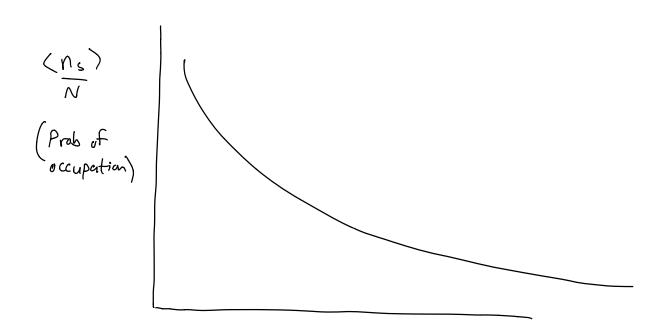
Consider an ideal gas of electrons (fermions)

Quantum States given by n_x, n_y, n_z $E_s = \frac{\hbar^2 \pi^2}{L^2} (n_x^2 + n_y^2 + n_z^2)$

| S | Es |
|---------|----|
| (1,1,1) | 3 |
| (2,1,1) | 6 |
| ζ1,z,/) | 6 |
| (1,1,2) | 6 |
| (3,1,17 | 11 |
| • | |

If I add an electron to the gas, which stak will it occupy?

At "normal" temperatures



-Particles don't necessarily fill the lowest available energy State

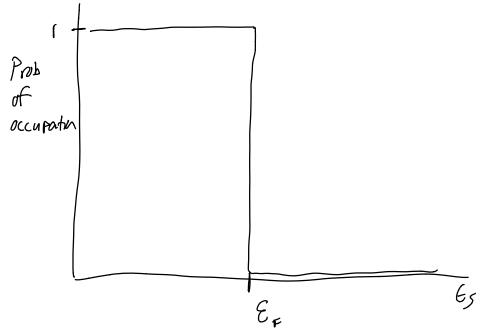
Particles will spontaneously jump to higher energies as long as SF(0 (JU (TUS) The gas wants to minimize energy while also maximizing entropy

This means there are some empty States between the lowest thighest energy electron

| S | es | Bring in first election, it |
|---------|----|--------------------------------|
| (1,1,1) | 3 | goes to ground State |
| (2,1,1) | 6 | Second electron Skips these |
| ζΙ,z,/) | 6 | t goes here instead |
| (1,1,2) | 6 | |
| (3,1,17 | 11 | |
| • | | |

How energetic is the highest energy electron? Cart say for Sure-

Compare this with the T=0
Scenario



States with Es (Ex are certain to be occupied

States $\omega/E_s > E_F$ are centainly empty

(no unoccupied states between gound + EF)

Particles will spontaneously jump
to higher energies as long
as SF(O (JU (TJS))

T -> O

SF(O => JU(O)

electrons will only move to states with lower energy; these are all occupied

Since all States below Ex are Filled, counting states > counting particles

| S | es | Ns |
|---------|----|---------------|
| (1,1,1) | 3 | 2 (11) |
| (2,1,1) | 6 | 2 |
| <1,2,1) | 6 | 2 |
| (1,1,2) | 6 | Z |
| (3,1,17 | 11 | 2 |
| | | |
| | EE | 2 |
| | • | U |
| | • | \bigcirc |
| | | \mathcal{O} |

The possible combinations of nx, ny, nz which result in energy Es

lie on the sphere with radius

$$\left(\frac{2mL^{2}}{\hbar^{2}\pi^{2}}e_{s}\right)^{\gamma_{2}}$$

$$E_{s} = \frac{\hbar^{2} \pi^{2}}{L^{2}} (n_{x}^{2} + n_{y}^{2} + n_{z}^{2})$$

$$\bigvee$$

$$n_x^2 + n_y^2 + n_z^2 = \frac{2mL^2}{\hbar^2 \pi^2} \epsilon_s$$

of states with energy Es

$$=2.\sum_{n_{x}}\sum_{n_{y}}\sum_{n_{z}}(1)$$

$$=2.\sum_{n_{x}}\sum_{n_{y}}\sum_{n_{x}^{2}+n_{y}^{2}+n_{z}^{2}}=\frac{2mL}{\kappa^{2}\pi^{2}}E_{s}$$

of states with energy below Es:

$$= 2 \cdot \sum_{n_{x}} \sum_{n_{y}} \sum_{n_{z}} (1)$$

$$= 1 \cdot \sum_{n_{x}} \sum_{n_{y}} \sum_{n_{z}} \sum_{n_{$$

$$= \frac{2 \cdot 1}{2n^3} \int \frac{dn_x dn_y dn_z}{n_x^2 + n_y^2 + n_z^2} \left(\frac{2mL^2}{\hbar^2 \eta^2} \epsilon_s \right)$$

$$\frac{1}{2} = \frac{1}{1^3} = 1$$

$$=2\cdot\int_{0}^{\frac{11}{2}}\sinh\theta\theta\int_{0}^{\frac{11}{2}}d\theta\int_{0}^{\frac{2mL^{2}E_{s}}{k^{2}\pi^{2}}}\eta^{2}d\eta$$

$$=2\cdot\frac{17}{2}\cdot\frac{1}{3}\left(\frac{2mL^{2}E_{s}}{k^{2}\pi^{2}}\right)^{\frac{3}{2}}$$

$$\eta\left(\varepsilon\left(\varepsilon_{s}\right)=2\cdot\frac{\pi}{6}\cdot\left(\frac{2mL^{2}}{\hbar^{2}\pi^{2}}\varepsilon_{s}\right)^{3/2}$$

$$\alpha + \mathcal{E}_{s} = \mathcal{E}_{F}, \quad n(\mathcal{E}(\mathcal{E}_{F}) = N)$$

$$N = \frac{11}{3} \cdot \left(\frac{2mL^{2}}{\hbar^{2}\pi^{2}}\mathcal{E}_{F}\right)^{3/2}$$

$$\mathcal{E}_{F} = \frac{\hbar^{2}}{2m} \left(\frac{3\pi^{2}N}{V} \right)^{2/3} \quad \forall = L^{3}$$

Energy of highest energy electron O T=0

What is U?

$$U = 2 \cdot \sum_{n_{x}} \sum_{n_{y}} \sum_{n_{z}} \frac{E(n_{x}, n_{y}, n_{z})}{\sum_{n_{x}} \frac{E(n_{x}, n_{y}, n_{z})}{$$

$$=2\int \frac{h^2\pi^2(n_x^2+n_y^2+n_z^2)\,dn_x\,dn_y\,dn_z}{L^2}$$

$$=2\int_{0}^{\frac{11}{2}}\sin\theta\,d\theta\int_{0}^{\frac{11}{2}}d\theta\int_{0}^{(\frac{2mL^{2}}{2})^{\frac{1}{2}}}\frac{1}{L^{2}}n^{2}\cdot n^{2}dn$$

$$=2\cdot 1\cdot \frac{\pi}{2}\cdot \frac{\hbar^{2}\pi^{2}}{L^{2}}\left(\frac{1}{5}\right)\left(\frac{2mL^{2}}{\hbar^{2}\pi^{2}}\right)^{\frac{2}{5}}$$

$$E_F \approx 7 \text{ eV}$$

$$LT \approx 3 \times 10^{-2} \text{ eV}$$

Since the energy of the gas
$$(\frac{3}{5}E_F)$$

 $U >> KT$ (the thermal energy)
we can approximate that $T \approx 0$

-So our gas gets a large amount of internal energy, even at
$$T=0$$

- The gas also exerts pressure
$$dU = T dS - P dV + M dN = P = -\frac{\partial U}{\partial V}$$

$$U = \frac{3}{5} \varepsilon_F = \left(\frac{3}{5}\right) \frac{h^2}{2m} \left(\frac{3\pi^2 N}{V}\right)^{2/3}$$

$$P = \frac{Z}{S} \frac{U}{V}$$

Example: Neutron Star

$$U_{K} = \frac{3}{5} \cdot N \cdot \frac{\hbar^{2}}{2m_{n}} \left(\frac{3\pi^{2}N}{V} \right)^{2/3}$$

$$M = Nm_n = N = \frac{M}{m_n}, V = \frac{4}{3}\pi R^3$$

$$U_{K} = \frac{3}{10} \frac{h^{2} \pi^{2}}{m_{n}} \left(\frac{M}{m_{n}}\right)^{5/3} \left(\frac{9}{4\pi^{2} R^{3}}\right)^{2/3}$$

$$U_{grav} = -\frac{3}{5} \frac{GM^2}{R}$$

$$\frac{3}{10} \frac{h^{2} \pi^{2}}{m_{n}} \left(\frac{M}{m_{n}}\right)^{\frac{5}{3}} \left(\frac{9}{4\pi^{2} R^{3}}\right)^{\frac{2}{3}} - \frac{3}{5} \frac{GM^{2}}{R}$$

$$\approx \frac{a}{R^2} - \frac{b}{R}$$

Equilibrium:
$$\frac{dU_{tot}}{dR} = 0$$
 Solve for R

$$M \approx 1 M_0$$

$$M_n = 1.67 \times 10^{-27} \text{ Kg}$$

$$g = \frac{M_{\odot}}{\frac{4}{3}\pi R^3} \approx 3 \times 10^{17} \frac{\text{kg}}{\text{m}^3}$$

$$\frac{6 \times 10^{24} \text{ kg}}{R^3} = 3 \times 10^{17}$$

$$T_{F} = \frac{\varepsilon_{F}}{12} \approx 7 \times 10^{11} \text{ K}$$