

(L-4: STATISTICAL MECHANICS || Kardar Notes

§ 4.1

→ $p_\mu(\mu)$ is calculated for eq. ensembles

Liouville's Thm justifies the assumption that

"all accessible microstates are equally likely in an equilibrium ensemble."

→ Stat mech (as opposed to the Kinetic Theory) doesn't explain how various systems evolve to eq.

§ 4.2

$$dQ=0 \quad \& \quad dW=0$$

MICRO-CANONICAL ENSEMBLE

$$\Rightarrow dE=0 \quad \& \quad dV=0$$

so, $M = (E, \underline{x}) \leftarrow$ Macroscopic State

$\mu \leftarrow$ Microstate

$\Omega(\mu) = E \rightarrow$ all microstates are confined to the surface specified.

Assumption: \exists no other conserved quantities, so that all points on the surface are mutually accessible.

Central Postulate:

$$P_{(E, \underline{x})}(\mu) = \frac{1}{\Omega(E, \underline{x})} \cdot \begin{cases} 1 & \text{for } \Omega(\mu) = E \\ 0 & \text{else} \end{cases}$$

Remarks

- 1) a) Points in the phase space are canonically conjugate (x, p)
b) Change by canonical transformation, $\mu \rightarrow \mu'$, leaves phase space volumes invariant.

2) (a) $\Omega(E, \underline{x})$ is the st of the const. E surface

(b) subtleties with densities non zero only at E is avoided by using

$$(i) \quad E - \Delta \leq \Omega(\mu) \leq E + \Delta$$

$$(ii) \quad \Omega' \approx (2\Delta)\Omega \quad \& \quad \Omega \text{ grows exponentially, } \Delta \sim \mathcal{O}(E^0) \text{ (or even } \mathcal{O}(E')) \quad \Omega' \sim \Omega$$

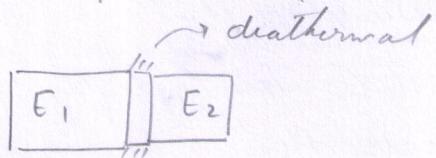
$$(c) (i) \quad S(E, \underline{x}) = k_B \ln \Omega(E, \underline{x})$$

$$(ii) \Sigma_{\text{tot}} = \prod_i \Sigma_i$$

for i independent sys.

\Rightarrow Entropy is additive - extensive

The zeroth law:



assumptions:

- 1) Interaction of the 2 systems are small

$$\mu = \mu_1 \otimes \mu_2$$

$$H(\mu_1 \otimes \mu_2) = H_1(\mu_1) + H_2(\mu_2)$$

Note: The joint sys is a micro-canonical ensemble

$$\rho(\mu_1 \otimes \mu_2) = \frac{1}{\mathcal{Z}(E)} \begin{cases} 1 & \text{for } H_1(\mu_1) + H_2(\mu_2) = E \\ 0 & \text{else} \end{cases}$$

Since E is fixed,

$$\Sigma(E) = \int dE, \Sigma_1(E_1) \Sigma_2(E-E_1) = \int dE, e^{\frac{S_1(E_1) + S_2(E-E_1)}{k_B}}$$

[We assume Σ_2 depends exponentially on E]
 Thus $S(E, x) \propto E$ & $E \propto N$ } Bad reasoning

$\Rightarrow S \propto \# \text{ particles}$. Thus Σ is exponentially large
 $(\because S \text{ is extensive})$

By saddle point; at $E_1^* \& E_2^* = E - E_1^*$, we have

$$S(E) = k_B \ln \Sigma(E) \approx k_B \ln e^{\frac{S_1(E_1^*) + S_2(E-E_1^*)}{k_B}} \quad \curvearrowright$$

$$= (k_B) \left(\frac{1}{k_B} \right) (S_1(E_1^*) + S_2(E-E_1^*))$$

E_1^* is evaluated by requiring (extremizing the exponent)

$$\frac{\partial S_1}{\partial E_1} \Big|_{x_1} = \frac{\partial S_2}{\partial E_2} \Big|_{x_2} \quad (\text{am I missing a minus sign?})$$

(No you idiot!)



Read the book for interpreting why the state is expected after eq to be E_1^* & E_2^*

From this we define

$$\left. \frac{\partial S}{\partial E} \right|_x = \frac{1}{T}$$

consistent with the earlier meaning of T .

First Law

Let's look at variations of $S(E, x)$ with x , reversibly.

$$\delta W = J \cdot S dx \text{ on the sys.}$$

Thus the internal energy changes to

$$E' = E + J \cdot S dx$$

$$\delta(S) = S(E + J \cdot S dx, x + S dx) - S(E, x)$$

$$= \left(\left. \frac{\partial S}{\partial E} \right|_x J + \left. \frac{\partial S}{\partial x} \right|_E \right) \cdot S dx$$

Since the prob depends exponentially on S , a first order change in δS would result in a spontaneous change.

$$\text{Thus } \left. \frac{\partial S}{\partial E} \right|_{x_i} J_i = - \left. \frac{\partial S}{\partial x_i} \right|_{E, x_{i+1}}$$

$$\Rightarrow -\frac{J_i}{T} = \left. \frac{\partial S}{\partial x_i} \right|_{E, x_{i+1}}$$

$$\begin{aligned} \text{Now } dS(E, x) &= \left. \frac{\partial S}{\partial x} \right|_E dx + \left. \frac{\partial S}{\partial E} \right|_x dE \\ &= -\frac{J}{T} \cdot dx + \frac{dE}{T} \end{aligned}$$

$$\Rightarrow dE = T dS + J \cdot dx$$

$$\text{Thus we define } dQ = T dS$$

Second Law



$$S_1(E_1^*, x_1) - S_2(E_2^*, x_2) > S_1(E_1, x_1) - S_2(x_2)$$

$$\Rightarrow \delta S = S_1(\epsilon_1^*) + S_2(\epsilon_2^*) - S_1(\epsilon_1) - S_2(\epsilon_2) \geq 0$$

consistent with the second law.

$$\delta S = \left(\frac{\partial S_1}{\partial E_1} \Big|_{x_1} - \frac{\partial S_2}{\partial E_2} \Big|_{x_2} \right) \delta E_1 = \left(\frac{1}{T_1} - \frac{1}{T_2} \right) \delta \epsilon_1 \geq 0$$

Shows that heat flows from hotter to colder.
(energy)

$\boxed{\square}$ stability condition

§ 4.3 Two-level System

$$\Omega(\{n_i\}) = e \sum_{i=1}^N n_i = e N,$$

$\boxed{\square}$ for def's

N : # impurity atoms n_i, e : energy normal/excited
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where N_1 is the total # of excited impurities

$$p(\{n_i\}) = \frac{1}{\Omega(e, N)} \delta e \sum_{i=1}^N n_i, e$$

$$\Omega(e, N) = {}^N C_{N_1} = \frac{N!}{N_1! (N-N_1)!}$$

$$S(e, N) = k_B \ln \Omega(e, N) = k_B \ln \frac{N!}{N_1! (N-N_1)!}$$

Stirling:

$$\begin{aligned} \Rightarrow S(e, N) &\approx k_B \left[N \ln N - N - N_1 \ln N_1 + N_1 \right. \\ &\quad \left. - (N-N_1) \ln (N-N_1) + (N-N_1) \right] \\ &= k_B \left[N \ln N - N_1 \ln N_1 - (N-N_1) \ln (N-N_1) \right] \\ &= k_B \left[(N-N_1) \ln N + N_1 \ln N - N_1 \ln N_1 - (N-N_1) \ln (N-N_1) \right] \\ &= k_B \left[N_1 \ln \frac{N}{N_1} + (N-N_1) \ln \frac{N}{N-N_1} \right] \end{aligned}$$

$$= -k_B \left[N_1 \ln \frac{N_1}{N} + (N - N_1) \ln \frac{N - N_1}{N} \right]$$

$$= -N k_B \left[\frac{N_1}{N} \ln \frac{N_1}{N} + \frac{N - N_1}{N} \ln \frac{N - N_1}{N} \right]$$

With $N_1 = E/\epsilon$

$$= -N k_B \left[\left(\frac{E}{\epsilon N} \right) \ln \left(\frac{E}{\epsilon N} \right) + \left(1 - \frac{E}{\epsilon N} \right) \ln \left(1 - \frac{E}{\epsilon N} \right) \right]$$

The eq. temp would be

$$\boxed{\frac{1}{T} = \frac{\partial S}{\partial E} \Big|_N = -\frac{k_B}{\epsilon} \ln \left(\frac{E}{N\epsilon - E} \right)}$$

Verification

$$\frac{\partial S}{\partial E} \Big|_N = -N k_B \left[\left\{ \frac{\cancel{E}}{\epsilon N} \frac{1}{\cancel{(E)}} \cdot \cancel{\frac{1}{\epsilon N}} + \left\{ \left(1 - \frac{E}{\epsilon N} \right) \frac{1}{\cancel{(1 - \frac{E}{\epsilon N})}} \cdot \cancel{\frac{1}{\epsilon N}} \right\} \right\} + \left(-\frac{1}{\epsilon N} \right) \ln \left(1 - \frac{E}{\epsilon N} \right) \right]$$

$$= -\frac{N k_B}{\epsilon N} \left[\ln \frac{\frac{E}{\epsilon N}}{1 - \frac{E}{\epsilon N}} \right] = -\frac{k_B}{\epsilon} \ln \left(\frac{E}{N\epsilon - E} \right)$$

$$\Rightarrow \frac{-t}{T k_B}$$

$$e^{\frac{-t}{T k_B}} = \frac{E}{N\epsilon - E}$$

$$\frac{N\epsilon}{1 + e^{-\frac{E}{T k_B}}} = \boxed{E = \frac{N\epsilon}{e^{E/T k_B} + 1}}$$

$$C = \frac{dE}{dT} = NK_B \left(\frac{E}{k_B T} \right)^2 e^{\frac{E}{k_B T}} \left[e^{\frac{E}{k_B T}} + 1 \right]^{-2}$$

Hint:
 Use $\frac{d(E^{-1})}{dT} = (-1) E^{-2} \frac{dE}{dT}$ to find $\frac{dE}{dT}$

 consequences & physics from the book.

Prob (Unconditional) for exciting a particular impurity is

$$p(n_1) = \sum_{\{n_2 \dots n_N\}} P(\{n_i\}) = \frac{\Omega(E - n_1, \epsilon, N-1)}{\Omega(\epsilon, N)}$$

once energy of one is specified,
 the rest can be calculated as
 shown.

$$\begin{aligned} p(n_1=0) &= \frac{\Omega(E, N-1)}{\Omega(E, N)} = \frac{(N-1)!}{N_1!(N-N_1-1)!} \cdot \frac{N_1! (N-N_1)!}{N!} \\ &= 1 - \frac{N_1}{N} \end{aligned}$$

and since $n_1 = \{1, 0\}$ are the only options,

$$p(n_1=1) = 1 - p(n_1=0) = \frac{N_1}{N}$$

putting $N_1 = E/\epsilon$ & $E = \frac{N\epsilon}{e^{\epsilon/k_B T} + 1}$

we have

$$p(n_1=0) = \frac{1}{1 + e^{-\epsilon/k_B T}}$$

$$p(n_1=1) = \frac{1}{1 + e^{\epsilon/k_B T}}$$

$$= \frac{e^{-\epsilon/k_B T}}{1 + e^{-\epsilon/k_B T}}$$

§4.4 The ideal gas

(will mostly skip all detail & write the eq's only)

$$\mu = \{ p_i, g_i \}$$

$$\mathcal{H} = \sum_{i=1}^N \left[\frac{\vec{p}_i^2}{2m} + U(\vec{g}_i) \right]$$

$$p(\mu) = \frac{1}{\Omega(E, V, N)} \cdot \begin{cases} 1 & \text{for } g_i \in \text{Box}, \sum_i \vec{p}_i^2 / 2m = E \\ 0 & \text{else} \end{cases}$$

constrained to move on the hypersurface

$$\sum_i \vec{p}_i^2 = 2mE ; \text{ sphere of radius } \delta = \sqrt{2mE}$$

$$\Delta s = \sqrt{2mE} \Delta E$$

Area of d -dim sphere is

$$A_d = S_d R^{d-1} \quad \text{where} \quad S_d \text{ is the solid angle.}$$

$$\text{where } S_d = \frac{2\pi^{d/2}}{\Gamma(d/2)}$$

Proof:

$$I_d = \int_{-\infty}^{\infty} (dx e^{-x^2})^d = \pi^{d/2}$$

$$\begin{aligned} I_d &= \int_{-\infty}^{\infty} \pi r_i dx_i e^{-x_i^2} = \int_{\text{all Volume}} dV e^{-\sum_i x_i^2} \\ &= \int_0^{\infty} dR S_d R^{d-1} e^{-R^2} = \frac{S_d}{2} \int_0^{\infty} dy y^{\frac{d}{2}-1} e^{-y} \end{aligned}$$

$$\text{(using } dV = \underbrace{S_d R^{d-1}}_{\text{Area of sphere}} dR \text{)}$$

$(4\pi r^2 \text{ in 3D})$

$$= \frac{S_d}{2} \Gamma(d/2)$$

equating I_d we get S_d as claimed.

\square notes on these
for details
of calculations

Now the vol. of phase space available

$$\Omega(\epsilon, v, N) = (v^N) \left(\underbrace{\int_0^r \int_0^p d\vec{q}_i d\vec{p}_i}_{\text{Area}} \right)^{d-1} \Delta R$$

$$= v^N \frac{2\pi^{3N/2}}{\left(\frac{3N}{2}-1\right)!} (2m\epsilon)^{\frac{(3N-1)}{2}} \Delta R$$

This can be ignored.

Neglecting terms of order 1, or $\ln E \sim \ln N$ (explained earlier) CLARIFY

$$S = k_B \ln \Omega \approx k_B \ln \left[N \ln v + \frac{3N}{2} \ln (2\pi m\epsilon) - \frac{3N \ln \frac{3N}{2}}{2} + \frac{3N}{2} \right]$$

$$= N k_B \ln \left[v \left(\frac{4\pi m\epsilon}{3N} \right)^{3/2} \right]$$

$$\text{Now, } T dS = dF + PdV - \mu dN$$

$$\frac{1}{T} = \left. \frac{\partial S}{\partial E} \right|_{N, V} = \frac{3}{2} \cdot \frac{N k_B}{E} \quad \left| \begin{array}{l} \text{in } \triangle \\ \text{Notes for next} \end{array} \right.$$

$$C_V = \frac{3N k_B}{T} \text{ is const.}$$

$$\frac{P}{T} = \left. \frac{\partial S}{\partial V} \right|_{N, E} = \frac{N k_B}{V} \Rightarrow PV = N k_B T$$

The unconditional prob. of finding a particle of momentum \vec{p}_i in the gas is $\Omega(\epsilon, v, N)$

$$\begin{aligned} p(\vec{p}_i) &= \int d^3 \vec{q}_i \underbrace{\prod_{i=2}^N d^3 q_i d^3 p_i}_{\Omega(\{\vec{q}_i, \vec{p}_i\})} \underbrace{p(\{\vec{q}_i, \vec{p}_i\})}_{\Omega(\epsilon, v, N)} \\ &= \frac{V \Omega(E - \vec{p}_i^2/2m, v, N-1)}{\Omega(\epsilon, v, N) \Delta} \end{aligned}$$

Upon using the formula for Ω , we have

$$\begin{aligned} p(\vec{p}_1) &= \frac{V^N \pi^{\frac{3(N-1)}{2}} (2mE - \vec{p}_1^2)^{\frac{(3N-4)}{2}}}{\left(\frac{3(N-1)}{2} - 1\right)!} \times \frac{\left(\frac{N}{2} - 1\right)!}{V^N \pi^{\frac{3N}{2}} (2mE)^{\frac{3N}{2}}} \\ &= \left(1 - \frac{\vec{p}_1^2}{2mE}\right)^{\frac{3N}{2}-2} \frac{1}{(2\pi mE)^{3/2}} \frac{\left(\frac{3N}{2} - 1\right)!}{\cancel{\left(\frac{3N-1}{2} - 1\right)!}} \\ &\quad \left. \begin{array}{l} \text{(from stirling's approx., you get)} \\ \left(\frac{3N}{2}\right)^{3/2} \end{array} \right] \end{aligned}$$

& in large E limit

$$p(\vec{p}_1) = \left(\frac{3N}{4\pi mE}\right)^{3/2} e^{-\frac{3N}{2} \frac{\vec{p}_1^2}{2mE}}$$

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Maxwell-Boltzmann Distribution ; with $E = 3NK_B T/2$

$$p(\vec{p}_1) = \frac{1}{(2\pi mK_B T)^{3/2}} e^{-\frac{\vec{p}_1^2}{2mK_B T}}$$

§ 4.5 MIXING ENTROPY AND THE GIBBS PARADOX

Entropy defined so far is not extensive.

$$(E, V, N) \rightarrow (\lambda E, \lambda V, \lambda N)$$

$$S \rightarrow \lambda(S + NK_B \ln \lambda)$$

→ 2 gasses at temp T , V_1 & V_2 volumes.

→ Mixing is clearly irreversible, & must be accompanied by an increase of entropy.

$$S_i = S_1 + S_2 = N_1 K_B (\ln v_1 + \sigma_1) + N_2 K_B (\ln v_2 + \sigma_2)$$

$$\text{where } \sigma_\alpha = \ln \left(\frac{4\pi e m_\alpha}{3} \frac{E_\alpha}{N_\alpha} \right)^{3/2}$$

↳ Momentum Contribution

$$\text{Since } \frac{E_\alpha}{N_\alpha} = \frac{3K_B T}{2} \text{ for monoatomic,}$$

$$\sigma_\alpha(T) = \frac{3}{2} \ln (2\pi e m_\alpha K_B T)$$

(don't know why) The temp. is unchanged since

$$\frac{3K_B T}{2} = \frac{E_1 + E_2}{N_1 + N_2} = \frac{\epsilon_1}{N_1} = \frac{\epsilon_2}{N_2} = \frac{3}{2} K_B T$$

$$\text{if } S_f = N_1 K_B \ln(v_1 + v_2) + N_2 K_B \ln(v_1 + v_2) + K_B (N_1 \sigma_1 + N_2 \sigma_2)$$

(haven't calculated this)

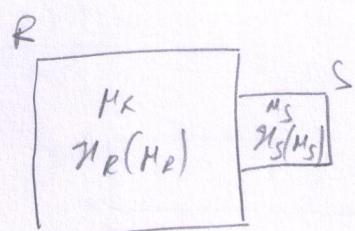
$$\text{↳ } (N_1 + N_2) K_B \ln(v_1 + v_2) + K_B (N_1 \sigma_1 + N_2 \sigma_2)$$

§ 4.6 THE CANONICAL ENSEMBLE

$$M = (\tau, \underline{x})$$

assumption:

$$E_{\text{tot}} \gg E_S$$



$$\rho(\mu_S \otimes M_R) = \frac{1}{\Omega_{S \otimes R}(E_{\text{tot}})} \cdot \begin{cases} 1 & \text{for } H_S(\mu_S) + H_R(H_R) = E_{\text{tot}} \\ 0 & \text{else} \end{cases}$$

$$\text{now } \rho(\mu_S) = \sum_{\{M_R\}} \rho(\mu_S \otimes M_R)$$

Once μ_S is specified, we have

$$\rho(\mu_S) = \frac{\Omega_R(E_{\text{tot}} - H_S(\mu_S))}{\Omega_{S \otimes R}(E_{\text{tot}})} \propto e^{\frac{1}{k_B} S_R(E_{\text{tot}} - H_S(\mu_S))}$$

Now since $H_S(\mu_S) \ll E_{\text{tot}}$, we have

$$S_R(E_{\text{tot}} - H_S(\mu_S)) \approx S_R(E_{\text{tot}}) - H_S(\mu_S) \frac{\partial S_R}{\partial E_R} \dots$$

$$= S_R(E_{\text{tot}}) - \frac{H_S(\mu_S)}{T}$$

Dropping the subscript (cause S_R disappears from the final expression)

$$\rho_{(\tau, \underline{x})}(\mu) = \frac{e^{-\beta H(\mu)}}{Z(\tau, \underline{x})}$$

$$\text{where } Z(\tau, \underline{x}) = \sum_{\{M\}} e^{-\beta H(\mu)}$$

ε of course now is a random variable.

$$\rho(\varepsilon) = \sum_{\{M\}} \rho(\mu) S(H(\mu) - \varepsilon) = \frac{e^{-\beta \varepsilon}}{Z} \sum_{\mu} \delta(H(\mu) - \varepsilon)$$

$$\Rightarrow p(\varepsilon) = \frac{s(\varepsilon)e^{-\beta\varepsilon}}{Z} = \frac{1}{Z} e^{\left[\frac{s(\varepsilon)}{k_B} - \frac{\varepsilon}{k_B T}\right]} \\ = \frac{1}{Z} e^{-\frac{F(\varepsilon)}{k_B T}}$$

with $F(\varepsilon) = \varepsilon - TS(\varepsilon)$

The dist. is peaked at ε^* that minimizes $F(\varepsilon)$

$$Z = \sum_{\{M\}} p(\mu) = \sum_{\{M\}} e^{-\beta H(\mu)} = \sum_{\varepsilon} p(\varepsilon) = \sum_{\varepsilon} e^{-\beta F(\varepsilon)} \\ \approx e^{-\beta F(\varepsilon^*)}$$

The average energy is

$$\langle H \rangle = \sum_{\{M\}} H(\mu) \frac{e^{-\beta H(\mu)}}{Z} = -\frac{1}{Z} \frac{\partial}{\partial \beta} \sum_{\{M\}} e^{-\beta H(\mu)}$$

$$\Rightarrow \langle H \rangle = -\frac{\partial \ln Z}{\partial \beta}$$

In thermodynamics,

$$E = F + TS = F - T \left. \frac{\partial F}{\partial T} \right|_x = -T^2 \left. \frac{\partial}{\partial T} \left(\frac{F}{T} \right) \right|_x = \frac{\partial (\beta F)}{\partial \beta}$$

$$= \frac{\partial (E)}{\partial (\frac{1}{T})} =$$

$$\text{so we get } E = \frac{\partial (\beta F)}{\partial \beta}$$

& we already had

so we may identify $-\ln Z = \beta F$

$$\Rightarrow \boxed{F(T, z) = -k_B T \ln Z(T, z)}$$

To find how close the average & most probable values are, we have find the variance $\langle H^2 \rangle$.

$$Z = \sum_{\mu} e^{-\beta E(\mu)}$$

$$\frac{\partial Z}{\partial \beta} = \sum_{\mu} \mu e^{-\beta E(\mu)}$$

$$+ \frac{\partial^2 Z}{\partial \beta^2} = \sum_{\mu} \mu^2 e^{-\beta E(\mu)}$$

$$\langle H \rangle_c = \frac{1}{Z} \sum_{\mu} \mu e^{-\beta E(\mu)} = - \frac{1}{Z} \frac{\partial Z}{\partial \beta} = - \frac{\partial \ln Z}{\partial \beta}$$

$$\langle H^2 \rangle_c = \langle H^2 \rangle - \langle H \rangle^2$$

$$= \frac{1}{Z} \sum_{\mu} \mu^2 e^{-\beta E(\mu)} - \frac{1}{Z^2} \left(\sum_{\mu} \mu e^{-\beta E(\mu)} \right)^2$$

$$= \frac{1}{Z} \frac{\partial^2 Z}{\partial \beta^2} - \frac{1}{Z^2} \left(\frac{\partial Z}{\partial \beta} \right)^2 = \frac{\partial^2 \ln Z}{\partial \beta^2}$$

$$= - \frac{\partial}{\partial \beta} \langle H \rangle$$

In general

$$\langle H^n \rangle_c = (-1)^n \frac{\partial^n \ln Z}{\partial \beta^n}$$

$$\langle H^2 \rangle_c = - \frac{\partial \langle H \rangle}{\partial (\gamma/k_B T)} = k_B T^2 \left. \frac{\partial \langle H \rangle}{\partial T} \right|_Z$$

$$\Rightarrow \langle H^2 \rangle_c = k_B T^2 C_x$$

◻ for details

$p(\epsilon)$ grows as $\sqrt{\langle H^2 \rangle_c} \propto N^{1/2}$, thus the relative

error $\frac{\sqrt{\langle H^2 \rangle_c}}{\langle H \rangle_c}$ goes as $\frac{1}{\sqrt{N}}$

Some detail is missing
(1 paragraph)

$$p(\varepsilon) = \frac{1}{Z} e^{-\frac{f(\varepsilon)}{k_B T}} \approx$$

§ 4.7 Canonical Examples

(1) Two-level system |

N impurities

$$M = (T, N)$$

$$H = \epsilon \sum_{i=1}^N n_i \quad \text{where } n_i = 0 \text{ or } 1$$

$$\mu = \{n_i\}$$

$$p(\{n_i\}) = \frac{1}{Z} e^{-\beta E \sum_i n_i}$$

$$\begin{aligned} Z(T, N) &= \sum_{\{n_i\}} e^{-\beta E \sum_{i=1}^N n_i} \\ &= \left(\sum_{n_1=0}^1 e^{-\beta E n_1} \right) \left(\sum_{n_2=0}^1 e^{-\beta E n_2} \right) \dots \left(\sum_{n_N=0}^1 e^{-\beta E n_N} \right) \\ &= (1 + e^{-\beta E})^N \end{aligned}$$

$$F(T, N) = -k_B T \ln Z$$

$$= -N k_B T \ln (1 + e^{-E/k_B T}) - F/T$$

$$S = -\left. \frac{\partial F}{\partial T} \right|_N = \underbrace{N k_B \ln (1 + e^{-E/k_B T})}_{N k_B \propto \left(\frac{e}{k_B T} \right)} + \frac{e^{-E/k_B T}}{1 + e^{-E/k_B T}}$$

$$E = F + TS = \frac{N E}{1 + e^{E/k_B T}}$$

$$\text{Also, } E = \frac{\partial \ln Z}{\partial \beta} = -\frac{\partial}{\partial \beta} N \ln (1 + e^{-\beta E}) = -\frac{N}{1 + e^{-\beta E}} \cdot e^{-\beta E} \cdot (-E) = \frac{N E}{1 + e^{-\beta E}}$$

$$p = \prod_i p_i \quad (\text{independent probabilities})$$

Thus the unconditional prob. of excitation is

$$p_i(n_i) = \frac{e^{-\beta E_i} (1 + e^{-\beta E})^{N-i}}{(1 + e^{-\beta E})^N}$$

6) The Ideal Gas

$$M = (\tau, v, N)$$

$$\mu = \{\vec{p}_i, \vec{q}_i\}$$

$$p(\{\vec{p}_i, \vec{q}_i\}) = \frac{1}{Z} e^{-\beta \sum_{i=1}^N \frac{p_i^2}{2m}} \cdot \begin{cases} 1 & \text{for } \{q_i\} \in \text{box} \\ 0 & \text{else.} \end{cases}$$

$$Z(\tau, v, N) = \int \frac{1}{N!} \prod_{i=1}^N \left(\frac{d^3 \vec{p}_i d^3 \vec{q}_i}{h^3} \right) e^{-\beta \sum_{i=1}^N \frac{p_i^2}{2m}}$$

$$= \frac{V^N}{N!} \frac{1}{h^{3N}} \int \prod_{i=1}^N \left(d^3 \vec{p}_i \right) \prod_{i=1}^N \left(e^{-\beta \frac{p_i^2}{2m}} \right)$$

$$= \frac{V^N}{N!} \frac{1}{h^{3N}} \int \prod_{i=1}^N \left(d^3 \vec{p}_i e^{-\beta \frac{p_i^2}{2m}} \right)$$

use $\int dx e^{-ax^2} = \sqrt{\pi/a}$
and $d^3 \vec{p}_i = dp_x dp_y dp_z$
 $\& p_i^2 = p_x^2 + p_y^2 + p_z^2$

$$\stackrel{?}{=} \frac{V^N}{N!} \left(\frac{2\pi m k_B T}{h^2} \right)^{3N/2} = \frac{1}{N!} \left(\frac{V}{\lambda(\tau)^3} \right)^N$$

$$\text{where } \lambda(\tau) = \frac{h}{\sqrt{2\pi m k_B T}}$$

Now

$$F = -k_B T \ln Z = -N k_B T \ln V + N k_B T \ln N \quad (\text{I didn't check this})$$

$$= -N k_B T \left[\ln \left(\frac{V}{N} \right) + \frac{5}{2} \ln \left(\frac{2\pi m k_B T}{h^2} \right) \right] - \frac{3N}{2} k_B T \ln \left(\frac{2\pi m k_B T}{h^2} \right)$$

Now using

$$dF = -SdT - PdV + \mu dN$$

We can calculate various thermodynamic quantities.

$$\begin{aligned} -S &= \left. \frac{\partial F}{\partial T} \right|_{V, N} = -NK_B \left[\ln \frac{Vc}{N} + \frac{3}{2} \ln \left(\frac{2\pi m k_B T}{h^2} \right) \right] \\ &\quad - NK_B T \left(\frac{3}{2} \frac{h^2}{2\pi m k_B T} \cdot \frac{2\pi m k_B T}{h^2} \right) \end{aligned}$$

any way $-S = \frac{F - E}{T}$

Thus

$$E = \frac{3}{2} N k_B T$$

For eqⁿ of state,

$$P = - \left. \frac{\partial F}{\partial V} \right|_{T, N} = \frac{N k_B T}{V} \Rightarrow PV = N k_B T$$

And $\mu = \left. \frac{\partial F}{\partial N} \right|_{T, V} = \frac{F}{N} + k_B T$

$$= \frac{E - TS}{N} + \cancel{K_B T} \frac{PV}{N}$$

$$= \frac{E - TS + PV}{N}$$

§ 4.8 The Gibbs Canonical Ensemble

$$\mu = (T, I)$$

Energy of the combined system is

$$H - J \cdot x - \beta H(\mu_s) + \beta J \cdot x$$

$$p(\mu_s, x) = \frac{e}{Z(T, N, J)}$$

$$\text{with } Z \text{ obviously} = \sum_{\mu_s, x} e^{\beta J \cdot x - \beta H(\mu_s)}$$

$$\text{Now } \langle x \rangle = k_B T \frac{\partial \ln Z}{\partial J}$$

$$\text{and } x = -\frac{\partial G}{\partial J} \quad \text{from Thermodynamics suggest}$$

$$G(N, T, J) = -k_B T \ln Z$$

$$\text{where } G = E - TS - x \cdot J$$

Also, H can be evaluated as

$$H = E - x \cdot J$$

$$-\frac{\partial \ln Z}{\partial \beta} = \langle H - J \cdot x \rangle = H$$

Heat capacities at const. T may be evaluated as

$$C_J = \frac{\partial H}{\partial T}$$

Examples

1) The Ideal Gas

isobaric ensemble

$$M = \{N, T, P\}$$

$$\mu = \{\vec{p}_i, \vec{q}_i\}$$

with vol. V

$$p(\{\vec{p}_i, \vec{q}_i\}, V) = \frac{1}{Z} e^{\left[-\beta \sum_{i=1}^N \frac{p_i^2}{2m} - \beta PV \right]} \cdot \begin{cases} 1 & \text{for } \{\vec{q}_i\} \in \\ & \text{of} \\ & \text{pressure, not momentum!} \\ 0 & \text{else} \end{cases}$$

$$\begin{aligned} Z(N, T, P) &= \int_0^\infty dV e^{-\beta PV} \underbrace{\int \frac{1}{N!} \prod_{i=1}^N \frac{d^3 \vec{q}_i d^3 \vec{p}_i}{h^3}}_{\infty} e^{-\beta \sum_{i=1}^N \frac{p_i^2}{2m}} \\ &= \int_0^\infty dV V^N e^{-\beta PV} \frac{1}{N!} \frac{1}{\lambda(T)^{3N}} \\ &= \frac{1}{N!} \frac{1}{\lambda(T)^{3N}} \int_0^\infty dV V^N e^{-\beta PV} \\ &= \frac{1}{N!} \frac{1}{\lambda(T)^{3N}} \frac{N!}{(PV)^{N+1}} = \frac{1}{(PV)^{N+1} \lambda(T)} \end{aligned}$$

Ignoring non-extensive contribution,

$$G = -K_B T \ln Z \approx N K_B T \left[\ln P - \frac{1}{2} \ln(K_B T) + \frac{3}{2} \ln \left(\frac{h}{2\pi m} \right) \right]$$

$$\text{Now again, from } dG = -SdT + VdP + \mu dN$$

$$V = \left. \frac{\partial G}{\partial P} \right|_{T, N} = \frac{N K_B T}{P} \Rightarrow PV = N K_B T$$

$$\text{Enthalpy } H = \langle E + PV \rangle$$

$$H = - \frac{\partial \ln Z}{\partial \beta} = \frac{5}{2} N K_B T$$

$$\& C_P = \frac{dH}{dT} = \frac{5}{2} N K_B$$

2) Spins +

Adding the work done against the external magnetic field we get

$$\mathcal{Z}(N, T, B) = \text{tr} \left[e^{-\beta H + \beta \vec{B} \cdot \vec{M}} \right] \quad \xrightarrow{\text{'em' not } \mu}$$

Assuming $\delta H = 0$ (non-interacting spins) we have

$$\text{where } \vec{M} = \mu_0 \sum_{i=1}^N \sigma_i$$

$\mu \equiv \{\sigma_i\}$ where σ_i is the i-th variable

$$\sigma_i = \{-1, 1\}$$

$$p(\{\sigma_i\}) = \frac{1}{Z} e^{\beta \mu_0 \vec{B} \cdot \sum_{i=1}^N \sigma_i}$$

$$\begin{aligned} Z &= \sum_{\{\sigma_i\}} e^{\beta \mu_0 \vec{B} \cdot \sum_{i=1}^N \sigma_i} = (e^{-\beta \mu_0 B} + e^{+\beta \mu_0 B})^N \\ &= [2 \cosh(\beta \mu_0 B)]^N \end{aligned}$$

$$\text{Thus } G = -k_B T \ln Z = -N k_B T \ln [2 \cosh(\beta \mu_0 B)]$$

Average magnetization is given by

$$\Rightarrow M = k_B T \frac{\partial \ln Z}{\partial B} \xrightarrow{\text{not } \beta} M = -\frac{\partial G}{\partial B} = N \mu_0 \tanh(\beta \mu_0 B)$$

$$\tanh(x) = x - \frac{x^3}{3} + \frac{2x^5}{45} + \dots$$

Thus, susceptibility $\Rightarrow M \approx N \mu_0 \beta \mu_0 B$

$$X(T) = \left. \frac{\partial M}{\partial B} \right|_{B=0} = \frac{N \mu_0^2}{k_B T}$$

The enthalpy is $H = \langle \mathcal{H} - BM \rangle = -BM$

$$C_B = \frac{\partial H}{\partial T} = -B \frac{\partial M}{\partial T}$$

§ 4.9 The grand canonical ensemble

9.1

$$M = (\tau, \mu, x)$$

$$p(M_s) = \frac{e^{[\beta \mu N(M_s) - \beta H(M_s)]}}{Q}$$

$$\begin{aligned} Q(\tau, \mu, x) &= \sum e^{\beta \mu N(M_s) - \beta H(M_s)} \\ &= \sum_{N=0}^{\infty} e^{\beta \mu N} \sum_{(M_s|N)} e^{-\beta H_N(M_s)} \\ &\stackrel{?}{=} \sum_{N=0}^{\infty} e^{\beta \mu N} Z(\tau, N, x) \end{aligned}$$

$$p(N) = \frac{e^{\beta \mu N} Z(\tau, N, x)}{Q}$$

$$\langle N \rangle = \frac{1}{Q} \frac{\partial Q}{\partial \beta \mu} = \frac{\partial \ln Q}{\partial \beta \mu}$$

$$\begin{aligned} \langle N^2 \rangle_c &= \langle N^2 \rangle - \langle N \rangle^2 \\ &= \frac{1}{Q} \cdot \frac{\partial^2 \ln Q}{\partial (\beta \mu)^2} - \left(\frac{\partial \ln Q}{\partial (\beta \mu)} \right)^2 \\ &\stackrel{?}{=} \cdot \frac{\partial^2}{\partial (\beta \mu)^2} \ln Q = \frac{\partial \langle N \rangle}{\partial (\beta \mu)} \end{aligned}$$

The sharpness of the distribution of N allows us to approximate the sum (Q) by its largest value $N = N^* \approx \langle N \rangle$

$$\begin{aligned} Q(\tau, \mu, x) &= \lim_{N \rightarrow \infty} \sum_{N=0}^{\infty} e^{\beta \mu N} Z(\tau, N, x) e^{-\beta(-NN^* + F)} \\ &= e^{-\beta(-NN^*)} Z(\tau, N^*, x) = e^{-\beta NN^*} \end{aligned}$$

$$= e^{-\beta(-\mu N^* + E - TS)} = e^{-\beta \bar{f}}$$

where

$$\boxed{\bar{f}(\bar{T}, \bar{N}, \bar{x}) = E - TS - \mu N = -k_B T \ln Q}$$

is the grand potential.

Using $d\bar{f} = -SdT - Nd\mu + \bar{T}dx$

we can evaluate $S, N, \delta T$ by appropriate partials.

Ideal gas of non-interacting particles

Final Example |

$$M = (T, \mu, V)$$

$$\mu = \{\vec{p}_i, \vec{q}_i, \dots\} \text{ c indefinite particle}$$

$$Q(T, \mu, V) = \sum_{N=0}^{\infty} e^{\beta \mu N} \underbrace{\frac{1}{N!} \int \left(\prod_{i=1}^N \frac{d^3 \vec{q}_i d^3 \vec{p}_i}{h^3} \right)}_{\text{volume element}} e^{-\beta \sum_i \frac{p_i^2}{2m}}$$

$$= \sum_{N=0}^{\infty} \frac{e^{\beta \mu N}}{N!} \left(\frac{V}{\lambda^3} \right)^N \quad \text{with } \lambda = \sqrt[3]{\frac{h}{2\pi m k_B T}}$$

$$= \sum_{N=0}^{\infty} \frac{1}{N!} \left(\frac{e^{\beta \mu} V}{\lambda^3} \right)^N$$

$$= \exp \left(\frac{e^{\beta \mu} V}{\lambda^3} \right)$$

$$f(T, \mu, V) = -k_B T \ln Q = -k_B T e^{\beta \mu} \frac{V}{\lambda^3}$$

? using $E = TS + J \cdot x + \mu \cdot N$

$$g = E - TS - \mu \cdot N = J \cdot x = -PV$$

thus $P = -\frac{g}{V} = -\frac{\delta g}{\delta V} \Big|_{\mu, T} = \frac{k_B T e^{\beta \mu}}{V}$

$$N = -\frac{\delta g}{\delta \mu} \Big|_{T, V} = \frac{e^{\beta \mu} V}{V}$$

comparing we get eqⁿ of state

$$P = k_B \frac{T N}{V}$$

also

$$k_B T \ln \left(\frac{N V}{V} \right) = \mu$$

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