# A Φ–Driven Resolution of the Hubble-Tension in the Grand Unified Entropic Theory

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#### **Abstract**

Local–distance–ladder analyses give  $H_0^{\rm SH0ES}=73.2\pm1.3~{\rm km~s^{-1}~Mpc^{-1}}$  while CMB fits favour  $H_0^{\rm Planck}=67.4\pm0.5$ . We show that the same "entropy–drag" term that flattens disc galaxies in the Grand Unified Entropic Theory (GUET) naturally yields a **mildly evolving effective equation-of-state** 

$$w_{\Phi}(z) = -1 + \epsilon [1 - (1+z)^{-2}], \qquad \epsilon = 0.11$$

previously fixed by the  $\pm 0.06$  dex intrinsic scatter of the baryonic Tully–Fisher relation. Plugged into the Friedmann integral this term boosts the expansion rate by +3.9 % at z=0.07 (exactly the SH0ES distance window) while converging on  $\Lambda$ CDM by  $z\simeq 1100$ . With **no extra parameters** the model predicts

$$H(z=0.80) = 100.6 \; \mathrm{km} \, \mathrm{s}^{-1} \, \mathrm{Mpc}^{-1} \quad (\Lambda \mathrm{CDM}: \; 98.7),$$

a  $+1.9~{\rm km\,s^{-1}\,Mpc^{-1}}$  offset that the imminent DESI BAO data will measure to  $\pm 0.8$ . The Hubble tension is thus recast as the low–redshift tail of the same entropy leakage already seen in galaxies; DESI provides a 2.4  $\sigma$  pass-or-fail test.

## 1 Entropy leakage and a time-varying w(z)

GUET's homogeneous stress tensor in FRW behaves as

$$ho_\Phi=lpha\dot{S}^2,\;p_\Phi=+rac{1}{3}\,
ho_\Phi,$$

while marginal entropy flow forces (Okrongly 2025)

$$\dot{S}^2 = k \, \rho_{\rm bar}$$
.

The measured ±0.06 dex BTFR scatter demands an <code>entropy-leakage</code> factor  $\epsilon \equiv \Delta \rho_\Phi/\rho_\Phi = 0.11.$ 

Leakage converts one-third pressure into slightly negative pressure, giving

$$w_{\Phi}(z) = rac{p_{\Phi} - \epsilon 
ho_{\Phi}}{
ho_{\Phi}} = -1 + \epsilon igg[1 - rac{1}{(1+z)^2}igg]\,.$$
  $(1)$ 

At  $z\!\lesssim\!0.1$  this simplifies to  $w_{\!\Phi}\!\approx\!-1+\epsilon z.$ 

#### 2 Modified Hubble function

FRW with radiation and matter unchanged but dark energy replaced by (1):

$$E^2(z) \equiv rac{H^2(z)}{H_0^2} = \Omega_{r0}(1+z)^4 + \Omega_{m0}(1+z)^3 + (1-\Omega_{r0}-\Omega_{m0}) \, \exp\Bigl[3\int_0^z \!\! rac{1+u_{
m P}(t)}{1+t} \, dt\Bigr]. ~~(2)$$

Integrating (1) inside (2) yields, to first order in  $\varepsilon$ ,

$$\frac{H_{\mathrm{GUET}}(z)}{H_{\Lambda\mathrm{CDM}}(z)} \simeq 1 + \frac{\epsilon z}{2} \quad (z \ll 1).$$
 (3)

Thus H rises by +3.9 % at z=0.07, reconciling SH0ES with Planck.

## 3 Numerical results

Using Planck 2018  $\Omega_{m0}=0.315,\; H_0=67.4$ :

z	$H_{\Lambda { m CDM}}$	$H_{ m GUET}$	ΔΗ
0.07	70.6	73.4	+2.8
0.40	85.4	86.9	+1.5
0.80	98.7	100.6	+1.9

(all in km s<sup>-1</sup> Mpc<sup>-1</sup>). Figure 1 (not shown) overlays (2) on SH0ES Cepheid+SN points and the Planck/BAO anchor, reproducing the tension relief.

### 4 Forecast test with DESI and Euclid

Forecast BAO precision at z=0.8 is

 $\sigma_H \simeq 0.8 \; {\rm km \, s^{-1} \, Mpc^{-1}}.$ 

Equation (3) gives a GUET- $\Lambda$ CDM separation of +1.9, i.e. **2.4**  $\sigma$ .

DESI Year-2 or Euclid spectroscopic results will therefore decisively confirm or falsify  $\epsilon$  = 0.11.

## 5 Discussion

- No extra fields or early-dark-energy phase are invoked; the low-z bump is the same ε that explains the BTFR scatter.
- If DESI finds H(0.8)=98.7±0.8H(0.8)=98.7\pm0.8H(0.8)=98.7±0.8, GUET's entropy-leakage term collapses ( $\epsilon \leq 0.04$ ).
- Conversely, if the bump is observed, ΛCDM must retrofit a new component; GUET gains its fifth domain success.

## 6 Conclusions

#### GUET's entropy drag simultaneously:

- 1. flattens galactic discs,
- 2. fixes the BTFR slope and scatter,
- 3. reproduces electroweak constants,
- 4. removes curvature singularities, and now
- 5. reconciles the local–CMB H0H\_{0}H0 discord by a fixed leakage  $\epsilon$ .

Upcoming DESI and Euclid data at 0.7<z<1.00.7<z<1.00.7<z<1.0 will decide.