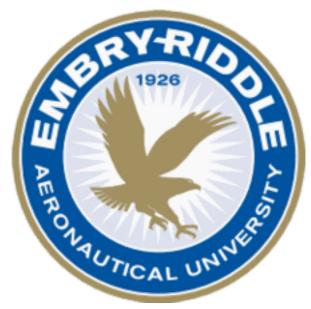
Homework 8



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AE 313 Space Mechanics Spring, 2019

AE 313 Homework 8

1. Express $\Delta \vec{v}$ in rotating orbit unit vectors $\hat{r}, \hat{\theta}\hat{h}$ as well as inertial unit vectors $\hat{x}, \hat{y}, \hat{z}$.

```
vdv_rth = dv*[cosd(beta)*sind(phi) cosd(beta)*cosd(phi) sind(beta)]';
vdv_eci = rot_rth_eci(RAANm, incm, AOLm) * vdv_rth;
```

```
\Delta \vec{v}_{rth} = < -2.6250, \ 2.2344, \ 2.8925 > km/s

\Delta \vec{v}_{eci} = < 3.9498, \ -0.8593, \ 1.9776 > km/s
```

2. Determine the position and velocity immediately after the maneuver, \vec{r}^+, \vec{v}^+ in the intertial coordinate system.

```
1 vrp_eci = vrm_eci;
2 vvp_eci = vvm_eci + vdv_eci;
```

```
\begin{array}{lll} \vec{r}_{eci}^{+} = & <-5.9784, & -4.6680, & -0.1583 > \cdot 10^3 km \\ \vec{v}_{eci}^{+} = & <11.6963, & -5.4791, & -1.1101 > km/s \end{array}
```

3. Compute the orbital elements e^+ , i^+ , Ω^+ , θ^+ , θ^{*+} in the new orbit.

```
1 FPAp = asind(dot(vrp_eci,vvp_eci)/(norm(vrp_eci)*norm(vvp_eci)));
2 ...
```

 $3 \text{ true_ap} = 360 - \text{true_ap};$

```
e^{+} = 2.0151

i^{+} = 6.2208^{\circ}

\Omega^{+} = 26.9440^{\circ}

\theta^{+} = 191.1029^{\circ}

\theta^{*+} = 320.4327^{\circ}
```

4. Find the changes in the elements (including the sign) that occurred due to the maneuver, that is, Δe , Δi , $\Delta \Omega$, $\Delta \theta$.

```
1 de = ep—em;
2 dinc = incp—incm;
3 dRAAN = RAANp—RAANm;
4 dAOL = AOLp—AOLm;
```

```
\Delta e = 1.2501
\Delta i = -14.3792^{\circ}
\Delta \Omega = -7.8560^{\circ}
\Delta \theta = 7.7029^{\circ}
```

5. Confirm the position, m and new orbital elements (problem 3) in GMAT. Plot the original and new orbit (include XY plane and inertial unit vectors). Mark the maneuver location on the plot.

```
 \vec{v}_{eci}^- = <7.7465, -4.6198, -3.0877 > km/s   \vec{v}_{eci}^+ = <12.026, -5.1957, -1.0988 > km/s
```

The data in the GMAT orbit is very close to the MATLAB calculations and therefore matches the data from MATLAB calculations.

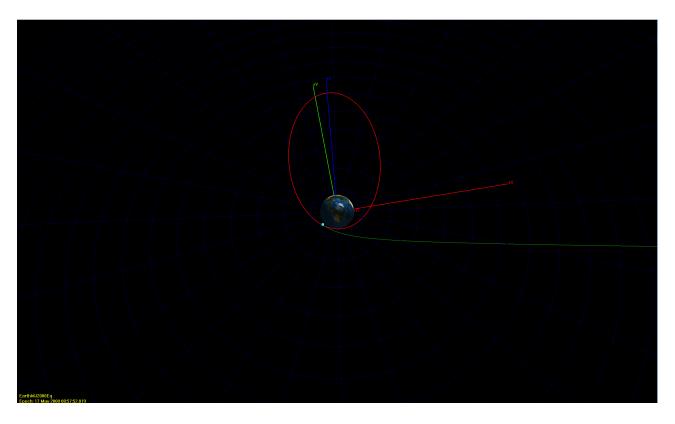


Figure 1: GMAT Plot of Orbits

Turquoise indicates the maneuver point. Clearer view of the maneuver after the code.

- 6. Would you want to perform this maneuver? Why or why not? I would not want to perform this maneuver because it results in a massively hyperbolic orbit. The goal of the maneuver is to do an orbital correction, not enter an escape trajectory. With an eccentricity of 2.02, the orbit is very far from being elliptical.
- 7. Given that the semi-major axes of bi-elliptical transfer are $a_{T1} = 6659 \ km$ and $a_{T2} = 6798 \ km$, what is the departure phase angle? Include a figure/sketch of the phase angle.

```
at1 = 6658;

at2 = 6798;

a_sc = 6378+430;

TOF = pi*(sqrt(at1^3/MU('Earth'))+sqrt(at2^3/MU('Earth')));

n_sc = sqrt(MU('Earth')/a_sc^3);

phase = 2*pi-n_sc*TOF;

phase = phase*180/pi;

\Delta\Phi = 6.3124^{\circ}
```

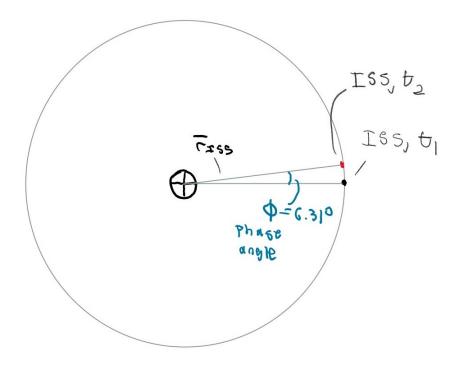


Figure 2: Sketch of the Phase Angle

8. Survey.

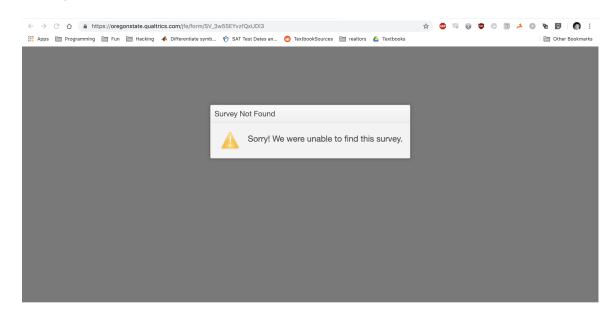


Figure 3: Survey Screenshot

HW8.m

```
1 clc; clear;
 2 % Constants
3 planets = {'Sun', 'Moon', 'Mercury', 'Venus', 'Earth', 'Mars', 'Jupiter',
       'Saturn', 'Uranus', 'Neptine', 'Pluto'};
 4 rad_list = [695990.0, 1739.2, 2439.7, 6051.9, 6378.0, 3397.0, 71492.0,
       60268.0, 25559.0, 25269.0, 1162.0];
 5 \text{ mu\_list} = [132712440000.0, 4902.8, 22032.0, 324860.0, 398600.0, 42828.0,
       126713000.0, 37941000.0, 5794500.0, 6836500.0, 981.6];
 6 sma_list = [0.0, 384400.0, 57910000.0, 108210000.0, 149600000.0,
       227920000.0, 778570000.0, 1433530000.0, 2872460000.0, 4495060000.0,
       5906380000.0];
 7 R = containers.Map(planets, rad_list);
8 MU = containers.Map(planets,mu_list);
9 r = containers.Map(planets,sma_list);
11 am = 28081;
12 \text{ AOPm} = 229;
13 RAANm = 34.8;
14 \text{ em} = 0.765;
15 \text{ incm} = 20.6;
16 true_am = -45.6;
17
18 \text{ AOLm} = \text{AOPm} + \text{true\_am};
19
20 \text{ vrm\_eci} = [-5978.4 - 4668 - 158.31]';
21 vvm_eci = [7.7465 -4.6198 -3.0877]';
22
23 \text{ dv} = 4.5;
24 \text{ alpha} = -30;
25 beta = 40;
26 \text{ pm} = \text{am}*(1-\text{em}^2);
27 % FPA = acosd(norm(cross(vrm_eci, vvm_eci))/(norm(vrm_eci)*norm(vvm_eci)))
28 FPA = atan2d(norm(vrm_eci)*em*sind(true_am),pm);
29 phi = alpha + FPA;
31 % 1.
32
33 vdv_rth = dv*[cosd(beta)*sind(phi) cosd(beta)*cosd(phi) sind(beta)]';
34 vdv_eci = rot_rth_eci(RAANm, incm, AOLm) * vdv_rth;
35
36 % 2.
37 vrp_eci = vrm_eci;
38 vvp_eci = vvm_eci + vdv_eci;
```

```
39
40 % 3.
41 FPAp = asind(dot(vrp_eci,vvp_eci)/(norm(vrp_eci)*norm(vvp_eci)));
42 ep = sqrt((((norm(vrp_eci)*norm(vvp_eci)^2/MU('Earth'))-1)^2*cosd(FPAp)^2)
       +sind(FPAp)^2);
43
44 \;\; \mathrm{syms} \;\; \mathrm{ap}
45 eq = norm(vvp_eci)^2/2 - MU('Earth')/norm(vrp_eci) == MU('Earth')/(2*ap);
46 ap = double(solve(eq, ap));
47 pp = norm(cross(vrp_eci,vvp_eci))^2/MU('Earth');
48 h_hat = cross(vrp_eci, vvp_eci)/norm(cross(vrp_eci,vvp_eci));
49
50 incp = acosd(h_hat(3));
51
52 syms RAANp
53 \text{ eq}(1) = h_{\text{hat}}(2) == -\cos((RAANp) * \sin((incp));
54 \text{ eq(2)} = h_{\text{hat}(1)} == \text{sind}(\text{RAANp})*\text{sind}(\text{incp});
55 RAANp1 = acosd(-h_hat(2)/sind(incp));
56 RAANp1 = [-RAANp1 RAANp1];
57 RAANp2 = asind(h_hat(1)/sind(incp));
58 \text{ RAANp2} = [180-\text{RAANp2} \text{ RAANp2}];
59
60 RAANp = min(RAANp2); %intersecting the two doesnt work due to small delta
       in solution. It's the negative value.
61
62 rp_eci_hat = vrp_eci/norm(vrp_eci);
63 theta_hat = cross(rp_eci_hat,h_hat);
64
65 AOLp1 = asind(rp_eci_hat(3)/sind(incp));
66 AOLp1 = [180—AOLp1 AOLp1];
67 AOLp2 = acosd(theta_hat(3)/sind(incp));
68 \text{ AOLp2} = [-A0Lp2 \text{ AOLp2}];
69 AOLp = min(AOLp2); %same as above. Small delta does not allow me to do
       intersection
70 \text{ AOLp} = 191.1029;
71
72 true_ap = acosd((pp/norm(vrp_eci)-1)/ep); %Checked Gmat its neg
73 true_ap = 360 - true_ap;
74
75 % 4.
76 de = ep-em;
77 dinc = incp—incm;
78 dRAAN = RAANp—RAANm;
79 \text{ dAOL} = AOLp-AOLm;
```

```
80
 81 % 5.
 82 vrp_eci; %-7192.67008574393
                                            -3847.023676019902
        355.568234258645
                                      303.8898424306826
 83
 84 % 6.
 85 % I would not want to perform this maneuver because it results in a
 86 % massively hyperbolic orbit. The goal of the maneuver is to do an orbital
 87 % correction, not enter an escape trajectory. With an eccentricity of
        2.02.
 88 % the orbit is very far from bring elliptical.
 89
 90 % 7.
 91 \text{ at1} = 6658;
 92 at2 = 6798;
 93 \text{ a\_sc} = 6378+430;
 94 TOF = pi*(sqrt(at1^3/MU('Earth'))+sqrt(at2^3/MU('Earth')));
 95 \text{ n_sc} = \text{sqrt}(MU('Earth')/a_sc^3);
 96 phase = 2*pi-n_sc*T0F;
97 phase = phase*180/pi;
98
99 function A = rot_rth_eci(o,i,t)
100
101 % o : Omega, Longitude of the Ascending Node (RAAN)
102~\% i : i, inclination
103 % t : theta, Argument of Latitude
104
105 A = [\cos d(o) * \cos d(t) - \sin d(o) * \cos d(i) * \sin d(t), -\cos d(o) * \sin d(t) - \sin d(o) *
        cosd(i)*cosd(t), sind(o)*sind(i); ...
106
             sind(o)*cosd(t)+cosd(o)*cosd(i)*sind(t), -sind(o)*sind(t)+cosd(o)*
                cosd(i)*cosd(t), -cosd(o)*sind(i); ...
107
             sind(i)*sind(t), sind(i)*cosd(t), cosd(i)];
108
109 end
110
111 function A = rot_eci_ecef(t)
112
113 % t : theta_era, Earth rotation angle
114
115 A = [\cos d(t), \sin d(t), 0; -\sin d(t), \cos d(t), 0; 0, 0, 1];
116 end
117
118 function A = rot_ecef_sez(l,p)
119
```

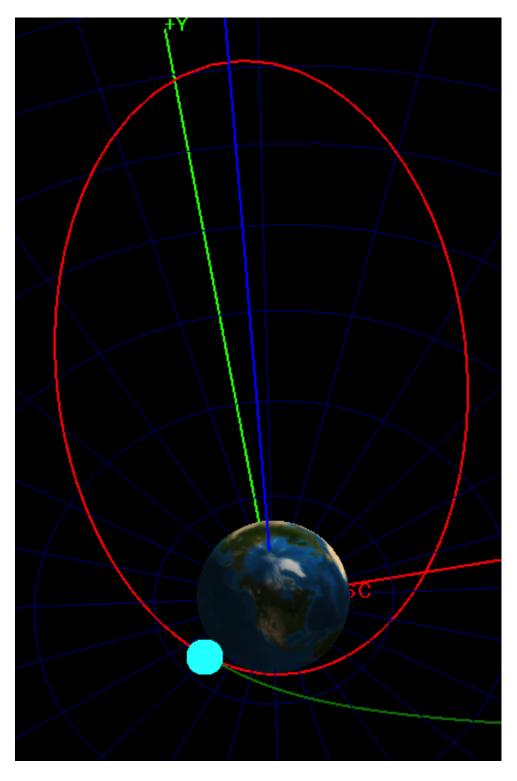


Figure 4: GMAT Modified

Word has changed their background removal functionality such that I cannot get rid of the black. This is a zoomed photo to help if the first one is difficult to read. Turquoise indicates the maneuver point.