## § 11. Motion in one dimension

Motion of a system with one degree of freedom is said to take place in one dimension. The most general Lagrangian for such a system would be

where of) is some of of g.

If q is the cartezian co-ordinate, then

We needn't even write the eg's of motion. Lets start with Energy eg' (1sting red)  $\frac{1}{2}m\dot{x}^2 + U(x) = E \Rightarrow \frac{dx}{dt} = \left(\frac{2(E-U(x))}{m}\right)^{\frac{1}{2}}$ , Thus

$$t = \sqrt{\frac{dx}{\sqrt{[E - U(x)]}}} + cont.$$

The two const are E & The "const".

Now since Imi2>0, Thus E>, U(x)

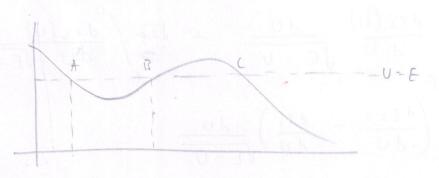
When U(x) = E, we get the limits of the motion. These points are

Finite = When the notion is bounded by Two such points, This is what The motion is said to be.

Infinite = If the region of motion is limited on only one side, or neither, then the motion is said to be this.

In 1-Dimension, a finite motion is oscillatory. The Time period then is trice the Time to go from x, l x2 (clear from 1) the integral by the Time reversibility) x, (E)

$$T(E) = \sqrt{\sum m} \int \frac{dx}{\sqrt{E - U(x)}} \quad \text{where } x, kx, \text{ are noots of } U(x) = E.$$



§ 12. Determination of the potential energy from the period of oscillation Given period of saillation T as a f' of E, we wish to find U(x). Mathemalically then, we want to solve  $T(E) = \sqrt{2m} \int \frac{dx}{\sqrt{E - U(x)}}$ , with T(E) as known  $\delta$  U(x) as unknown. We start with assuming that U(x) has only one root. minima in the region  $(x_1, x_2)$ . (We aren't considering solor. That may exist which don't satisfy the said condition) - Further, for convenience, we choose the origin at the position of the minima. (Why can we do this? THINK)

Instrictively its clear; Mothematically, so long as the field is not dependent on time, you can write the energy conservation in the x, x, x, x handlited co-ordinalis also) Further are take this minima to be zero. The eg's of motion don't depend on arkitary constants added to U. ( but T(E) seems to ... THINK) -> In the integral, we regard x as a f" of U (unlike earlier, where us regarded U as a for of 1). Also,  $\chi(0)$  is two-valued as is clear. Thus, we write  $\frac{\chi(0)}{\chi(0)}$  $x(0) = \begin{cases} x_1(0) & x_1 < x < 0 \\ x_2(0) & 0 < x < x_2 \end{cases}$ Further, we replace dx by  $\left(\frac{dx}{dU}\right)dU$ The limits of integration in terms of U, endently become E to 0

& 0 to E. (seems more like E to E to me, but thats cause using E tracker is to E. O then O, E)  $\Rightarrow T(E) = \sqrt{2m} \int \frac{dx_2(U)}{dU} \frac{dU}{\sqrt{E-U}} + \sqrt{2m} \int \frac{dx_1(U)}{dU} \frac{dU}{\sqrt{E-U}}$  $= \sqrt{2m} \int \left( \frac{dx_2}{dU} - \frac{dx_1}{dU} \right) \frac{dU}{\sqrt{E-U}}$ 

Both sides are now divided by  $\sqrt{x-E}$  ( $\alpha$  is a parameter) of integrated wit E from 0 to  $\alpha$ .

 $\Rightarrow \int_{0}^{\infty} \frac{T(E)}{\sqrt{\alpha - E}} dE = \sqrt{2m} \int_{0}^{\infty} \left[ \frac{dx_{\perp}}{dv} - \frac{dx_{1}}{dv} \right] \frac{dv dE}{\sqrt{(\alpha - E)(E - v)}}$ 

or changing the order of integrations (how stills)

 $= \sqrt{2}m \int_{0}^{\infty} \left[ \frac{dx_{1}}{dv} - \frac{dx_{1}}{dv} \right] dV \int_{0}^{\infty} \frac{dE}{\sqrt{(\alpha - E)(E - U)}}$ 

Thus we get  $\int_{-\infty}^{\infty} \frac{T(E) dE}{\sqrt{x_2 - E}} = \frac{11\sqrt{2}m}{\left(x_2(x) - x_1(x)\right)} - \left(x_2(x) - x_1(x)\right)$ 

Replacing & with U (x, & x, are f's of U)

 $x_{i}(u) - x_{i}(u) = \frac{1}{77\sqrt{2m}} \int_{0}^{U} \frac{T(\varepsilon) d\varepsilon}{\sqrt{U-\varepsilon}}$ 

Note | Remark:  $x_{\perp}(v)$  &  $x_{\parallel}(v)$  are still inditerminate. If we thus there are infinitely many U=U(x) which satisfy the time period dependence on F.

If we assume that the cure U=U(x) is symmetric about x=0,  $\Rightarrow x_{\parallel}(v)=-x_{\perp}(v)\equiv x(v)$ 

Then we get  $x(U) = \frac{1}{2\pi \sqrt{2m}} \int \frac{T(E)dE}{\sqrt{U-E}}$ questions I'm

(space for remarks on questions I'm

The potential energy of the interaction of 2 particle depends only on the distance b/w Them; in the Lagrangian is therefore

$$L = \frac{1}{2}m, \vec{3},^{2} + \frac{1}{2}m_{1}\vec{3}_{2}^{2} - U(1\vec{3}, -\vec{3}_{1})$$

 $\rightarrow$  We let  $\vec{x} = \vec{x}_1 - \vec{x}_2$ .

-> Let the origin be at the COM. Thus m, I, + m2 12 = 0

use there to get  $\vec{X}_1 = \frac{m_2 \vec{\lambda}}{m_1 + m_2}$ substitute in the lagrangian  $m_1 + m_2$   $\vec{\lambda}_1 = \frac{m_1 \vec{\lambda}}{m_1 + m_2}$   $\vec{\lambda}_2 = \frac{m_1 \vec{\lambda}}{m_1 + m_2}$ m, x, + m, x, = m, x

| Remark: This reduces the 2

body problem's into Lagrangian to where  $m \equiv m_1 m_2 = reduced mass$ that of a single particle in an or cotional field, symmetric about

a fixed origin. Pemark 2: The soln.  $\vec{\chi}(t)$  entails solns:  $\vec{\chi}_{i}(t) = \underbrace{m_{i}\vec{\chi}(t)}_{m_{i}+m_{i}} \hat{\vec{\chi}}_{i}(t)$ .

## \$14. Motion in a central field

The central field is one in which the potential energy depends only on the distance of a of so the particle from some fixed origin.

The force is thus

$$\vec{F} = -\frac{\partial u(x)}{\partial \vec{x}} = -\frac{\partial x}{\partial u(x)} \left(\frac{\vec{x}}{\vec{x}}\right) \qquad \frac{\partial x}{\partial u(x)} = \frac{\partial x}{\partial u(x)} \frac{\partial x}{\partial u(x)} = \frac{\partial x}{\partial u($$

I As we'd already seen, angular momentum defined wit the centre of the field, will have be conserved.

The angular momentum of a single particle is  $M = \vec{x} \times \vec{p}$ . Since M is const, 3) the radius victor is always in a plane I M. Thus the motion is always constrained to a plane.

Ilsing polar co-ordinates for a plane,

$$L = \frac{1}{2}m(\dot{x}^2 + x^2\dot{p}^2) - U(x)$$

Note: L'is independent of of (explicitly)  $= \frac{d}{dt} \left( \frac{\partial b}{\partial \Gamma} \right) = \frac{\partial b}{\partial \Gamma} = 0$ 

lydie = Any generalized a-ordinate that doesn't appear explicitly in the Lagrangian is called this. \* The corresponding generalized momentum is an integral of motion. Mere, 10 = Mz (Think: This & know that M2 = 21 From the law of conservation of y momentum, por them else can I show the law of = const.

What is the general relation from the general relation (1 2) (10) = df Jad Indd Y base (height)  $\Rightarrow M = 2mf$ where f = the sectorial velocity. + is const. (Keppler's second law) Constancy of M implies The soln to the central field problem We use it Conservation of 1) Energy ( 2) Angular Momentum  $E = \frac{1}{2} m(\dot{x}^2 + \dot{x}^2 \dot{\phi}^2) + U(4) = \frac{1}{2} m \dot{\dot{x}}^2 + \frac{1}{2} \frac{\dot{M}^2}{m \lambda^2} + U(2)$  $\dot{S} = \frac{dS}{dt} = \int \frac{2}{m} \left[ E - U(S) \right] - \frac{M^2}{m^2 \chi^2}$  $t = \int \frac{dt}{\int \frac{2}{m} \left[E - U(\Lambda)\right] - \frac{M^2}{m^2 R^2}} + const.$ M=mgg, dø = Mdt Also, from

 $= \frac{M du}{\sqrt{\frac{2m[E-U(x)]-M^2}{x^2}}}$ 

Remark: \$ varies monotonically with time, since \$ = conet > \$ news changes sign.

Further, the radial part of the motion can be taken as the motion in one dimension with

Veff = U(1) + M2
2m22

M² is called = centrique energy.

Posts of  $E = U(x) + \frac{M^2}{2ma^2}$  yield the limits of the

radial motion; the dutance from the concentre.

Note: this \$ that the particle is at rest when i = 0 -:
\$\phi \phi 0.

lase 1: & has only 1 bound, &>, min.
The particle comes from & goes to infinity.

lase 2: 3 is bounded: 9 min < 2 < 4 max

The path is bounded by corresponding wieles but  $\not\Rightarrow$  the path is closed. That happens in specific when  $\Delta \phi = 2 \overline{\Lambda} m \pmod{m, n \in I}$ 

where  $\Delta \phi = \frac{2\pi m}{n}$   $(m, n \in I)$  why the factor of 2? where  $\Delta \phi = \frac{2\pi m}{n}$  and  $\frac{1}{n}$  for a complete verolution,  $\frac{1}{n}$   $\frac{1}{n}$ 

m surdutions, the radial part would'se completed look at the image jum in the text)

Claim: I only two types of central fields in which all finite motions take place in closed paths.

Remark on puth construction from repeating segments: Read from the text. (It fun)

Remark: When  $M \neq 0$ , (rather obvious)

From energy conservation, we have  $\frac{1}{2}m\dot{x}^2 = E - U(x) - \frac{M^2}{2ma^2} > 0$   $\Rightarrow 3^2U(x) + \frac{M^2}{2m} \geq Ex^2$ Now a can take values tending to zero if  $\left[x^2U(x)\right]_{x\to 0} < -\frac{M^2}{2m}$ viz. U(x) must tend to  $-\infty$  either if as  $-\frac{x}{2}$  with  $x > \frac{M^2}{2m}$ or proportionally to  $-\frac{1}{2}$ , x > 2.

§ 15. Kepler's Problem