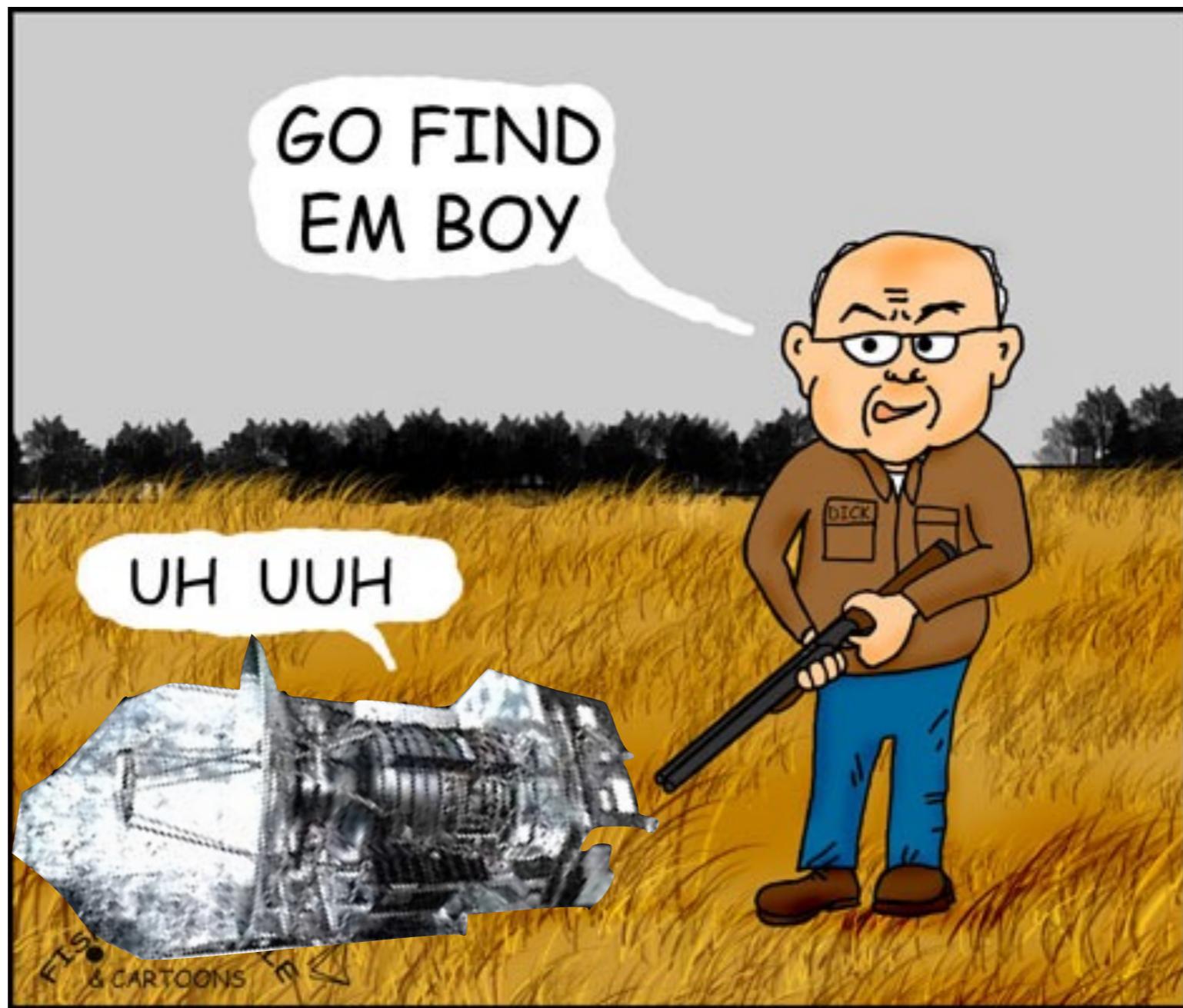
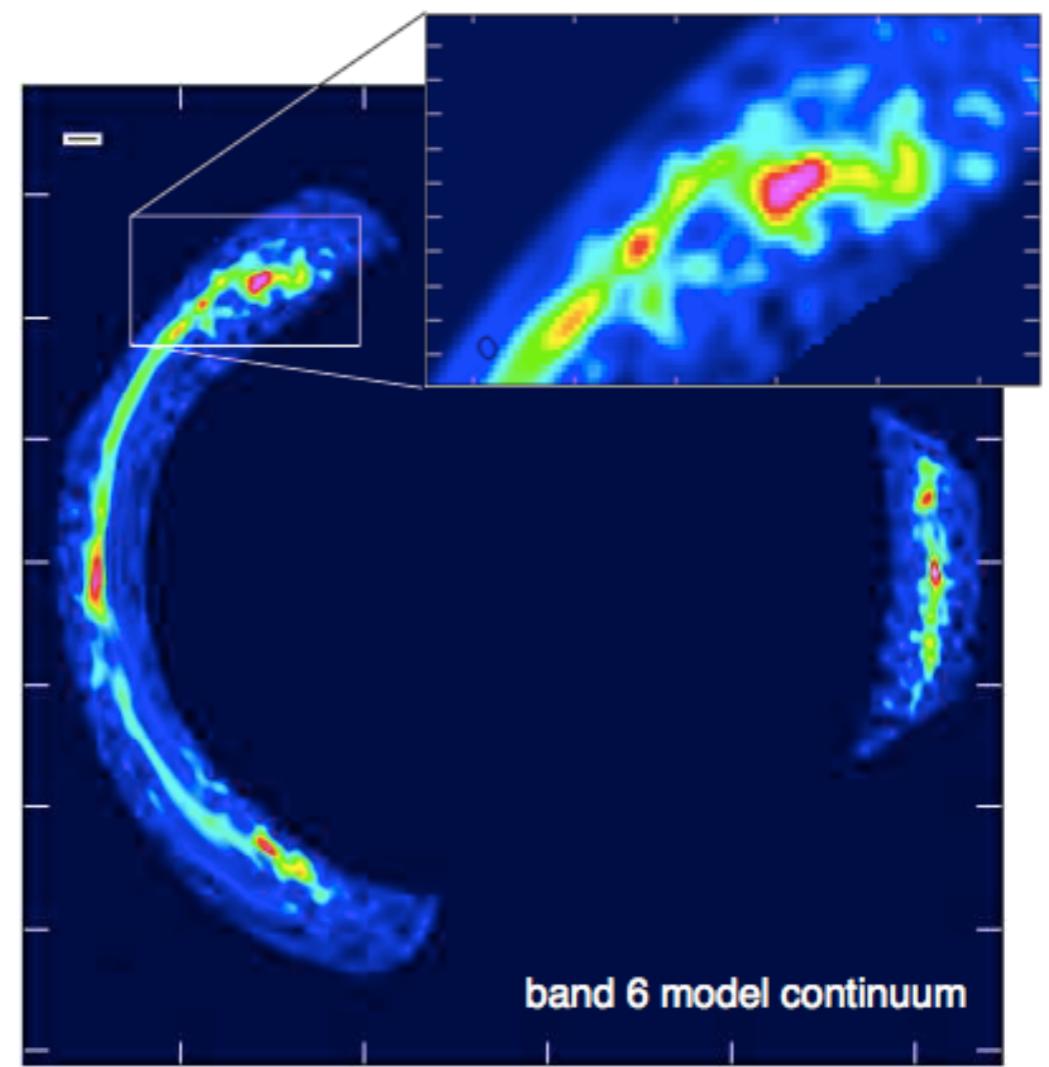


Lens hunting with *Herschel*

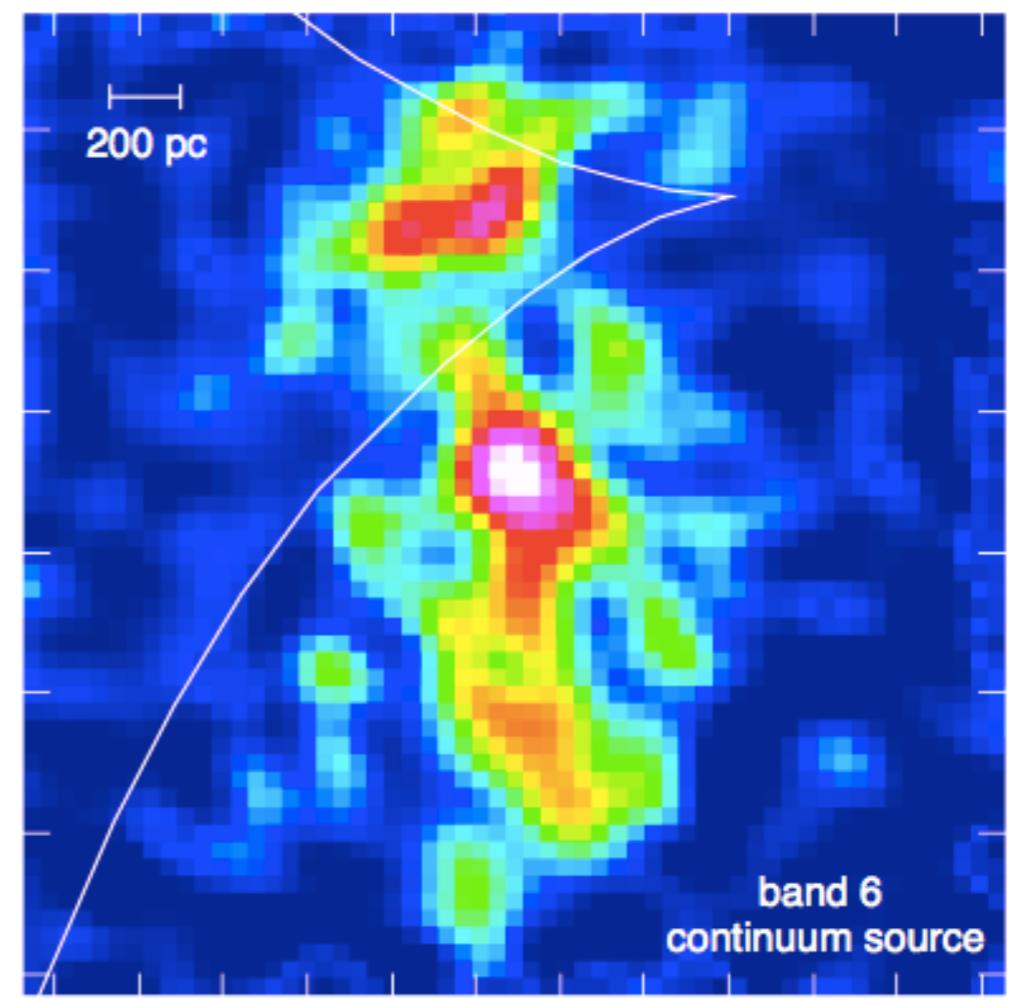
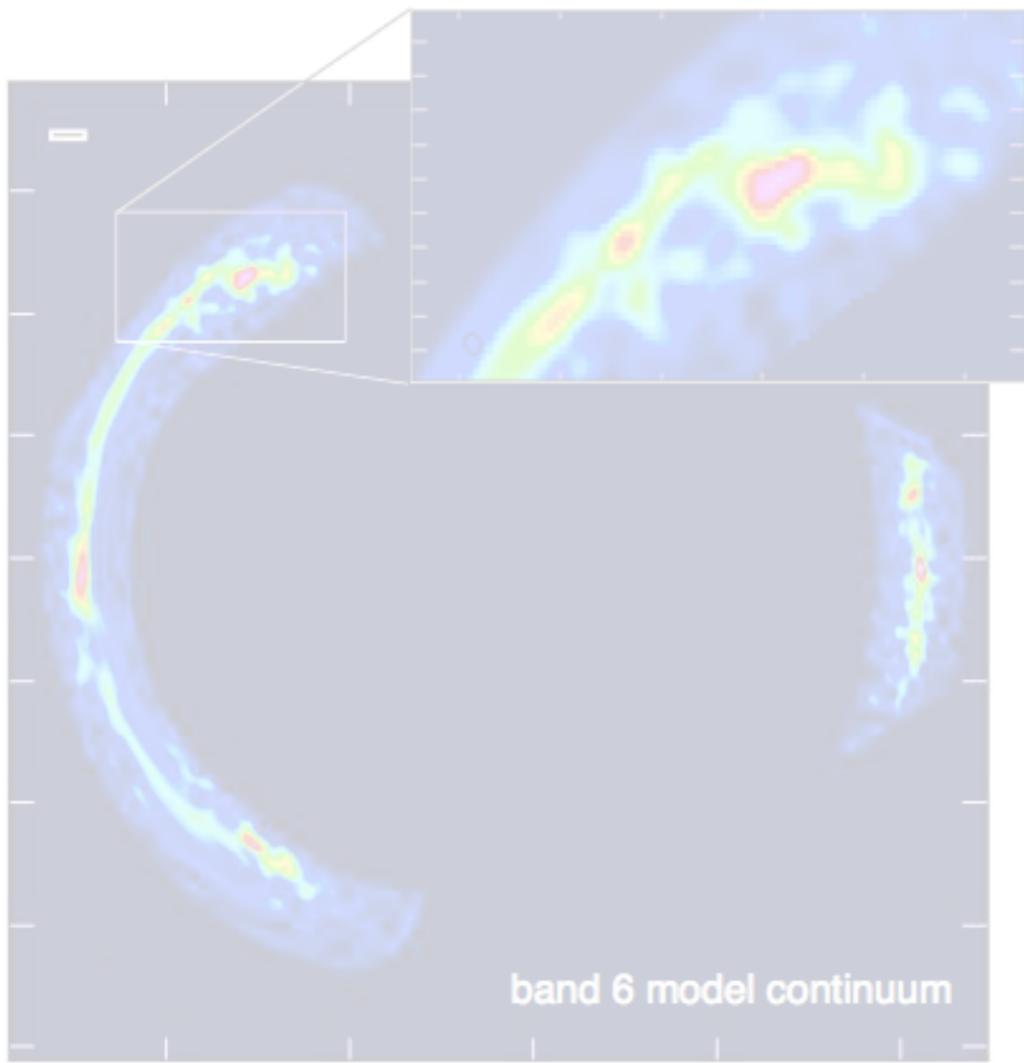
Tom Bakx



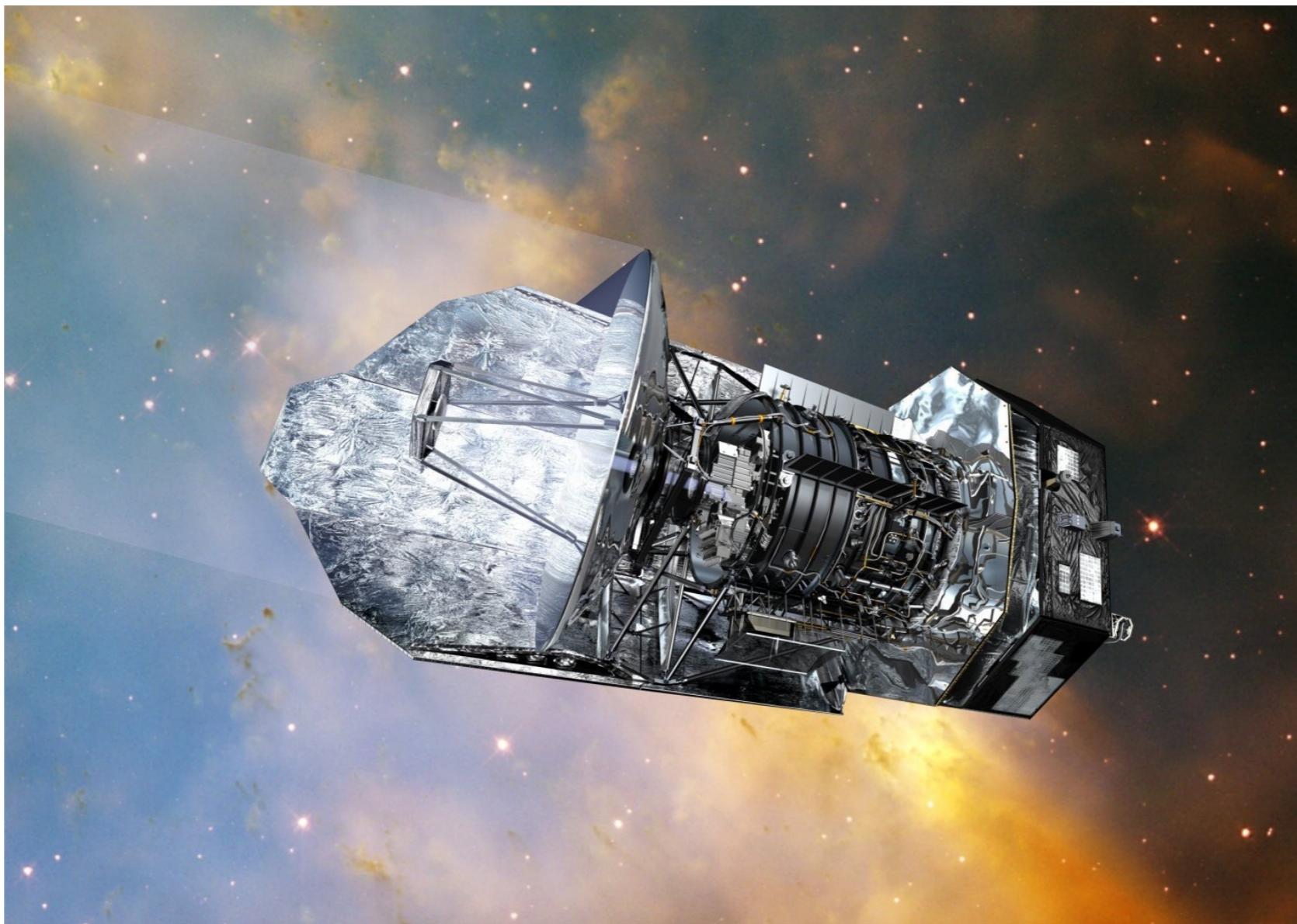
Gravitational lenses provide otherwise unreachable levels of detail



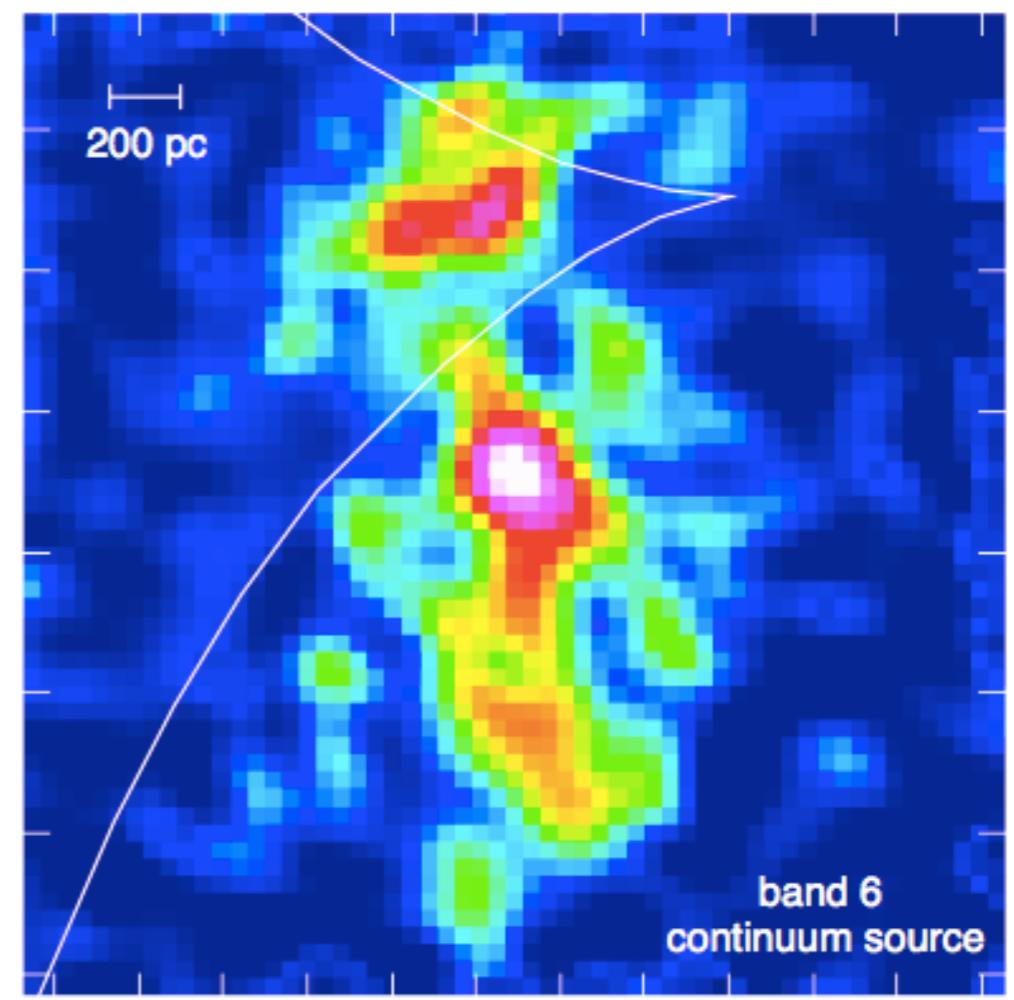
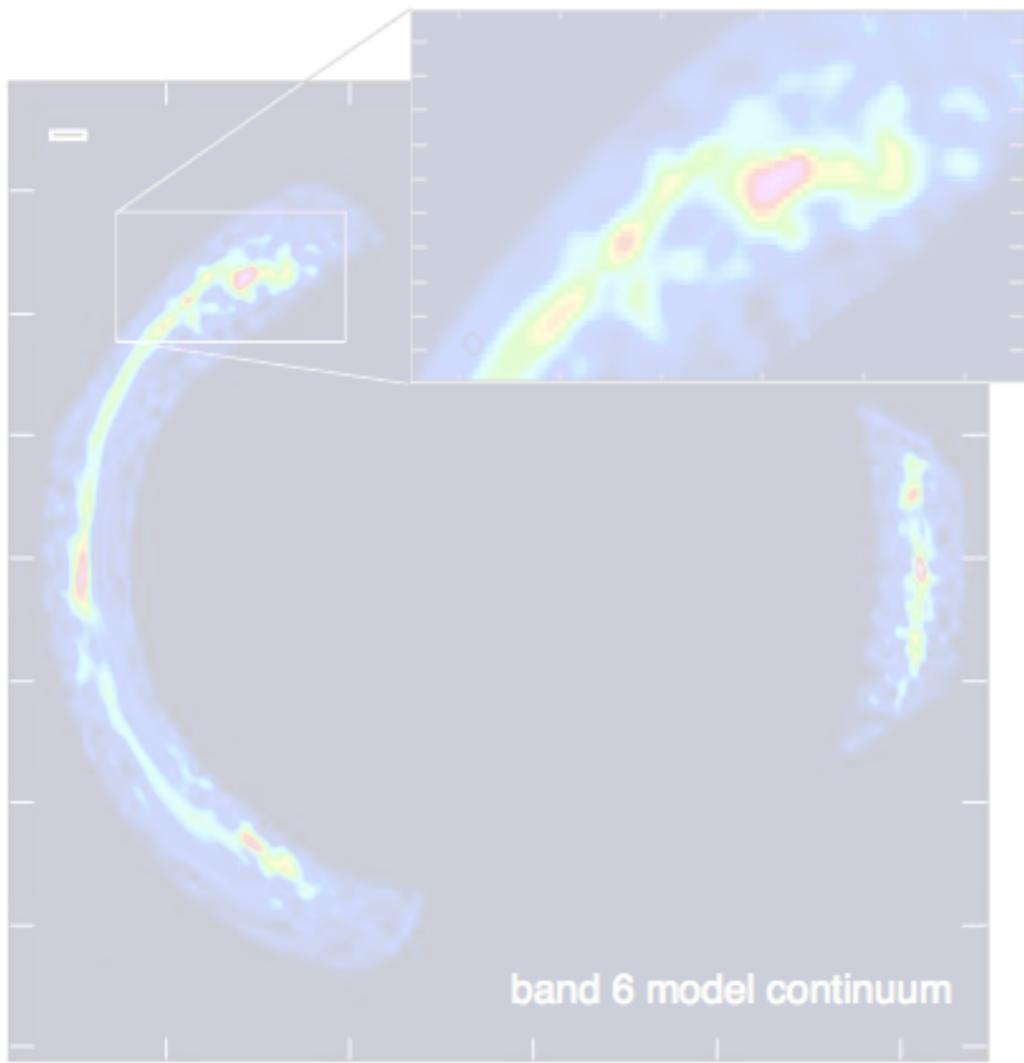
Gravitational lenses provide otherwise unreachable levels of detail



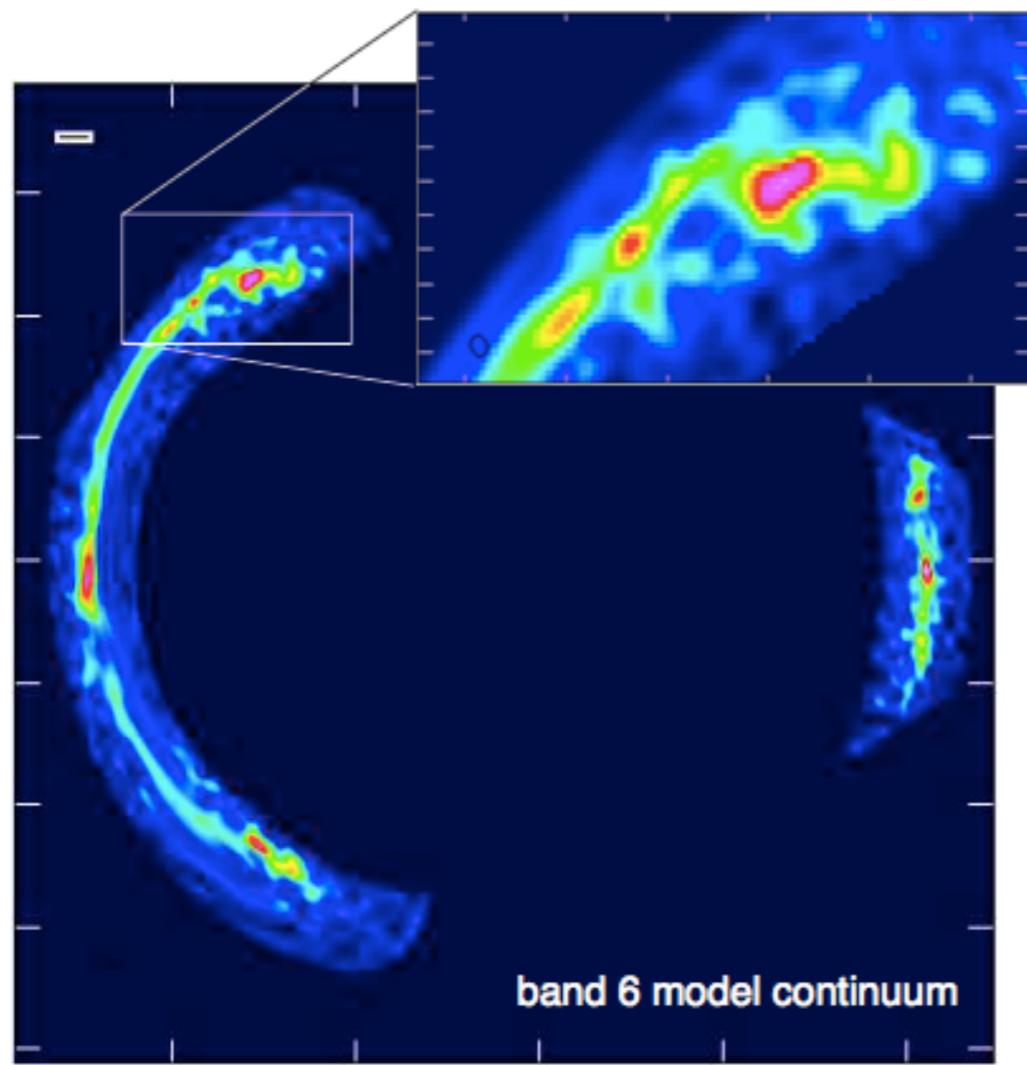
The *H*-ATLAS survey is used to find large samples of lenses



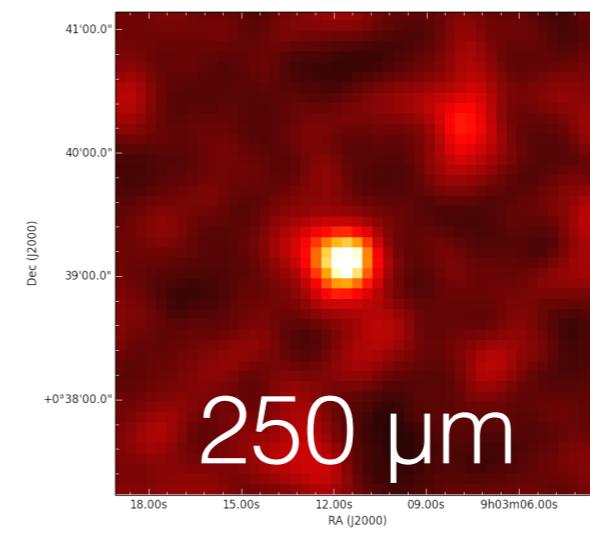
Gravitational lenses provide otherwise unreachable levels of detail



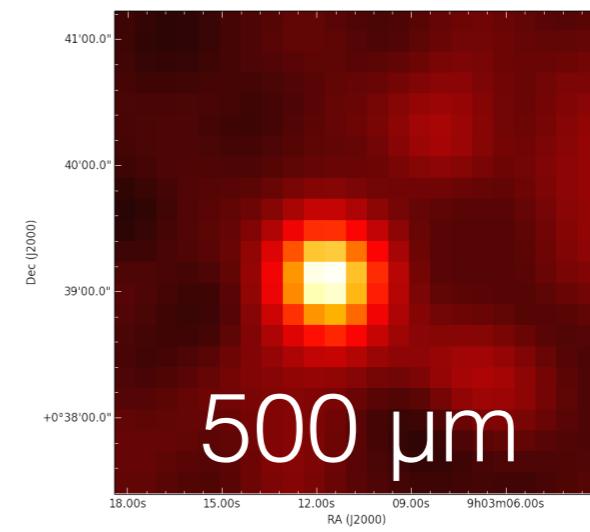
Large surveys don't have the detail of ALMA observations



ALMA



Herschel



Herschel

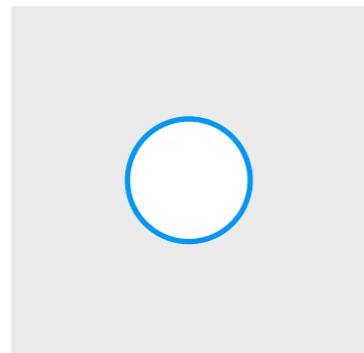
Herschel

Herschel

Source confusion effects need to be measured

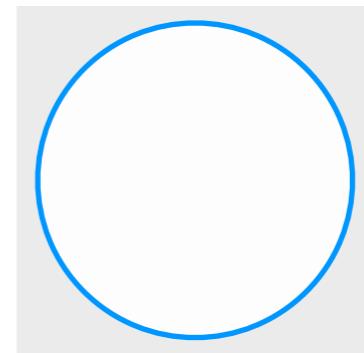
	<i>Herschel</i>		
λ [μm]	250	350	500
Angular size	18"	25"	36"
Surface	158%	306%	634%
Beam size			

JCMT's luminosities at different resolutions estimate source confusion



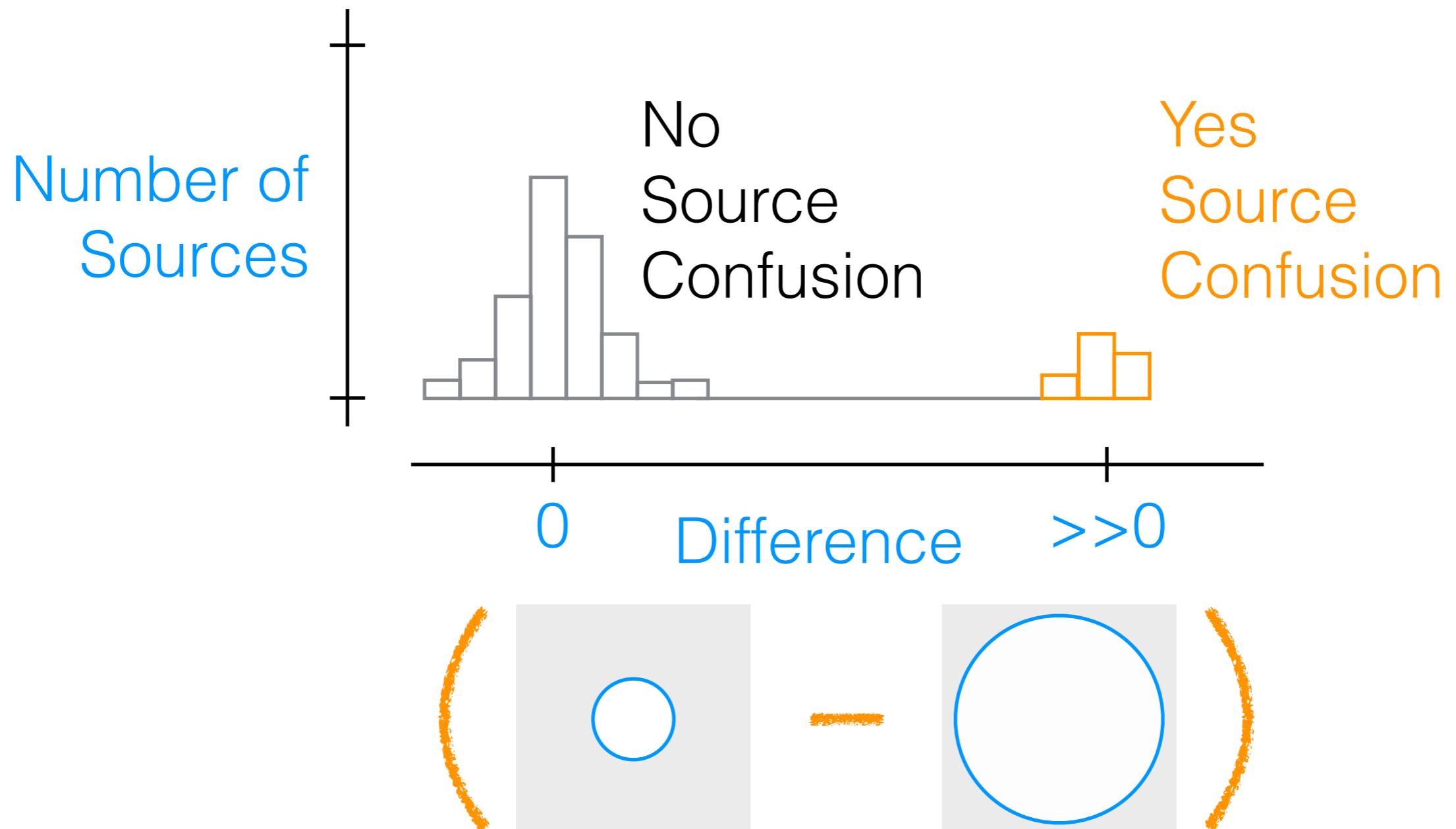
JCMT
14''

Convolution



JCMT
36''

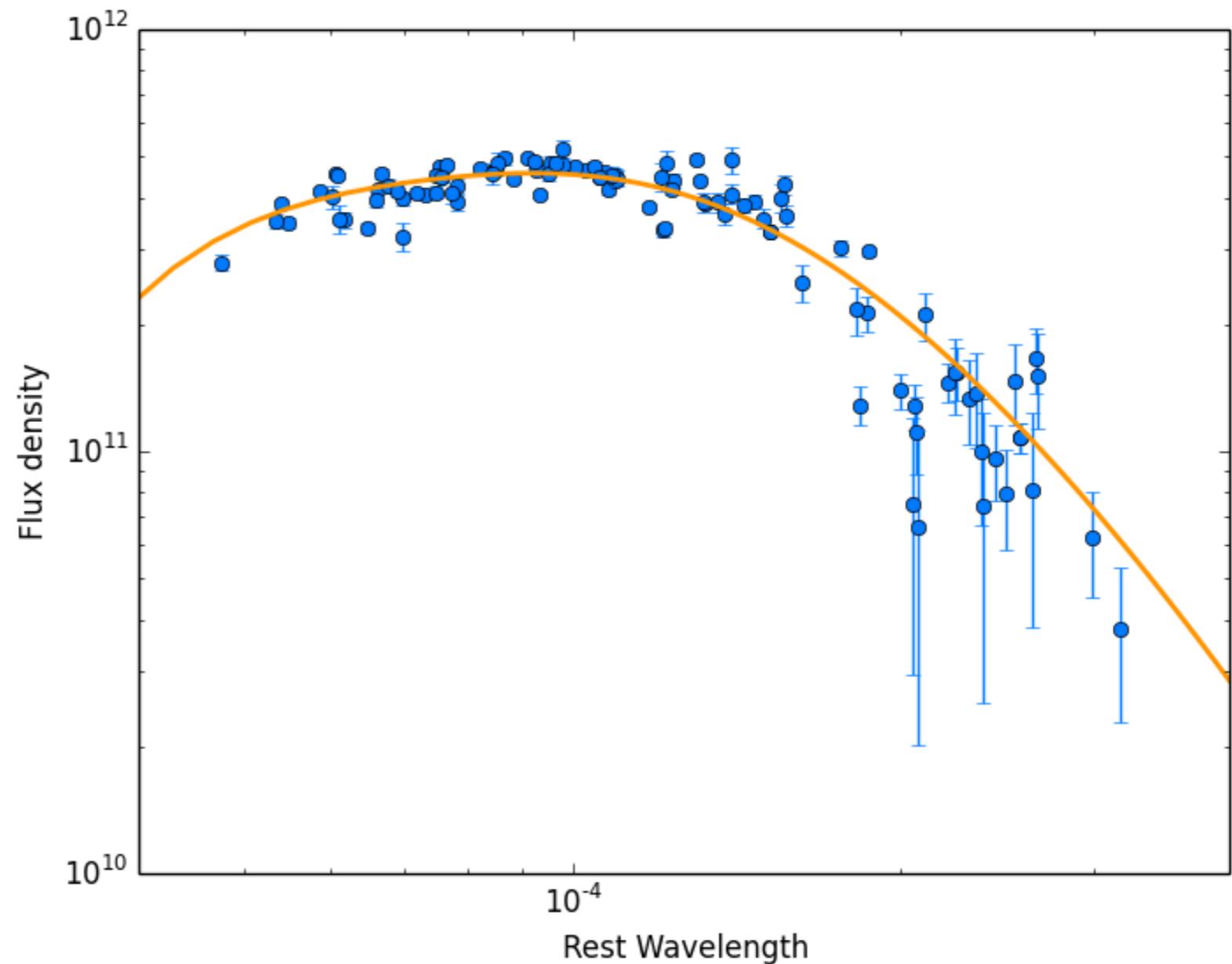
JCMT's luminosities at different resolutions estimate source confusion



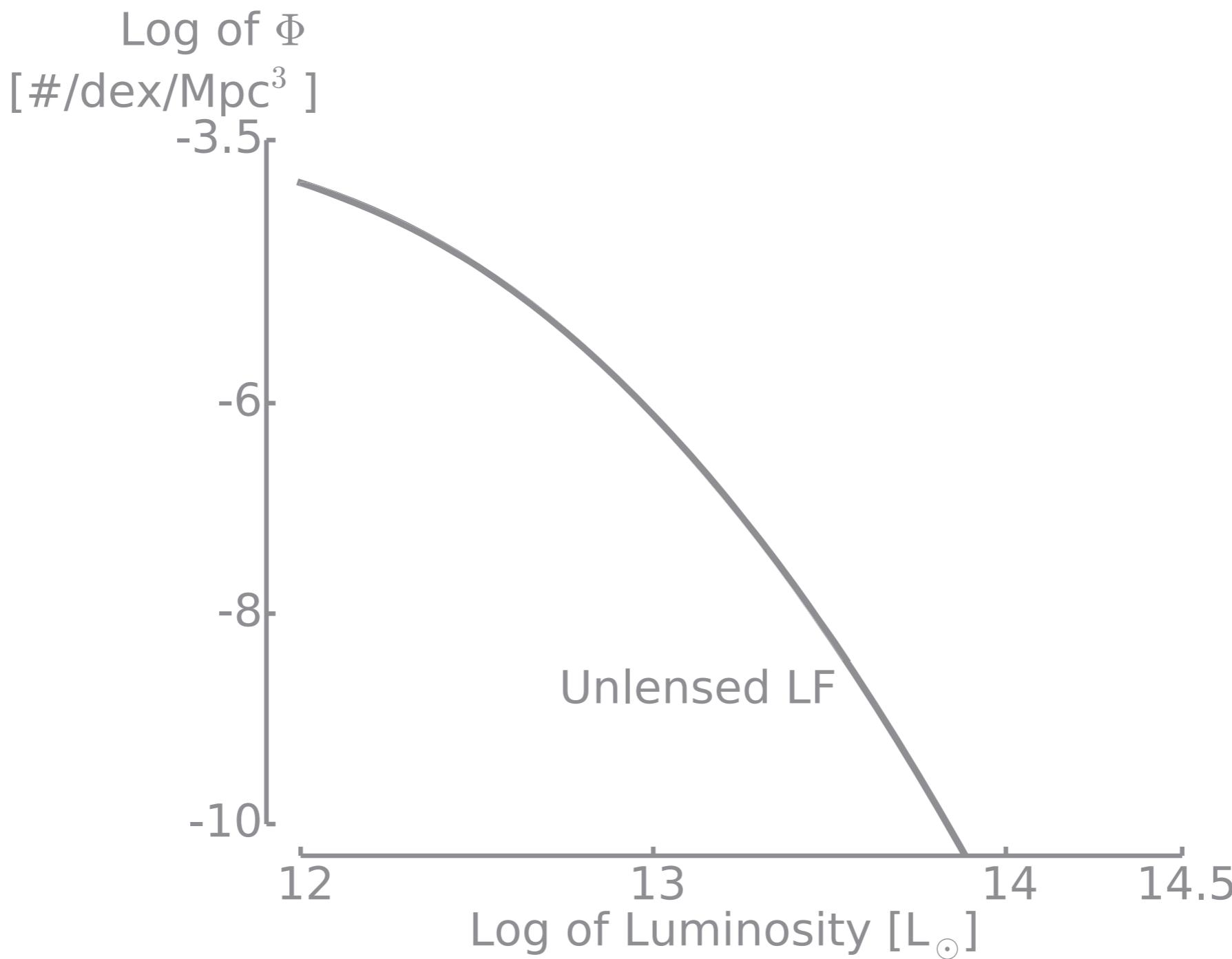
Template SED made from 26 spectroscopic sources

Two-temperature
grey-body SED fit

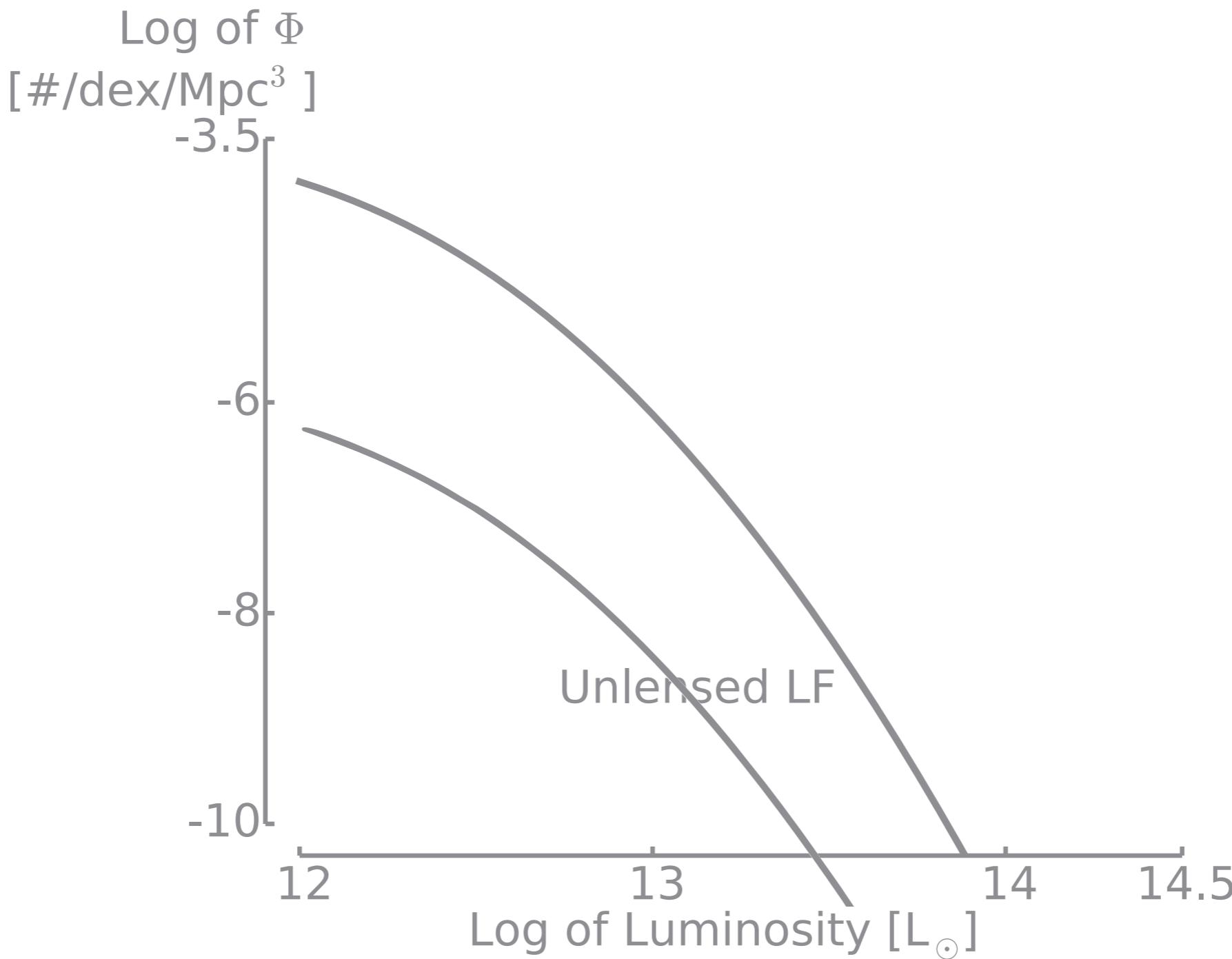
$$S_\nu = A \left[B_\nu(T_h) \nu^\beta + \alpha B_\nu(T_c) \nu^\beta \right]$$



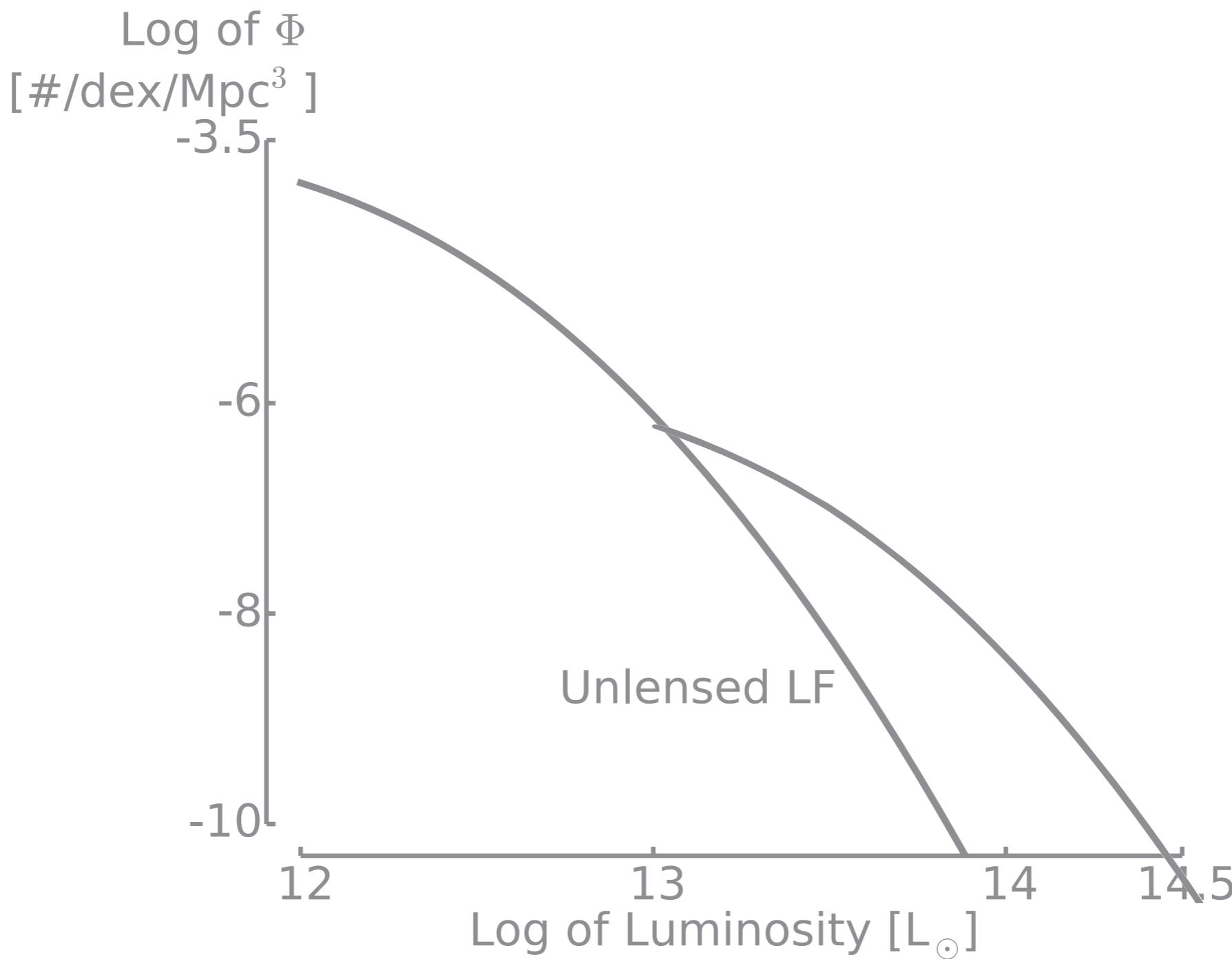
Luminosity functions provide global information on lensing



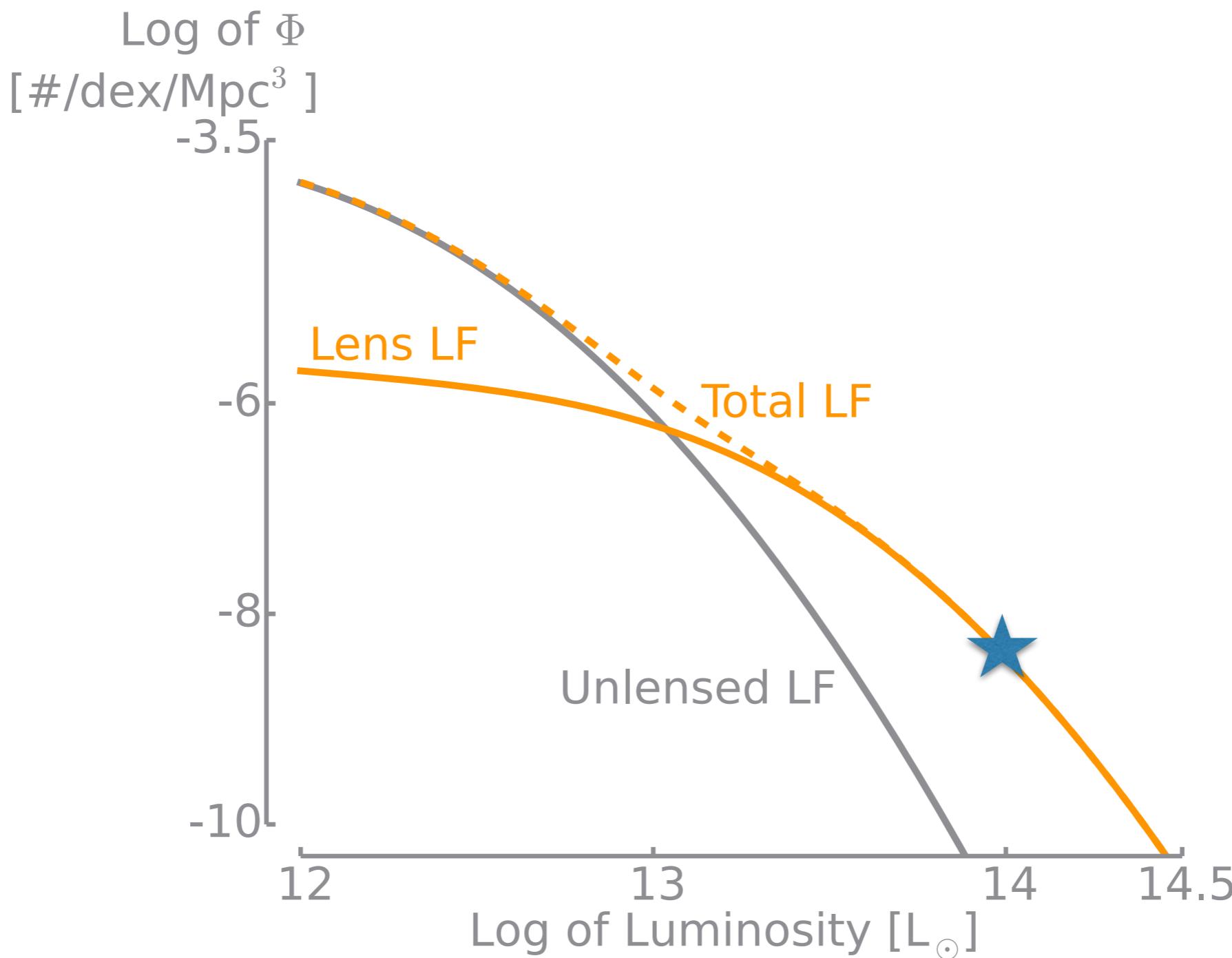
Luminosity functions provide global information on lensing



Luminosity functions provide global information on lensing



Luminosity functions provide global information on lensing

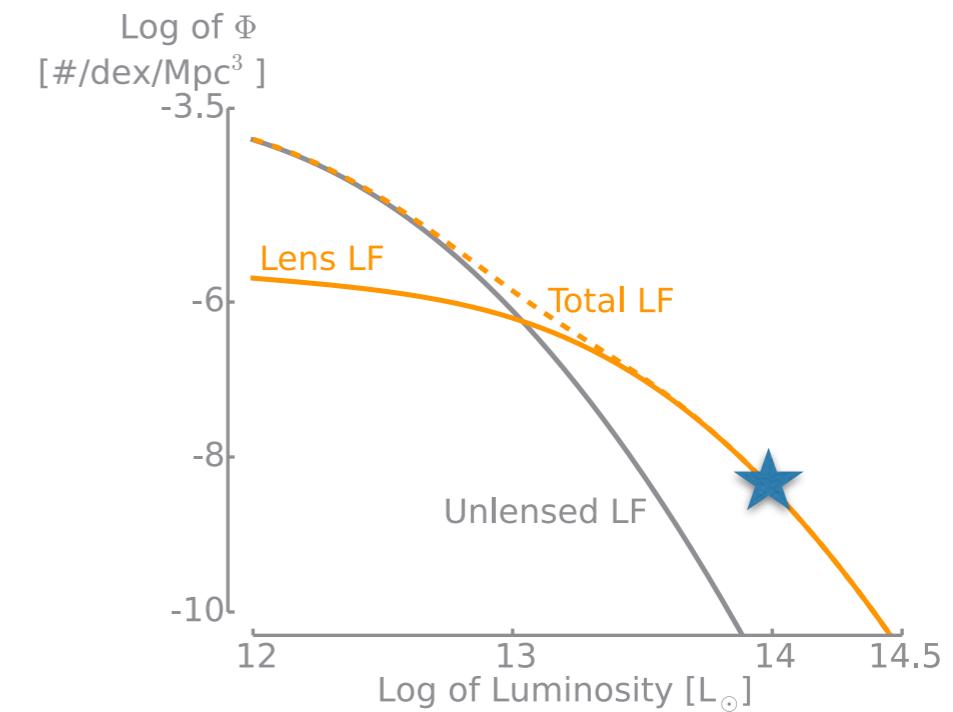


First analysis of the sources suggest a lensed sample bias

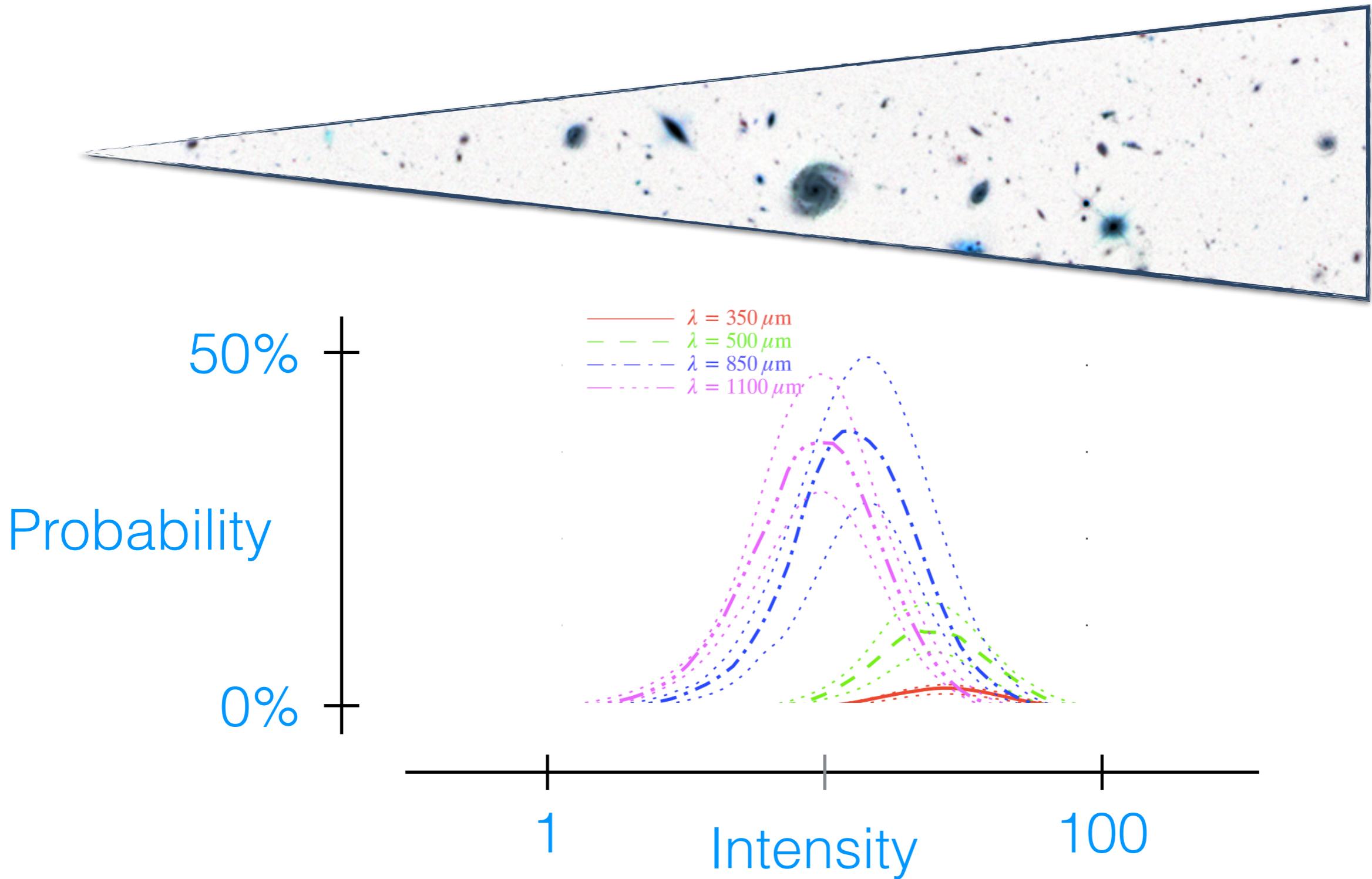
Galaxies

35
30
25
20
15
10
5
0

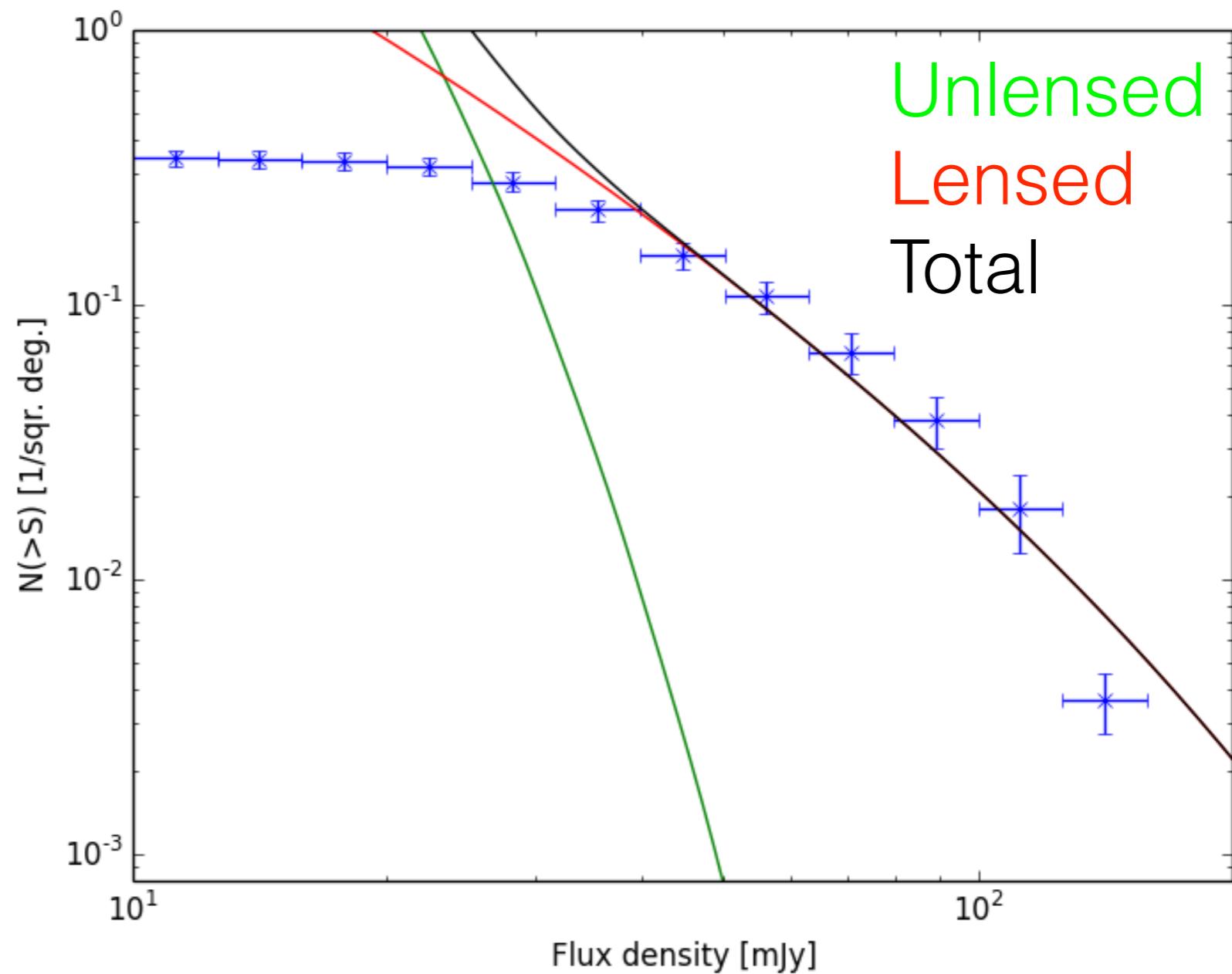
2.0 2.5 3.0 3.5 4.0 4.5
Redshift



Cosmological models provide lensing probabilities



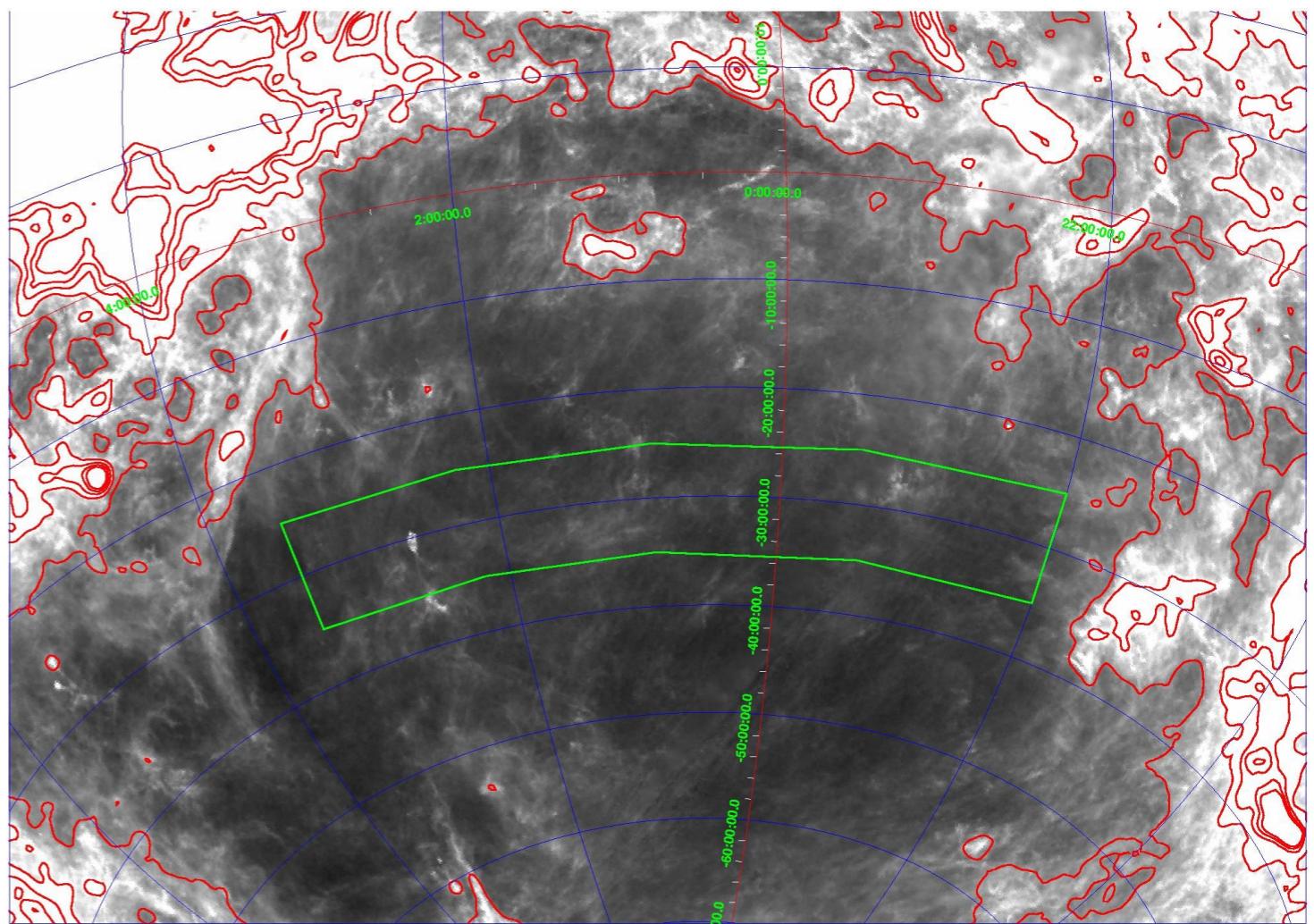
Preliminary results show a lensing-biased survey



Resolved images are necessary for definitive lensing evidence

KIDS and VIKINGS surveys
will provide information
on the lenses

This will help us
improve lens-finding
algorithms



Lens hunting with *Herschel*

Tom Bakx

