

Supplement: Jaynes-Bretthorst algorithm, and Bayesian periodograms for event data

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For event data details see Loredo (1993): “Bayesian Inference With Log-Fourier Arrival Time Models and Event Location Data” (unpublished report)

Agenda

- ① Jaynes-Bretthorst algorithm
- ② Periodograms for event time series

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② Periodograms for event time series

Bayesian Curve Fitting & Least Squares

Setup

Data $D = \{d_i\}$ are measurements of an underlying function $f(x; \theta)$ at N sample points $\{x_i\}$. Let $f_i(\theta) \equiv f(x_i; \theta)$:

$$d_i = f_i(\theta) + \epsilon_i, \quad \epsilon_i \sim N(0, \sigma_i^2)$$

We seek learn θ , or to compare different functional forms (model choice, M).

Likelihood

$$\begin{aligned} p(D|\theta, M) &= \prod_{i=1}^N \frac{1}{\sigma_i \sqrt{2\pi}} \exp \left[-\frac{1}{2} \left(\frac{d_i - f_i(\theta)}{\sigma_i} \right)^2 \right] \\ &\propto \exp \left[-\frac{1}{2} \sum_i \left(\frac{d_i - f_i(\theta)}{\sigma_i} \right)^2 \right] \\ &= \exp \left[-\frac{\chi^2(\theta)}{2} \right] \end{aligned}$$

Posterior

For prior density $\pi(\theta)$,

$$p(\theta|D, M) \propto \pi(\theta) \exp \left[-\frac{\chi^2(\theta)}{2} \right]$$

If you have a least-squares or χ^2 code:

- Think of $\chi^2(\theta)$ as $-2 \log \mathcal{L}(\theta)$.
- Bayesian inference amounts to exploration and numerical integration of $\pi(\theta)e^{-\chi^2(\theta)/2}$.
- If noise level is uncertain, keep the $1/\sigma_i$ factors (dropped above!) and include noise parameters in inference (e.g., scale all σ_i by a parameter, α)

Important Case: Separable Nonlinear Models

A (linearly) separable model has parameters $\theta = (A, \psi)$:

- Linear amplitudes $A = \{A_\alpha\}$
- Nonlinear parameters ψ

$f(x; \theta)$ is a linear superposition of M nonlinear components $g_\alpha(x; \psi)$:

$$d_i = \sum_{\alpha=1}^M A_\alpha g_\alpha(x_i; \psi) + \epsilon_i$$

or

$$\vec{d} = \sum_{\alpha} A_\alpha \vec{g}_\alpha(\psi) + \vec{\epsilon}.$$

Note: “linear/nonlinear” refers to how the predictions depend on the *parameters*, not how they depend on the sample location!

Examples

Polynomials (simple or orthogonal); $\psi = \emptyset$:

$$\begin{aligned}f(x) &= A_0 + A_1x + A_2x^2 + A_3x^3 \\&= A_0 + A_1x + A'_2(2x^2 - 1) + A'_3(4x^3 - 3x), \quad x \in [-1, 1] \\&= A_0g_0(x) + A_1g_1(x) + A'_2g_2(x) + A'_3g_3(x)\end{aligned}$$

Sinusoids; $\psi = \omega$:

$$\begin{aligned}f(x) &= A \cos(\omega x + \phi) \\&= A_1 \cos \omega x + A_2 \sin \omega x \\&= A_1g_1(x, \omega) + A_2g_2(x, \omega)\end{aligned}$$

Chirps; $\psi = (\omega, \alpha)$:

$$\begin{aligned}f(x) &= A \cos(\alpha x^2 + \omega x + \phi), \quad \text{inst. freq.} = \omega + 2\alpha x \\&= A_1 \cos(\alpha x^2 + \omega x) + A_2 \sin(\alpha x^2 + \omega x)\end{aligned}$$

Spectral continuum + lines; multipoles/spherical harmonics. . .

The Jaynes-Bretthorst Algorithm

Why separable structure is important: You can marginalize over A *analytically* \rightarrow *Jaynes-Bretthorst algorithm* (“Bayesian Spectrum Analysis & Param. Estimation” 1988)

Algorithm is closely related to linear least squares, diagonalization (eigenvectors/values), and SVD

Goals:

- Estimate the nonlinear parameters ψ
- Estimate amplitudes
- Compare rival models

The log-likelihood is a quadratic form in A_α ,

$$\mathcal{L}(A, \psi) \propto \frac{1}{\sigma^N} \exp \left[-\frac{Q(A, \psi)}{2\sigma^2} \right]$$

$$\begin{aligned} \text{with } Q &= \left[\vec{d} - \sum_{\alpha} A_{\alpha} \vec{g}_{\alpha} \right]^2 \\ &= \left[\vec{d} - \sum_{\alpha} A_{\alpha} \vec{g}_{\alpha} \right] \cdot \left[\vec{d} - \sum_{\beta} A_{\beta} \vec{g}_{\beta} \right] \\ &= d^2 - 2 \sum_{\alpha} A_{\alpha} \vec{d} \cdot \vec{g}_{\alpha} + \sum_{\alpha, \beta} A_{\alpha} A_{\beta} \eta_{\alpha\beta} \\ &\quad \text{where } \eta_{\alpha\beta}(\psi) = \vec{g}_{\alpha}(\psi) \cdot \vec{g}_{\beta}(\psi) \end{aligned}$$

We seek to integrate out the amplitudes, but completing the square is complicated because of the nontrivial metric $\eta_{\alpha\beta}$

Change the basis for \vec{f} from \vec{g}_α to an *orthonormal basis* \vec{h}_μ :

$$\vec{g}_\alpha = \sum_{\mu} a_{\alpha\mu} \vec{h}_\mu \quad \text{with } \vec{h}_\mu \cdot \vec{h}_\nu = \delta_{\mu\nu}$$

which implies $\vec{h}_\mu = \sum_{\alpha} (a^{-1})_{\mu\alpha} \vec{g}_\alpha$. Note $a = a(\psi)$.

Rewriting \vec{f} ,

$$\vec{f}(\theta) = \sum_{\alpha=1}^M A_{\alpha} \vec{g}_{\alpha}(\psi) = \sum_{\mu=1}^M B_{\mu}(A, \psi) \vec{h}_{\mu}(\psi)$$

with orthonormal amplitudes $B_{\mu}(A, \psi) = \sum_{\alpha} A_{\alpha} a_{\alpha\mu}(\psi)$

Some linear algebra shows that $\eta = aa^T$, so we can get a from η via Cholesky/eigen/QR decomposition.

Now write the quadratic form in terms of the B s instead of the A s:

$$\begin{aligned}
 Q &= d^2 - 2 \sum_{\alpha} A_{\alpha} \vec{d} \cdot \vec{g}_{\alpha} + \sum_{\alpha, \beta} A_{\alpha} A_{\beta} \eta_{\alpha\beta} \\
 &= d^2 - 2 \sum_{\mu} B_{\mu} \vec{d} \cdot \vec{h}_{\mu} + \sum_{\mu} B_{\mu}^2 \\
 &= \sum_{\mu} \left[B_{\mu} - \hat{B}_{\mu}(\psi) \right]^2 + r^2(\psi)
 \end{aligned}$$

with $\hat{B}_{\mu}(\psi) \equiv \vec{d} \cdot \vec{h}_{\mu}(\psi)$ and the residual $\vec{r}(\psi) \equiv \vec{d} - \sum_{\mu} \hat{B}_{\mu} \vec{h}_{\mu}$

The posterior in terms of B s is

$$p(B, \psi | D, I) \propto \frac{\pi(\psi) J(\psi)}{\sigma^N} \exp \left[-\frac{r^2(\psi)}{2\sigma^2} \right] \exp \left[\frac{-1}{2\sigma^2} \sum_{\mu} [B_{\mu} - \hat{B}_{\mu}(\psi)]^2 \right]$$

$J(\psi) = (\det \eta)^{1/2}$ comes from changing variables from A s to B s

Marginalize B 's analytically (*check the range!*):

$$p(\psi|D, I) \propto \frac{\pi(\psi)J(\psi)}{\sigma^{N-M}} \exp \left[-\frac{r^2(\psi)}{2\sigma^2} \right]$$

If σ unknown, marginalize using $p(\sigma|I) \propto \frac{1}{\sigma}$.

$$p(\psi|D, I) \propto \pi(\psi)J(\psi) [r^2(\psi)]^{\frac{M-N}{2}}$$

For given ψ , r^2 is just the residual sum of squares from a least squares fit to the basis functions. We can write

$$\begin{aligned} r^2(\psi) &= d^2 - \sum_{\mu} \hat{B}_{\mu}^2(\psi) \\ &= d^2 - S(\psi) \end{aligned}$$

with $S(\psi) = \sum_{\mu} [\vec{d} \cdot \vec{h}_{\mu}(\psi)]^2$, the sum of squared projections

Bayesian Spectrum Analysis

Adopt a sinusoid periodic signal model:

$$\begin{aligned} f(t) &= A \cos(\omega t - \phi) && \text{parameters } \omega, A, \phi \\ &= A_1 \cos \omega t + A_2 \sin \omega t && \text{parameters } \omega, A_1, A_2 \end{aligned}$$

$$d_i = f(t_i) + e_i \quad \text{Gaussian error pdfs; rms} = \sigma$$

Estimate ω :

$$\begin{aligned} p(\omega|D) &\propto \int dA_1 \int dA_2 p(\omega, A_1, A_2) \mathcal{L}(\omega, A_1, A_2) \\ &\propto p(\omega) J(\omega) \exp \left[\frac{S(\omega)}{\sigma^2} \right] \end{aligned}$$

- Equally-spaced samples: $S(\omega) \rightarrow P(\omega)$ for large N (when η is nearly diagonal)
- Unequally-spaced samples: $S(\omega) \approx$ Lomb-Scargle periodogram

Two Sinusoids

Adopt a model with two sinusoids at distinct frequencies:

$$\begin{aligned}f(t) &= A \cos(\omega_1 t - \phi_1) + B \cos(\omega_2 t - \phi_2) \\&= A_1 \cos \omega_1 t + A_2 \sin \omega_1 t \\&\quad + A_3 \cos \omega_2 t + A_4 \sin \omega_2 t\end{aligned}$$

Use JB algorithm to find $p(\omega_1, \omega_2 | D)$ or $p(\omega_1, \delta\omega | D)$

Call $\log p$ the *doublet periodogram*

If the frequencies are well-separated, can show that

$$p(\omega_1, \omega_2 | D) \propto p(\omega_1, \omega_2) J(\omega_1, \omega_2) \exp \left[\frac{S(\omega_1)}{\sigma^2} \right] \exp \left[\frac{S(\omega_2)}{\sigma^2} \right]$$

a product of two *independent* dist'ns, each determined by the periodogram

If the frequencies are close, there is no such simplification; the real and imaginary parts of the DTFT at each frequency (projections!) combine in a more complicated way

Exoplanets

Data are radial velocities; model as due to Keplerian motion.

For single planet:

$$V(t) = A_1 + A_2[e + \cos v(t)] + A_3 \sin v(t)$$

$$v(t) = f(t; \tau, e, M_0) \quad \text{via Kepler's eqn}$$

Period τ

3 linear amplitudes (COM velocity, orbital velocity, arg. of periastron)

2 other nonlinear parameters (e , M_0)

Follow the recipe! For $e = 0 \rightarrow$ LSP, but for nonzero eccentricity it generalizes to a *Kepler periodogram*

For astrometry, 2D data require $x(t)$, $y(t)$.

Extra parameters: inclination, parallax, proper motion.

Terminology for Generalized Periodograms

Fourier models

$\ln p(\omega|D, M) \propto$ periodogram (Schuster, Lomb-Scargle)

$\ln p(\omega_1, \omega_2|D, M) \propto$ doublet periodogram

Chirp model

$$A_1 \cos(\omega t + \alpha t^2) + A_2 \sin(\omega t + \alpha t^2)$$

$\ln p(\omega, \alpha|D, M) \propto$ Chirpogram (Jaynes)

Keplerian reflex motion model

$\ln p(\tau, e, M_p|D, M_p) =$ (Radial) Keplerogram

$\log p(\tau|D) \propto$ (Radial) Kepler periodogram

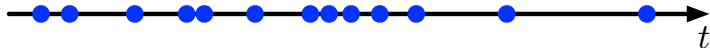
Cf. Box and other periodograms. . .

Agenda

① Jaynes-Bretthorst algorithm

② Periodograms for event time series

Time-Tagged Event/Arrival Time Series



Data D : $\{t_i\}$ for $i = 1$ to N

Nonhomogeneous Poisson point process sampling distribution, rate $r(t; \theta)$:

$$p(D|\theta, M) = \exp \left[- \int_T dt r(t) \right] \prod_{i=1}^N r(t_i)$$

Parametric models: $r(t; \theta)$; θ fixed dimension, e.g. (A, ω, ϕ)

Semiparametric models: $r(t; \theta, \psi)$

- θ of fixed dimension
- ψ of variable dimension (e.g., adaptive shape, # of components. . .)

Conventional period detection approaches

Try to reject a “null” (constant rate) hypothesis with an omnibus test; report p -value.

- Rayleigh statistic:

$$R^2(\omega) = \frac{1}{N} \left[\left(\sum_{i=1}^N \cos \omega t_i \right)^2 + \left(\sum_{i=1}^N \sin \omega t_i \right)^2 \right]$$

- Z_n^2 statistic:

$$Z_n^2(\omega) = \sum_{j=1}^n R^2(j\omega)$$

- χ^2 -Epoch folding:

- Fold data with trial period \rightarrow phases $\theta_i = \omega t_i + \phi \bmod 2\pi$;
bin $\rightarrow n_j, j = 1$ to M
- Calculate Pearson's $\chi^2(\omega, \phi)$ vs. $n_j = N/M$; *average over phase*

Bayesian Setup

Rate = avg. rate $A \times$ periodic shape $\rho(\theta)$, shape params \mathcal{S} :

$$r(t) = A\rho(\omega t - \phi; \mathcal{S})$$

Non-homogeneous point process likelihood (for $T \gg$ period):

$$\mathcal{L}(A, \omega, \phi, \mathcal{S}) = \left[A^N e^{-AT} \right] \prod_i \rho(\omega t_i - \phi; \mathcal{S})$$

Marginal likelihood for $\omega, \phi, \mathcal{S}$:

$$\mathcal{L}_m(\omega, \phi, \mathcal{S}) \propto \prod_i \rho(\omega t_i - \phi; \mathcal{S})$$

Goal: Find shapes where some Bayesian integrals are *analytical*

Bayesian Harmonic Modeling

Analytical phase marginalization

Log-sinusoid model

$$\rho(\theta) \propto e^{\kappa \cos \theta}, \quad \text{Von Mises dist'n}$$

Can analytically marginalize ϕ :

$$\mathcal{L}_m(\omega) \propto I_0[\kappa N R(\omega)] / I_0^N(\kappa)$$

Rayleigh statistic “appears;” κ provides sensitivity to narrow pulses (vs. Rayleigh test)

Log-Fourier models

For multi-peak light curves, consider log-Fourier models:

$$\rho(\theta) \propto \exp \left[\sum_j \kappa_j \cos(j\theta) \right]$$

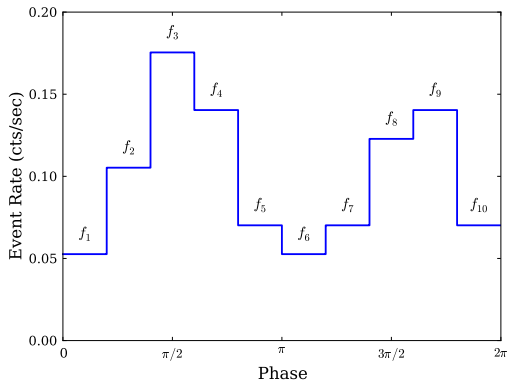
Issues: cannot analytically integrate ϕ ; dimensionality of κ

Laplace approximation for ϕ :

$$\mathcal{L}_m(\omega, \kappa) \approx \text{weighted } Z_n^2 + \text{interference terms}$$

Bayesian Stepwise Modeling

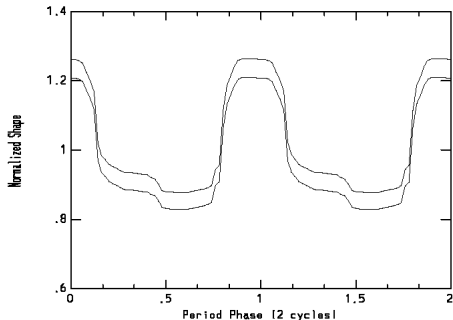
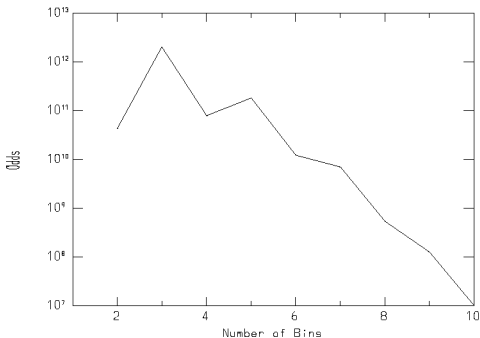
Analytical shape marginalization



Flat prior on $\mathbf{f} \rightarrow$ Can marginalize \mathbf{f} ; for M bins:

$$\mathcal{L}(\omega, \phi, \mathbf{f}) \propto \frac{(M-1)!}{(N+M-1)!} \left[\frac{n_1! n_2! \dots n_M!}{N!} \right] \sim \exp [\chi^2(\omega, \phi)]$$

X-Ray Pulsar PSR 0540-693 (Gregory & TL 1996)
3300 events over 10^5 s, many gaps, Rayleigh test fails

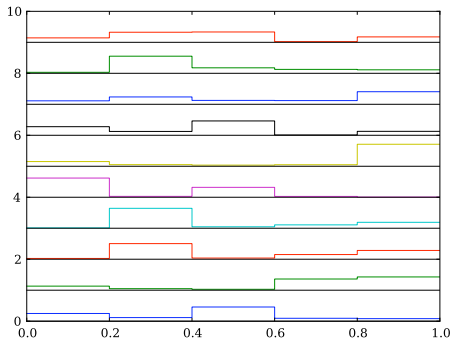


Other applications by Arnold Rots; see IAU 285 poster

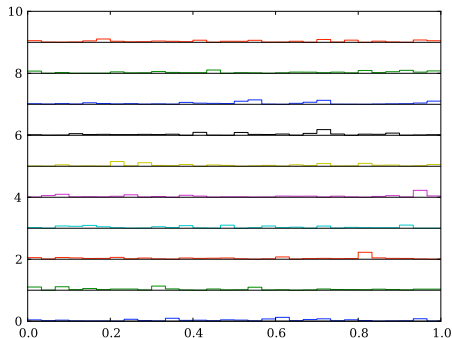
How Might We Do Better?

The flat stepwise shape prior is... *flat!*

Flat prior, $m=5$



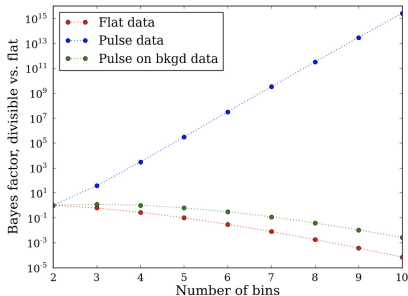
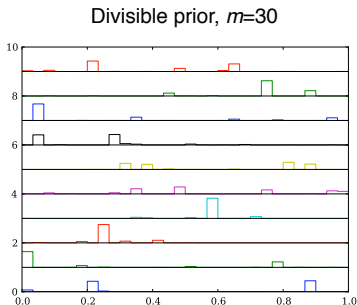
Flat prior, $m=30$



A simple example of a general phenomenon:
Concentration of measure

Promising direction

- Cross-model consistency requirement: 4-bin prior should become 2-bin prior when binned up, etc.
- → Form of prior should depend on dimension,
$$p(f) \propto \prod_{i=1}^M f^{\alpha-1} \text{ with } \alpha = \text{Const}/M$$
- This is called *divisibility*



Best solutions combine divisibility & structure, e.g., mixtures