# Introduction to Bayesian Inference: Supplemental Topics

Tom Loredo
Cornell Center for Astrophysics and Planetary Science
http://hosting.astro.cornell.edu/~loredo/

CASt Summer School — 5-9 June 2023

## **Supplemental Topics**

- Estimation and model comparison for binary outcomes; probability & frequency
- 2 Basic inference with normal errors
- 3 Poisson distribution; the on/off problem
- **4** Model uncertainty
- **5** Measurement error & hierarchical/graphical models
- **6** Typical sets (from MOO to MOE)
- Assigning priors

# **Supplemental Topics**

- Estimation and model comparison for binary outcomes; probability & frequency
- Basic inference with normal errors
- 3 Poisson distribution; the on/off problem
- 4 Model uncertainty
- **6** Measurement error & hierarchical/graphical models
- **1** Typical sets (from MOO to MOE)
- Assigning priors

# Binary Outcomes: Parameter Estimation

M= Existence of two outcomes, S and F; for each case or trial, the probability for S is  $\alpha$ ; for F it is  $(1-\alpha)$ 

 $H_i$  = Statements about  $\alpha$ , the probability for success on the next trial  $\rightarrow$  seek  $p(\alpha|D,M)$ 

D =Sequence of results from N observed trials:

FFSSSFSSFS (
$$n = 8$$
 successes in  $N = 12$  trials)

Likelihood:

$$p(D|\alpha, M) = p(\text{failure}|\alpha, M) \times p(\text{failure}|\alpha, M) \times \cdots$$
  
=  $\alpha^{n} (1 - \alpha)^{N-n}$   
=  $\mathcal{L}(\alpha)$ 

#### Prior

Starting with no information about  $\alpha$  beyond its definition, use as an "uninformative" prior  $p(\alpha|M) = 1$ . Justifications:

- Intuition: Don't prefer any  $\alpha$  interval to any other of same size
- Bayes's justification: "Ignorance" means that before doing the N trials, we have no preference for how many will be successes:

$$P(n \operatorname{success}|M) = \frac{1}{N+1} \longrightarrow p(\alpha|M) = 1$$

Consider this a *convention*—an assumption added to M to make the problem well posed.

#### Prior Predictive

$$p(D|M) = \int d\alpha \, \alpha^{n} (1-\alpha)^{N-n}$$

$$= B(n+1, N-n+1) = \frac{n!(N-n)!}{(N+1)!}$$

A Beta integral, 
$$B(a,b) \equiv \int dx \, x^{a-1} (1-x)^{b-1} = \frac{\Gamma(a)\Gamma(b)}{\Gamma(a+b)}$$
.

#### Posterior

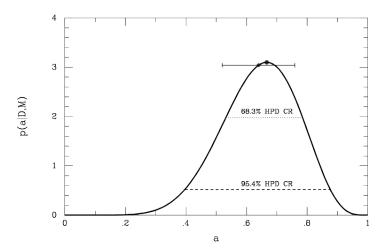
$$p(\alpha|D,M) = \frac{(N+1)!}{n!(N-n)!}\alpha^n(1-\alpha)^{N-n}$$

A Beta distribution. Summaries:

- Best-fit:  $\hat{\alpha} = \frac{n}{N} = 2/3$ ;  $\langle \alpha \rangle = \frac{n+1}{N+2} \approx 0.64$
- Uncertainty:  $\sigma_{\alpha} = \sqrt{\frac{(n+1)(N-n+1)}{(N+2)^2(N+3)}} \approx 0.12$ Find credible regions numerically, or with incomplete beta function

Note that the posterior depends on the data only through n, not the N binary numbers describing the sequence.

n is a (minimal) sufficient statistic.



# **Binary Outcomes: Model Comparison**

**Equal Probabilities?** 

 $M_1$ :  $\alpha = 1/2$ 

$$M_2$$
:  $\alpha \in [0,1]$  with flat prior.

### Maximum Likelihoods

$$M_1: p(D|M_1) = \frac{1}{2^N} = 2.44 \times 10^{-4}$$

$$M_2: \quad \mathcal{L}(\hat{\alpha}) = \left(\frac{2}{3}\right)^n \left(\frac{1}{3}\right)^{N-n} = 4.82 \times 10^{-4}$$

$$\frac{p(D|M_1)}{p(D|\hat{\alpha}, M_2)} = 0.51$$

Maximum likelihoods favor  $M_2$  (failures more probable).

### Bayes Factor (ratio of model likelihoods)

$$p(D|M_1) = \frac{1}{2^N};$$
 and  $p(D|M_2) = \frac{n!(N-n)!}{(N+1)!}$ 

Bayes factor (odds) favors  $M_1$  (equiprobable).

Note that for n = 6,  $B_{12} = 2.93$ ; for this small amount of data, we can never be very sure results are equiprobable.

If n=0,  $B_{12}\approx 1/315$ ; if n=2,  $B_{12}\approx 1/4.8$ ; for extreme data, 12 flips *can* be enough to lead us to strongly suspect outcomes have different probabilities.

(Frequentist significance tests can reject null for any sample size.)

## **Binary Outcomes: Binomial Distribution**

Suppose D = n (number of heads in N trials), rather than the actual sequence. What is  $p(\alpha|n, M)$ ?

#### Likelihood

Let S = a sequence of flips with n heads.

$$p(n|\alpha, M) = \sum_{S} p(S|\alpha, M) p(n|S, \alpha, M)$$

$$= \alpha^{n} (1 - \alpha)^{N-n}$$

$$= \alpha^{n} (1 - \alpha)^{N-n} C_{n,N}$$

$$= \alpha^{n} (1 - \alpha)^{N-n} C_{n,N}$$

 $C_{n,N} = \#$  of sequences of length N with n heads.

$$\to p(n|\alpha,M) = \frac{N!}{n!(N-n)!}\alpha^n(1-\alpha)^{N-n}$$

The *binomial distribution* for n given  $\alpha$ , N.

#### Posterior

$$p(\alpha|n, M) = \frac{\frac{N!}{n!(N-n)!}\alpha^n(1-\alpha)^{N-n}}{p(n|M)}$$

$$p(n|M) = \frac{N!}{n!(N-n)!} \int d\alpha \, \alpha^n(1-\alpha)^{N-n}$$

$$= \frac{1}{N+1}$$

$$\to p(\alpha|n, M) = \frac{(N+1)!}{n!(N-n)!}\alpha^n(1-\alpha)^{N-n}$$

Same result as when data specified the actual sequence.

# **Probability & frequency**

Frequencies are relevant when modeling repeated trials, or repeated sampling from a population or ensemble.

### Frequencies are observables

- When available, can be used to *infer* probabilities for next trial
- When unavailable, can be predicted

### Bayesian/Frequentist relationships

- Relationships between probability and frequency
- Long-run performance of Bayesian procedures

# Probability & frequency in IID settings

### Frequency from probability

Bernoulli's law of large numbers: In repeated i.i.d. trials, given  $P(\text{success}|...) = \alpha$ , predict

$$rac{ extit{n}_{ ext{success}}}{ extit{N}_{ ext{total}}} 
ightarrow lpha \quad ext{as} \quad extit{N}_{ ext{total}} 
ightarrow \infty$$

If p(x) does not change from sample to sample, it may be interpreted as a frequency distribution.

### Probability from frequency

Bayes's "An Essay Towards Solving a Problem in the Doctrine of Chances"  $\rightarrow$  First use of Bayes's theorem:

Probability for success in next trial of i.i.d. sequence:

$$\mathsf{E}(lpha) o rac{n_{
m success}}{N_{
m total}} \quad {\sf as} \quad N_{
m total} o \infty$$

If p(x) does not change from sample to sample, it may be estimated from a frequency distribution.

# The weather forecaster (Jaynes 1976)

Joint Frequencies of Actual & Predicted Weather

	Actual		
Prediction	Rain	Sun	
Rain	1/4	1/2	3/4
Sun	0	1/4	1/4
	1/4	3/4	,

Forecaster is right only 50% of the time

Observer notes a prediction of 'Sun' every day would be right 75% of the time, and applies for the forecaster's job

Should the observer get the job?

	Actual	
Prediction	Rain	Sun
Rain	1/4	1/2
Sun	0	1/4

Forecaster: You'll never be in an unpredicted rain

Observer: You'll be in an unpredicted rain 1 day out of 4

### Bayesian viewpoint

The value of an inference lies in its usefulness in the individual case

Long run performance is not an adequate criterion for assessing the usefulness of of an inference procedure

When long run performance is deemed important, it needs to be separately evaluated

# **Supplemental Topics**

- Estimation and model comparison for binary outcomes; probability & frequency
- 2 Basic inference with normal errors
- 3 Poisson distribution; the on/off problem
- 4 Model uncertainty
- **5** Measurement error & hierarchical/graphical models
- **1** Typical sets (from MOO to MOE)
- Assigning priors

# Inference With Normals/Gaussians

#### Gaussian PDF

$$p(x|\mu,\sigma) = \frac{1}{\sigma\sqrt{2\pi}}e^{-\frac{(x-\mu)^2}{2\sigma^2}}$$
 over  $[-\infty,\infty]$ 

Common abbreviated notation:  $x \sim N(\mu, \sigma^2)$ 

### **Parameters**

$$\mu = \langle x \rangle \equiv \int dx \, x \, p(x|\mu,\sigma)$$

$$\sigma^2 = \langle (x-\mu)^2 \rangle \equiv \int dx \, (x-\mu)^2 \, p(x|\mu,\sigma)$$

# Gauss's Observation: Sufficiency

Suppose our data consist of N measurements,  $d_i = \mu + \epsilon_i$ . Suppose the noise contributions are independent, and  $\epsilon_i \sim \mathcal{N}(0, \sigma^2)$ .

$$p(D|\mu, \sigma, M) = \prod_{i} p(d_{i}|\mu, \sigma, M)$$

$$= \prod_{i} p(\epsilon_{i} = d_{i} - \mu|\mu, \sigma, M)$$

$$= \prod_{i} \frac{1}{\sigma\sqrt{2\pi}} \exp\left[-\frac{(d_{i} - \mu)^{2}}{2\sigma^{2}}\right]$$

$$= \frac{1}{\sigma^{N}(2\pi)^{N/2}} e^{-Q(\mu)/2\sigma^{2}}$$

Find dependence of Q on  $\mu$  by completing the square:

$$Q = \sum_{i} (d_{i} - \mu)^{2} \qquad [\text{Note: } Q/\sigma^{2} = \chi^{2}(\mu)]$$

$$= \sum_{i} d_{i}^{2} + \sum_{i} \mu^{2} - 2 \sum_{i} d_{i}\mu$$

$$= \left(\sum_{i} d_{i}^{2}\right) + N\mu^{2} - 2N\mu\overline{d} \qquad \text{where } \overline{d} \equiv \frac{1}{N} \sum_{i} d_{i}$$

$$= N(\mu - \overline{d})^{2} + \left(\sum_{i} d_{i}^{2}\right) - N\overline{d}^{2}$$

$$= N(\mu - \overline{d})^{2} + Nr^{2} \quad \text{where } r^{2} \equiv \frac{1}{N} \sum_{i} (d_{i} - \overline{d})^{2}$$

Likelihood depends on  $\{d_i\}$  only through  $\overline{d}$  and r:

$$\mathcal{L}(\mu, \sigma) = \frac{1}{\sigma^N (2\pi)^{N/2}} \exp\left(-\frac{Nr^2}{2\sigma^2}\right) \exp\left(-\frac{N(\mu - \overline{d})^2}{2\sigma^2}\right)$$

The sample mean and variance are sufficient statistics.

This is a miraculous compression of information—the normal dist'n is highly *abnormal* in this respect!

# **Estimating a Normal Mean**

### Problem specification

Model:  $d_i = \mu + \epsilon_i$ ,  $\epsilon_i \sim N(0, \sigma^2)$ ,  $\sigma$  is known  $\to I = (\sigma, M)$ .

Parameter space:  $\mu$ ; seek  $p(\mu|D, \sigma, M)$ 

#### Likelihood

$$p(D|\mu, \sigma, M) = \frac{1}{\sigma^{N}(2\pi)^{N/2}} \exp\left(-\frac{Nr^{2}}{2\sigma^{2}}\right) \exp\left(-\frac{N(\mu - \overline{d})^{2}}{2\sigma^{2}}\right)$$

$$\propto \exp\left(-\frac{N(\mu - \overline{d})^{2}}{2\sigma^{2}}\right)$$

### "Uninformative" prior

Translation invariance  $\Rightarrow p(\mu) \propto C$ , a constant. This prior is *improper* unless bounded.

### Prior predictive/normalization

$$p(D|\sigma, M) = \int d\mu \ C \exp\left(-\frac{N(\mu - \overline{d})^2}{2\sigma^2}\right)$$
$$= C(\sigma/\sqrt{N})\sqrt{2\pi}$$

... minus a tiny bit from tails, using a proper prior.

#### Posterior

$$p(\mu|D,\sigma,M) = \frac{1}{(\sigma/\sqrt{N})\sqrt{2\pi}} \exp\left(-\frac{N(\mu-\overline{d})^2}{2\sigma^2}\right)$$

Posterior is  $N(\overline{d}, w^2)$ , with standard deviation  $w = \sigma/\sqrt{N}$ .

68.3% HPD credible region for  $\mu$  is  $\overline{d} \pm \sigma/\sqrt{N}$ .

Note that C drops out  $\rightarrow$  limit of infinite prior range is well behaved.

### Informative Conjugate Prior

Use a normal prior,  $\mu \sim N(\mu_0, w_0^2)$ .

Conjugate because the posterior turns out also to be normal.

#### Posterior

Normal  $N(\tilde{\mu}, \tilde{w}^2)$ , but mean, std. deviation "shrink" towards prior.

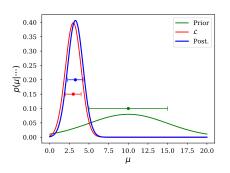
Define  $B = \frac{w^2}{w^2 + w_0^2}$ , so B < 1 and B = 0 when  $w_0$  is large. Then

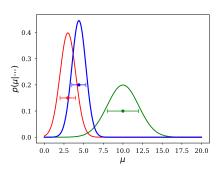
$$\widetilde{\mu} = \overline{d} + B \cdot (\mu_0 - \overline{d})$$
 $\widetilde{w} = w \cdot \sqrt{1 - B}$ 

"Principle of stable estimation" — The prior affects estimates only when data are not informative relative to prior.

### Conjugate normal examples:

- Data have  $\overline{d} = 3$ ,  $\sigma/\sqrt{N} = 1$
- Priors at  $\mu_0 = 10$ , with  $w = \{5, 2\}$





# Estimating a Normal Mean: Unknown $\sigma$

### Problem specification

Model:  $d_i = \mu + \epsilon_i$ ,  $\epsilon_i \sim N(0, \sigma^2)$ ,  $\sigma$  is unknown

Parameter space:  $(\mu, \sigma)$ ; seek  $p(\mu|D, M)$ 

#### Likelihood

$$p(D|\mu,\sigma,M) = \frac{1}{\sigma^N(2\pi)^{N/2}} \exp\left(-\frac{Nr^2}{2\sigma^2}\right) \exp\left(-\frac{N(\mu-\overline{d})^2}{2\sigma^2}\right)$$
 $\propto \frac{1}{\sigma^N} e^{-Q/2\sigma^2}$ 
where  $Q = N\left[r^2 + (\mu-\overline{d})^2\right]$ 

#### Uninformative Priors

Assume priors for  $\mu$  and  $\sigma$  are independent.

Translation invariance  $\Rightarrow p(\mu) \propto C$ , a constant.

Scale invariance  $\Rightarrow p(\sigma) \propto 1/\sigma$  (flat in log  $\sigma$ ).

Joint Posterior for  $\mu$ ,  $\sigma$ 

$$p(\mu, \sigma|D, M) \propto \frac{1}{\sigma^{N+1}} e^{-Q(\mu)/2\sigma^2}$$

### Marginal Posterior

$$p(\mu|D,M) \propto \int d\sigma \; rac{1}{\sigma^{N+1}} e^{-Q/2\sigma^2}$$
Let  $au = rac{Q}{2\sigma^2}$  so  $\sigma = \sqrt{rac{Q}{2 au}}$  and  $|d\sigma| = au^{-3/2} \sqrt{rac{Q}{2}} \; d au$ 

$$\Rightarrow p(\mu|D,M) \; \propto \; 2^{N/2} Q^{-N/2} \int d au \; au^{rac{N}{2}-1} e^{- au}$$

$$\propto \; Q^{-N/2}$$

Write 
$$Q = Nr^2 \left[ 1 + \left( \frac{\mu - \overline{d}}{r} \right)^2 \right]$$
 and normalize:

$$p(\mu|D,M) = \frac{\left(\frac{N}{2}-1\right)!}{\left(\frac{N}{2}-\frac{3}{2}\right)!\sqrt{\pi}}\frac{1}{r}\left[1+\frac{1}{N}\left(\frac{\mu-\overline{d}}{r/\sqrt{N}}\right)^2\right]^{-N/2}$$

"Student's t distribution," with  $t = \frac{(\mu - d)}{r/\sqrt{N}}$  A "bell curve," but with power-law tails Large N:

$$p(\mu|D,M) \sim e^{-N(\mu-\overline{d})^2/2r^2}$$

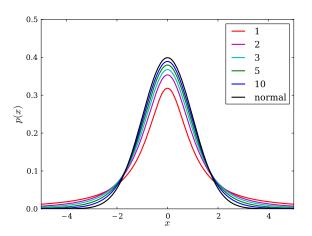
This is the rigorous way to "adjust  $\sigma$  so  $\chi^2/\text{dof} = 1$ ."

It doesn't just plug in a best  $\sigma$ ; it slightly broadens posterior to account for  $\sigma$  uncertainty.

### Student t examples:

• 
$$p(x) \propto \frac{1}{\left(1+\frac{x^2}{n}\right)^{\frac{n+1}{2}}}$$

- Location = 0, scale = 1
- Degrees of freedom =  $\{1, 2, 3, 5, 10, \infty\}$



# **Gaussian Background Subtraction**

Measure background rate  $b = \hat{b} \pm \sigma_b$  with source off.

Measure total rate  $r = \hat{r} \pm \sigma_r$  with source on.

Infer signal source strength s, where r = s + b.

With flat priors,

$$p(s, b|D, M) \propto \exp\left[-\frac{(b-\hat{b})^2}{2\sigma_b^2}\right] \times \exp\left[-\frac{(s+b-\hat{r})^2}{2\sigma_r^2}\right]$$

Marginalize b to summarize the results for s (complete the square to isolate b dependence; then do a simple Gaussian integral over b):

⇒ Background *subtraction* is a special case of background *marginalization*; i.e., marginalization "told us" to subtract a background estimate.

Recall the standard derivation of background uncertainty via "propagation of errors" based on Taylor expansion (statistician's *Delta-method*).

Marginalization provides a generalization of error propagation—without approximation!

# **Supplemental Topics**

- ① Estimation and model comparison for binary outcomes; probability & frequency
- Basic inference with normal errors
- 3 Poisson distribution; the on/off problem
- 4 Model uncertainty
- **5** Measurement error & hierarchical/graphical models
- **(5)** Typical sets (from MOO to MOE)
- Assigning priors

### Poisson Dist'n: Infer a Rate from Counts

#### Problem:

Observe n counts in T; infer rate, r

#### Likelihood

$$\mathcal{L}(r) \equiv p(n|r, M) = p(n|r, M) = \frac{(rT)^n}{n!}e^{-rT}$$

### Prior

Two simple standard choices (or conjugate gamma dist'n):

• r known to be nonzero; it is a scale parameter:

$$p(r|M) = \frac{1}{\ln(r_u/r_l)} \frac{1}{r}$$

• r may vanish; require  $p(n|M) \sim \text{Const}$ :

$$p(r|M) = \frac{1}{r_u}$$

### Prior predictive

$$p(n|M) = \frac{1}{r_u} \frac{1}{n!} \int_0^{r_u} dr (rT)^n e^{-rT}$$

$$= \frac{1}{r_u T} \frac{1}{n!} \int_0^{r_u T} d(rT) (rT)^n e^{-rT}$$

$$\approx \frac{1}{r_u T} \text{ for } r_u \gg \frac{n}{T}$$

#### **Posterior**

A gamma distribution:

$$p(r|n,M) = \frac{T(rT)^n}{n!}e^{-rT}$$

#### **Gamma Distributions**

A 2-parameter family of distributions over nonnegative x, with shape parameter  $\alpha$  and scale parameter s:

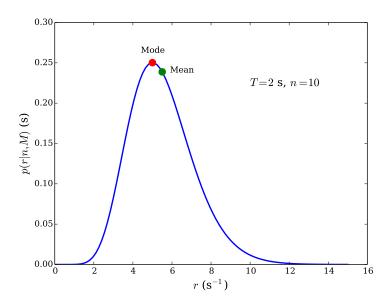
$$p_{\Gamma}(x|\alpha,s) = \frac{1}{s\Gamma(\alpha)} \left(\frac{x}{s}\right)^{\alpha-1} e^{-x/s}$$

Moments:

$$\mathsf{E}(x) = s\alpha \qquad \mathsf{Var}(x) = s^2 \alpha$$

Our posterior corresponds to  $\alpha = n + 1$ , s = 1/T.

- Mode  $\hat{r} = \frac{n}{T}$ ; mean  $\langle r \rangle = \frac{n+1}{T}$  (shift down 1 with 1/r prior)
- Std. dev'n  $\sigma_r = \frac{\sqrt{n+1}}{T}$ ; credible regions found by integrating (can use incomplete gamma function)



#### Conjugate prior

Note that a gamma distribution prior is the conjugate prior for the Poisson sampling distribution:

$$p(r|n, M') \propto \operatorname{Gamma}(r|\alpha, s) \times \operatorname{Pois}(n|rT)$$
  
 $\propto r^{\alpha-1}e^{-r/s} \times r^n e^{-rT}$   
 $\propto r^{\alpha+n-1} \exp[-r(T+1/s)]$ 

For  $\alpha=1,\ s\to\infty$ , the gamma prior becomes an "uninformative" flat prior; posterior is proper even for n=0

#### Useful conventions

- Use a flat prior for a rate that may be zero
- Use a log-flat prior  $(\propto 1/r)$  for a nonzero scale parameter
- Use proper (normalized, bounded) priors
- Plot posterior with abscissa that makes prior flat (use log r abscissa for scale parameter case)

# Infer a Signal in a Known Background

#### Problem:

As before, but r = s + b with b known; infer s

$$p(s|n,b,M) = C \frac{T \left[ (s+b)T \right]^n}{n!} e^{-(s+b)T}$$

$$C^{-1} = \frac{e^{-bT}}{n!} \int_0^\infty d(sT) (s+b)^n T^n e^{-sT}$$
$$= \sum_{i=0}^n \frac{(bT)^i}{i!} e^{-bT}$$

A sum of Poisson probabilities for background events; it can be evaluated using the incomplete gamma function. (Helene 1983)

# The On/Off Problem

#### Basic problem

- Look off-source; unknown background rate b
   Count N<sub>off</sub> photons in interval T<sub>off</sub>
- Look on-source; rate is r = s + b with unknown signal s Count  $N_{\rm on}$  photons in interval  $T_{\rm on}$
- Infer s

#### Conventional solution

$$\begin{split} \hat{b} &= N_{\rm off}/T_{\rm off}; & \sigma_b = \sqrt{N_{\rm off}}/T_{\rm off} \\ \hat{r} &= N_{\rm on}/T_{\rm on}; & \sigma_r = \sqrt{N_{\rm on}}/T_{\rm on} \\ \hat{s} &= \hat{r} - \hat{b}; & \sigma_s = \sqrt{\sigma_r^2 + \sigma_b^2} \end{split}$$

But  $\hat{s}$  can be *negative!* 

# Bayesian Solution to On/Off Problem

First consider off-source data; use it to estimate *b*:

$$p(b|N_{\rm off},I_{\rm off}) = \frac{T_{\rm off}(bT_{\rm off})^{N_{\rm off}}e^{-bT_{\rm off}}}{N_{\rm off}!}$$

Use this as a prior for b to analyze on-source data. For on-source analysis  $I_{\rm all} = (I_{\rm on}, N_{\rm off}, I_{\rm off})$ :

$$p(s,b|N_{
m on}) \propto p(s)p(b)[(s+b)T_{
m on}]^{N_{
m on}}e^{-(s+b)T_{
m on}} \quad || I_{
m all}$$
  $p(s|I_{
m all})$  is flat, but  $p(b|I_{
m all}) = p(b|N_{
m off},I_{
m off})$ , so  $p(s,b|N_{
m on},I_{
m all}) \propto (s+b)^{N_{
m on}}b^{N_{
m off}}e^{-sT_{
m on}}e^{-b(T_{
m on}+T_{
m off})}$ 

Now marginalize over b;

$$p(s|N_{\rm on}, I_{\rm all}) = \int db \ p(s, b \mid N_{\rm on}, I_{\rm all})$$

$$\propto \int db \ (s+b)^{N_{\rm on}} b^{N_{\rm off}} e^{-sT_{\rm on}} e^{-b(T_{\rm on}+T_{\rm off})}$$

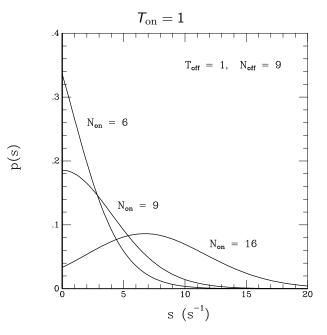
Expand  $(s+b)^{N_{\rm on}}$  and do the resulting  $\Gamma$  integrals:

$$p(s|N_{\rm on},I_{\rm all}) = \sum_{i=0}^{N_{\rm on}} C_i \frac{T_{\rm on}(sT_{\rm on})^i e^{-sT_{\rm on}}}{i!}$$

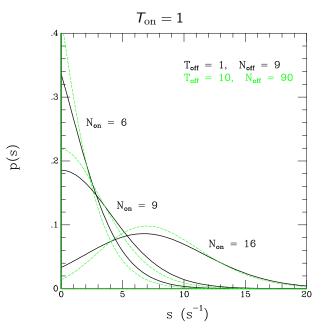
$$C_i \propto \left(1 + \frac{T_{\rm off}}{T_{\rm on}}\right)^i \frac{(N_{\rm on} + N_{\rm off} - i)!}{(N_{\rm on} - i)!}$$

Posterior is a weighted sum of Gamma distributions, each assigning a different number of on-source counts to the source. (Evaluate via recursive algorithm or confluent hypergeometric function.)

### Example On/Off Posteriors—Short Integrations



### Example On/Off Posteriors—Long Background Integrations



# Second Solution of the On/Off Problem

Consider all the data at once; the likelihood is a product of Poisson distributions for the on- and off-source counts:

$$\begin{array}{lll} \mathcal{L}(s,b) & \equiv & p(N_{\mathrm{on}},N_{\mathrm{off}}|s,b,I) \\ & \propto & \left[ (s+b)T_{\mathrm{on}} \right]^{N_{\mathrm{on}}} e^{-(s+b)T_{\mathrm{on}}} \times (bT_{\mathrm{off}})^{N_{\mathrm{off}}} e^{-bT_{\mathrm{off}}} \end{array}$$

Take joint prior to be flat; find the joint posterior and marginalize over *b*;

$$p(s|N_{\rm on}, I_{\rm on}) = \int db \ p(s, b|I) \mathcal{L}(s, b)$$

$$\propto \int db \ (s+b)^{N_{\rm on}} b^{N_{\rm off}} e^{-sT_{\rm on}} e^{-b(T_{\rm on}+T_{\rm off})}$$

 $\rightarrow$  same result as before.

### A profound consistency

We solved the on/off problem in multiple ways, finding the same final results.

This reflects something fundamental about Bayesian inference.

R. T. Cox proposed two necessary conditions for a quantification of uncertainty:

- It should duplicate deductive logic when there is no uncertainty
- Different decompositions of arguments should produce the same final quantifications (internal consistency)

Surprising result: These conditions (formalized) are *sufficient*; they require uncertainty quantification to follow the rules of Bayesian probability theory. E. T. Jaynes and others refined and simplified Cox's analysis.

# Multibin On/Off

The more typical on/off scenario:

Data = spectrum or image with counts in many bins

Model M gives signal rate  $s_k(\theta)$  in bin k, parameters  $\theta$ 

To infer  $\theta$ , we need the likelihood:

$$\mathcal{L}(\theta) = \prod_{k} \rho(N_{\text{on }k}, N_{\text{off }k} | s_k(\theta), M)$$

For each k, we have an on/off problem as before, only we just need the marginal likelihood for  $s_k$  (not the posterior). The same  $C_i$  coefficients arise.

XSPEC and CIAO/Sherpa provide this as an option (maybe not the latest Sherpa)

van  $Dyk^+(2001)$  does the same thing via data augmentation (Monte Carlo)

### **Supplemental Topics**

- ① Estimation and model comparison for binary outcomes; probability & frequency
- Basic inference with normal errors
- 3 Poisson distribution; the on/off problem
- **4** Model uncertainty
- **6** Measurement error & hierarchical/graphical models
- **1** Typical sets (from MOO to MOE)
- Assigning priors

### Model comparison

#### Problem statement

$$C = (M_1 \vee M_2 \vee ...)$$
 — Specify a set of models.  $H_i = M_i$  — Hypothesis chooses a model.

### Posterior probability for a model

$$p(M_i|D,C) = p(M_i|C)\frac{p(D|M_i,C)}{p(D|C)}$$

$$\propto p(M_i|C)\mathcal{L}(M_i)$$

$$\mathcal{L}(M_i) \equiv p(D|M_i) = \int d\theta_i \, p(\theta_i|M_i) p(D|\theta_i, M_i)$$

Likelihood for model = Average likelihood for its parameters

$$\mathcal{L}(M_i) = \langle \mathcal{L}(\theta_i) \rangle$$

Varied terminology: Prior predictive =  $\frac{Marginal\ likelihood}{Marginal\ likelihood}$  = Average likelihood = Global likelihood = (Weight of) Evidence for model

### **Odds and Bayes factors**

A ratio of probabilities for two propositions using the same premises is called the *odds* favoring one over the other:

$$O_{ij} \equiv \frac{p(M_i|D,C)}{p(M_j|D,C)}$$
$$= \frac{p(M_i|C)}{p(M_i|C)} \times \frac{p(D|M_i,C)}{p(D|M_i,C)}$$

The data-dependent part is called the Bayes factor.

$$B_{ij} \equiv \frac{p(D|M_i,C)}{p(D|M_j,C)}$$

It is a *likelihood ratio*; the BF terminology is usually reserved for cases when the likelihoods are marginal/average likelihoods for *composite hypotheses* 

#### An automatic Ockham's razor

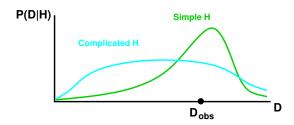
#### Consider nested models:

- Simpler model  $M_1$  with parameters  $\theta_1$
- "Larger" rival  $M_2$  with parameters  $\theta_2 = (\theta_1, \eta)$

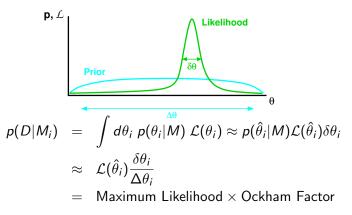
$$\Rightarrow \mathcal{L}(\hat{\theta}_2) \geq \mathcal{L}(\hat{\theta}_1)$$
But what about  $p(D|M_i) = \int d\theta_i \ p(\theta_i|M) \ \mathcal{L}(\theta_i)$ ?

#### Prior predictive distributions

Normalization implies there must be data that favor  $M_1$ :



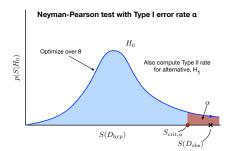
#### The Ockham factor

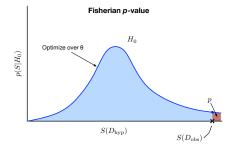


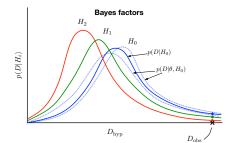
Models with more parameters often make the data more probable — for the best fit

Ockham factor penalizes models for "wasted" volume of parameter space

Quantifies intuition that models shouldn't require fine-tuning







• NP & Fisher give H<sub>0</sub> a special role

they don't measure "goodness-of-fit"

- NP & Fisher optimize over  $\theta$ , integrate over  $D_{\mbox{\scriptsize hyp}}$
- · Bayes considers rival H; symmetrically
- Bayes integrates over  $\theta$ , uses only  $D_{\mbox{obs}}$

Bayes factors can only compare rival models;

Posterior predictive p-values are a BDA alternative for measuring "suprisingness" of data for model checking; they integrate over both data and parameter spaces

See "p-value note" online for 2016 CASt summer; or 2018 Sagan workshop slides

### Model averaging

#### Problem statement

 $I = (M_1 \vee M_2 \vee \ldots)$  — Specify a set of models Models all share a set of "interesting" parameters,  $\phi$ Each has different set of nuisance parameters  $\eta_i$  (or different prior info about them)

 $H_i = \text{statements about } \phi$ 

### Model averaging

Calculate posterior PDF for  $\phi$ :

$$p(\phi|D,C) = \sum_{i} p(M_{i}|D,C) p(\phi|D,M_{i})$$

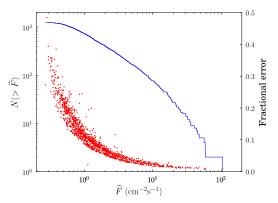
$$\propto \sum_{i} \mathcal{L}(M_{i}) \int d\eta_{i} p(\phi,\eta_{i}|D,M_{i})$$

The model choice is a (discrete) nuisance parameter here Useful for handling *systematic error* in estimation & prediction

### **Supplemental Topics**

- ① Estimation and model comparison for binary outcomes; probability & frequency
- Basic inference with normal errors
- 3 Poisson distribution; the on/off problem
- **4** Model uncertainty
- **6** Measurement error & hierarchical/graphical models
- **6** Typical sets (from MOO to MOE)
- Assigning priors

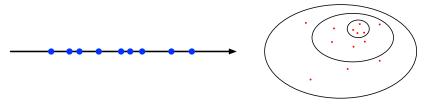
# Motivation: Complications with survey data



- Selection effects (truncation, censoring) obvious (usually)
   Typically treated by "correcting" data
   Most sophisticated: product-limit estimators
- "Scatter" effects (measurement error, etc.) insidious
   Typically ignored (average out? No—Eddington bias!)

### **Accounting for measurement error**

Suppose  $f(x|\theta)$  is a distribution for an observable, x (scalar or vector,  $\vec{x} = (x, y, ...)$ ); and  $\theta$  is unknown



From N precisely measured samples,  $\{x_i\}$ , we can infer  $\theta$  from

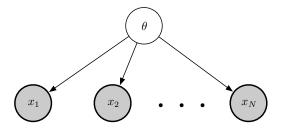
$$\mathcal{L}(\theta) \equiv p(\{x_i\}|\theta) = \prod_i f(x_i|\theta)$$
(A binomial point process)

$$p(\theta|\{x_i\}) \propto p(\theta)\mathcal{L}(\theta) = p(\theta,\{x_i\})$$

Posterior ∝ joint for params & data

#### Graphical representation

- Nodes/vertices = uncertain quantities (gray → known)
- Edges specify conditional dependence
- Absence of an edge denotes conditional independence

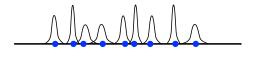


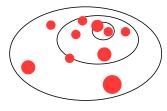
Graph specifies the form of the joint distribution:

$$p(\theta, \{x_i\}) = p(\theta) p(\{x_i\}|\theta) = p(\theta) \prod_i f(x_i|\theta)$$

Posterior from BT:  $p(\theta|\{x_i\}) = p(\theta,\{x_i\})/p(\{x_i\})$ 

But what if the x data are *noisy*,  $D_i = \{x_i + \epsilon_i\}$ ?





 $\{x_i\}$  are now uncertain (latent/hidden/incidental) parameters

*Member/item likelihoods* quantify uncertainties:  $\ell_i(x_i) = p(D_i|x_i)$ 

The joint PDF for everything is

$$p(\theta, \{x_i\}, \{D_i\}) = p(\theta) p(\{x_i\}|\theta) p(\{D_i\}|\{x_i\})$$
  
=  $p(\theta) \prod_i f(x_i|\theta) \ell_i(x_i)$ 

The conditional (posterior) PDF for the unknowns is

$$p(\theta, \{x_i\} | \{D_i\}) = \frac{p(\theta, \{x_i\}, \{D_i\})}{p(\{D_i\})} \propto p(\theta, \{x_i\}, \{D_i\})$$

$$p(\theta, \{x_i\}|\{D_i\}) \propto p(\theta, \{x_i\}, \{D_i\})$$

$$= p(\theta) \prod_i f(x_i|\theta) \ell_i(x_i)$$

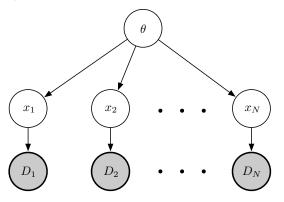
*Marginalize over*  $\{x_i\}$  to summarize inferences for  $\theta$ 

*Marginalize over*  $\theta$  to summarize inferences for  $\{x_i\}$ 

Key point: Maximizing over  $x_i$  (i.e., just using best-fit/MLE  $\hat{x}_i$ ) and integrating over  $x_i$  can give very different results!

(See Loredo (2004) for tutorial examples)

#### Graphical representation



$$p(\theta, \{x_i\}, \{D_i\}) = p(\theta) p(\{x_i\}|\theta) p(\{D_i\}|\{x_i\})$$

$$= p(\theta) \prod_i f(x_i|\theta) p(D_i|x_i) = p(\theta) \prod_i f(x_i|\theta) \ell_i(x_i)$$

A two-level hierarchical Bayes model, multi-level model (MLM), or probabilistic graphical model (PGM)

# Hierarchical Bayes/MLMs/PGMs in astronomy

A few entry points (see the *Resources* document for books with HB content):

- Joel Leja's tutorial at the 2022 Astroinformatics Summer School (next Wednesday!)
- Chapter by Loredo & Hendry: Multilevel and hierarchical Bayesian modeling of cosmic populations
- Survey of MLMs in astronomy: Bayesian astrostatistics: A backward look to the future (TJL 2013)
- CASt 2014 Supplement Session Includes discussion of selection effects
- TL's AAS workshop at AAS 231 (2018): Hierarchical modeling of cosmic populations
- CUDAHM C++ GPU software (Szalai-Gindl, Budavari, Kelly, TL); see Astron. & Comp. paper (longer: arXiv:2105.08026)

### **Supplemental Topics**

- ① Estimation and model comparison for binary outcomes; probability & frequency
- Basic inference with normal errors
- 3 Poisson distribution; the on/off problem
- 4 Model uncertainty
- **6** Measurement error & hierarchical/graphical models
- **6** Typical sets (from MOO to MOE)
- Assigning priors

#### Priors do more than shift the mode

Frequentist *penalized maximum likelihood* methods optimize the product of the likelihood and a penalty function,  $r(\theta)$  (e.g., a regularizer):

$$\hat{\theta} = \arg\max[r(\theta) \mathcal{L}(\theta)]$$

Equivalently, they add a penalty function to the log likelihood:

$$\hat{\theta} = \arg\max\left[\log r(\theta) + \log \mathcal{L}(\theta)\right]$$

The penalty function shifts the location of the maximum.

Many machine learning methods work similarly, modifying a default "goodness-of-fit" objective by adding a regularizer term.

The penalty *looks* like a prior, but because Bayesian calculations *integrate rather than maximize* over  $\theta$ , a prior can do much more than shift the location of the mode.

Perhaps surprisingly, even a flat (constant, uniform) prior can make you focus on parameter values away from the mode when working with high-dimensional models.

#### Relevant ideas:

- Curse of dimensionality (hi-D geometry)
- Concentration of measure (measure theory)
- Typical sets (information theory)

These all indicate that, in hi-D spaces with a kind of symmetry (product spaces), volume (probability!) can accumulate in unanticipated ways

### Typical sets

The mode of a discrete PMF or continuous PDF can be highly *atypical* compared to samples drawn from the dist'n.

This sounds surprising, but we have solid intuition about it for *sample spaces*.

### Coin flips (Bernoulli trials)

- Consider N = 1000 flips of a coin with  $\alpha = P(\text{heads}) = 0.8$ ; predict the number of heads
- The sequence of outcomes with the highest probability is 1000 heads
- Most of us would predict  $\approx 800 \pm 28$  heads
- The sequence with 1000 heads is  $\left(\frac{\alpha}{1-\alpha}\right)^{200}\approx 2\times 10^{120}$  more probable than any sequence with 800 heads
- But there are *many* sequences with  $\approx$  800 heads; the probability of that *typical set* is  $\approx$  1

### Gaussian samples and $\chi^2$

- Consider curve fitting with N = 1000  $(x_i, y_i)$  datapoints with  $y_i = f(x_i) + \epsilon_i$ , with standard normal errors  $\epsilon_i$
- The most probable sample has  $\epsilon_i = 0$  for all 1000 samples
- That sample would have  $\chi^2 = \sum_i [y_i f_i]^2 = \sum_i \epsilon_i^2 = 0$
- But we know that we expect  $\chi^2 \approx N \pm \sqrt{2N}$
- Note that  $\chi^2$  is just the squared distance of the  $\epsilon$  vector from the origin
- A set of random ε vectors will lie in a thin shell far from the origin when N is large (even though the marginal for every component peaks at the origin)
- The density of points is maximized at the origin, but volume grows so quickly with radius in hi-D that it becomes likely a random vector will point far from the origin

# In high-D spaces the mode of a distribution is very atypical

The math doesn't care if we call a variable "data" or "parameter"—this holds for hi-D parameter spaces as well as for large sample spaces.

MCMC *explores* (samples) rather than optimizes the posterior; as a result, the posterior samples lie in the typical set, and can be far from the mode.

If uncertainty quantification matters for your application, rather than "Model + Objective + Optimizer" (MOO), adopt "Model + Objective + Explorer" (MOE):



### **Supplemental Topics**

- Estimation and model comparison for binary outcomes; probability & frequency
- 2 Basic inference with normal errors
- 3 Poisson distribution; the on/off problem
- 4 Model uncertainty
- **6** Measurement error & hierarchical/graphical models
- **1** Typical sets (from MOO to MOE)
- Assigning priors

#### **Well-Posed Problems**

The rules (BT, LTP, ...) express desired probabilities in terms of other probabilities

To get a numerical value *out*, at some point we have to put numerical values *in* 

*Direct probabilities* are probabilities with numerical values determined directly by premises (via modeling assumptions, symmetry arguments, previous calculations, desperate presumption . . . )

An inference problem is *well posed* only if all the needed probabilities are assignable based on the context. We may need to add new assumptions as we see what needs to be assigned. We may not be entirely comfortable with what we need to assume! (Remember Euclid's fifth postulate!)

Should explore how results depend on uncomfortable assumptions ("robustness")

# Contextual/prior/background information

Bayes's theorem moves the data and hypothesis propositions wrt the solidus:

$$P(H_i|D_{ ext{obs}}, C) = P(H_i|C) \frac{P(D_{ ext{obs}}|H_i, C)}{P(D_{ ext{obs}}|C)}$$

It lets us add to the premises (here adding  $D_{obs}$  to C)

"Prior information" or "background information" or "context" = information that is **always** a premise (for the current calculation)

Notation:  $P(\cdot|\cdot, I)$  or  $P(\cdot|\cdot, C)$  or  $P(\cdot|\cdot, M)$  or ...

The context can be a notational nuisance! "Skilling conditional":

$$P(H_i|D_{\text{obs}}) = P(H_i) \frac{P(D_{\text{obs}}|H_i)}{P(D_{\text{obs}})} \qquad || \mathcal{C}$$

### **Essential contextual information**

We can only be uncertain about A if there are alternatives; what they are will bear on our uncertainty. We must explicitly specify relevant alternatives.

**Hypothesis space:** The set of alternative hypotheses of interest (and auxiliary hypotheses needed to predict the data)

**Data/sample space:** The set of possible data we may have predicted before learning of the observed data

**Predictive model:** Information specifying the likelihood function (e.g., the conditional predictive dist'n/sampling dist'n)

**Other prior information:** Any further information available or necessary to assume to make the problem *well posed* 

Bayesian literature often uses **model** to refer to *all* of the contextual information used to study a particular dataset and predictive model

### Directly assigned sampling distributions

Some examples of reasoning leading to sampling distributions:

- Binomial distribution:
  - ightharpoonup Ansatz: Probability for a Bernoulli trial,  $\alpha$
  - ightharpoonup LTP  $\Rightarrow$  binomial for *n* successes in *N* trials
- Poisson distribution:
  - Ansatz:  $P(\text{event in } dt | \lambda) \propto \lambda dt$ ; probabilities for events in disjoint intervals independent
  - ▶ Product & sum rules  $\Rightarrow$  Poisson for *n* in *T*
- Gaussian distribution:
  - ► CLT: Probability theory for sum of many quantities with independent, finite-variance PDFs
  - ➤ Sufficiency (Gauss): Seek distribution with sample mean as sufficient statistic (also sample variance)
  - ▶ Asymptotic limits: large *n* Binomial, Poisson
  - ▶ Others: Herschel's invariance argument (2-D), maximum entropy...

# **Assigning priors**

### Sources of prior information

- Analysis of previous experimental/observational data (but begs the question of what prior to use for the first such analysis)
- Subjective priors: Elicit a prior from an expert in the problem domain, e.g., via ranges, moments, quantiles, histograms
- Population priors: When it's meaningful to pool observations, we potentially can learn a shared prior—hierarchical/multilevel modeling does this

#### "Non-informative" priors

- Seek a prior that in some sense (TBD!) expresses a lack of information prior to considering the data
- No universal solution—this notion must be problem-specific, e.g., exploiting symmetries

# Priors derived from sampling distributions

Location/scale problems often admit a transformation group argument identifing good "noninformative" priors. In other settings we need a more general approach to formal assignment of priors that express "ignorance" in some sense.

There is no universal consensus on how to do this (yet? ever?).

A common underlying idea: The same  $\mathcal C$  appears in the prior,  $p(\theta|\mathcal C)$ , and the likelihood,  $p(D|\theta,\mathcal C)$ —the prior "knows" about the likelihood function, although it doesn't know what data values will be plugged into the sampling dist'n to get it.

**Jeffreys priors**: Use Fisher information to define a (parameter-dependent) scale defining a prior; parameterization invariant, but strange behavior in many dimensions.

**Reference priors**: Use information theory to define a prior that (asymptotically) has the least effect on the posterior; complicated algorithm; gives good frequentist behavior to Bayesian inferences.

# "Objective" priors

#### Heuristic motivation:

- Dimensionally,  $\pi(\theta) \propto 1/(\theta \text{ scale})$
- Use the likelihood function to determine a (relative) scale at each  $\theta$ , say,  $s(\theta) \to \pi(\theta) \propto 1/s(\theta)$
- Seek a scale definition that is consistent WRT reparameterization

Such a prior essentially specifies a way to slice-and-dice the  $\theta$  axis so assigning equal probability to intervals reflects intrinsic scales in the problem, and is consistent WRT reparameterization

### **Jeffreys priors**

#### Heuristic motivation:

 If we have data D, a natural inverse scale at θ, from the likelihood function, is the square root of the observed Fisher information (recall Laplace approximation):

$$I_D(\theta) \equiv -rac{\mathrm{d}^2 \log \mathcal{L}_D(\theta)}{\mathrm{d}\theta^2}$$

• For a prior, we don't know D; for each  $\theta$ , average over D predicted by the sampling distribution; this defines the (expected) Fisher information:

$$I(\theta) \equiv -\mathbb{E}_{D|\theta} \left[ rac{\mathrm{d}^2 \log \mathcal{L}_D(\theta)}{\mathrm{d}\theta^2} 
ight]$$

• Invariance: Can show for  $\phi = \Phi(\theta)$ , and  $\theta = \Theta(\phi)$ :

$$I(\phi) = I(\theta) \left(\frac{\mathrm{d}\Theta}{\mathrm{d}\phi}\right)^2$$

#### Jeffreys' prior.

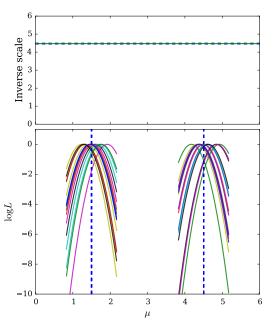
$$\pi(\theta) \propto [I(\theta)]^{1/2}$$

- Note the proportionality—the prior scale depends on how much the likelihood function scale *changes* vs.  $\theta$
- Puts more weight in regions of parameter space where the data are expected to be more informative
- Parameterization invariant, due to use of derivatives and vanishing expectation of the score function  $d \log \mathcal{L}_D(\theta)$

$$S_D(\theta) = \frac{\mathrm{d} \log \mathcal{L}_D(\theta)}{\mathrm{d} \theta}$$

- Typically improper when parameter space is non-compact
- Improves frequentist performance of posterior intervals w.r.t. intervals based on flat priors
- Only considered sound for a single parameter (or considering a single parameter at a time in some multiparameter problems)

### Jeffreys prior for normal mean



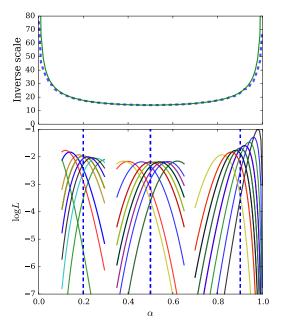
N=20 samples from normals with  $\sigma=1$ 

Likelihood width is independent of  $\mu \Rightarrow$ 

$$\pi(\mu) = \mathsf{Const}$$

Another justification of the uniform prior Prior is improper without prior limits on the range

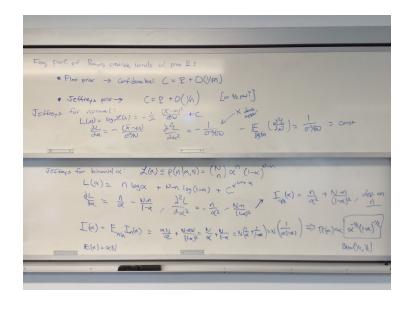
### Jeffreys prior for binomial probability



Binomial success counts n from N = 50 trials

$$\pi(\mu) = \frac{1}{\pi \alpha^{1/2} (1 - \alpha)^{1/2}}$$
$$= \text{Beta}(1/2, 1/2)$$

### Analytical calculations for normal, binomial



# **Limitations of the Jeffreys prior**

- Only considered sound for a single parameter (or considering a single parameter at a time in some multiparameter problems) E.g., for  $Norm(\mu, \sigma)$ , the Jeffreys prior is  $\propto 1/\sigma^2$ , not the product of separate Jeffreys  $\mu$ ,  $\sigma$  priors
- Only applicable to continuous spaces
- ightarrow Seek more general notions of "objective" or "uninformative" that reproduce good things about the Jeffreys prior

# Uncertainty, information, and entropy

Other rules for assigning "non-informative" priors rely on a more formal measure of the *information content* (or its complement, amount of *uncertainty*) in a probability distribution

Intuitively appealing metric-based measures, like standard deviation or interval size, are not general enough; e.g., they don't apply to categorical distributions

Desiderata for an *uncertainty functional*  $S_N[\vec{p}]$ —a map from a PMF  $\vec{p} = (p_1, p_2, \dots, p_N)$  to a single scalar quantifying its uncertainty (treat PDFs later):

- $S_N[\vec{p}]$  should be continuous w.r.t. the  $p_i$ s
- Uncertainty grows with multiplicity: When the  $p_i$  are all equal,  $s(N) = S_N[\vec{p}]$  should grow monotonically with N
- Additivity over subgroups

# **Information Gain as Entropy Change**

#### Entropy and uncertainty

Shannon entropy = a scalar measure of the degree of uncertainty expressed by a probability distribution

$$S = \sum_{i} p_{i} \log \frac{1}{p_{i}}$$
 "Average surprisal"  
=  $-\sum_{i} p_{i} \log p_{i}$ 

#### Information gain

Information gain upon learning D = decrease in uncertainty:

$$\mathcal{I}(D) = \mathcal{S}[\{p(H_i)\}] - \mathcal{S}[\{p(H_i|D)\}]$$

$$= \sum_{i} p(H_i|D) \log p(H_i|D) - \sum_{i} p(H_i) \log p(H_i)$$

### A 'Bit' About Entropy

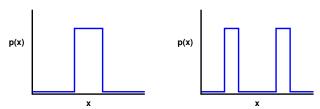
### Entropy of a Gaussian

$$p(x) \propto e^{-(x-\mu)^2/2\sigma^2} \rightarrow S \propto \log(\sigma)$$

$$\textit{p}(\vec{x}) \propto \exp\left[-\frac{1}{2}\vec{x} \cdot \mathsf{V}^{-1} \cdot \vec{x}\right] \ \, \rightarrow \ \, \mathcal{S} \propto \log(\det \mathsf{V})$$

→ Asymptotically like log Fisher matrix

A log-measure of "volume" or "spread," not range



These distributions have the same entropy/amount of information.

#### Expected information gain

When the data are yet to be considered, the *expected* information gain averages over D; straightforward use of the product rule/Bayes's theorem gives:

$$\mathbb{E}\mathcal{I} = \int dD \, p(D) \, \mathcal{I}(D)$$

$$= \int dD \, p(D) \, \sum_{i} p(H_{i}|D) \log \left[ \frac{p(H_{i}|D)}{p(H_{i})} \right]$$

For a continuous hypothesis space labeled by parameter(s)  $\theta$ ,

$$\mathbb{E}\mathcal{I} = \int dD \, p(D) \, \int d\theta \, p(\theta|D) \log \left[ rac{p(\theta|D)}{p(\theta)} 
ight]$$

This is the expectation value of the (data-dependent) Kullback-Leibler divergence between the prior and posterior:

$$\mathcal{D} \equiv \int d heta \, p( heta|D) \log \left[rac{p( heta|D)}{p( heta)}
ight]$$

### Reference priors

Bernardo (later joined by Berger & Sun) advocates *reference priors*, priors chosen to maximize the KLD between prior and posterior, as an "objective" expression of the idea of a "non-informative" prior: reference priors let the data most strongly dominate the prior (on average)

- Rigorous definition invokes asymptotics and delicate handling of non-compact parameter spaces to make sure posteriors are proper
- For 1-D problems, the reference prior is the Jeffreys prior
- In higher dimensions, the reference prior is not the Jeffreys prior; it behaves better
- The construction in higher dimensions is complicated and depends on separating interesting vs. nuisance parameters (but see Berger, Bernardo & Sun 2015, "Overall objective priors")
- Reference priors are typically improper on non-compact spaces
- They give Bayesian inferences good frequentist properties
- A constructive numerical algorithm exists (from particle physicists)

# Stan team's informal recommendations for priors

#### 5 levels of priors

- Flat prior (not usually recommended)
- Super-vague but proper prior: normal(0, 1e6) (not usually recommended)
- Weakly informative prior, very weak: normal(0, 10)
- Generic weakly informative prior: normal(0, 1)
- Specific informative prior: normal(0.4, 0.2) or whatever. Sometimes this can be expressed as a scaling followed by a generic prior:  $\theta = 0.4 + 0.2 \times z$ ;  $z \sim \text{normal}(0, 1)$

The above numbers assume that parameters are roughly on unit scale.

For details and further advice, see Stan's "Prior Choice Recommendations" Wiki