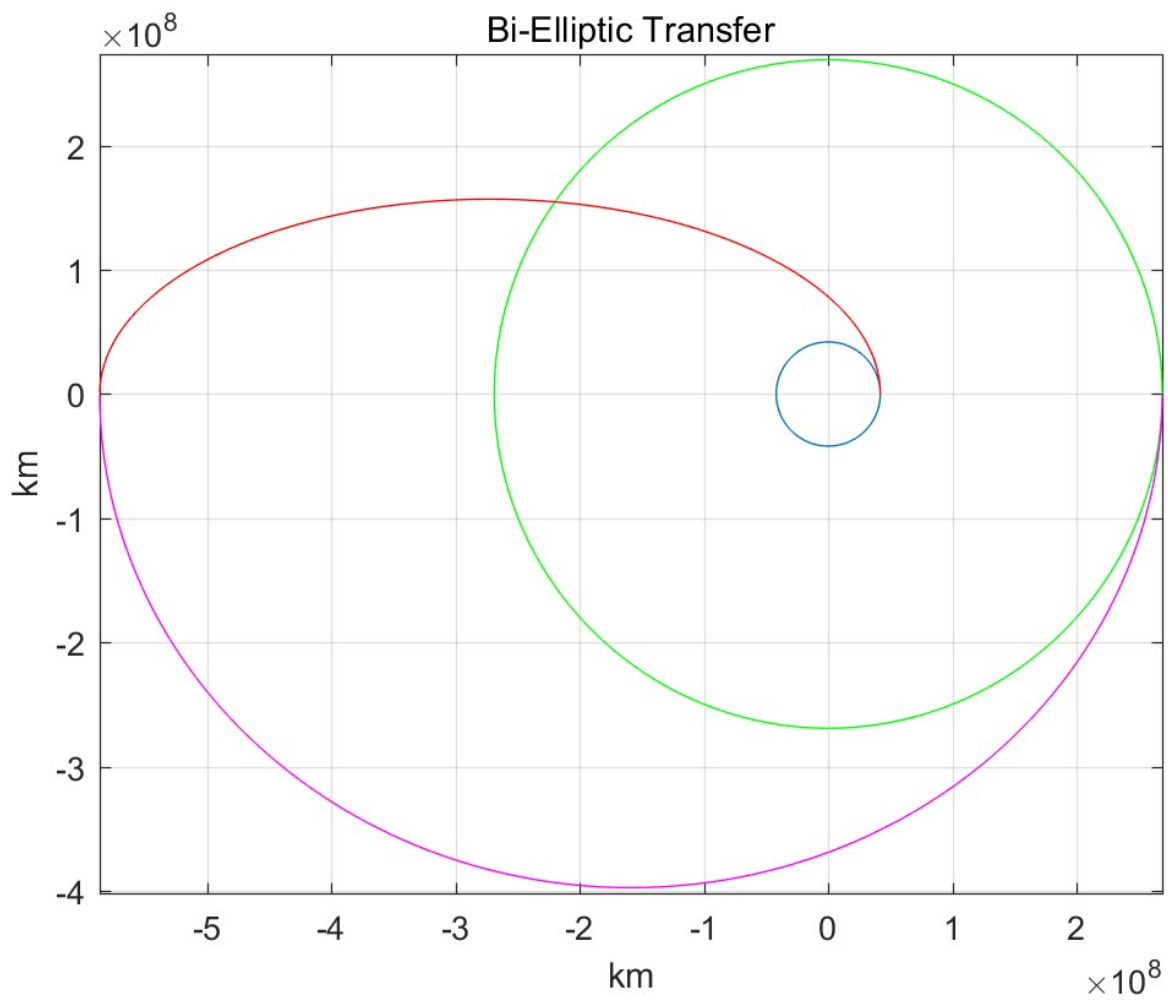
$$\begin{cases} \Delta V_a = V_{trans,a} - V_{initial} \\ \Delta V_b = V_{final} - V_{trans,b} \\ \Delta V_{total} = \Delta V_a + \Delta V_b \end{cases}$$



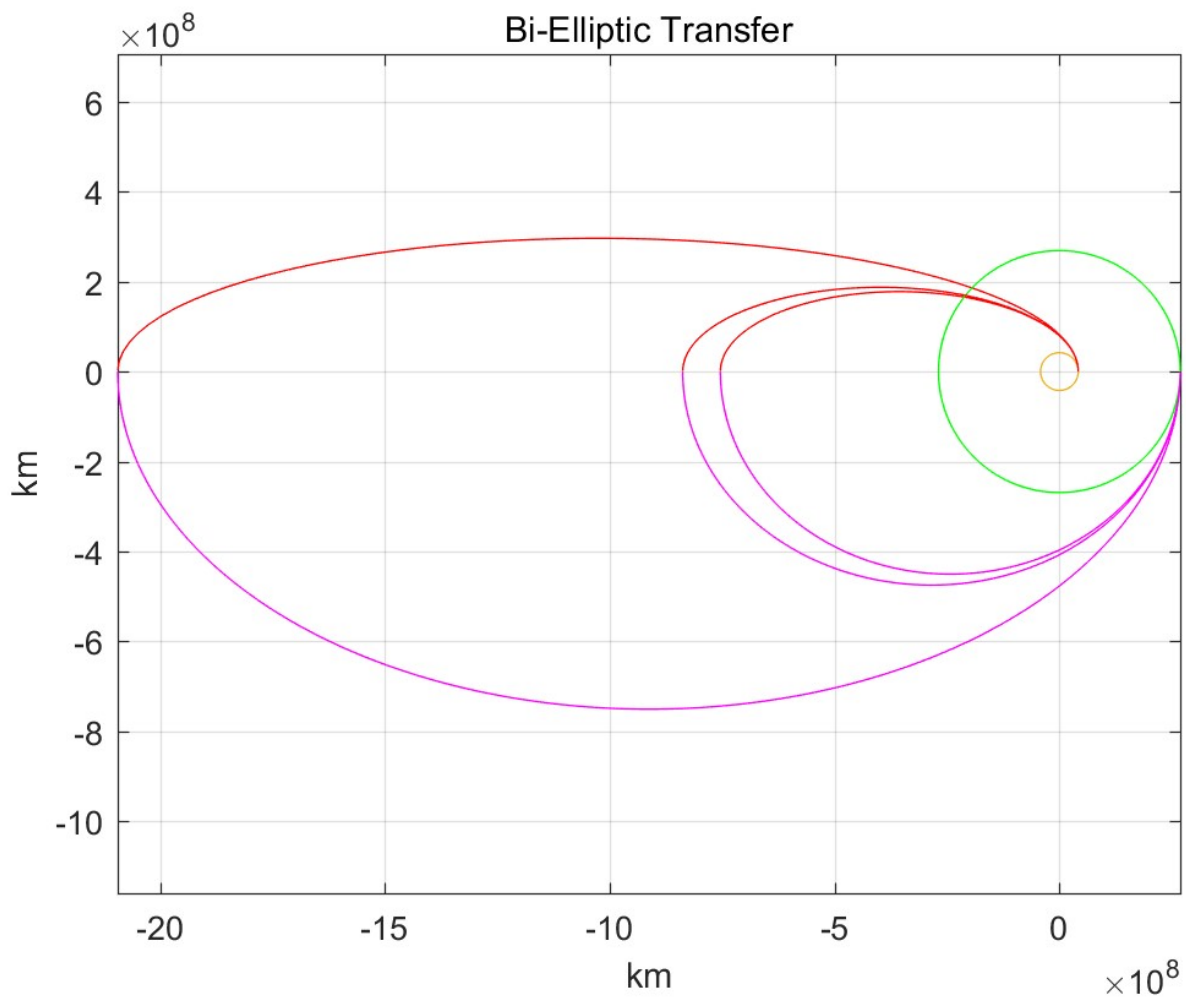
```
biellip(191.344+re,105112,35781.35+re,0, 0, 0, pi)
```



```
biellip(191.344+re,91973,35781.35+re,0, 0, 0, pi)
```

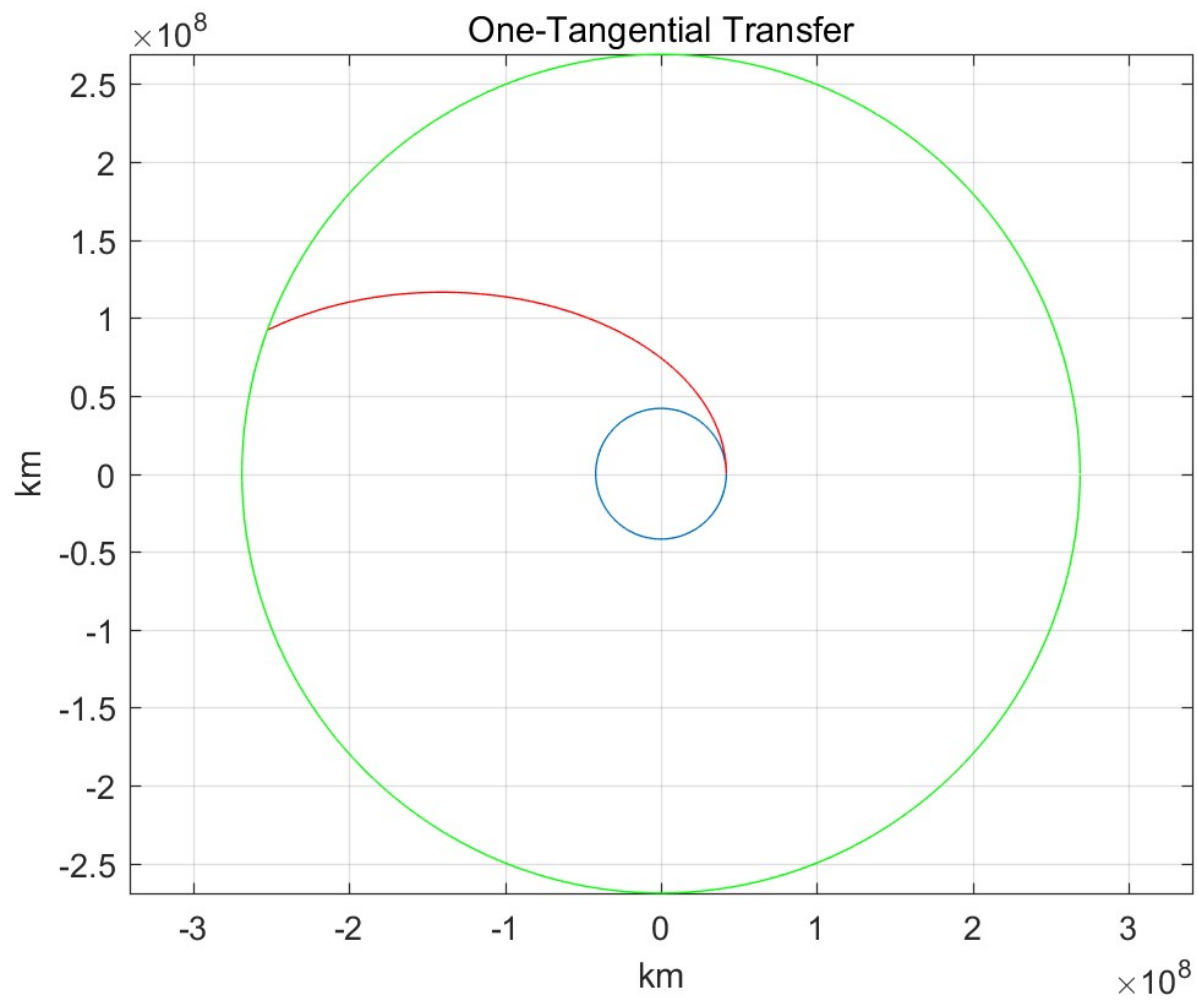


```
biellip(191.344+re,328475,35781.35+re,0, 0, 0, pi)
```



onetang(191.344+re,35781.35+re,0,0,0,160/180\*pi)





In [ ]: