

Scattering Theory

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Part I

Quantum Mechanics

This article is designed for leading the reader from old scattering theory of quantum mechanics to nowadays scattering theory of quantum field theory. The quantum mechanics part would be based on *Lectures on Quantum Mechanics* by S.Weinberg.

1 Potential Scattering

In this chapter we would study the theory of scattering in a simple but important case, the elastic scattering of a non-relativistic particle in a local potential, but using modern techniques that could easily be extended to more general problems.

1.1 In-States

Consider a non-relativistic particle of mass μ in a potential $V(\mathbf{x})$, whose Hamiltonian is,

$$H = H_0 + V(\mathbf{x}) \quad (1.1)$$

Here $H_0 = \mathbf{p}^2/2\mu$ is the kinetic energy operator. The potential is assumed to be function of position operator, and tends to zero as $|\mathbf{x}| = r \rightarrow \infty$. Then it is concerned here with a positive-energy particle¹. The particle comes into facing the potential from great distances regarded to have no interaction with momentum $\hbar\mathbf{k}$, and is scattered, going out again to infinity, generally along a different direction.

In Heisenberg picture, for a particle with momentum $\hbar\mathbf{k}$ far from the scattering center if the measurement of the particle is made at very early times, and this situation is represented by a time-independent state vector $|\Psi_{\mathbf{k}}^{\text{in}}\rangle$. As mentioned, at very early times the particle is at a location that is far from the scattering center

¹Actually, there could possibly be negative energy if it is in bound state, that is, the energy is less than potential at boundaries. But now, we are considering scattering state, which is opposite with bound state.

so that the potential is negligible there, so the state has an energy $\hbar^2 \mathbf{k}^2 / 2\mu$, with the following relation,

$$H |\Psi_{\mathbf{k}}^{\text{in}}\rangle = \frac{\hbar^2 \mathbf{k}^2}{2\mu} |\Psi_{\mathbf{k}}^{\text{in}}\rangle \quad (1.2)$$

Now switch to Schrödinger picture, the state would evolve as time goes on. As mentioned, the scattering state would be continuous *superposition*² of states with a spread of energies,

$$|\Psi_g(t)\rangle = \int d^3k / (2\pi)^3 g(\mathbf{k}) \exp\left(-i \frac{\hbar t \mathbf{k}^2}{2\mu}\right) |\Psi_{\mathbf{k}}^{\text{in}}\rangle \quad (1.3)$$

where $g(\mathbf{k})$ is a smooth function that is peaked around some wave number \mathbf{k}_0 . The eigenstate $|\Psi_{\mathbf{k}}^{\text{in}}\rangle$ satisfies the further condition that for any sufficiently smooth function $g(\mathbf{k})$, in the limit $t \rightarrow -\infty$,

$$|\Psi_g(t)\rangle \rightarrow \int d^3k / (2\pi)^3 g(\mathbf{k}) \exp\left(-i \frac{\hbar t \mathbf{k}^2}{2\mu}\right) |\Phi_{\mathbf{k}}\rangle \quad (1.4)$$

where $|\Phi_{\mathbf{k}}\rangle$ are orthonormal eigenstates of the momentum operator \mathbf{P} with eigenvalue $\hbar \mathbf{k}$,

$$\begin{aligned} \mathbf{P} |\Phi_{\mathbf{k}}\rangle &= \hbar \mathbf{k} |\Phi_{\mathbf{k}}\rangle \\ \langle \Phi_{\mathbf{k}} | \Phi_{\mathbf{k}'} \rangle &= (2\pi)^3 \delta^{(3)}(\hbar \mathbf{k} - \hbar \mathbf{k}') \end{aligned} \quad (1.5)$$

and hence eigenstates of H_0 (not H), with eigenvalue $E(|\mathbf{k}|) = \hbar^2 \mathbf{k}^2 / 2\mu$. Note that $|\Psi_{\mathbf{k}}^{\text{in}}\rangle$ and $|\Phi_{\mathbf{k}}\rangle$ belong to different Hilbert spaces. The incident wave packet $|\Psi_g(t)\rangle$ should satisfy the normalization condition $\langle \Psi_g(t) | \Psi_g(t) \rangle = 1$, which is equivalent to the condition at time limit,

$$\begin{aligned} &\int \int d^3k d^3k' / (2\pi)^3 g^*(\mathbf{k}) g(\mathbf{k}') \exp\left(-i \frac{\hbar t}{2\mu} (\mathbf{k}'^2 - \mathbf{k}^2)\right) \delta^{(3)}(\hbar \mathbf{k} - \hbar \mathbf{k}') \\ &= \hbar^{-3} \int d^3k / (2\pi)^3 |g(\mathbf{k})|^2 = 1 \end{aligned} \quad (1.6)$$

Rewrite the equation (1.2) as,

$$(E(|\mathbf{k}|) - H_0) |\Psi_{\mathbf{k}}^{\text{in}}\rangle = V |\Psi_{\mathbf{k}}^{\text{in}}\rangle \quad (1.7)$$

²It would be discrete sum for a bound state.

This has a formal solution,

$$|\Psi_{\mathbf{k}}^{\text{in}}\rangle = |\Phi_{\mathbf{k}}\rangle + (E(|\mathbf{k}|) - H_0 + i\epsilon)^{-1} |\Psi_{\mathbf{k}}^{\text{in}}\rangle \quad (1.8)$$

The first term on RHS could always be added up because it is the *homogeneous* solution of equation (1.7), and ϵ is a positive infinitesimal quantity, which is inserted to give meaning to the operator $(E(|\mathbf{k}|) - H_0 + i\epsilon)^{-1}$ when we integrate over the eigenvalues of H_0 ³. It is known as the *Lippmann-Schwinger equation*. The special feature of the particular 'solution' is that it also satisfies the additional initial condition.

To see this, we could expand $V|\Psi_{\mathbf{k}}^{\text{in}}\rangle$ in the orthonormal free-particle states $|\Phi_{\mathbf{q}}\rangle$:

$$V|\Psi_{\mathbf{k}}^{\text{in}}\rangle = \int d^3q/(2\pi)^3 |\Phi_{\mathbf{q}}\rangle \langle\Phi_{\mathbf{q}}| V |\Psi_{\mathbf{k}}^{\text{in}}\rangle \quad (1.9)$$

Thus equation (1.8) becomes,

$$\begin{aligned} |\Psi_{\mathbf{k}}^{\text{in}}\rangle &= |\Phi_{\mathbf{k}}\rangle + \hbar^3 \int d^3q/(2\pi)^3 (E(|\mathbf{k}|) - H_0 + i\epsilon)^{-1} |\Phi_{\mathbf{q}}\rangle \langle\Phi_{\mathbf{q}}| V |\Psi_{\mathbf{k}}^{\text{in}}\rangle \\ &= |\Phi_{\mathbf{k}}\rangle + \hbar^3 \int d^3q/(2\pi)^3 (E(|\mathbf{k}|) - E(|\mathbf{q}|) + i\epsilon)^{-1} |\Phi_{\mathbf{q}}\rangle \langle\Phi_{\mathbf{q}}| V |\Psi_{\mathbf{k}}^{\text{in}}\rangle \end{aligned} \quad (1.10)$$

In calculating the integral over \mathbf{k} in equation (1.3),

$$\begin{aligned} &\int d^3k/(2\pi)^3 g(\mathbf{k}) \frac{\exp(-i\hbar t \mathbf{k}^2/2\mu)}{E(|\mathbf{k}|) - E(q) + i\epsilon} \langle\Phi_{\mathbf{q}}| V |\Psi_{\mathbf{k}}^{\text{in}}\rangle \\ &= \int d\Omega \int_0^\infty dk/(2\pi)^3 k^2 g(\mathbf{k}) \frac{\exp(-i\hbar t \mathbf{k}^2/2\mu)}{E(k) - E(q) + i\epsilon} \langle\Phi_{\mathbf{q}}| V |\Psi_{\mathbf{k}}^{\text{in}}\rangle \end{aligned}$$

where $d\Omega = \sin\theta d\theta d\phi$. We could convert the integral over k to an integral over the kinetic energy, using $dk = \mu dE/k\hbar^2$. Now, when $t \rightarrow -\infty$, the exponential term oscillates very rapidly, so that the only value of E that contribute are *those very near* $E(q)$, where the denominator also varies very rapidly.

$$= \int d\Omega/(2\pi)^3 \int_0^\infty dE \frac{\mu k}{\hbar^2} g(\mathbf{k}) \frac{\exp(-iEt/\hbar)}{E - E(q) + i\epsilon} \langle\Phi_{\mathbf{q}}| V |\Psi_{\mathbf{k}}^{\text{in}}\rangle$$

³It is more reasonable in path integral formalism that the additional term must be added to prevent the integral diverges at large value of wave function.

Hence, for the time limit $t \rightarrow -\infty$, we could set $k = q$ everywhere except in the rapidly varying exponential and denominator,

$$= \int d\Omega / (2\pi)^3 g(\Omega) \frac{\mu q}{\hbar^2} \langle \Phi_{\mathbf{q}} | V | \Psi_{\mathbf{q}}^{\text{in}} \rangle \int_{\text{around } E(q)} dE \frac{\exp(-iEt/\hbar)}{E - E(q) + i\epsilon}$$

We could extend the integration range to the whole real axis, which is permissible because the integral receives no appreciable contributions anywhere that is far from $E(q)$,

$$\propto \int_{-\infty}^{\infty} dE \frac{\exp(-iEt/\hbar)}{E - E(q) + i\epsilon} \quad (1.11)$$

For $t \rightarrow -\infty$, we can close the contour of the integration with a very large semi-circle in the upper half of the complex plane, on which the integration is negligible because, for $\text{Im}(E) > 0$ and $t \rightarrow -\infty$, the numerator $\exp(-iEt/\hbar)$ is exponentially small. The only singularity of the integration is a pole at $E = E(q) - i\epsilon$, which is always in the lower half plane because ϵ is positive infinitesimal parameter. Hence, the integral vanishes for $t \rightarrow -\infty$. Only the first term of equation (1.10) is left, which gives the correct condition for $t \rightarrow -\infty$

1.2 Scattering Amplitudes

In the previous section, we defined a state that at early times has the appearance of a particle traveling toward a collision with a scattering center. Now we must consider what this state looks like after the collision.

For this purpose, we consider the coordinate-space wave function of the state $|\Psi_{\mathbf{k}}^{\text{in}}\rangle$ ⁴.

$$V |\Psi_{\mathbf{k}}^{\text{in}}\rangle = \int d^3x |\Phi_{\mathbf{x}}\rangle \langle \Phi_{\mathbf{x}} | V | \Psi_{\mathbf{k}}^{\text{in}} \rangle = \int d^3x |\Phi_{\mathbf{x}}\rangle V(\mathbf{x}) \psi_{\mathbf{k}}(\mathbf{x}) \quad (1.12)$$

where $\psi_{\mathbf{k}}(\mathbf{x})$ is the coordinate-space wave function of the in-state,

$$\psi_{\mathbf{k}}(\mathbf{x}) = \langle \Phi_{\mathbf{x}} | \Psi_{\mathbf{k}}^{\text{in}} \rangle \quad (1.13)$$

Then, by taking the scalar product of the Lippmann-Schwinger equation (1.8), and using the fact that the scalar product of state of definite momentum and state of definite position would be plane-wave function,

$$\langle \Phi_{\mathbf{x}} | \Phi_{\mathbf{k}} \rangle = e^{i\mathbf{k} \cdot \mathbf{x}}$$

⁴By this way, it would give us a physically vivid picture about scattering.

we have,

$$\begin{aligned}\psi_{\mathbf{k}}(\mathbf{x}) &= e^{i\mathbf{k}\cdot\mathbf{x}} + \langle \Phi_{\mathbf{x}} | (E(|\mathbf{k}|) - H_0 + i\epsilon)^{-1} V | \Psi_{\mathbf{k}}^{\text{in}} \rangle \\ \psi_{\mathbf{k}}(\mathbf{x}) &= e^{i\mathbf{k}\cdot\mathbf{x}} + \int d^3y \langle \Phi_{\mathbf{x}} | [E(|\mathbf{k}|) - H_0 + i\epsilon]^{-1} | \Phi_{\mathbf{y}} \rangle V(y) \psi_{\mathbf{k}}(\mathbf{y})\end{aligned}\quad (1.14)$$

where we define the Green function and then evaluate it by Fourier transform,

$$\begin{aligned}G_k(\mathbf{x} - \mathbf{y}) &= \langle \Phi_{\mathbf{x}} | [E(|\mathbf{k}|) - H_0 + i\epsilon]^{-1} | \Phi_{\mathbf{y}} \rangle \\ &= \int d^3q / (2\pi)^3 \frac{e^{i\mathbf{q}\cdot(\mathbf{x}-\mathbf{y})}}{E(|\mathbf{k}|) - E(|\mathbf{q}|) + i\epsilon} \\ &= \int -d\cos(\theta) d\phi \int q^2 dq / (2\pi)^3 \frac{e^{iq|\mathbf{x}-\mathbf{y}|\cos(\theta)}}{\hbar^2 k^2 / 2\mu - \hbar^2 q^2 / 2\mu + i\epsilon} \\ &= \frac{2\pi}{(2\pi)^3} \int_0^\infty q^2 dq \frac{2i \sin(q|\mathbf{x} - \mathbf{y}|)}{iq|\mathbf{x} - \mathbf{y}|} \frac{2\mu/\hbar^2}{k^2 - q^2 + i\epsilon} \\ &= -i \frac{2\mu}{\hbar^2} \frac{1}{4\pi^2 |\mathbf{x} - \mathbf{y}|} \int_{-\infty}^\infty dq \frac{q e^{iq|\mathbf{x}-\mathbf{y}|}}{k^2 - q^2 + i\epsilon} \\ &= -\frac{2\mu}{\hbar^2} \frac{1}{4\pi |\mathbf{x} - \mathbf{y}|} e^{ik|\mathbf{x}-\mathbf{y}|}\end{aligned}\quad (1.15)$$

The last line could be evaluated by completing the contour integral with a large semi-circle in the upper half plane. The poles are $q = \pm\sqrt{k^2 - i\epsilon} = k \pm i\epsilon$, and picking up the contribution of pole $q = k + i\epsilon$.

$$2\pi i \lim_{\epsilon \rightarrow 0} \frac{(k + i\epsilon) e^{i(k+i\epsilon)|\mathbf{x}-\mathbf{y}|}}{-(k + q + i\epsilon)} = -\pi i e^{ik|\mathbf{x}-\mathbf{y}|}$$

For a potential $V(\mathbf{y})$ that vanishes sufficiently rapidly as $|\mathbf{y}| \rightarrow \infty$. Write equation (1.14)

$$\psi_{\mathbf{k}}(\mathbf{x}) = e^{i\mathbf{k}\cdot\mathbf{x}} + \int d^3y G_{\mathbf{k}}(\mathbf{x} - \mathbf{y}) V(\mathbf{y}) \psi_{\mathbf{k}}(\mathbf{y}) \quad (1.16)$$

For $|\mathbf{x}| \rightarrow \infty$,

$$\psi_{\mathbf{k}}(\mathbf{x}) \rightarrow e^{i\mathbf{k}\cdot\mathbf{x}} + f_{\mathbf{k}}(\hat{x}) e^{ikr} / r \quad (1.17)$$

where $|\mathbf{x}| \equiv r$ and $f_{\mathbf{k}}(\hat{x})$ is *scattering amplitude*.

$$f_{\mathbf{k}}(\hat{x}) = -\frac{\mu}{2\pi\hbar^2} \int d^3y e^{-ik\hat{x}\cdot\mathbf{y}} V(\mathbf{y}) \psi_{\mathbf{k}}(\mathbf{y}) \quad (1.18)$$

There is an interesting result that the coordinate-space wave function would have incident-coming wave part and scattering wave part, and the scattering wave part would act like spherical wave.

Now let's consider how the superposition equation (1.3),

$$\psi_g(\mathbf{x}, t) \equiv \langle \Phi_{\mathbf{x}} | \Psi_g(t) \rangle = \int d^3k / (2\pi)^3 g(\mathbf{k}) \psi_{\mathbf{k}}(\mathbf{x}) \exp(-i\hbar t \mathbf{k}^2 / 2\mu) \quad (1.19)$$

in the limit $t \rightarrow +\infty$, with r/t held fixed, and \mathbf{x} off the 3-axis⁵, equation (1.19) gives⁶,

$$\begin{aligned} \psi_g(\mathbf{x}, t) &\rightarrow \frac{1}{(2\pi)^3 r} \int d^2k_{\perp} \int_{-\infty}^{\infty} dk_3 g(\mathbf{k}_{\perp}, k_3) \\ &\times \exp(ik_3 r - i\hbar t k_3^2 / 2\mu) f_{\mathbf{k}_0}(\hat{x}) \end{aligned} \quad (1.20)$$

We have taken the subscript on the scattering amplitude to be \mathbf{k}_0 , because the function g is sharply peaked at this value of \mathbf{k} , and we have approximated $k = \sqrt{k_3^2 + \mathbf{k}_{\perp}^2}$ as $k \approx k_3$ in the exponent, because $g(\mathbf{k}_{\perp}, k_3)$ is assumed to be negligible except for $|\mathbf{k}_{\perp}| \ll k_3$. That is, we assume that the function $g(\mathbf{k}_{\perp}, k_3)$, though smooth, is strongly peaked at $k_3 = k_0$ and $\mathbf{k}_{\perp} = 0$, so we could set k_3 in $g(\mathbf{k}_{\perp}, k_3)$ equal to the value $\mu r / \hbar t$, so that,

$$\begin{aligned} \psi_g(\mathbf{x}, t) &\rightarrow \frac{1}{(2\pi)^3 r} f_{\mathbf{k}_0}(\hat{x}) \int d^2k_{\perp} g(\mathbf{k}_{\perp}, \mu r / \hbar t) \\ &\times \int_{-\infty}^{\infty} dk_3 \exp(ik_3 r - i\hbar t k_3^2 / 2\mu) \\ &= \frac{1}{(2\pi)^3 r} f_{\mathbf{k}_0} \int d^2k_{\perp} g(\mathbf{k}_{\perp}, \mu r / \hbar t) \exp(i\mu r^2 / 2\hbar t) \sqrt{\frac{2\pi \mu}{i\hbar t}} \end{aligned} \quad (1.21)$$

Where we reach the last line by Gaussian integral.

The probability $dP(\hat{x})$ that the particle at late times is somewhere within the cone of infinitesimal solid angle $d\omega$ around the direction \hat{x} is then the integral of $|\psi_g(\mathbf{x}, t)|^2$ over this cone:

$$dP(\hat{x}, \mathbf{k}_0) = d\Omega \int_0^{\infty} r^2 dr |\psi_g(\mathbf{x}, t)|^2 \quad (1.22)$$

⁵Which means the first term in equation (1.17) vanishes.

⁶We would assume that the particle comes in from a great distance along the negative 3-axis, so we are interested in the limit of very large negative t and x_3 , but with x_3/t held finite. And also assume the particle velocity is sufficiently closely confined to the 3-direction, where the function $g(\mathbf{k})$ is not negligible.

$$\rightarrow d\Omega \frac{2\mu\pi}{(2\pi)^6 \hbar t} |f_{\mathbf{k}_0}(\hat{x})|^2 \int_0^\infty dr \left| \int d^2 k_\perp g(\mathbf{k}_\perp, \mu r / \hbar t) \right|^2$$

changing the variable of integration r to $k_3 \equiv \mu r / \hbar t$,

$$\frac{dP(\hat{x}, \mathbf{k}_0)}{d\Omega} = |f_{\mathbf{k}_0}(\hat{x})|^2 \int_0^\infty dk_3 / (2\pi)^5 \left| \int d^2 k_\perp g(\mathbf{k}_\perp, k_3) \right|^2 \quad (1.23)$$

Now, the coefficient of $|f_{\mathbf{k}_0}(\hat{x})|^2$ in equation (1.23) has the dimensions of an inverse area. In fact, it is precisely the probability per unit area that the particle is in a small area centered on the 3-axis and normal to that axis:

$$\rho_\perp \equiv \lim_{t \rightarrow -\infty} \int_{-\infty}^\infty dx_3 |\psi_g(0, x_3, t)|^2 \quad (1.24)$$

To see this, recall equation (1.4) and its scalar product with state of definite position,

$$\langle \Phi_{\mathbf{x}} | \Psi_g(t) \rangle \rightarrow \int d^3 k / (2\pi)^3 g(\mathbf{k}) \exp(i\mathbf{k} \cdot \mathbf{x} - i\hbar t \mathbf{k}^2 / 2\mu)$$

As what we've done in equation (1.21), we assume that the particle moves along the 3-axis,

$$\begin{aligned} & \langle \Phi_{\mathbf{x}} | \Psi_g(t) \rangle \rightarrow \\ & \frac{1}{(2\pi)^3} \int d^2 k_\perp \int dk_3 g(\mathbf{k}_\perp, k_3) \exp(i\mathbf{k}_\perp \cdot \mathbf{x}_\perp + ik_3 x_3 - i\hbar t (k_3 - \mu x_3 / \hbar t)^2 / 2\mu) \end{aligned}$$

Replace k_3 with $k_3 = \mu x_3 / \hbar t$ except the exponent term.

$$\begin{aligned} & \langle \Phi_{\mathbf{x}} | \Psi_g(t) \rangle \rightarrow \\ & \frac{1}{(2\pi)^3} \int d^2 k_\perp g(\mathbf{k}_\perp, \mu x_3 / \hbar t) \exp(i\mathbf{k}_\perp \cdot \mathbf{x}_\perp) \\ & \times \exp(ix_3^2 \mu / 2\hbar t) \int_{-\infty}^\infty dk_3 \exp(-i\hbar t (k_3 - \mu x_3 / \hbar t)^2 / 2\mu) \\ & = \frac{1}{(2\pi)^3} \exp(ix_3^2 \mu / 2\hbar t) \sqrt{\frac{2\mu\pi}{i\hbar t}} \times \int d^2 k_\perp g(\mathbf{k}_\perp, \mu x_3 / \hbar t) \exp(i\mathbf{k}_\perp \cdot \mathbf{x}_\perp) \quad (1.25) \end{aligned}$$

Therefore, in particular as t tends to infinity far (∞), the spatial probability distribution is,

$$|\langle \Phi_{\mathbf{x}} | \Psi_g(t) \rangle|^2 \rightarrow \frac{\mu}{\hbar t (2\pi)^5} \left| \int d^2 k_\perp g(\mathbf{k}_\perp, \mu x_3 / \hbar t) \exp(i\mathbf{k}_\perp \cdot \mathbf{x}_\perp) \right|^2 \quad (1.26)$$

Integrate over spatial with equation (1.26)⁷,

$$\begin{aligned}\rho_{\perp} &= \int dx_3 |\langle \Phi_{\mathbf{x}} | \Psi_g(t) \rangle|^2 \rightarrow \frac{\mu}{\hbar t (2\pi)^5} \int dx_3 \left| \int d^2 k_{\perp} g(\mathbf{k}_{\perp}, \mu x_3 / \hbar t) \right|^2 \\ &\rightarrow \frac{1}{(2\pi)^5} \int dk_3 \left| \int d^2 k_{\perp} g(\mathbf{k}_{\perp}, \mu x_3 / \hbar t) \right|^2\end{aligned}$$

Hence, equation (1.23) may be written,

$$\frac{dP(\hat{x}, \mathbf{k}_0)}{d\Omega} = \rho_{\perp} |f_{\mathbf{k}_0}(\hat{x})|^2 \quad (1.27)$$

We define the *differential cross-section* as the ratio,

$$\frac{d\sigma(\hat{x}, \mathbf{k}_0)}{d\Omega} \equiv \frac{1}{\rho_{\perp}} \frac{dP(\hat{x}, \mathbf{k}_0)}{d\Omega} = |f_{\mathbf{k}_0}(\hat{x})|^2 \quad (1.28)$$

We could think of $d\sigma(\hat{x}, \mathbf{k}_0)$ as a tiny area normal to the 3-axis, where the incoming particle must hit in order for it to be scattered into a solid angle $d\Omega$ around the direction \hat{x} . The equation (1.28) then says that the probability of the above process to happen equals the ratio of $d\omega$ to the effective cross-sectional area $1/\rho_{\perp}$ of the beam.

Of course, to measure $d\sigma/d\Omega$, experimenters do not actually send a particle or particles toward a single target. Instead, they direct a beam of particles toward a thin slab containing some large number N_T of targets. Scattering into some particular range of angles can occur only if a particle from the beam hits a tiny area $d\sigma$ around one of the targets, then the number of particles that are scattered into this range of angles⁸ is the number of beam particles per unit transverse area \mathcal{N}_B , times the total area $N_T d\sigma$ that the beam particles have to hit upon.

2 General Scattering Theory

There are much more general circumstances to which scattering theory is applicable. The scattering can produce additional particles; the interaction may not be a local potential; some or all of the particles involved may be moving at relativistic velocities; some may be photon (massless); and the initial state may even contain more than two particles. This section would describe scattering theory at a level of generality that encompasses all these possibilities.

⁷Note that we only integrate over 3-axis, so the dimension is the probability per unit area.

⁸Or says, the number of scattering event.

2.1 The S-Matrix

2.2 Rates

2.3 The General Optical Theorem

Part II

Quantum Field Theory

3 Introduction to Perturbation Theory and Scattering

4 Perturbation I. Wick Diagrams

5 Perturbation II. Divergence and Counterterms

6 Feynman Diagrams

7 Cross-Sections S-Matrix

8 Computing S-Matrix Elements from Feynman Diagrams