

# Homework 9



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Thorne Wolfenbarger  
wolfent1@my.erau.edu

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Professor Kaela Martin  
College of Engineering

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## AE 313 Homework 9

2. Draw the heliocentric trajectory of Insight.

On separate page.

3. Determine Earth's and InSight's heliocentric properties  $(r, v, \gamma)$  at departure.

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```
1 r_earth = r('Earth');
2 r_mars = r('Mars');
3
4 v_earth = sqrt(MU('Sun')/r_earth);
5 FPA_earth = 0;
6 ...
```

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$$\begin{aligned} r_{Earth} = r_{InSight} &= 149600000 \text{ km} \\ v_{Earth} &= 29.7845 \text{ km/s} \\ v_{InSight} &= 32.7285 \text{ km/s} \\ \gamma_{Earth} = \gamma_{InSight} &= 0^\circ \end{aligned}$$

4. Draw the vector diagram in the heliocentric orbit and determine  $v_{\infty/Earth}^+$ .

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```
1 v_inf_earth = v_sc1 - v_earth;
```

---

Diagram on separate page.

$$v_{\infty/Earth}^+ = 2.944 \text{ km/s}$$

5. What is the  $\Delta v_1$  to depart Earth? Don't forget the vector diagram!

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```
1 energy_inf_earth = v_inf_earth^2/2;
2 r_sc_earth = 300 + R('Earth');
3 ...
```

---

Diagram on separate page.

$$\Delta v_1 = 3.5898 \text{ km/s}$$

6. Now move to the capture portion of the orbit. Determine Mars' and InSight's heliocentric properties at arrival  $(r, v, \gamma)$ .

---

```
1 r_mars = r('Mars');
2 v_mars = sqrt(MU('Sun')/r_mars);
3 FPA_mars = 0;
4 ...
```

---

$$\begin{aligned} r_{Mars} = r_{InSight} &= 227920000 \text{ km} \\ v_{Mars} &= 24.1304 \text{ km/s} \\ v_{InSight} &= 21.4920 \text{ km/s} \\ \gamma_{Mars} = \gamma_{InSight} &= 0^\circ \end{aligned}$$

7. Draw the vector diagram in the heliocentric orbit and determine  $v_{\infty/Mars}^-$ .

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1 `v_inf_mars = abs(v_sc2 - v_mars);`

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Diagram on separate page.

$$v_{\infty/Mars}^- = 2.6484 \text{ km/s}$$

8. What is the required  $\Delta v_2$  to capture at Mars? Don't forget the vector diagram!

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1 `dv_capture = v_mars_circ - v_mars_peri;`

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Diagram on separate page.

$$\Delta v_2 = 2.1075 \text{ km/s (slowing down)}$$

9. Are the required  $\Delta v$  values reasonable? Why or why not?

The  $\Delta v$  values are quite small compared to many transfer maneuvers. Given the small size of a lander these  $\Delta v$  values are not unreasonable for the resource costs associated with them. Therefore they are reasonable.

```
1 clc; clear;
2 % Constants
3 planets = {'Sun', 'Moon', 'Mercury', 'Venus', 'Earth', 'Mars', 'Jupiter',
             'Saturn', 'Uranus', 'Neptune', 'Pluto'};
4 rad_list = [695990.0, 1739.2, 2439.7, 6051.9, 6378.0, 3397.0, 71492.0,
             60268.0, 25559.0, 25269.0, 1162.0];
5 mu_list = [132712440000.0, 4902.8, 22032.0, 324860.0, 398600.0, 42828.0,
            126713000.0, 37941000.0, 5794500.0, 6836500.0, 981.6];
6 sma_list = [0.0, 384400.0, 57910000.0, 108210000.0, 149600000.0,
            227920000.0, 778570000.0, 1433530000.0, 2872460000.0, 4495060000.0,
            5906380000.0];
7 R = containers.Map(planets,rad_list);
8 MU = containers.Map(planets,mu_list);
9 r = containers.Map(planets,sma_list);
10 %%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%%
11 % 3. Determine Earth and InSight heliocentric properties (r,v GPA) at
12 % departure
13 r_earth = r('Earth');
14 r_mars = r('Mars');
15
16 v_earth = sqrt(MU('Sun')/r_earth);
17 FPA_earth = 0;
18
19 r_sc1 = r_earth;
20 FPA_sc1 = 0;
21
22 r_sc2 = r_mars;
23 a_t = (r_sc1 + r_sc2)/2;
24
25 syms v_sc1
26 eq = -MU('Sun')/(2*a_t) == v_sc1^2/2 - MU('Sun')/r_sc1;
27 v_sc1 = max(double(solve(eq, v_sc1)));
28 % v_sc1 = sqrt(2*MU('Sun')/r_sc1 - MU('Sun')/a_t)
29
30 % 4. Determine v_inf_earth
31
32 v_inf_earth = v_sc1 - v_earth;
33
34 % 5. Required dv to depart earth
35 energy_inf_earth = v_inf_earth^2/2;
36 r_sc_earth = 300 + R('Earth');
37 % dv = sqrt(v_inf_earth^2 + 2*MU('Earth')/r_sc_earth) - sqrt(MU('Earth')/
    r_sc_earth);
```

```

38 syms v_earth_peri
39 eq = energy_inf_earth == v_earth_peri^2/2 - MU('Earth')/r_sc_earth;
40 v_earth_peri = max(double(solve(eq, v_earth_peri)));
41 v_earth_circ = sqrt(MU('Earth')/r_sc_earth);
42 dv_escape = v_earth_peri - v_earth_circ;
43
44 % 6. Mars and insight properties at arrival (r, v, FPA)
45 r_mars = r('Mars');
46 v_mars = sqrt(MU('Sun')/r_mars);
47 FPA_mars = 0;
48
49 r_sc2 = r_mars;
50 syms v_sc2
51 eq = -MU('Sun')/(2*a_t) == v_sc2^2/2 - MU('Sun')/r_sc2;
52 v_sc2 = max(double(solve(eq, v_sc2)));
53 FPA_sc2 = 0;
54
55 % determine v_inf_mars
56 v_inf_mars = abs(v_sc2 - v_mars); % Positive by convention
57 energy_inf_mars = v_inf_mars^2/2;
58 r_sc_mars = 150 + R('Mars');
59
60 syms v_mars_peri
61 eq = energy_inf_mars == v_mars_peri^2/2 - MU('Mars')/r_sc_mars;
62 v_mars_peri = max(double(solve(eq, v_mars_peri)));
63 v_mars_circ = sqrt(MU('Mars')/r_sc_mars);
64 dv_capture = v_mars_circ - v_mars_peri;

```

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