

§ 7.1.



$$(-1)^P = \begin{cases} +1 & \text{if } P \text{ involves even \# of exchanges} \\ -1 & \text{if } P \text{ involves odd \# of exchanges} \end{cases}$$

$$|\Psi(1, 2)\rangle = + |\Psi(2, 1)\rangle \quad \text{or}$$

$$|\Psi(1, 2)\rangle = - |\Psi(2, 1)\rangle$$

(1) Bosons

$$P |\Psi(1, \dots, N)\rangle = + |\Psi(1, \dots, N)\rangle$$

(2) Fermions

$$P |\Psi(1, \dots, N)\rangle = (-1)^P |\Psi(1, \dots, N)\rangle$$

$\mathcal{H}$  itself is symmetric for identical particles

$$P \mathcal{H} = \mathcal{H}$$

Eg. Consider  $N$  non-interacting particles, in a box of vol.  $V$ .

$$\mathcal{H} = \sum_{\alpha=1}^N \mathcal{H}_\alpha = \sum_{\alpha=1}^N \left( -\frac{\hbar^2}{2m} \nabla_\alpha^2 \right)$$

Each  $\mathcal{H}_\alpha$  can be diagonalized separately with plane wave states  $\{|\vec{k}\rangle\}$  & corresponding energies  $\epsilon(\vec{k}) = \frac{\hbar^2 k^2}{2m}$

Using sums & products of these states we can construct

(1) the product Hilbert space:

$$|\vec{k}_1, \dots, \vec{k}_N\rangle_{\otimes} = |\vec{k}_1\rangle \dots |\vec{k}_N\rangle$$
$$= |\vec{k}_1, \dots, \vec{k}_N\rangle$$

In coordinate representation,

$$\langle \vec{x}_1, \dots, \vec{x}_N | \vec{k}_1, \dots, \vec{k}_N \rangle = \frac{1}{\sqrt{N!}} \exp\left(i \sum_{\alpha=1}^N \vec{k}_{\alpha} \cdot \vec{x}_{\alpha}\right)$$

$$\delta \quad \mathcal{H} |\vec{k}_1, \dots, \vec{k}_N\rangle = \left( \sum_{\alpha=1}^N \frac{\epsilon^2}{2m} k_{\alpha}^2 \right) |\vec{k}_1, \dots, \vec{k}_N\rangle$$

However the product states don't satisfy the (anti)-symmetrization properties.

(2) Fermionic subspace:

$$|\vec{k}_1, \dots, \vec{k}_N\rangle_- = \frac{1}{\sqrt{N_-}} \sum_P (-1)^P P |\vec{k}_1, \dots, \vec{k}_N\rangle$$

where the sum is over all  $N!$  permutations

$$\rightarrow N_- = N!$$

$\rightarrow$  if two  $\vec{k}$ 's are equal, the state vanishes  
 $\rightarrow n_{\vec{k}} = 0 \text{ or } 1$  (# states with momentum  $\vec{k}$ )

(3) Bosonic Subspace:

$$|\vec{k}_1, \dots, \vec{k}_N\rangle_+ = \frac{1}{\sqrt{N_+}} \sum_P P (|\vec{k}_1, \dots, \vec{k}_N\rangle)$$

$$\Gamma_{+\{\vec{k}\} | \{\vec{k}\}} = \frac{1}{N_+} \sum_{P, P'} \langle P \{\vec{k}\} | P' \{\vec{k}\} \rangle$$

$$(\text{1Dk now}) = \frac{N!}{N_+} \sum_P \langle \{\vec{k}\} | P \{\vec{k}\} \rangle = \frac{N! \prod_{\vec{k}} n_{\vec{k}}!}{N_+} = 1$$

$$\Rightarrow N_+ = N! \prod_{k=1}^n k! \quad | \text{ Also } N_- = N! \prod_{k=1}^n k! \rightarrow !$$

$$\Rightarrow N^- = N!$$

-2-

$$\rightarrow N_+ = N! \prod_k (h_k!)$$

E.g. A bosonic state for 2 particles in  $|K\rangle$ ,  
1 particle in  $|P\rangle$ ,

we'll have

$$|\alpha\alpha\beta\rangle_+ = \frac{|\alpha\alpha\beta\rangle + |\alpha\beta\alpha\rangle + |\beta\alpha\alpha\rangle + |\alpha\alpha\beta\rangle + |\beta\alpha\alpha\rangle + |\alpha\beta\alpha\rangle}{\sqrt{12}}$$

$$= \frac{|\alpha\alpha\beta\rangle + |\alpha\beta\alpha\rangle + |\beta\alpha\alpha\rangle}{\sqrt{3}}$$

$\rightarrow$   $n \in$  can be any int  $> 0$ .

$$|\{\vec{k}\}\rangle_n = \frac{1}{\sqrt{N_\eta}} \sum_P \eta^P P |\{\vec{k}\}\rangle, \text{ with } \eta = \begin{cases} +1 & \text{for Bosons} \\ -1 & \text{for Fermions} \end{cases}$$

## § 7.3 GRAND CANONICAL FORMULATION

N particles in a box,  
volume V.

$$\sum_{\{K\}} = \sum_{K_1, K_2, K_3 \dots K_N} \text{ without going from } 0 \text{ to } \infty$$

$$Z_N = \text{Tr} (e^{-\beta H}) = \sum_{\{\vec{k}_\alpha\}} \langle \{\vec{k}_\alpha\} | e^{-\beta H} | \{\vec{k}_\alpha\} \rangle$$

DOUBT  
what when all  
are zero

$$= \sum_{\{\vec{k}_\alpha\}} e^{-\beta \sum_{\alpha=1}^N \epsilon(\vec{k}_\alpha)}$$

tensor product state  
 $\epsilon(\vec{k}) = \frac{k^2 R^2}{2m}$

$$= \sum_{\{\vec{k}_\alpha\}} e^{-\beta \sum_{\alpha=1}^N \varepsilon(\vec{k}_\alpha)} \quad \text{tensor product state} \quad \varepsilon(\vec{k}) = \frac{k^2 k^2}{2m}$$

: the particles are =  
identical, I can  
only say how many  
have what momentum

have what momentum.  
 Remarks: Hard to find,  $\therefore$  of restriction of  $\{n\}$ .

$$1) \quad \sum_{k=1}^n h_k = N$$

2) (A)  $n_{\bar{c}} = 0 \text{ or } 1$  (Fermion)  
 (B)  $n_{\bar{c}} = 0, 1, 2 \dots$  (Bosons)

These constraints can be removed by looking at

$$Q_\eta(T, \mu) = \sum_{N=0}^{\infty} e^{+\beta M N} \sum_{\substack{\{n_k\} \\ \sum_k n_k = N}} e^{-\left(\beta \sum_k \varepsilon(k) n_k\right)}$$

s.t.  $\sum_k n_k = N$

(claim)  $= \sum_{\{n_k\}} \prod_k e^{-\beta (\varepsilon(k) - \mu) n_k}$

Proof:

$$\begin{aligned} \sum_{N=0}^{\infty} \sum_{\substack{\{n_k\} \\ \sum_k n_k = N}} &= \sum_{\substack{\{n_k\} \\ \sum_k n_k = 0}} + \sum_{\substack{\{n_k\} \\ \sum_k n_k = 1}} + \dots \\ &= \sum_{\substack{\{n_k\} \\ \sum_k n_k = 1}} \\ \text{So } Q_\eta(T, \mu) &= \sum_{\{n_k\}} e^{+\beta M \sum_k n_k - \beta \sum_k \varepsilon(k) n_k} \\ &= \sum_{\{n_k\}} e^{\sum_k \beta (\mu - \varepsilon(k)) n_k} \\ &= \sum_{\{n_k\}} \prod_k e^{-\beta (\varepsilon(k) - \mu) n_k} \\ \text{Not sure} &\quad = \prod_k \sum_{n_k} e^{-\beta (\varepsilon(k) - \mu) n_k} \end{aligned}$$

Remarks:

1) Fermions:  $n_k = 0 \text{ or } 1$ .

$$Q_- = \prod_k [1 + e^{-\beta (\varepsilon(k) - \mu)}]$$

2) Bosons:  $n_k = 0, 1, 2, \dots$

$$Q_+ = \prod_k \frac{1}{[1 + e^{-\beta (\varepsilon(k) - \mu)}]}$$

Combined

$$\ln Q_\eta = -\eta \sum_k \ln [1 - n e^{-\beta (\varepsilon(k) - \mu)}]$$

$$\begin{aligned} f(x) &= a + ax + ax^2 \dots \\ a + xf(x) &= f(x) \\ \Rightarrow \frac{a}{1-x} &= f(x) \end{aligned}$$

In the grand canonical formulation, different one-particle states are occupied independently. Thus we have the following

$$P_\eta(\{n(\vec{k})\}) = \frac{1}{Q_\eta} \prod_{\vec{k}} \left[ e^{-\beta(\varepsilon(\vec{k}) - \mu)/k_B T} \right]$$

$\uparrow$   
we just calculated this

joint probability.

The average occupation number of a state with energy  $\varepsilon(\vec{k})$  is given by

$$\boxed{\langle n_{\vec{k}} \rangle_\eta = - \frac{\partial \ln Q_\eta}{\partial (\beta \varepsilon(\vec{k}))} = \frac{1}{\Xi^{-1} e^{\beta \varepsilon(\vec{k})} - \eta}}$$

where  $\Xi = e^{\beta \mu}$

[Proof]

$$\begin{aligned} & \sum_{\vec{k}} \ln [1 - \eta e^{-\beta(\varepsilon(\vec{k}) - \mu)}] \\ &= \sum_{\vec{k}} \frac{1}{\Xi^{-1} e^{\beta(\varepsilon(\vec{k}) - \mu)}} \cdot (-\beta(\varepsilon(\vec{k}) - \mu)) \\ &= \chi \frac{1}{1 - \eta e^{-\beta(\varepsilon(\vec{k}) - \mu)}} \cdot (-\beta(\varepsilon(\vec{k}) - \mu)) \\ &= \frac{1}{e^{\beta(\varepsilon(\vec{k}) - \mu)} - \eta} \end{aligned}$$

□

The average values of particle # l internal energy are then given by

$$\boxed{\begin{aligned} N_\eta &= \sum_{\vec{k}} \langle n_{\vec{k}} \rangle_\eta = \sum_{\vec{k}} \frac{1}{\Xi^{-1} e^{\beta \varepsilon(\vec{k})} - \eta} \\ E_\eta &= \sum_{\vec{k}} \varepsilon(\vec{k}) \langle n_{\vec{k}} \rangle_\eta = \sum_{\vec{k}} \frac{\varepsilon(\vec{k})}{\Xi^{-1} e^{\beta \varepsilon(\vec{k})} - \eta} \end{aligned}}$$

## § 7.4 | Non-relativistic gas

- Quantum Particles are further characterized by spin.
- In the absence of a magnetic field, different spin states have the same energy.  
Thus a spin degeneracy factor

$$g = 2S + 1$$

is multiplied to, in the  $\ln \Omega_1$ , e.g.

- Further, taking  $V$  to be large, we have  $\sum_k \rightarrow \frac{V}{(2\pi)^3} d^3 k$

also anyway,  $\epsilon(\vec{k}) = \frac{\hbar^2 k^2}{2m}$

$$\beta P_1^{\text{measure}} = \frac{\ln \Omega_1}{V} = -\eta g \int \frac{d^3 k}{(2\pi)^3} \ln \left[ 1 - \eta e^{-\frac{\beta \hbar^2 k^2}{2m}} \right]$$

$$\Gamma - \beta f = \ln \Omega$$

$$E = TS + \mu N - PV$$

$$f = E - TS - \mu N$$

$$\hookrightarrow f = -PV$$

$$n_1 = \frac{N_1}{V} = g \int \frac{d^3 k}{(2\pi)^3} \frac{1}{\left( e^{-\frac{\beta \hbar^2 k^2}{2m}} - \eta \right)}$$

density of states?

$$\varepsilon_1 = \frac{E_1}{V} = g \int \frac{d^3 k}{(2\pi)^3} \xrightarrow{\epsilon(k) \rightarrow \frac{\hbar^2 k^2}{2m}} \frac{1}{\left( e^{-\frac{\beta \hbar^2 k^2}{2m}} - \eta \right)}$$

Energy density?

To simplify, we change variables to  $x = \frac{\beta \hbar^2 k^2}{2m}$

so that  $k = \sqrt{\frac{2m k_B T}{\hbar^2}} x^{1/2} = \frac{2\pi x^{1/2}}{\lambda} \Rightarrow dk = \frac{\pi x^{1/2}}{\lambda} dx$

Then

$$\begin{aligned} \beta P_n &\stackrel{\text{claim}}{=} -\eta \frac{g}{2\pi^2} \cdot \frac{4\pi^{3/2}}{\lambda^3} \int_0^\infty dx x^{y_2} \ln(1 - \eta \leq e^{-x}) \\ &= \frac{g}{\lambda^3} \cdot \frac{4}{3\sqrt{\pi}} \int_0^\infty \frac{dx x^{y_2}}{\xi^{-1}e^x - \eta} \quad (\text{Integration by parts}) \end{aligned}$$

$$\begin{aligned} \Gamma_{\beta P_n} &= -\eta g \int \frac{d^3 k}{(2\pi)^3} \ln \left[ 1 - \eta \leq e^{-\frac{\beta k^2 \lambda^2}{2m}} \right] \\ &= -\eta g 4\pi \int \frac{k^2 dk}{(2\pi)^3} \ln \left[ 1 - \eta \leq e^{-\frac{\beta k^2 \lambda^2}{2m}} \right] \\ &= -\eta g \frac{4\pi}{(2\pi)^3} \int \left( \frac{4\pi}{\lambda^2} \int x^{y_2} dx \cdot \frac{\pi^{y_2}}{\lambda} \right) \ln \left[ 1 - \eta \leq e^{-\frac{\beta k^2 \lambda^2}{2m}} \right] \\ \text{claim 1 } \square &= -\frac{\eta g}{2\pi^2} \cdot \frac{4\pi^{3/2}}{\lambda^3} \int_0^\infty dx x^{y_2} \ln \left[ 1 - \eta \leq e^{-x} \right] \end{aligned}$$

$$\begin{aligned} \Gamma &\int_0^\infty dx x^{y_2} \ln \left[ 1 - \eta \leq e^{-x} \right] \\ &= dx \left[ \ln \left[ 1 - \eta \leq e^{-x} \right] \cdot \frac{x^{3/2}}{3/2} \right]_0^\infty \quad (\text{Claim, not justified, can try (Hospital)} \\ &\quad - \int_0^\infty dx \frac{x^{3/2}}{3/2} \cdot \frac{1}{(1 - \eta \leq e^{-x})} \cdot \eta \leq e^{-x} \\ &= + \frac{\eta g}{2\pi^2 \lambda^{-3/2}} \cdot \frac{4\pi^{3/2}}{\lambda^3} \cdot \frac{\chi}{3} \quad \chi \stackrel{\infty}{=} \int_0^\infty \frac{dx x^{3/2}}{\eta \leq^{-1} e^x - \eta} \end{aligned}$$

 $\square$ 

$$n_2 \stackrel{\text{claim}}{=} \frac{g}{\lambda^3} \frac{2}{\sqrt{\pi}} \int_0^\infty \frac{dx x^{y_2}}{\xi^{-1}e^x - \eta}$$

$$\begin{aligned} \Gamma &= \frac{g}{(2\pi)^3} \int_0^\infty 4\pi k^2 dk \frac{1}{\xi^{-1}e^x - \eta} \\ &= \frac{2g}{2 \cdot 4\pi \lambda^2 \sqrt{\pi}} \int_0^\infty 4\pi \left( \frac{\sqrt{\pi}}{\lambda} x^{-y_2} dx \right) \left( \frac{4\pi x}{\lambda^2} \right) \frac{1}{\xi^{-1}e^x - \eta} \\ &= \frac{2g}{\lambda^3 \sqrt{\pi}} \int_0^\infty \frac{dx x^{y_2}}{\xi^{-1}e^x - \eta} \end{aligned}$$

 $\square$

$$\beta \varepsilon_1 = \frac{g}{\lambda^3} \frac{2}{\sqrt{\pi}} \int_0^\infty \frac{dx}{\xi^{-1} e^x - \eta} x^{3/2}$$

$\square$  obvious

We define

$$f_m^n(z) = \frac{1}{(m-1)!} \int_0^\infty \frac{dx}{\xi^{-1} e^x - z} x^{m-1}$$

$$\Gamma_{m!} \equiv \Gamma(m+1) = \int_0^\infty dx x^m e^{-x}$$

$$\text{claim: } \left(\frac{1}{2}\right)! = \Gamma\left(\frac{3}{2}\right) = \frac{\sqrt{\pi}}{2}$$

$$\Rightarrow \left(\frac{3}{2}\right)! = \frac{3}{2} \frac{\sqrt{\pi}}{2}$$

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$$\Rightarrow \beta P_2 = \frac{g}{\lambda^3} f_{5/2}^n (\approx)$$

$$n_2 = \frac{g}{\lambda^3} f_{3/2}^n (\approx)$$

$$\varepsilon_2 = \frac{3}{2} P_2$$

To find the eq<sup>n</sup> of state explicitly, we need to understand

$$f_m^n(\xi) =$$

sum of gp  
↓ or else  
binomial expand

$$\frac{1}{(m-1)!} \int_0^\infty \frac{dx}{\xi^{-1} e^x - \eta} x^{m-1} = \frac{1}{(m-1)!} \int_0^\infty dx x^{m-1} (\xi e^{-x}) (1 - \eta \xi e^{-x})^{-1}$$

$$= \frac{1}{(m-1)!} \int_0^\infty dx x^{m-1} (\xi e^{-x}) \sum_{\alpha=0}^{\infty} (\eta)^{\alpha} (\xi e^{-x})^{\alpha} = \sum_{\alpha=0}^{\infty} \eta^{\alpha+1} (\xi)^{\alpha} \left[ \frac{1}{(m-1)!} \int_0^\infty dx x^{m-1} e^{-\alpha x} \right]$$

$$\Gamma \xi e^{-x} \sum_{\alpha=0}^{\infty} (\eta)^{\alpha} (\xi e^{-x})^{\alpha} = \frac{1}{\alpha^m} \int_0^\infty dx (\alpha x)^{m-1} e^{-\alpha x}$$

$$= \sum_{\alpha=0}^{\infty} (\eta)^{\alpha} (\xi e^{-x})^{\alpha+1} \Big| = \frac{(m-1)!}{\alpha^m}$$

$$= \sum_{\alpha=1}^{\infty} (\eta)^{\alpha-1} (\xi e^{-x})^{\alpha} \cdot (\eta)^2 \Big| = \int_0^\infty dx x^{m-1} e^{-\alpha x}$$

$$\boxed{= \sum_{\alpha=1}^{\infty} (\eta)^{\alpha+1} (\xi e^{-x})^{\alpha}}$$

$$= \sum_{\alpha=1}^{\infty} \eta \frac{\gamma^{\alpha+1}}{\alpha} \frac{\Sigma^\alpha}{m} \quad \frac{(m-1)!}{(m-1)!}$$

$$= \Sigma + \eta \frac{\Sigma^2}{2^{\frac{3}{2}}} + \frac{\Sigma^3}{3^{\frac{3}{2}}} + \eta \frac{\Sigma^4}{4^{\frac{3}{2}}} + \dots$$

We thus have

$$\frac{n_1 \lambda^3}{g} = f_{3/2}^n(\Sigma) = \Sigma + \eta \frac{\Sigma^2}{2^{3/2}} + \frac{\Sigma^3}{3^{3/2}} + \eta \frac{\Sigma^4}{4^{3/2}} + \dots$$

$$\beta \frac{P_n \lambda^3}{g} = f_{5/2}^n(\Sigma) = \Sigma + \eta \frac{\Sigma^2}{2^{5/2}} + \frac{\Sigma^3}{3^{5/2}} + \eta \frac{\Sigma^4}{4^{5/2}} + \dots$$

We solve perturbatively using recursion.

$$\Sigma = \frac{n_1 \lambda^3}{g} - \eta \frac{\Sigma^2}{2^{3/2}} - \frac{\Sigma^3}{3^{3/2}} + \mathcal{O}(\Sigma^4)$$

$$= \left( \frac{n_1 \lambda^3}{g} \right) - \frac{\eta}{2^{3/2}} \left( \frac{n_1 \lambda^3}{g} \right)^2 - \frac{\Sigma^3}{3^{3/2}} + \mathcal{O}(\Sigma^4)$$

$$= \left( \frac{n_1 \lambda^3}{g} \right) - \frac{\eta}{2^{3/2}} \left( \frac{n_1 \lambda^3}{g} \right)^2 + \left( \frac{1}{4} - \frac{1}{3^{3/2}} \right) \left( \frac{n_1 \lambda^3}{g} \right)^3$$

$$\Gamma \Sigma^2 \approx \left( \frac{n_1 \lambda^3}{g} - \eta \frac{\Sigma^2}{2^{3/2}} \right)^2$$

$$\approx \left( \frac{n_1 \lambda^3}{g} \right)^2 - 2 \left( \eta \frac{\Sigma^2}{2^{3/2}} \right) \left( \frac{n_1 \lambda^3}{g} \right) \approx \left( \frac{n_1 \lambda^3}{g} \right)^2 - 2 \left( \eta \frac{\Sigma^2}{2^{3/2}} \left( \frac{n_1 \lambda^3}{g} \right)^2 \right) \left( \frac{n_1 \lambda^3}{g} \right)$$

$$\begin{aligned} \Sigma^3 &\approx \left( \frac{n_1 \lambda^3}{g} + \eta \frac{\Sigma^2}{2^{3/2}} \right)^3 \\ &\approx \left( \frac{n_1 \lambda^3}{g} \right)^3 + \beta \left( \frac{n_1 \lambda^3}{g} \right)^2 \frac{\eta \Sigma^2}{2^{3/2}} + \mathcal{O}(\Sigma^3) \end{aligned}$$

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$$\text{Note: } \Sigma^2 \approx \left( \frac{n_1 \lambda^3}{g} \right)^2$$

$$\Sigma^3 \approx - \left( \frac{1}{4} - \frac{1}{3^{3/2}} \right) \left( \frac{n_1 \lambda^3}{g} \right)^3 \cdot 3^{3/2}$$

(just look at the series we obtained & look at the contribution now in 4 where its coming from)

$$\frac{\beta P_n \lambda^3}{g} = \Sigma + \frac{n}{2^{5/2}} \Sigma^2 + \frac{n}{3^{5/2}} \Sigma^3 + \dots$$

we substitute for  $\Sigma$ ,  $\Sigma^2$  &  $\Sigma^3$  as just calculated.

$$\frac{\beta P_n \lambda^3}{g} = \left[ \left( \frac{n_1 \lambda^3}{g} \right) - \frac{1}{2^{3/2}} \left( \frac{n_1 \lambda^3}{g} \right)^2 + \left( \frac{1}{4} - \frac{1}{3^{3/2}} \right) \left( \frac{n_1 \lambda^3}{g} \right)^3 \dots \right]$$

$$+ \frac{n}{2^{5/2}} \left[ \left( \frac{n_1 \lambda^3}{g} \right)^2 + \dots \right] + \frac{1}{3^{5/2}} \left[ - \left( \frac{1}{4} - \frac{1}{3^{3/2}} \right) \left( \frac{n_1 \lambda^3}{g} \right)^3 + \dots \right]$$

$$P = k_B T n_1 \left( \frac{n_1 \lambda^3}{g} \right)^{-1} \left\{ \left[ \left( \frac{n_1 \lambda^3}{g} \right) - \dots \right] \right\}$$

$$P = k_B T n_1 \left\{ \left[ 1 - \frac{n}{2^{3/2}} \left( \frac{n_1 \lambda^3}{g} \right)^2 + \left( \frac{1}{4} - \frac{1}{3^{3/2}} \right) \left( \frac{n_1 \lambda^3}{g} \right)^3 \dots \right] \right\}$$

$$+ \frac{n}{2^{5/2}} \left[ \left( \frac{n_1 \lambda^3}{g} \right)^2 + \dots \right] + \frac{1}{3^{5/2}} \left[ \left( \frac{1}{3^{3/2}} - \frac{1}{4} \right) \cdot 3^{3/2} \left( \frac{n_1 \lambda^3}{g} \right)^3 + \dots \right]$$

just add coefficients we get

$$P = n_1 k_B T \left[ 1 - \frac{n}{2^{5/2}} \left( \frac{n_1 \lambda^3}{g} \right)^2 + \left( \frac{1}{8} - \frac{2}{3^{5/2}} \right) \left( \frac{n_1 \lambda^3}{g} \right)^3 + \dots \right]$$

$$\frac{n_1 \lambda^3}{g} \geq 1 \quad (\text{the quantum degenerate limit})$$

this is when quantum fix becomes important.

## § 7.5 THE DEGENERATE FERMI GAS

At zero temperature limit, the fermi occupation #

$$\langle n_{\vec{k}} \rangle_- = \frac{1}{e^{\beta(\varepsilon(\vec{k}) - \mu)} + 1} = \begin{cases} 1 & \varepsilon(\vec{k}) - \mu < 0 \\ 0 & \text{else} \end{cases}$$

The limiting value of  $\mu$  at zero temperature is called the fermi energy,  $E_F$

All states with energy less than  $E_F$  are occupied, forming a fermi sea.

For the ideal gas with  $\varepsilon(\vec{k}) = \frac{\hbar^2 k^2}{2m}$ , there's a corresponding Fermi Wavenumber  $k_F$

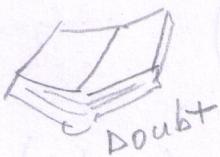
$$N = \sum_{|\vec{k}| \leq k_F} (2s+1) = gV \int_{\substack{K < k_F \\ |\vec{k}| \leq k_F}} \frac{d^3 \vec{k}}{(2\pi)^3} = g \frac{V}{6\pi^2} k_F^3$$

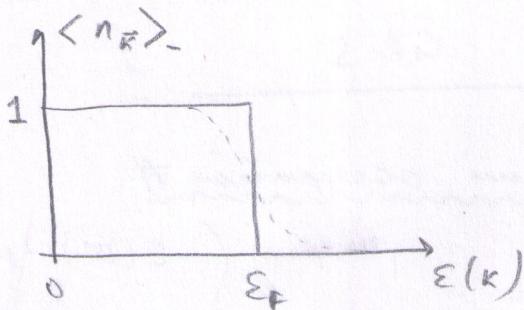
$$\left[ \frac{gV}{2(2\pi)^3} \int_{k^2}^{k_F^2} k^2 dk \right]$$

$$= \frac{gV}{2\pi^2} \cdot \frac{k_F^3}{3}$$

In terms of # density  $n = \frac{N}{V}$ ,

$$k_F = \left( \frac{6\pi^2 n}{g} \right)^{1/3} \Rightarrow E_F(n) = \frac{\hbar^2 k_F^2}{2m} = \frac{\hbar^2}{2m} \left( \frac{6\pi^2 n}{g} \right)^{2/3}$$





To see how the fermi sea is modified at low temperatures, we need to know the behaviour of  $f_m^-(\Xi)$  for large  $\Xi$ , which is

$$f_m^-(\Xi) \stackrel{\text{claim}}{=} \frac{1}{m!} \int_0^\infty dx x^m \frac{d}{dx} \left( \frac{-1}{\Xi^{-1} e^x + 1} \right)$$

$$\begin{aligned} f_m^n(\Xi) &= \frac{1}{(m-1)!} \int_0^\infty \frac{dx}{\Xi^{-1} e^x - n} x^{m-1} \\ &= \frac{1}{(m-1)!} \left[ \left[ \frac{1}{\Xi^{-1} e^x - 1} - \frac{x^m}{m} \right]_0^\infty - \int_0^\infty \frac{d}{dx} \left( \frac{1}{\Xi^{-1} e^x - 1} \right) \frac{x^m}{m} dx \right] \end{aligned}$$

□

Since the derivative in the above equation is sharply peaked, we can expand around this peak.

Set  $x = \ln \Xi + t$  &  $t$  ranges from  $-\infty$  to  $\infty$ .

$$f_m^-(\Xi) \approx \frac{1}{m!} \int_{-\infty}^\infty dt (\ln \Xi + t)^m \frac{d}{dt} \left( \frac{-1}{e^t + 1} \right)$$

$$\begin{aligned} &\left[ \frac{e^x}{\infty} = \frac{e^t \cdot \Xi}{e^t + 1} \right] \text{ on the generalized binomial expansion} \\ &= \frac{1}{m!} \int_{-\infty}^\infty dt \sum_{\alpha=0}^m \binom{m}{\alpha} t^\alpha (\ln \Xi)^{m-\alpha} \frac{d}{dt} \left( \frac{-1}{e^t + 1} \right) \\ &= \frac{(\ln \Xi)^m}{m!} \sum_{\alpha=0}^\infty \frac{m!}{\alpha! (m-\alpha)!} (\ln \Xi)^{-\alpha} \int_{-\infty}^\infty dt t^\alpha \frac{d}{dt} \left( \frac{-1}{e^t + 1} \right) \end{aligned}$$

$$\int_a^b uv \, dx = [uv]_a^b - \int_a^b u' (v \, dx) \, dx$$

When  $[uv]_a^b = 0$ ,

$$\int_a^b uv \, dx = - \int_a^b u' (v \, dx) \, dx$$

$$\Gamma \int_{-\infty}^{\infty} dt \quad \frac{d}{dt} \left( \frac{-1}{e^{t+1}} \right) = t^\alpha$$

$$\Gamma u = \frac{-1}{e^{t+1}}$$

$$v = xt^{\alpha-1}$$

$$= \int_{-\infty}^{\infty} dt \frac{xt^{\alpha-1}}{e^{t+1}}$$

Using the (anti-)symmetry of the integral under  $t \mapsto -t$ , & undoing integration by parts, we have

$$\frac{1}{\alpha!} \int_{-\infty}^{\infty} dt \quad t^\alpha \frac{d}{dt} \left( \frac{-1}{e^{t+1}} \right) = \begin{cases} 0 & \text{for odd } \alpha \\ \frac{2}{(\alpha-1)!} \int_0^{\infty} dt \frac{t^{\alpha-1}}{e^{t+1}} & \text{for even } \alpha \end{cases} = 2f_\alpha(1)$$

Plugging this in we get

$$\lim_{z \rightarrow \infty} f_m(z) = \frac{(\ln z)^m}{m!} \sum_{\substack{\alpha=0 \\ (\text{even})}}^m 2f_\alpha(1) \frac{m!}{(m-\alpha)!} (\ln z)^{-\alpha}$$

looking up integrals (tabulated)

$$= \frac{(\ln z)^m}{m!} \left[ 1 + \frac{\pi^2}{6} \frac{m(m-1)}{(\ln z)^2} + \frac{7\pi^4}{360} \frac{m(m-1)(m-2)(m-3)}{(\ln z)^4} + \dots \right]$$

Now we have in the "degenerate et",

$$\frac{n\lambda^3}{g} = f_{3/2}(z) = \frac{(\ln z)^{3/2}}{(3/2)!} \left[ 1 + \frac{\pi^2}{6} \cdot \frac{3}{2} \cdot \frac{1}{2} (\ln z)^{-2} + \dots \right] \gg 1$$

The lowest order result produces the eq<sup>7</sup> for Fermi energy.

$$\lim_{T \rightarrow 0} \ln \Xi = \left( \frac{3}{2}! \frac{n \lambda^3}{g} \right)^{2/3} = \left( \frac{3}{4\sqrt{\pi}} x^3 n \frac{1}{g} \right)^{2/3}$$

Using  $\lambda = \frac{h}{\sqrt{2\pi m k_B T}}$

$$= \frac{\beta \lambda^2}{2m} \left( \frac{6\pi^2 n}{g} \right)^{2/3} = \beta \epsilon_F$$

$$\begin{aligned} &= \left( \frac{3}{4\sqrt{\pi}} \frac{(2\pi\lambda)^3}{(2\pi m k_B T)^{3/2}} \frac{n}{g} \right)^{2/3} \\ &= \frac{\beta \lambda^2}{2m} \left( \frac{3}{4\sqrt{\pi}} \frac{2^2 (\lambda)^{-3/2} n}{(\lambda)^{3/2}} \frac{1}{g} \right)^{2/3} \\ &= \frac{\beta \lambda^2}{2m} \left( \frac{6\pi^2 n}{g} \right)^{2/3} \end{aligned}$$

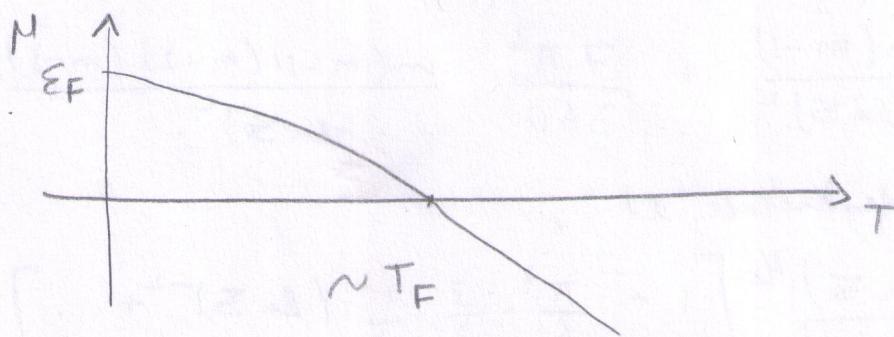
□

Inserting the zero-temp limit, we have the first order

$$\begin{aligned} \ln \Xi &= \beta \epsilon_F \left[ 1 + \frac{\pi^2}{8} \left( \frac{k_B T}{\epsilon_F} \right)^2 + \dots \right]^{-2/3} \\ &\approx \beta \epsilon_F \left[ 1 - \frac{\pi^2}{12} \left( \frac{k_B T}{\epsilon_F} \right)^2 + \dots \right] \end{aligned}$$

$$\begin{aligned} \text{Now since } \Xi &= e^{\beta M} \Rightarrow \ln \Xi = \beta M \\ &\Rightarrow k_B T \ln \Xi = M \end{aligned}$$

We thus have (with  $T_F = \frac{\epsilon_F}{k_B}$ )



At high temperatures  $\ln \Xi$  is -ve.

Doubt  
Why/How?

Similarly, the low temp expansion for the pressure is

$$\beta P = \frac{g}{\lambda^3} + \frac{5}{2} (\approx)$$

$$= \frac{g}{\lambda^3} \frac{(\ln \Sigma)^{5/2}}{(5/2)!} \left[ 1 + \frac{\pi^2}{8} \frac{5}{2} \frac{3}{2} (\ln \Sigma)^{-2} + \dots \right]$$

$$= \frac{g}{\lambda^3} \frac{8}{15} \frac{(\beta \epsilon_F)^{5/2}}{\sqrt{\pi}} \left[ 1 - \frac{5}{2} \frac{\pi^2}{12} \left( \frac{k_B T}{\epsilon_F} \right)^2 + \dots \right] \cdot \\ \left[ 1 + \frac{5\pi^2}{8} \left( \frac{k_B T}{\epsilon_F} \right)^2 + \dots \right]$$

using these

$$\Gamma_{\frac{5}{2}}! = \frac{5}{2} \cdot \frac{3}{2} \cdot \left( \frac{1}{2} \right)!$$

$$= \frac{15}{4} \frac{\sqrt{\pi}}{2} = \frac{15}{8} \sqrt{\pi}$$

$$\Gamma_{\ln \Sigma} = \beta \epsilon_F \left[ 1 - \frac{\pi^2}{12} \left( \frac{k_B T}{\epsilon_F} \right)^2 + \dots \right]$$

$$(\ln \Sigma)^{5/2} = \beta \epsilon_F \left[ 1 - \frac{5}{2} \frac{\pi^2}{12} \left( \frac{k_B T}{\epsilon_F} \right)^2 + \dots \right]$$

$$\text{Next} \quad \Gamma = \beta \frac{g}{(2\pi\hbar)^3} \frac{(2\pi k_B T m)^{3/2}}{(4\pi)^{3/2}} \frac{(\beta \epsilon_F)^{5/2}}{\sqrt{\pi}} \cdot \frac{8}{15} \cdot \left[ 1 + \underbrace{\left( -\frac{5}{2} \frac{\pi^2}{12} + \frac{5\pi^2}{8} \right)}_{\frac{5}{12}\pi^2} \left( \frac{k_B T}{\epsilon_F} \right)^2 + \dots \right]$$

$$\ln \beta \frac{\pi^2 \hbar^{3/2} \left( g \left( \frac{2m}{\hbar} \right)^{3/2} \right) \epsilon_F^{5/2}}{\sqrt{\pi} \cdot (2\pi)^3 (6\hbar^3 \pi^2)} \Gamma \epsilon_F = \frac{\hbar^2}{2m} \left( \frac{6\pi^2 n}{g} \right)^{2/3}$$

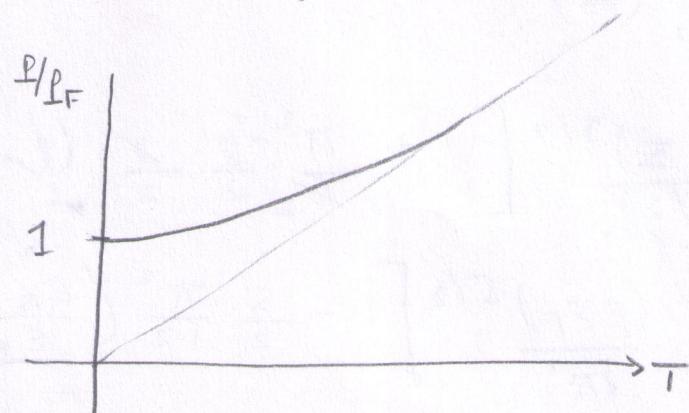
$$= n \beta \frac{\epsilon_F^{5/2}}{\epsilon_F^{3/2}} \cdot \frac{2}{5} \quad \left| \epsilon_F^{3/2} = \frac{\hbar^3}{(2m)^{3/2}} \frac{6\pi^2 n}{g} \right.$$

$$= \beta \left( \frac{2}{5} n \epsilon_F \right) \quad \left| \frac{\pi^{3/2} - \gamma_2^{-3}}{\pi \gamma_2 \hbar^3} = \pi^{1/3} \right.$$

$$= \beta P_F \left[ 1 + \frac{5}{12} \pi^2 \left( \frac{k_B T}{\epsilon_F} \right)^2 + \dots \right]$$

where  $P_F = \frac{2}{5} n \epsilon_F$  is the Fermi Pressure

Unlike its classical counterpart, the Fermi gas at zero temperature has finite pressure & internal energy.



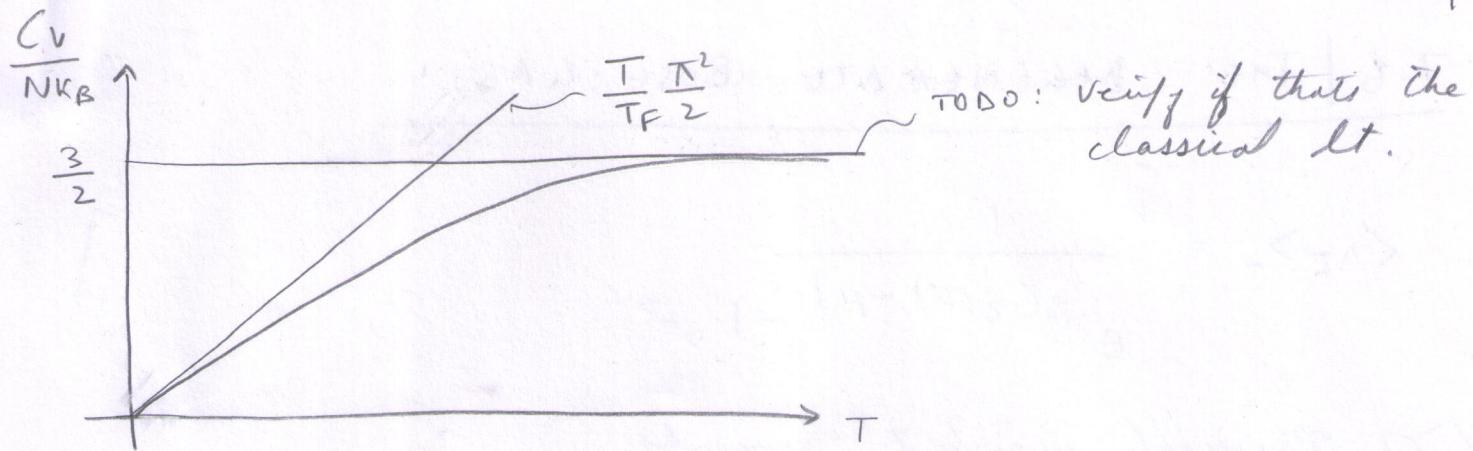
As has already been derived,

$$\begin{aligned} \varepsilon = \frac{E}{V} &= \frac{3}{2} \frac{P}{V} = \frac{3}{2} \cdot \frac{P_F}{V} \left[ 1 + \frac{5}{12} \pi^2 \left( \frac{k_B T}{\varepsilon_F} \right)^2 + \dots \right] \\ &\quad \left. \vphantom{\left[ 1 + \frac{5}{12} \pi^2 \left( \frac{k_B T}{\varepsilon_F} \right)^2 + \dots \right]} \right\} \frac{2}{5} n \varepsilon_F \\ &= \frac{3}{5} n k_B \left( \frac{\varepsilon_F}{k_B} \right) \left[ 1 + \frac{5}{12} \pi^2 \left( \frac{k_B T}{\varepsilon_F} \right)^2 + \dots \right] \\ &= \frac{3}{5} n k_B T_F \left[ 1 + \frac{5}{12} \pi^2 \left( \frac{k_B T}{\varepsilon_F} \right)^2 + \dots \right] \end{aligned}$$

where  $T_F = \frac{\varepsilon_F}{k_B}$

thus,  $C_V = \frac{dE}{dT} = \frac{\pi^2}{2} N k_B \left( \frac{T}{T_F} \right) + O\left(\frac{T}{T_F}\right)^3$

$$\begin{aligned} E &= \frac{3}{5} V n k_B T_F \left[ 1 + \frac{5}{12} \pi^2 \left( \frac{k_B T}{\varepsilon_F} \right)^2 + \dots \right] \\ \frac{dE}{dT} &= \frac{3}{5} V n k_B T_F \left[ \frac{5 \pi^2 k_B^2}{12 \cdot 4 \varepsilon_F^2} \cdot \frac{2T}{T_F} + \dots \right] \\ &= \frac{\pi^2}{T_F} N k_B T \end{aligned}$$



$$\left[ \frac{C_V}{Nk_B} = \frac{\pi^2}{2} \left( \frac{T}{T_F} \right) + \mathcal{O}\left(\frac{T}{T_F}\right)^3 \right]$$

There are important physical remarks



TODO: complete this

## § 7.6 | THE DEGENERATE BOSE GAS

$$\langle n_{\vec{k}} \rangle_+ = \frac{1}{e^{\beta(\varepsilon(\vec{k}) - \mu)} - 1}$$

 Physically important remarks

→  $\langle n_{\vec{k}} \rangle_+$  must be positive.

$$\text{Thus } \varepsilon(\vec{k}) - \mu \geq 0 \quad + \overline{k}$$

$$\Rightarrow \mu \leq \min[\varepsilon(\vec{k})]_{\vec{k}} \quad [\varepsilon(\vec{k}) = \frac{\hbar^2 k^2}{2m}]$$

Claim / IDK Y

ASK → At high temperatures (classical limit),

↳  $N$  is large &  $\rightarrow \infty$

↳ increases towards 0 like  $k_B T \ln\left(\frac{n\lambda^3}{g}\right)$   
as the temperature is dropped.

→ In the degenerate limit,  $\mu$  approaches its limiting value of zero.

To see that, we look at  $f_m^+(\xi)$

→ This increases monotonically with  $\xi$  in the interval  $0 \leq \xi \leq 1$ .

DOUBT  
NOT SURE IF  
THIS IS VAIN

$$f_m^+(\xi) = \sum_{\alpha=1}^{\infty} \frac{\xi^\alpha}{\alpha^m}$$

DOUBT

→ The integrand attains the maximum (is it global) at  $\xi = 1$ , given by

$$\zeta_m = f_m^+(1) = \frac{1}{(m-1)!} \int_0^\infty \frac{dx}{e^x - 1} x^{m-1}$$

$$\left[ \lim_{x \rightarrow 0} \frac{x^{m-1}}{e^x - 1} \right] = \lim_{x \rightarrow 0} \frac{(m-1)x^{m-2}}{e^x} \sim x^{m-2}$$

→ When  $x \rightarrow 0$ , the integrand behaves like  $\int dx x^{m-2}$

↳ Thus,  $\zeta_m$  is finite for  $m > 1$ ,

$$\left[ \int dx x^{m-2} \sim x^{m-1} \right]$$

$$\begin{aligned} \left[ m-1 > 0 \rightarrow m > 1 \rightarrow \text{finite} \right] \\ \left[ m \leq 1 \rightarrow \text{infinite} \right] \end{aligned}$$

& infinite else.

→ A useful recursive property (for  $m > 1$ )

$$\frac{d}{d\zeta} f_m^+(\zeta) = \int_0^\infty dx \frac{x^{m-1}}{(m-1)!} \frac{d}{d\zeta} \left( \frac{1}{\zeta^{-1} e^x - 1} \right)$$

$$\left[ f_m^+(\zeta) = \frac{1}{(m-1)!} \int_0^\infty \frac{dx}{\zeta^{-1} e^x - 1} x^{m-1} \right]$$

$$\begin{aligned} \left[ \text{we have } \frac{d}{d\zeta} f(\zeta^{-1} e^x) = f \cdot \frac{-1}{\zeta^2} \cdot e^x \frac{df}{d\zeta}(\zeta^{-1} e^x) \right] \\ = -\frac{1}{\zeta} \frac{df}{dx}(\zeta^{-1} e^x) \end{aligned}$$

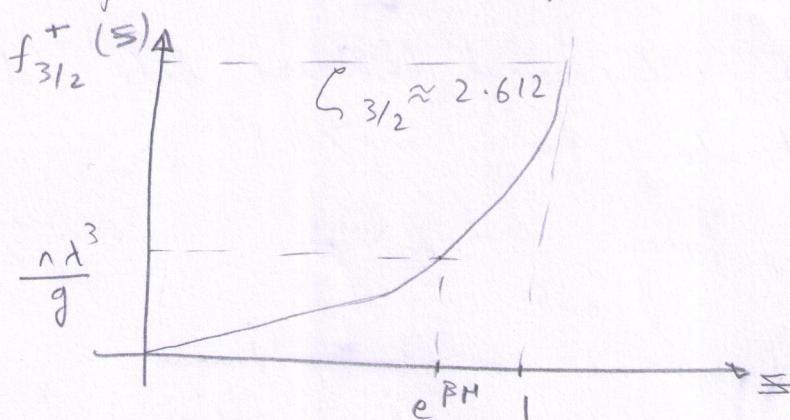
$$= \int_0^\infty dx \frac{x^{m-1}}{(m-1)!} \left( -\frac{1}{\zeta} \right) \frac{d}{dx} \left( \frac{1}{\zeta^{-1} e^x - 1} \right)$$

$$\left[ \text{by parts, we get } \int_a^b uv dx = [uv]_a^b - \int_a^b (u' v) dx \right]$$

$$\left[ = -\frac{1}{\zeta} \left[ \frac{x^{m-1}}{(m-1)!} \cdot \frac{1}{\zeta^{-1} e^x - 1} \right]_0^\infty - \left( -\frac{1}{\zeta} \right)_0^\infty \int_0^\infty \frac{x^{m-2}}{(m-2)!} \cdot \frac{1}{\zeta^{-1} e^x - 1} dx \right]$$

$$= \frac{1}{\Xi} \int_0^\infty dx \frac{x^{m-2}}{(m-2)!} \frac{1}{\Xi^{-1} e^x - 1} = \frac{1}{\Xi} f_{m-1}^+(\Xi)$$

↳ Hence, a sufficiently high derivative of  $f_m^+(\Xi)$  will be divergent at  $\Xi = 1$  for all  $m$ .  
 (Justify the following graph!)



$$\text{using } n^+ = \frac{g}{\lambda^3} f_{3/2}^+(\Xi)$$

where  $\Xi = e^{\beta\mu}$

→ The density of excited states for the non-relativistic bose gas (in 3D) is thus bounded by

$$n_x = \frac{g}{\lambda^3} f_{3/2}^+(\Xi) \leq n^* = \frac{g}{\lambda^3} \zeta_{3/2} \approx 2.612$$

↳ At sufficiently high temperatures, s.t.

$$\frac{n\lambda^3}{g} = \frac{n}{g} \left( \frac{\hbar}{\sqrt{2\pi m k_B T}} \right)^3 \leq \zeta_{3/2} \approx 2.612 \dots$$

The bound has no effect.

↳  $n_x = n$

↳ At lowering the temperature, the limit density of excited states is

$$\left[ \left( \frac{n}{g} \right)^{2/3} \frac{(\hbar)^2}{2\pi m k_B T_c} \right]^{2/3} = (\zeta_{3/2})^{2/3}$$

$$\frac{\hbar^2}{2\pi m k_B T} \left( \frac{n}{g} \right)^{2/3} = T_c(n)$$

↳ For  $T \leq T_c$ ,

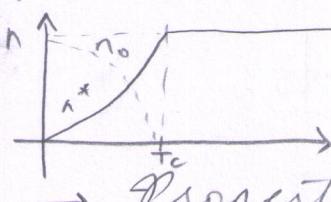
(a)  $\Rightarrow \mu \rightarrow 0$  stuck to ( $\mu = 0$ ) (can't get smaller for maintaining purity etc.)

(b)  $n^* = \frac{g L_{3/2}}{\lambda^3} \propto T^{3/2}$  the density of excited states depends only on temperature.

(c)  $\sim n^*$  when  $T \leq T_c$

DOUBT |  $\Rightarrow n_0 = n - n^*$  occupy the lowest  
why? & wouldn't the energy state with  $k=0$   
vanish disappear? This phenomenon of macroscopic occupation  
for  $k=0$ : of a single one-particle state, is called

Bose-Einstein condensate.



Properties (Unusual) of the condensate

↳ gas pressure for  $T < T_c$

DOUBT  
↓

$$\beta P = \frac{g}{\lambda^3} f_{5/2}^+(1) = \frac{g}{\lambda^3} L_{5/2} \approx 1.341 \frac{g}{\lambda^3} \propto T^{3/2}$$

This is independent of density!

It's excited because only the excited  $n^*$  fraction have non-zero momentum.

↳ <missing, wasn't done in class>

→ Latent Heat Calculation <missing, not done in class>

$\rightarrow C_V$

$$\varepsilon = \frac{3}{2} P \Rightarrow E = \frac{3}{2} PV$$

$$= \frac{3}{2} V \frac{g}{\lambda^3} k_B T f_{5/2}^+(\xi) \propto T^{5/2} f_{5/2}^+(\xi)$$

Now using  $\frac{d}{d\xi} f_m^+(\xi) = \frac{1}{\Xi} f_{m-1}^+(\xi)$  we have

$$C_{V,N} = \left. \frac{dE}{dT} \right|_{V,N} = \frac{3}{2} V \frac{g}{\lambda^3} k_B T \left[ \frac{5}{2T} f_{5/2}^+(\xi) + \frac{1}{\Xi} f_{1/2}^+(\xi) \left. \frac{d\xi}{dT} \right|_{V,N} \right]$$

$$\boxed{\Gamma} = \frac{3}{2} V g k_B \left[ \frac{d}{dT} \left( \frac{T}{\lambda^3} \right) f_{5/2}^+(\xi) + \frac{T}{\lambda^3} \frac{d}{dT} f_{3/2}^+(\xi) \right]$$

$$\begin{aligned} & \boxed{\frac{d}{dT} \left( \frac{T}{\lambda^3} (2\pi m k_B T)^{3/2} \right)} \cdot \boxed{\frac{d}{d\xi} f_{5/2}^+(\xi) \cdot \frac{d\xi}{dT}} \\ &= \frac{5}{2} \frac{(2\pi m k_B)^{3/2}}{\lambda^3} T^{3/2-1} \quad \boxed{= \frac{1}{\Xi} f_{3/2}^+(\xi) \cdot \left. \frac{d\xi}{dT} \right|_{V,N}} \\ &= \frac{5}{2T} \left( \frac{1}{\lambda^3} \right) \end{aligned}$$

$\square$

$\left. \frac{d\xi}{dT} \right|_{V,N}$  is found by fixing the # particles, viz.

$$\boxed{\left. \frac{dN}{dT} \right|_V = 0 = \frac{g}{\lambda^3} V \left[ \frac{3}{2T} f_{3/2}^+(\xi) + \frac{1}{\Xi} f_{1/2}^+(\xi) \left. \frac{d\xi}{dT} \right|_{V,N} \right]}$$

$$\boxed{N = \frac{g}{\lambda^3} V f_{3/2}^+(\xi)}$$

$$\boxed{\left. \frac{dN}{dT} \right|_V = gV \left[ \frac{d}{dT} \left( \frac{1}{\lambda^3} \right) f_{3/2}^+(\xi) + \frac{1}{\lambda^3} df_{3/2}^+(\xi) \right]}$$

$$\boxed{= gV \left[ \frac{3}{2T} \left( \frac{1}{\lambda^3} \right) f_{3/2}^+(\xi) + \frac{1}{\lambda^3} \frac{1}{\Xi} f_{1/2}^+(\xi) \left. \frac{d\xi}{dT} \right|_{V,N} \right]}$$

$\square$

Substitution of  $\zeta$

$$-\frac{3}{2} \frac{f_{3/2}^+(\xi)}{f_{1/2}^+(\xi)} = \frac{T}{\pi} \left. \frac{d\xi}{dT} \right|_{V,N}$$

yields

$$\frac{C_V}{Nk_B} = \frac{3}{2} \frac{g}{\lambda^3} \left[ \frac{5}{2} f_{5/2}^+(\xi) - \frac{3}{2} \frac{f_{3/2}^+(\xi)^2}{f_{1/2}^+(\xi)} \right]$$

<TODO: Expand in powers of  $\xi$ >

This is greater than the classical value at high temp.

$$\frac{C_V}{Nk_B} = \frac{3}{2} \left[ 1 + n \frac{\lambda^3}{2f_{1/2}} + \dots \right]$$

At low temperatures,  $z=1$  &  
(10K low)

$$\frac{C_V}{Nk_B} = \frac{15}{4} \frac{g}{n\lambda^3} \zeta_{5/2} = \frac{15}{4} \frac{\zeta_{5/2}}{\zeta_{3/2}} \left( \frac{T}{T_c} \right)^{3/2}$$



<discussion ASK>