

## 1 Constants and units

$$m_e \text{ (mass of an electron)} = 9.10938291(40) \times 10^{-31} kg \quad (1)$$

$$= 510.9989 \text{ keV}/c^2 \quad (2)$$

$$m_p \text{ (mass of a proton)} = 0.938272 \text{ GeV}/c^2 \quad (3)$$

## 2 Mathematics

### Trigonometric Identities

$$\sin(\theta \pm \phi) = \sin(\theta) \cos(\phi) \pm \cos(\theta) \sin(\phi) \quad (4)$$

$$\cos(\theta \pm \phi) = \cos(\theta) \cos(\phi) \mp \sin(\theta) \sin(\phi) \quad (5)$$

### Series expansions.

$$\text{[Def]: } f(x) = f(a) + f'(a)(x-a) + \frac{1}{2!}f''(a)(x-a)^2 + \frac{1}{3!}f'''(a)(x-a)^3 + \dots \quad (6)$$

$$f(x) = e^x \implies f(x) = 1 + x + \frac{1}{2!}x^2 + \frac{1}{3!}x^3 + \dots \quad (7)$$

$$f(x) = \ln(1+x) \implies f(x) = x - \frac{1}{2}x^2 + \frac{1}{3}x^3 - \dots [|x| < 1] \quad (8)$$

$$f(x) = \sin(x) \implies f(x) = x - \frac{x^3}{3!} + \frac{x^5}{5!} - \dots \quad (9)$$

$$f(x) = \cos(x) \implies f(x) = 1 - \frac{x^2}{2!} + \frac{x^4}{4!} - \dots \quad (10)$$

$$f(x) = \sinh(x) \implies f(x) = x + \frac{x^3}{3!} + \frac{x^5}{5!} + \dots \quad (11)$$

$$f(x) = \cosh(x) \implies f(x) = 1 + \frac{x^2}{2!} + \frac{x^4}{4!} + \dots \quad (12)$$

$$(1+x)^n = 1 + nx + \frac{n(n-1)}{2!}x^2 + \dots [|x| < 1] \quad (13)$$

### 3 Differential Equations

$$\ddot{x} + Ax = B \implies x(t) = \frac{B}{A} + C_1 \cos(\sqrt{A}t) + C_2 \sin(\sqrt{A}t) \quad (14)$$

$$\ddot{x} - Ax = B \implies x(t) = -\frac{B}{A} + C_1 e^{\sqrt{A}t} + C_2 e^{-\sqrt{A}t} \quad (15)$$

$$\text{or } \dots \implies x(t) = -\frac{B}{A} + C_1 \sinh(\sqrt{A}t) + C_2 \cosh(\sqrt{A}t) \quad (16)$$

## 4 Statistics

The binomial coefficient

$$\binom{n}{k} = \frac{n!}{(n-k)!k!} \quad (17)$$

A probability distribution

$$\mathcal{P}_n(m; p) = \binom{n}{m} p^m (1-p)^{n-m} \quad (18)$$

The Poisson Distribution

$$\mathcal{P}(m, \lambda) = \frac{\lambda^m}{m!} e^{-\lambda} \quad (19)$$

The mean number of events is

$$\langle m \rangle = \sum_{m=0}^{\infty} m \frac{\lambda^m}{m!} e^{-\lambda} = \lambda \quad (20)$$

And the standard deviation is

$$\sigma = \sqrt{\lambda} \quad (21)$$

The normal, or Gaussian distribution

$$\mathcal{P}(x; \mu, \sigma) = \frac{1}{\sqrt{2\pi}\sigma} \exp\left[-\frac{(x-\mu)^2}{2\sigma^2}\right] \quad (22)$$

Given some function  $f(x_1, x_2, \dots, x_n)$ , the error of a calculation can be determined by

$$\sigma_f^2 = \left(\frac{\partial f}{\partial x_1}\right)^2 \sigma_{x_1}^2 + \left(\frac{\partial f}{\partial x_2}\right)^2 \sigma_{x_2}^2 + \dots + \left(\frac{\partial f}{\partial x_n}\right)^2 \sigma_{x_n}^2 \quad (23)$$

## 5 Classical Mechanics

Dot and cross products

$$\vec{r} \cdot \vec{s} = rs \cos(\theta) = r_x s_x + r_y s_y + r_z s_z \quad (24)$$

$$\vec{r} \times \vec{s} = (r_y s_z - r_z s_y, r_z s_x - r_x s_z, r_x s_y - r_y s_x) = \det \begin{vmatrix} \hat{x} & \hat{y} & \hat{z} \\ r_x & r_y & r_z \\ s_x & s_y & s_z \end{vmatrix} \quad (25)$$

Newtons Second Law in Cartesian coordinates

$$\vec{F} = m\vec{a} = m\ddot{\vec{r}} \iff \begin{cases} F_x = m\ddot{x} \\ F_y = m\ddot{y} \\ F_z = m\ddot{z} \end{cases} \quad (26)$$

Newtons Second Law in 2D polar coordinates

$$\vec{F} = m\vec{a} \iff \begin{cases} F_r = m(\ddot{r} - r\dot{\phi}^2) \\ F_\phi = m(r\ddot{\phi} + 2\dot{r}\dot{\phi}) \end{cases} \quad (27)$$

Newtons Second Law in cylindrical polar coordinates

$$\vec{F} = m\vec{a} \iff \begin{cases} F_r = m(\ddot{\rho} - \rho\dot{\phi}^2) \\ F_\phi = m(\rho\ddot{\phi} + 2\dot{\rho}\dot{\phi}) \\ F_z = m\ddot{z} \end{cases} \quad (28)$$

The Lorentz Force on a charged particle.

$$\vec{F} = q(\vec{E} + \vec{v} \times \vec{B}) \quad (29)$$

Equation of motion for a rocket

$$m\dot{v} = -\dot{m}v_{ex} + F^{external} \quad (30)$$

The center of mass of several particles

$$\vec{R} = \frac{1}{M} \sum_{\alpha=1}^N m_\alpha \vec{r}_\alpha = \frac{m_1 \vec{r}_1 + \cdots + m_N \vec{r}_N}{M} \quad (31)$$

$$\vec{R} = \frac{1}{M} \int \vec{r} dm = \frac{1}{M} \int \rho \vec{r} dV \quad (32)$$

## 6 Special Relativity

Relativistic time dilation and length contraction.

$$\Delta t = \frac{\Delta t_o}{\sqrt{1 - \beta^2}} = \gamma \Delta t_o \quad (33)$$

$$\Delta l = \Delta l_o \sqrt{1 - \beta^2} = \frac{\Delta l_o}{\gamma} \quad (34)$$

$$\beta = \frac{v}{c} \quad (35)$$

$$\gamma = \frac{1}{\sqrt{1 - \beta^2}} \quad (36)$$

Lorentz Transformations for space and time coordinates.

$$x' = \gamma(x - vt) \quad (37)$$

$$y' = y \quad (38)$$

$$z' = z \quad (39)$$

$$t' = \gamma(t - vx/c^2) \quad (40)$$

The relativistic velocity transformation is.

$$u' = \frac{u - v}{1 - vu/c^2} \quad (41)$$

$$u = \frac{u' + v}{1 + vu'/c^2} \quad (42)$$

The rest energy of a particle

$$E_0 = mc^2 \quad (43)$$

the lorentz transformation for momentum and energy is.

$$p'_x = \gamma(p_x - vE/c^2) \quad (44)$$

$$p'_y = p_y \quad (45)$$

$$p'_z = p_z \quad (46)$$

$$E' = \gamma(E - vp_x) \quad (47)$$

Relativistic mass and momentum.

$$E = \gamma mc^2 \quad (48)$$

$$p = \gamma mv \quad (49)$$

Combining the above equations give

$$\frac{E}{p} = \frac{c^2}{v} \implies E = \frac{pc^2}{v} \quad (50)$$

Mass-energy equivalence

$$E^2 = (mc^2)^2 + (pc)^2 \quad (51)$$

$$E = K_E + E_0 \quad (52)$$

$K_E =$  Kinetic Energy

Combining the above equation gives

$$p = \frac{1}{c} \sqrt{K_E^2 + 2K_E E_0} \quad (53)$$

Invariant dot product in c=1 notation

$$A \cdot B = (E, \vec{p}) \cdot (U, \vec{q}) = EU - \vec{p} \cdot \vec{q} \quad (54)$$

Relativistic frequency and wavelength shifts

$$f = f_0 \sqrt{\frac{c \pm v}{c \mp v}} \quad (55)$$

$$\lambda = \lambda_0 \sqrt{\frac{c \mp v}{c \pm v}} \quad (56)$$

Space-time equivalence (same in all reference frames)

$$S \equiv (c\Delta t)^2 - (\Delta x)^2 \equiv E^2 - (pc)^2 \quad (57)$$

## 7 Thermodynamics

Useful constants:

$$\text{The specific heat of water is } c = 4186 \text{ J/(kg}\cdot\text{K).} \quad (58)$$

$$1 \text{ cal} = 4.186 \text{ J} \quad (59)$$

Temperature relationships.

$$^{\circ}\text{F} = \frac{9}{5} C + 32 \quad (60)$$

$$^{\circ}\text{C} = \frac{5}{9} (^{\circ}\text{F} - 32) \quad (61)$$

$$^{\circ}\text{K} = ^{\circ}\text{C} + 273.15 \quad (62)$$

The heat required to raise the temperature of a mass  $m$  by  $\Delta T$  is

$$Q = cm\Delta T \quad (63)$$

The temperature of an object determines the radiated power of the object, which is given by the **Stefan-Boltzmann equation**

$$P_{\text{radiated}} = \sigma \epsilon A T^4 \quad (64)$$

$$\sigma = 5.67 \times 10^{-8} \text{ W/K}^4\text{m}^2 \quad (65)$$

$$\epsilon = \text{emissivity, and } 0 \leq \epsilon \leq 1 \quad (66)$$

The work done on a system in going from initial volume ( $V_i$ ) to a final volume ( $V_f$ ) is

$$W = \int dW = \int_{V_i}^{V_f} p dV. \quad (67)$$

The first law of thermodynamics

$$\Delta E_{\text{internal}} = Q - W \quad (68)$$

different processes include

- (i) An adiabatic process is one in which  $Q = 0$ .
- (ii) In a constant-volume process,  $W = 0$ .
- (iii) In a closed-loop process,  $Q = W$ .
- (iv) In an adiabatic free expansion,  $Q = W = \Delta E_{\text{internal}} = 0$ .

If heat is added to an object, its change in temperature is given by

$$\Delta T = \frac{Q}{C} \quad (69)$$

$$C = \text{heat capacity of the object} \quad (70)$$



If heat is added to an object with mass  $m$ , its change in temperature is given by

$$\Delta T = \frac{Q}{cm} \quad (71)$$

$$c = \text{specific heat of the object} \quad (72)$$

The ideal gas law

$$PV = nRT \quad (73)$$

$$R = 1.38106504(24) \times 10^{-23} \text{ J/K} \quad (74)$$

With a constant number of moles we get from the ideal gas law the following relation:

$$\frac{P_1 V_1}{T_1} = \frac{P_2 V_2}{T_2} \quad (75)$$

Dalton's law - The total pressure exerted by a mixture of gases is equal to the sum of the partial pressures of the gases in the mixture.

$$P_{total} = P_1 + P_2 + P_3 + \cdots + P_n \quad (76)$$

The work done by an ideal gas at constant temperature is

$$W = nRT \ln \left( \frac{V_f}{V_i} \right) \quad (77)$$

The average kinetic energy of an ideal gas

$$K_{ave} = \frac{1}{N} \sum_{i=1}^N K_i = \frac{1}{N} \sum_{i=1}^N \frac{1}{2} m v_i^2 = \frac{1}{2} m v_{rms}^2 \quad (78)$$

The root-mean-square speed of gas molecules is

$$v_{rms} = \sqrt{\frac{1}{N} \sum_{i=1}^N v_i^2} = \sqrt{\frac{3RT}{m}} \quad (79)$$

For an adiabatic process, we have

$$dE_{internal} = -PdV = nC_V dT \quad (80)$$

$$C_V = \text{specific heat at constant volume} \quad (81)$$

$$C_P = \text{specific heat at constant pressure} \quad (82)$$

$$PV^\gamma = \text{constant} \quad (83)$$

$$\gamma = \frac{C_P}{C_V} \quad (84)$$

$$P_f V_f^\gamma = P_i V_i^\gamma \quad (85)$$

$$T_f V_f^{\gamma-1} = T_i V_i^{\gamma-1} \quad (86)$$

## 8 Quantum Physics

Electromagnetic wave frequency and wavelength

$$c = \nu\lambda \implies \nu = \frac{c}{\lambda} \implies \lambda = \frac{c}{\nu} \quad (87)$$

$$\nu = \text{frequency} \quad (88)$$

The energy in a photon (packet of light)

$$E = h\nu = \frac{hc}{\lambda} \quad (89)$$

The proportionality factor (Planck's constant) is

$$h = 6.6261 \times 10^{-34} J \cdot s \quad (90)$$

$$= 4.1357 \times 10^{-15} eV \cdot s$$

$$hc = 1240 eV \cdot nm \quad (91)$$

$$\hbar = \frac{h}{2\pi} = 1.05 \times 10^{-34} J \cdot s \quad (92)$$

Wien's Displacement Law

$$\lambda_{MAX} T = 2.898 \times 10^{-3} m \cdot K \quad (93)$$

Total Power Stefan-Boltzmann Law

$$R(T) = \int_0^\infty I(\lambda, T) d\lambda = \epsilon \sigma T^4 \quad (94)$$

$$\epsilon = \text{emissivity (unitless)} \quad (95)$$

$$\sigma = 5.67 \times 10^{-8} \frac{W}{m^2 K^4} \quad (96)$$

Max Planck's Radiation Law:

$$I(\lambda, T) = \frac{2\pi c^2 h}{\lambda^5} \frac{1}{e^{hc/\lambda kT} - 1} \quad (97)$$

The kinetic energy of an emitted photoelectron is

$$KE = h\nu - \phi \quad (98)$$

$$E_{\text{photon}} = KE_{\text{electron}} + \phi \quad (99)$$

$$KE_{\text{electrons}} = 0 \text{ (at threshold)} \quad (100)$$

Where  $\phi$  = binding energy of electron to metal surface (the work function).

Rutherford Scattering Formula

Any particle hitting an area  $\sigma$  around the nucleus will be scattered through an angle of  $\theta$  or greater.

$$b = (r_{min}/2) \cot(\theta/2) \quad (101)$$

$$r_{min} = \frac{Z_1 Z_2 e^2}{4\pi\epsilon_0 K} \quad (102)$$

$$\sigma = \pi b^2 = \text{cross sectional area} \quad (103)$$

$$\frac{e^2}{4\pi\epsilon_0} = 1.44 \times 10^{-9} \text{eV}\cdot\text{m} \quad (104)$$

A common unit of  $\sigma$  is one barn.

$$\text{barn (unit)} = 10^{-28} \text{m}^2 = 100 \text{fm}^2 \quad (105)$$

Number of atoms per area = (atoms/volume)\*thickness

$$n = \left( N_A \frac{\text{atoms}}{\text{mole}} \right) \left( \frac{1}{A} \frac{\text{mole}}{\text{gm}} \right) \left( \rho \frac{\text{gm}}{\text{cm}^3} \right) = \frac{\rho N_A}{A} \quad (106)$$

The Compton effect describes the photon wavelength  $\lambda'$  after a photon of wavelength  $\lambda$  scatters off an electron.

$$\lambda' = \lambda + \frac{h}{m_e c} (1 - \cos(\theta)) \quad (107)$$

The Compton wavelength of an electron is

$$\lambda_e = \frac{h}{m_e c} = 2.426 \times 10^{-12} \text{m} \quad (108)$$

Heisenberg Uncertainty relation

$$\Delta x \cdot \Delta p_x \geq \frac{1}{2} \hbar \quad (109)$$

Uncertainty relation of energy and the measurement of time.

$$\Delta E \cdot \Delta t \geq \frac{1}{2} \hbar \quad (110)$$

The de Broglie wavelength is defined as

$$\lambda = \frac{h}{p} = \frac{h}{mv\gamma} = \frac{h}{mv} \sqrt{1 - \frac{v^2}{c^2}} = \frac{hc}{\sqrt{K_E^2 + 2K_E E_0}} \quad (111)$$

Rutherford Scattering.

$$K = \frac{1}{4\pi\epsilon_0} \frac{Z_1 Z_2 e^2}{R_{min}} \iff R_{min} = \frac{1}{4\pi\epsilon_0} \frac{Z_1 Z_2 e^2}{K} \quad (112)$$

$Z$ 's are the atomic masses of the particles within the interaction and  $R_{min}$  is the minimum distance they reach (from center to center), and  $e$  is

$$e = 1.602177 \times 10^{-19} C \quad (113)$$

$$\epsilon \approx 8.854 \times 10^{-12} F/m \quad (114)$$

The Rutherford Scattering Formula

$$N(\theta) = \frac{N_i n t}{16 r^2} (R_{min})^2 \frac{1}{\sin^4(\theta/2)} \quad (115)$$

Centripetal force due to coulomb attraction

$$\frac{1}{4\pi\epsilon_0} \frac{e^2}{r^2} = m a_c = m \frac{v^2}{r} \implies v^2 = \frac{1}{4\pi\epsilon_0} \frac{e^2}{mr} \quad (116)$$

$$\implies r = 4\pi\epsilon_0 \frac{n^2 \hbar^2}{m e^2} \quad (117)$$

Energy levels

$$E = KE + PE = -\frac{1}{2} \frac{e^2}{4\pi\epsilon_0 r} \implies E = \frac{-E_0}{n^2}, \quad (118)$$

$$\text{where } E_0 = \alpha^2 m c^2 / 2 = 13.6 \text{ eV}. \quad (119)$$

Energy of emitted radiation

$$E = E_n - E_m = E_0 \left( \frac{1}{m^2} - \frac{1}{n^2} \right) \quad (120)$$

**Note.** Using the Planck formula in the above equation leads to the Rydberg formula.

The Rydberg formula: Wavelength of the spectral lines in Hydrogen:

$$\frac{1}{\lambda} = R_H \left( \frac{1}{n_f^2} - \frac{1}{n_i^2} \right) \quad (121)$$

$$n \in \mathbb{N} = 1, 2, 3, 4, 5, \dots \quad (122)$$

$$R_H = \frac{E_0}{hc} = \frac{13.6 \text{ eV}}{1240 \text{ eV} \cdot \text{nm}} \quad (123)$$

$$= 10,967,760 \text{ m}^{-1} \quad (124)$$

$$= 1.096776 \times 10^7 \text{ m}^{-1} \text{ (Rydberg's constant)} \quad (125)$$

**Note.** ZnS (Zinc Sulfide) emits a faint flash of light when struck by an  $\alpha$ -ray.

L quantized

$$L = mvr = n\hbar \quad (126)$$

Stationary state orbits

$$r = a_0 n^2 \quad (127)$$

$$a_0 = \text{Bohr Radius} \quad (128)$$

Stationary state energies

$$E_n = -Z^2 \frac{E_0}{n^2} \quad (129)$$

$$\Delta E = \frac{hc}{\lambda} \quad (130)$$

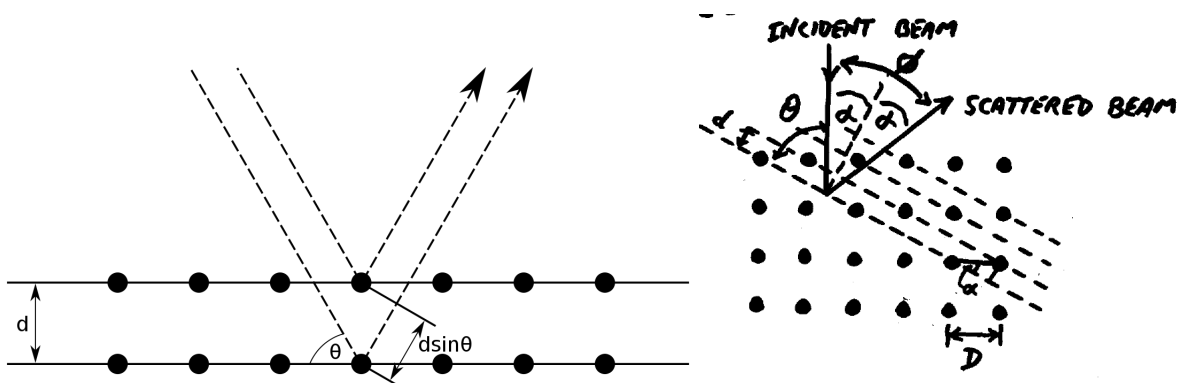
Bragg's Law: When scattering off of crystal structures, the wavelengths will peak at specific angles determined by the diagrams below

$$n\lambda = 2d \sin(\theta) = 2d \cos(\alpha) = 2D \sin(\alpha) \cos(\alpha) = D \sin(2\alpha) = D \sin(\phi) \quad (131)$$

$$d = D \sin(\alpha) \quad (132)$$

$$\phi = 2\alpha \quad (133)$$

$$\theta = 90^\circ - \alpha \quad (134)$$



## 9 Quantum Mechanics

A plane wave

$$\psi(x, t) = A \cos[2\pi(x - ct)/\lambda] \quad (135)$$

$$f = c/\lambda \quad (136)$$

$$T = 1/f \quad (137)$$

$$\psi(x, t) = A \cos(kx - \omega t) \quad (138)$$

$$k = 2\pi/\lambda \quad (139)$$

$$\omega = 2\pi f = 2\pi/T \quad (140)$$

A periodic wave can be constructed from a sum of plane waves

$$\psi(x, t) = \sum_{i=1}^n A_i \cos(k_i x_i - \omega_i t) \quad (141)$$

Fourier Transform: A wave packet can be constructed as a continuous sum of plane waves

$$\psi(x, t) = \int A(k) \cos(kx - \omega t) dk \quad (142)$$

The wave equation

$$\frac{1}{c^2} \frac{\partial^2 \psi}{\partial t^2} = \frac{\partial^2 \psi}{\partial x^2} \quad (143)$$

The Schrödinger Equation

$$i\hbar \frac{\partial \psi}{\partial t} = \frac{-\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} + V(x)\psi \quad (144)$$