

# Lecture 9 | Gravitational Redshift + EM

01 October 2017 01:00 PM

## Gravitational Redshift

$$\vec{x}_{(1)}^0$$

$$\vec{x}_{(1)}^0 + c$$

$$\vec{x}_{(1)}^0$$

$$\vec{x}_{(1)}^0 + c$$

$$\frac{1}{\vec{x}_1}$$

$$\frac{1}{\vec{x}_2}$$

consider: a clock sitting on the point  $\vec{x}_1$  in space.

$$\rightarrow \frac{dx^i}{dt} = 0$$

Assume: Time independent metric  $g_{\mu\nu} = g_{\mu\nu}(\vec{x})$   
(I think we're taking the non-relativistic limit  $\rightarrow$  weak field approx.)

Recall: For a physical particle,  $\frac{ds^2}{du^2} < 0$  & since  $u$  is an affine parameter  $\frac{ds}{du} = \text{const}$

This means I can always rescale to get  $\frac{ds^2}{du^2} = -1$  (for time like)

$\rightarrow du^2 = -ds^2 = dz^2$ . So I can write

$$-\frac{ds^2}{du^2} = \frac{dz^2}{du^2} = -g^{\mu\nu} \frac{dx^\mu}{du} \frac{dx^\nu}{du} = 1$$

(In the weak field limit, one ignores  $\frac{dx^i}{du}$  compared to  $\frac{dx^0}{du}$ . That's what I think is being done here but I need to check this)

$$-g^{00} \left( \frac{dx^0}{du} \right)^2 = 1 \Rightarrow \Delta x^0 = \Delta \vec{x}_1 (-g^{00}(\vec{x}_{(1)}))^{-1/2}$$

$\approx \frac{\text{not speed of light, const}}{g}$

Story: The clock emits two signals, at  $\vec{x}_{(1)}^0$  &  $\vec{x}_{(1)}^0 + dx^0$ . Say it took 'ct length of time' for the first signal to reach point  $\vec{x}_2$ . Since  $g^{\mu\nu}$  is time independent, the second signal will also take as long (see the figure). The interval between the two signals at  $\vec{x}_2$  will again be  $dx^0$ .

I now want to find the proper time b/w two clock ticks, as seen by an observer sitting on the point  $\vec{x}_{(2)}$ .  
Why bother? Because the clock of observer 2 is ticking according to his proper time.

Calc:  $\Delta x^0 = \Delta \vec{x}_{(1)} (-g^{00}(\vec{x}_{(1)}))^{-1/2}$  I wrote the eq. for a clock sitting at  $\vec{x}_{(1)}$  observing the time difference of the signals that reached  $\vec{x}_{(2)}$ .

$$\Rightarrow \Delta \vec{x}_{(1)} g^{00}(\vec{x}_{(1)})^{-1/2} = \Delta \vec{x}_{(2)} g^{00}(\vec{x}_{(2)})^{-1/2}$$

$$\Rightarrow \frac{\Delta \vec{x}_{(1)}}{\Delta \vec{x}_{(2)}} = \left[ \frac{g^{00}(\vec{x}_{(1)})}{g^{00}(\vec{x}_{(2)})} \right]^{1/2}$$

It seems that sis didn't actually make the weak field approximation but he does it now.

Weak Field:  $g^{00} = \eta^{00} + h^{00}$  where  $h^{00} = -\frac{2\phi}{c^2}$  &  $\eta^{00} = -1$

$$\Rightarrow \frac{\Delta \vec{x}_{(1)}}{\Delta \vec{x}_{(2)}} = \left[ \frac{1 + \frac{2\phi(\vec{x}_{(1)})}{c^2}}{1 + \frac{2\phi(\vec{x}_{(2)})}{c^2}} \right]^{1/2} \approx 1 + \frac{1}{c^2} [\phi(\vec{x}_{(1)}) - \phi(\vec{x}_{(2)})]$$

$$\text{If } \phi(\vec{x}_{(1)}) \geq \phi(\vec{x}_{(2)}), \text{ then } \Delta \vec{x}_{(1)} \geq \Delta \vec{x}_{(2)}.$$

The clock at  $\vec{x}_{(1)}$  will be slower than the one at  $\vec{x}_{(2)}$ .

I am, however, still a little confused about why we're using 'proper time'.  
What happens if I take the clock at  $\vec{x}_{(1)}$  & bring it to  $\vec{x}_{(2)}$ ?  
Shouldn't their proper times be the same?  
 $\rightarrow$  clock already at  $\vec{x}_{(1)}$  & the one I got from  $\vec{x}_{(2)}$ .

# Forces + Electromagnetism in GR introduced (cont.)

22 October 2017 06:19 PM

Recall: A free particle moves along a geodesic  $\frac{d^2 x^\mu}{d\tau^2} + \Gamma^\mu_{\nu\rho} \frac{dx^\nu}{d\tau} \frac{dx^\rho}{d\tau} = 0$  where  $g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} = -1$

Question: If there's a "force" how do we evaluate the LHS?

Idea: We can use the principle of equivalence. While this is not "sacred" it is correct to a very good approximation.

: go to a frame where  $\Gamma^\mu_{\nu\rho} = 0$  so that  $g^{\mu\nu} = \eta^{\mu\nu}$  by the equiv. principle.

NB: In this frame, the eq<sup>n</sup> of motion becomes  $\frac{d^2 x^\mu}{d\tau^2} = f^\mu$  force in the absence of gravity

: in local inertial frame, the principle of equivalence says that the eq<sup>n</sup> of motion should look exactly as though there's no gravity

More Bubble: Consider having a rocket you're tested in the absence of gravity & know  $f$ , the acceleration of the rocket. The principle of equivalence tells us that the eq<sup>n</sup> will be the same in this special local inertial frame.

NB: It isn't convenient to keep changing frames so

consider: A general coordinate system  $x$ , where

(TODO: This I couldn't figure; need to derive the parallel transport eq<sup>n</sup>/geodesic eq<sup>n</sup>)  
(Turns out it was an exercise there also which I remember doing)

$$\frac{d^2 x^\mu}{d\tau^2} = \partial_\nu x^\mu \left( \frac{d^2 x^\nu}{d\tau^2} + \Gamma^\nu_{\rho\sigma} \frac{dx^\rho}{d\tau} \frac{dx^\sigma}{d\tau} \right) \Rightarrow \underbrace{\left( \partial'_\mu x^\alpha \right)}_{\delta_\mu^\alpha} \left( \frac{d^2 x^\alpha}{d\tau^2} + \Gamma^\alpha_{\rho\sigma} \frac{dx^\rho}{d\tau} \frac{dx^\sigma}{d\tau} \right) = f^\mu (\partial'_\mu x^\alpha)$$

$$\Rightarrow \frac{d^2 x^\alpha}{d\tau^2} + \Gamma^\alpha_{\rho\sigma} \frac{dx^\rho}{d\tau} \frac{dx^\sigma}{d\tau} = \underbrace{\partial'_\mu x^\alpha f^\mu}_{f^\alpha} =: f^\alpha$$

NB: We didn't assume that  $f$  is a vector. We just determined that  $f'$  can be computed

Def<sup>n</sup>:

consider:  $x''$  is another coordinate system.

$$\text{Eq<sup>n</sup> in the } x'' \text{ coordinate will be } \frac{d^2 x''^\alpha}{d\tau^2} + \Gamma''^\alpha_{\rho\sigma} \frac{dx''^\rho}{d\tau} \frac{dx''^\sigma}{d\tau} = f''^\alpha = \partial'_\mu x''^\alpha f'^\mu$$

Ex:  $f''^\alpha = \partial'_\mu x''^\alpha f'^\mu$  (easy exercise) essentially saying that  $f$  indeed transforms like a vector.

consider: Force due to an external field, e.g. Electromagnetic force.

Idea: Same as earlier, use principle of equivalence;

Recall:  $m \frac{d^2 x^\mu}{d\tau^2} = q \eta^{\mu\nu} F_{\nu\rho} \frac{dx^\rho}{d\tau}$ ;  $c=1$  unit where  $F_{\nu\rho} = \partial_\nu A_\rho - \partial_\rho A_\nu$ ;  $A_\nu = \{A_0, A_1, A_2, A_3\}$   
↳ Electrostatic Potential Vector potential

$$E_i = -F_{0i} = \partial_i A_0 - \partial_0 A_i = -\nabla\phi - \frac{\partial A}{\partial t}$$

$$B_i = \frac{1}{2} \sum_{j,k} \epsilon_{ijk} F_{jk} \text{ which componentwise is } \left. \begin{aligned} B_1 &= \partial_2 A_3 - \partial_3 A_2 \\ B_2 &= \partial_3 A_1 - \partial_1 A_3 \\ B_3 &= \partial_1 A_2 - \partial_2 A_1 \end{aligned} \right\} \text{ essentially is } B = \nabla \times A$$

consider:  $\mu=i$  case,  $\frac{d}{d\tau} \left( m \frac{dx^i}{d\tau} \right) = q \eta^{ij} \left( F_{j\rho} \frac{dx^\rho}{d\tau} \right) = q \left( F_{i0} \frac{dx^0}{d\tau} + F_{ik} \frac{dx^k}{d\tau} \right)$  using  $F_{i0} = -F_{0i}$  & justify this properly.  
 $= q \left( E_i \frac{dx^0}{d\tau} + \epsilon_{ikl} B_l \frac{dx^k}{d\tau} \right)$

$$\frac{dp^i}{d\tau} = \frac{dp^i}{dt} \frac{dt}{d\tau} \Rightarrow \frac{dp^i}{dt} = q \left( E_i + \epsilon_{ikl} B_l \frac{dx^k}{dt} \right) \text{ NB: The } p^i \text{ has } \tau \text{ still}$$

↳  $\frac{dx^i}{d\tau} = \frac{dx^i}{dt} \frac{dt}{d\tau}$  which is the Lorentz force law.

# Electromagnetism in GR (cont.)

28 October 2017 03:01 PM

So in a local inertial frame, one can write (using the principle of equivalence)

$$m \frac{d^2 x'^\mu}{d\tau^2} = g^{\mu\nu} F'_{\nu\rho} \frac{dx'^\rho}{d\tau}$$

NB: The meaning of  $F'$  would be the same as that in flat space; use test charges to measure the  $E$  &  $B$  fields.

$$m \partial_\nu x'^\mu \left( \frac{d^2 x'^\nu}{d\tau^2} + \Gamma_{\rho\sigma}^\nu \frac{dx'^\rho}{d\tau} \frac{dx'^\sigma}{d\tau} \right) = g^{\mu\nu} \partial_\alpha x'^\mu \partial_\beta x'^\nu g^{\rho\sigma} F'_{\nu\rho} \partial_\gamma x'^\sigma \frac{dx'^\gamma}{d\tau} \} \sim \partial'_\mu x^\delta$$

$$\Rightarrow m \frac{d^2 x^\delta}{d\tau^2} + \Gamma_{\rho\sigma}^\delta \frac{dx^\rho}{d\tau} \frac{dx^\sigma}{d\tau} = g^{\delta\mu} \partial_\alpha x'^\mu \partial_\beta x'^\nu g^{\rho\sigma} F'_{\nu\rho} \partial_\gamma x'^\sigma \frac{dx'^\gamma}{d\tau} = g^{\delta\mu} \left( \partial_\beta x'^\nu \partial_\gamma x'^\rho F'_{\nu\rho} \right) \frac{dx'^\gamma}{d\tau}$$

Def<sup>n</sup>:  $F_{\beta\gamma} := \partial_\beta x'^\nu \partial_\gamma x'^\rho F'_{\nu\rho}$

$$= g^{\delta\mu} F_{\beta\gamma} \frac{dx'^\gamma}{d\tau}$$

Recall:  $F'_{\nu\rho} = \partial'_\nu A'_\rho - \partial'_\rho A'_\nu$  ; Claim: Show that (using the recall) that  $F_{\beta\gamma} = \partial_\beta A_\gamma - \partial_\gamma A_\beta$

where  $A_\beta = \partial_\beta x'^\mu A'_\mu$ .

Remark: Repeat the calc for another arbitrary  $x''$  coordinate sys. & conclude

$$A''_\beta = \partial_\beta x''^\mu A'_\mu$$

which can be combined as

$$A''_\beta = \partial_\beta x''^\alpha A_\alpha$$

so this shows that  $A_\beta$  transforms as a rank 0,1 tensor

$$\begin{aligned} \text{Proof: } F_{\beta\gamma} &= \partial_\beta x'^\nu \partial_\gamma x'^\rho (\partial'_\nu A'_\rho - \partial'_\rho A'_\nu) \\ &= \partial_\beta x'^\rho \partial_\gamma x'^\nu \partial'_\nu A'_\rho - \partial_\beta x'^\nu \partial_\gamma x'^\rho \partial'_\rho A'_\nu \\ &= \partial_\beta x'^\rho \underbrace{\frac{\partial x'^\nu}{\partial x'^\rho} \frac{\partial A'_\rho}{\partial x'^\nu}}_{\partial_\beta A'_\rho} - \partial_\beta x'^\nu \partial_\gamma A'_\nu \\ &= \left[ \partial_\beta (\partial_\gamma x'^\rho A'_\rho) \right] - \left[ \partial_\gamma (\partial_\beta x'^\nu A'_\nu) \right] \\ &= \left[ \partial_\beta \partial_\gamma x'^\rho A'_\rho \right] - \left[ \partial_\gamma \partial_\beta x'^\nu A'_\nu \right] \quad \square \end{aligned}$$

Bubble:  $F^\mu$  earlier was a force field.

This makes sense only along the trajectory of the particle. The  $A^\mu$  field however is a proper vector field defined without invoking any special trajectory.

Claim:  $F_{\beta\gamma}$  is a rank (0,2) tensor.

Proof:  $F_{\alpha\beta} = \partial_\alpha A_\beta - \partial_\beta A_\alpha$

$$F'_{\alpha'\beta'} = \partial'_{\alpha'} A'_{\beta'} - \partial'_{\beta'} A'_{\alpha'}$$

$$F'_{\alpha'\beta'} \stackrel{\text{claim}}{=} \partial'_{\alpha'} x^\mu \partial'_{\beta'} x^\nu F_{\mu\nu}$$

$$= \partial'_{\alpha'} x^\mu \partial'_{\beta'} x^\nu (\partial_\mu A_\nu - \partial_\nu A_\mu)$$

$$= \partial'_{\alpha'} x^\mu \partial_\mu A_\nu \partial'_{\beta'} x^\nu - \partial'_{\alpha'} x^\mu \partial'_{\beta'} x^\nu \partial_\nu A_\mu$$

$$= \partial'_{\beta'} x^\nu \partial'_{\alpha'} A_\nu - \partial'_{\alpha'} x^\mu \partial'_{\beta'} A_\mu$$

$$= \partial'_{\alpha'} (\partial'_{\beta'} x^\nu A_\nu) - \partial'_{\beta'} (\partial'_{\alpha'} x^\mu A_\mu)$$

$$= \partial'_{\alpha'} A'_{\beta'} - \partial'_{\beta'} A'_{\alpha'} \quad \square$$

Remark: How did they know this would work out?

Conclusion:

$$m \frac{d^2 x^\delta}{d\tau^2} + \Gamma_{\rho\sigma}^\delta \frac{dx^\rho}{d\tau} \frac{dx^\sigma}{d\tau} = g^{\delta\mu} F_{\beta\gamma} \frac{dx'^\gamma}{d\tau}$$

from the equivalence principle.

Further: To the RHS one can add

$$a g^{\delta\mu} F_{\beta\gamma} \frac{dx'^\gamma}{d\tau} F^{\beta\gamma}$$

but it is small (as argued earlier)

Point: Principle of equivalence is not sacred. Any law that respects general coordinate invariance might be present.

# Rough (disambiguation)

07 October 2017

01:56 PM

$-dc^2$  is proper time.

clock

$$\frac{1}{x_{(1)}}$$

$$\frac{1}{x_{(2)}}$$

(1) Assume Time Independent Metric  $g_{\mu\nu}(\vec{x})$ .

(2) The object is at rest in this metric  $\rightarrow \frac{dx^i}{dt} = 0$ .

$$g_{\mu\nu} \frac{dx^\mu}{du} \frac{dx^\nu}{du} = \frac{ds^2}{du^2} = -1$$

In the weak field limit,

$$g_{00}(\vec{x}) \left( \frac{dx^0}{dz} \right)^2 = -1$$

$$\frac{ds^2}{du^2} = -1$$

$$\frac{d^2}{du^2} = \frac{d^2}{dz^2}$$

$$\frac{ds^2}{du^2} = 1$$

$$dz = du$$

$$-ds^2 = c^2 dt^2 - dx^2 - dy^2 - dz^2$$

$$\frac{ds^2}{du^2} = \text{const.}$$

Let  $\Delta z$  be the period of the clock (intrinsic  $\Rightarrow$  in its rest frame).

$$[g_{00}(\vec{x}_0)] \cdot (dx^0)^2 = -dz^2$$

$$dx^0 = dz \left( -g_{00}(\vec{x}_0) \right)^{1/2}$$

$$\eta^{\mu\nu} = \text{diag}(1, +1, +1, +1)$$

$$\frac{d^2 x^\mu}{dz^2} = \frac{d}{dz} \left( \frac{dx^\mu}{dz} \right)$$

$$\frac{\partial f}{\partial x_i} dx_i = df$$

$$= \frac{d}{dz} \left( \frac{dx^\mu}{dz} \right) = \frac{d}{dz} \left( \partial_\nu x^\mu \frac{dx^\nu}{dz} \right)$$

$$= \frac{d}{dz} \left[ \partial_\nu x^\mu \right] \frac{dx^\nu}{dz} + \partial_\nu x^\mu \frac{d^2 x^\nu}{dz^2}$$

=