



cherenkov
telescope
array



CTA Software and dependencies

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University of Rijeka, 16/02/21



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Curie FELLINI - Grant 754496



Analysis Software



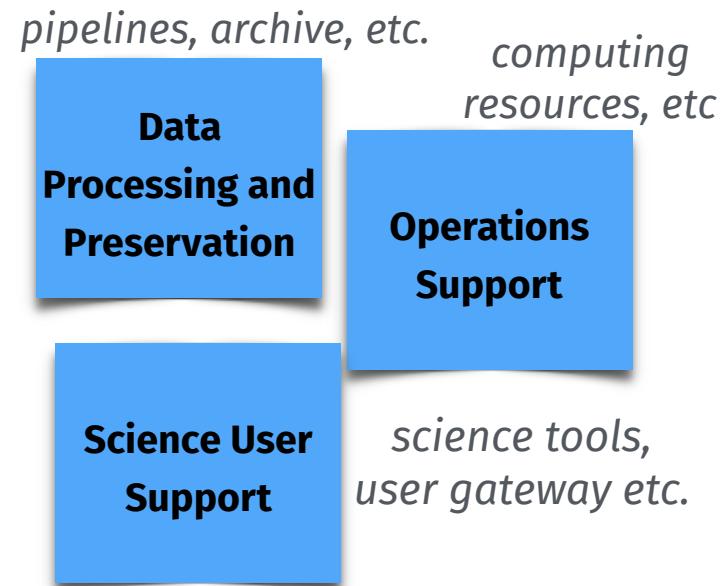
- The **low level** analysis software within the CTA philosophy falls within the so called Data Preparation and Preservation System (**DPPS**)
- **High level** analysis falls within the Science User Support (**SUSS**)

Previously...

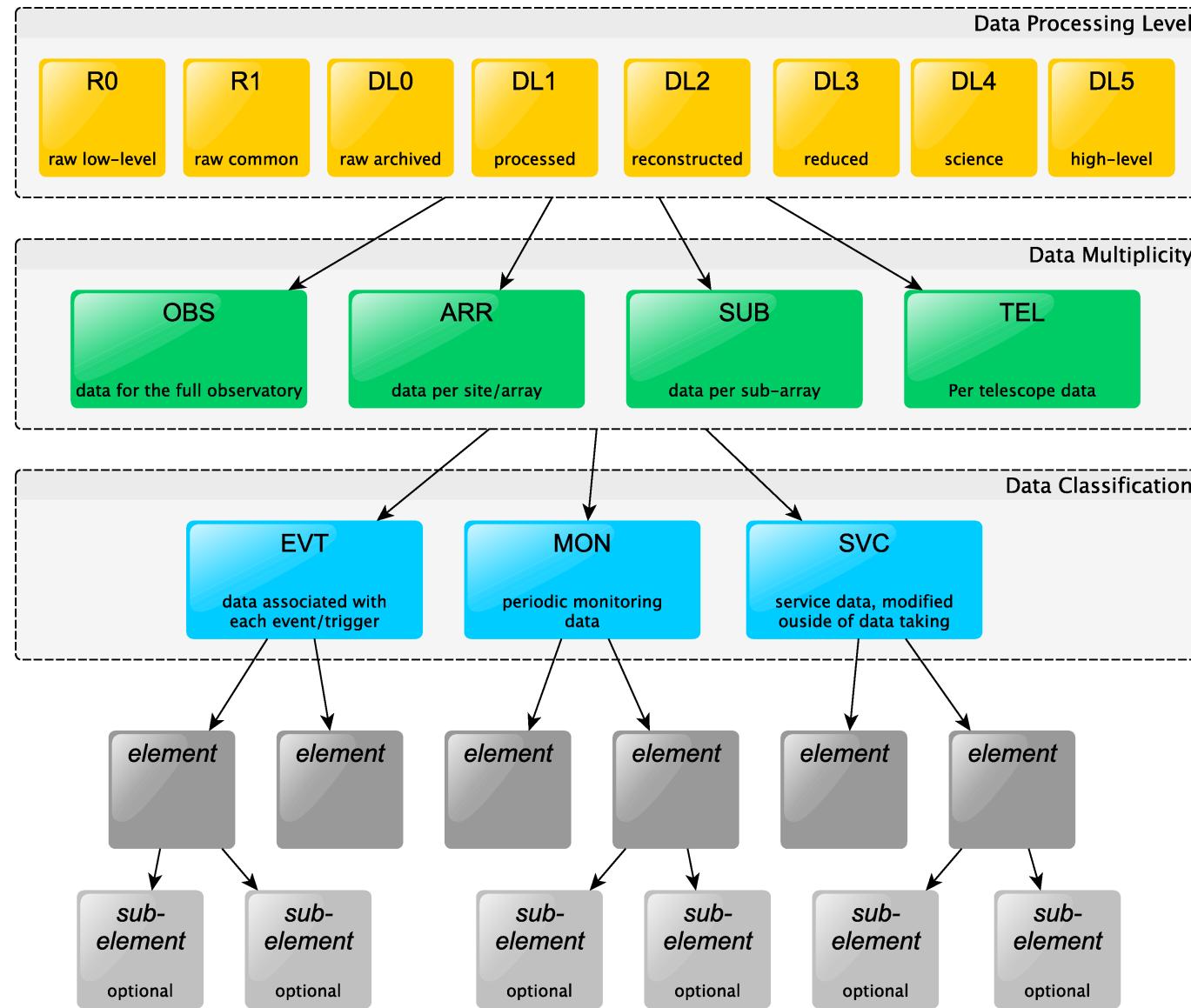
Data
management



Now:



Data Level definition



Data Levels: amount of processing



MCs are somewhere here right now



R0 (*raw low-level*) camera data transmitted from telescope to central servers. R0 content and format is internal to each camera and is specified and coordinated between individual camera teams.

Pipeline starts here
Will need to eventually produce data here to be compatible with "real" CTA data



R1 (*raw common*) data output by an individual camera functional unit to the camera DAQ functional unit. This is the first level of data seen by the ACTL system and is therefore as common as possible between all cameras/hardware. Exceptionally, some R1 data may be stored for engineering purposes.

Reconstructed Events
(many reconstructions and parameters, no more telescopes)



DL1 (*processed*) processed DL0 data that may still include some TEL data and parameters derived from them. For example this includes calibrated image charge, Hillas parameters, and a usable telescope pattern. This is only optionally stored in the archive.

Science Data:
Classified Events
(final reconstruction),
IRFs



DL2 (*reconstructed*) reconstructed shower parameters such as energy, direction, particle ID, and related signal discrimination parameters. At this point, no TEL information is stored. For each event this information may be repeated for multiple reconstruction and discrimination methods. This is only optionally stored in the archive. At this point, telescope-wise info is generally dropped.

DL3 (*reduced*) Sets of selected (e.g. gamma-ray candidates, electron candidates, selected hadron candidates, etc.) events with a single final set of reconstruction and discrimination parameters, along with associated instrumental response characterizations and any technical data needed for science analysis.

DL4 (*science*) binned data products like spectra, sky maps, or light curves, along with associated data (source models, fit results, etc).

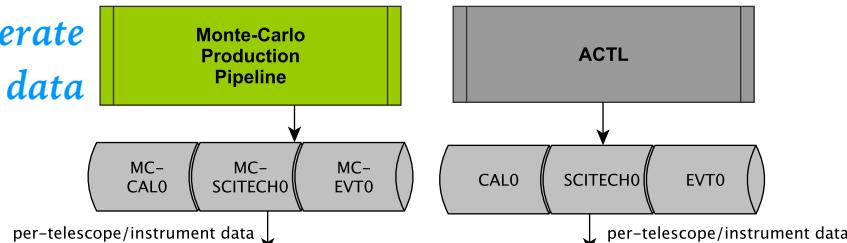
DL5 (*high-level*) high-level or "legacy" observatory data, such as CTA survey sky maps or the CTA source catalog.

high-level functionality

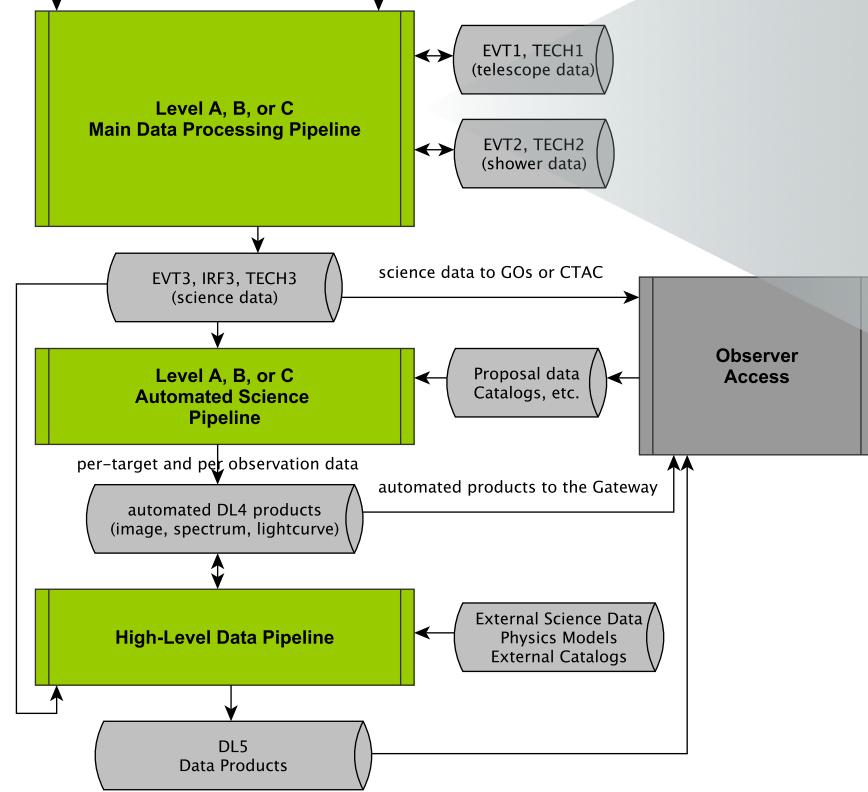
source: TDR, names a bit out of date



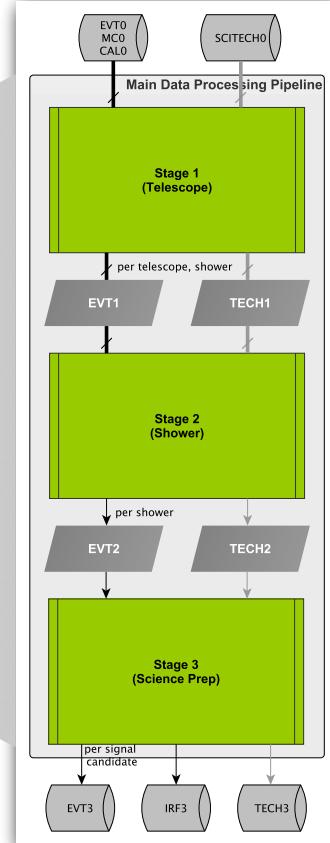
*generate
MC data*



*make event lists
and IRFs*



*run science
tools, detect
sources, make
sky plots, etc.*



See the DATA TDR:

2.1.2.3 Main Data Processing
Pipeline Common Functional Design

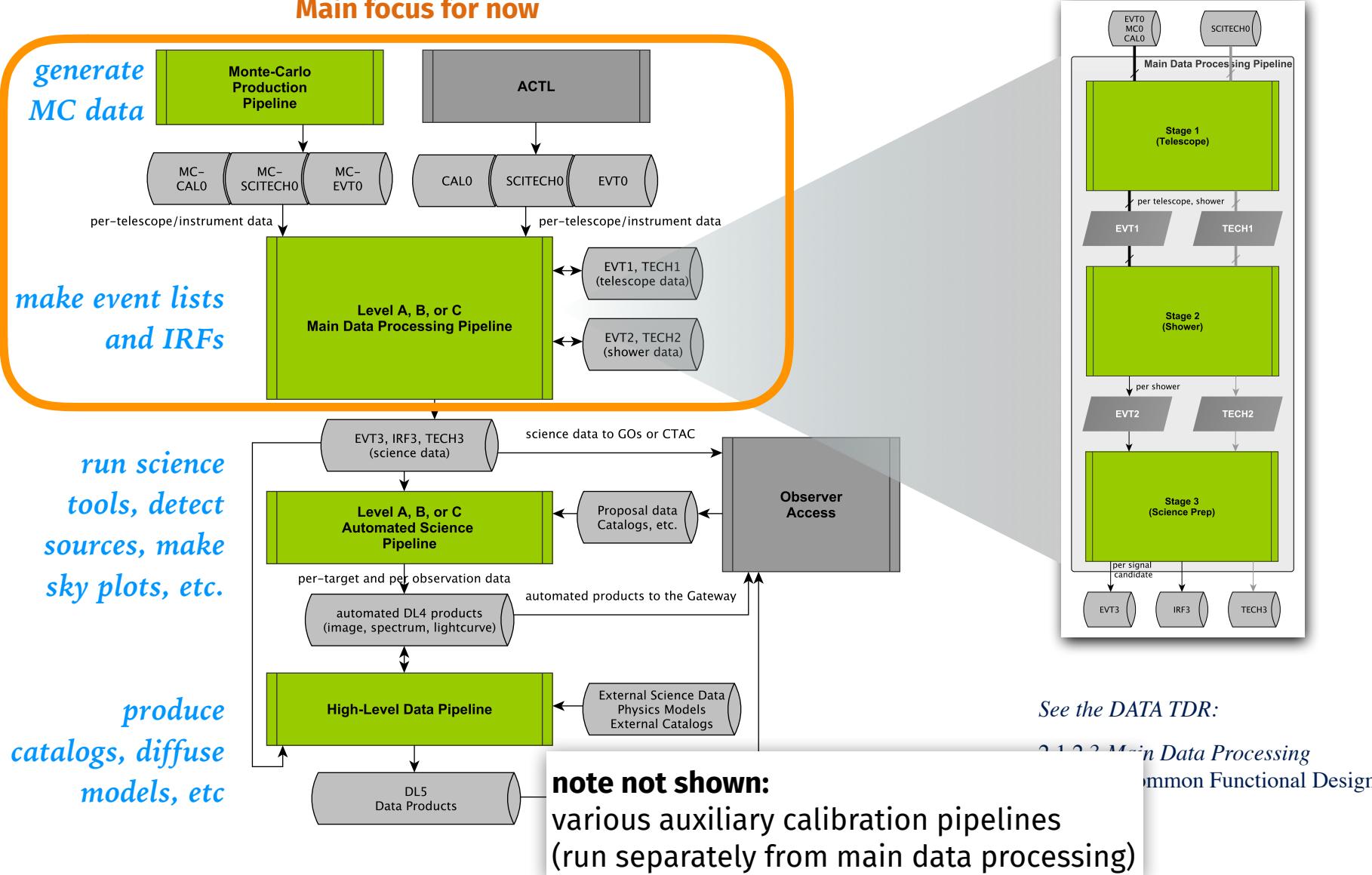
*produce
catalogs, diffuse
models, etc*

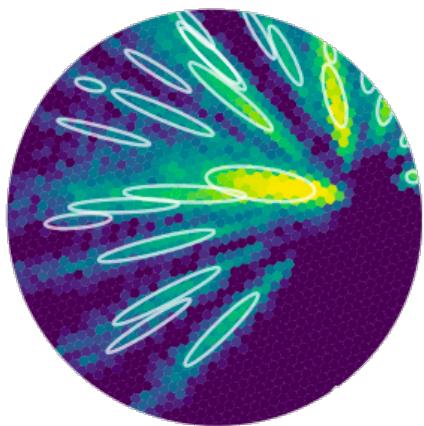
high-level functionality

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Main focus for now



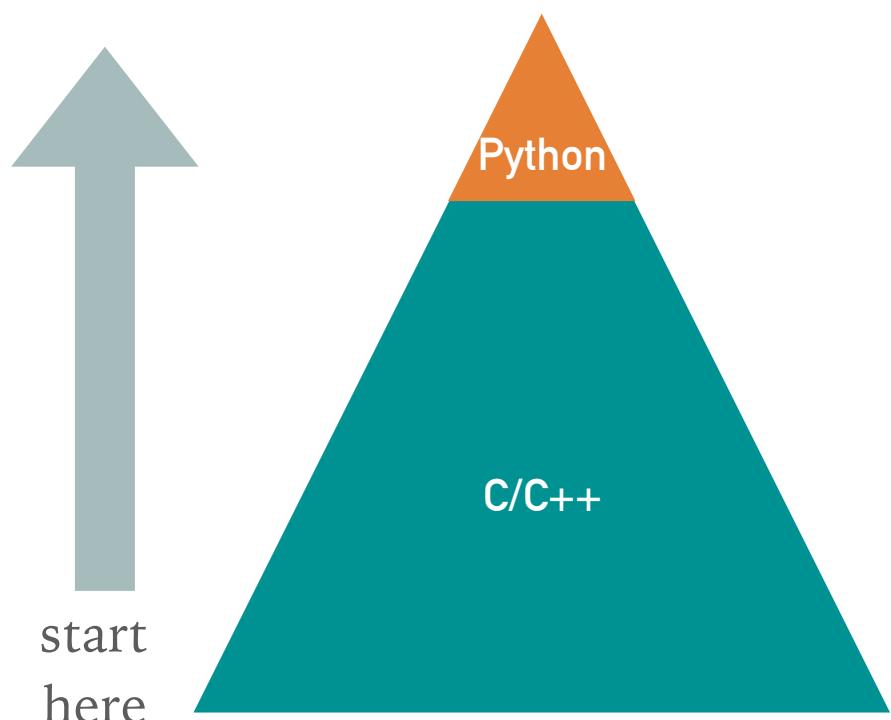


ctapipe

Building a Framework

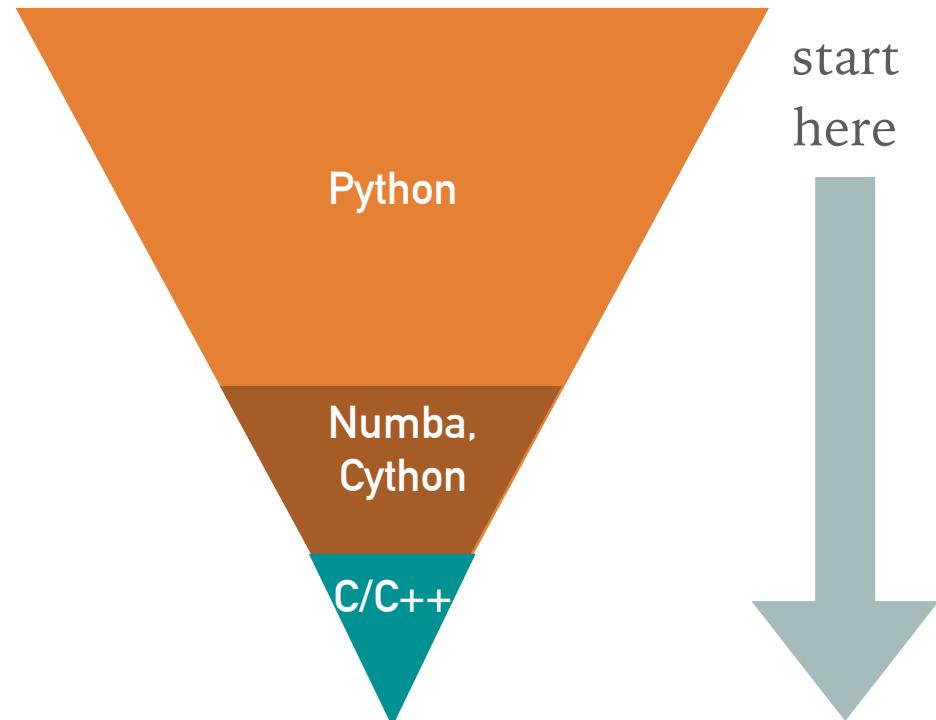


Bottom-Up approach



Most current frameworks did it this way (if they use python at all)

Top-Down approach

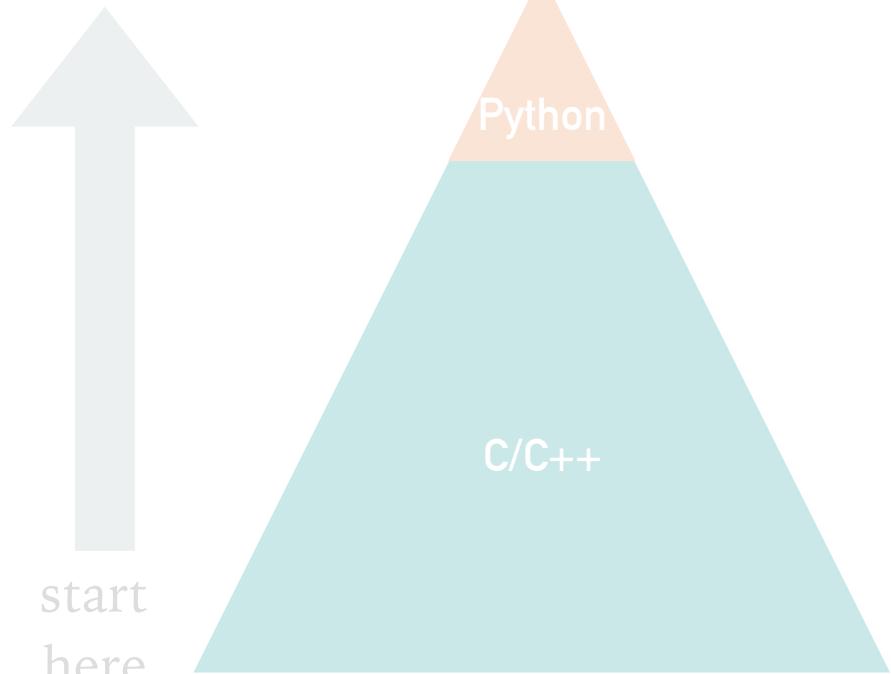


Our approach: start early with python and high-level API

Building a Framework

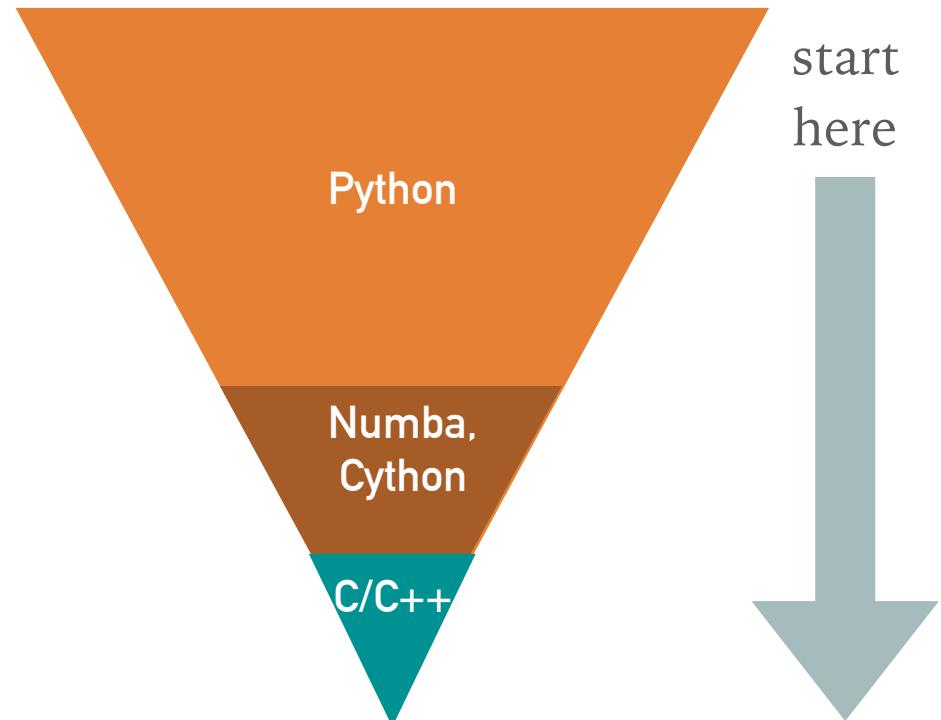


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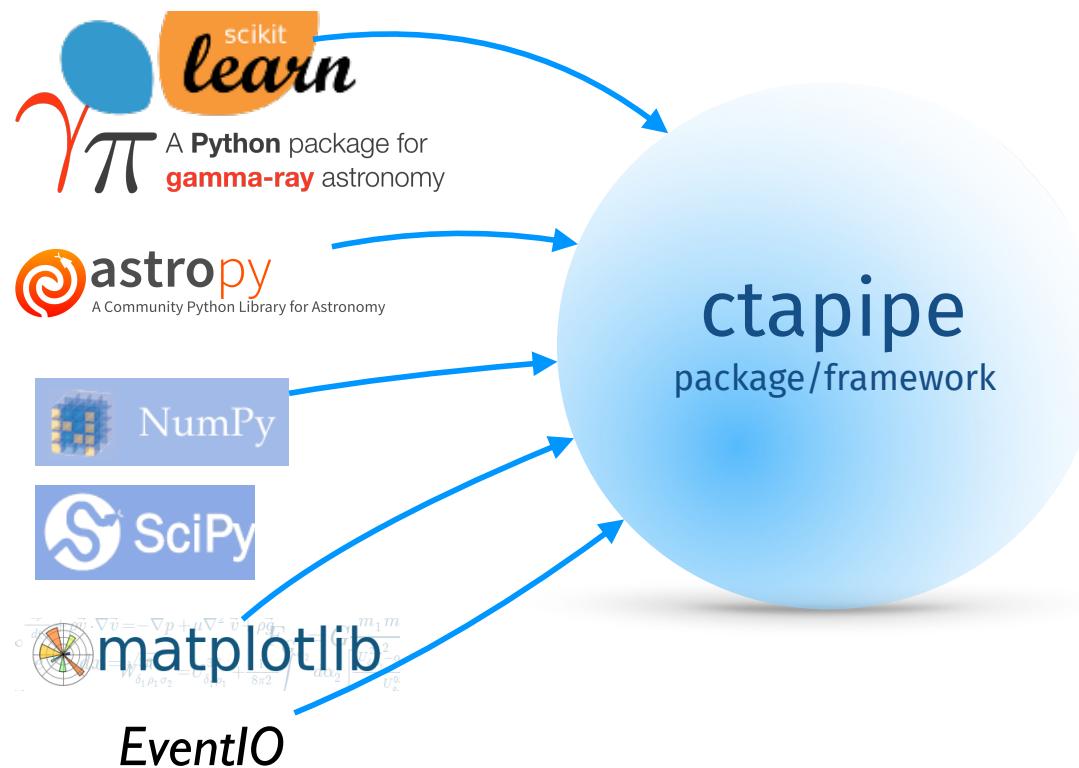


Our approach: start early with python and high-level API

common “core” package



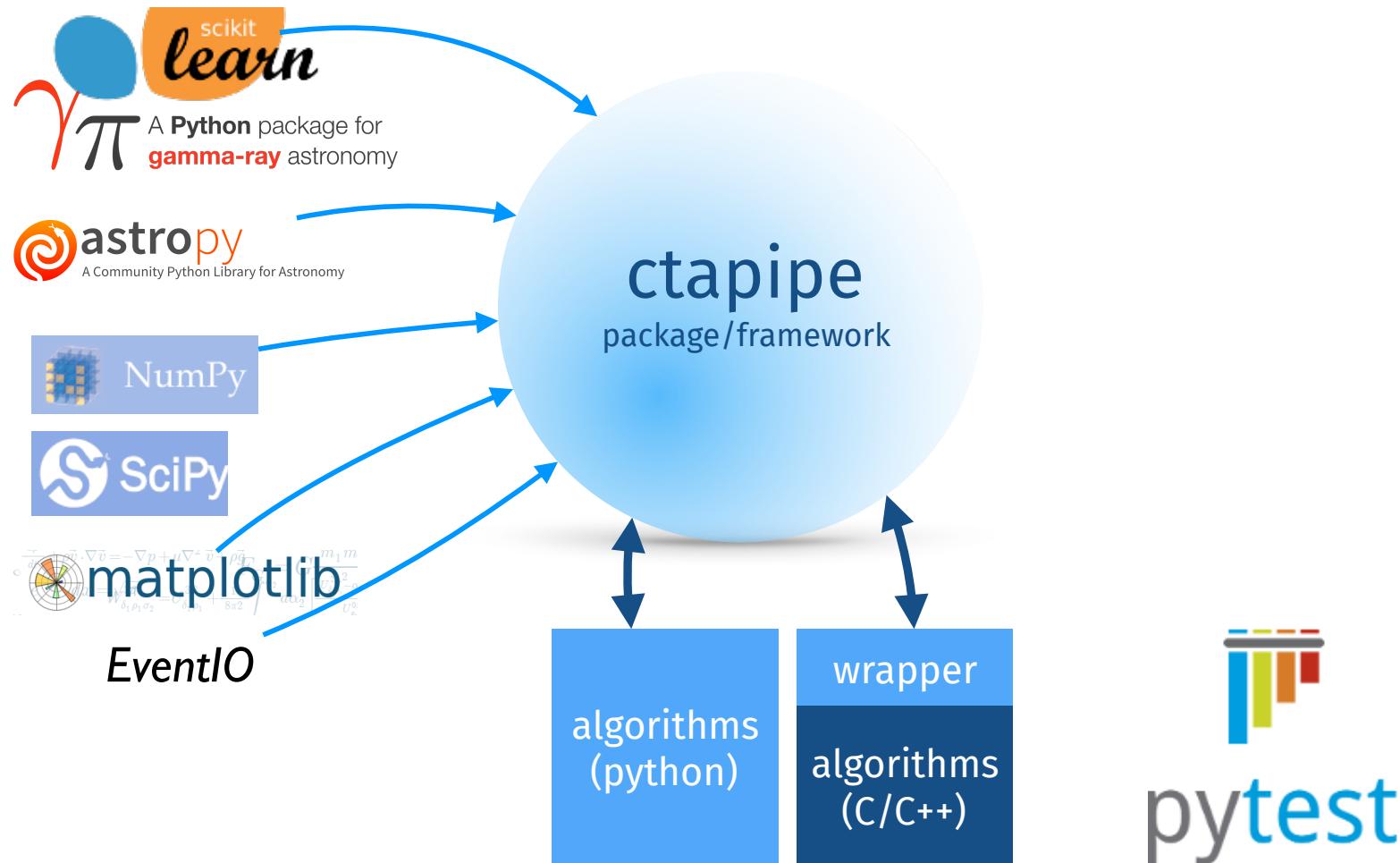
ctapipe will be **glue** between various components.
Provides common APIs and user interfaces
packaging, etc.



common “core” package



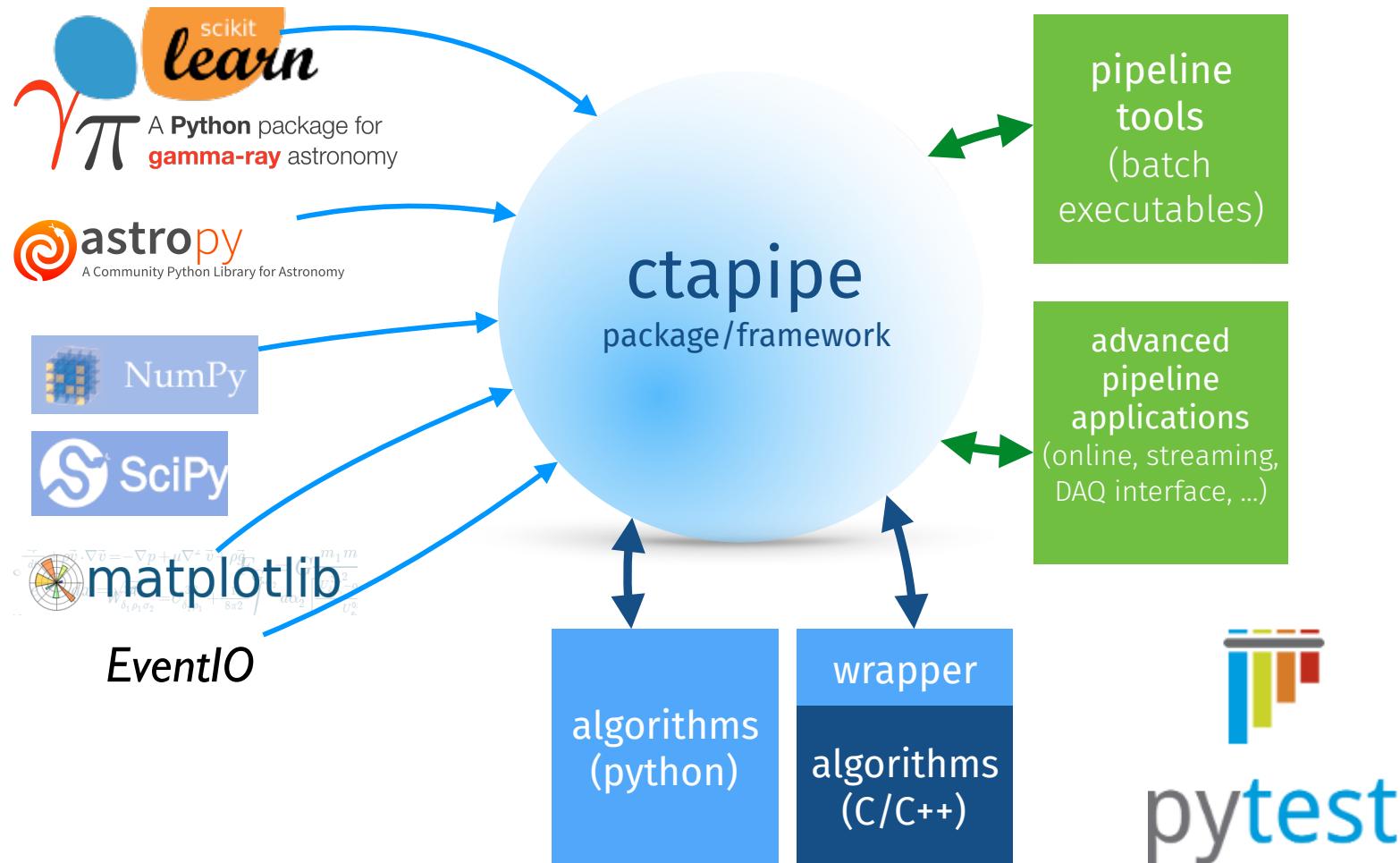
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common “core” package



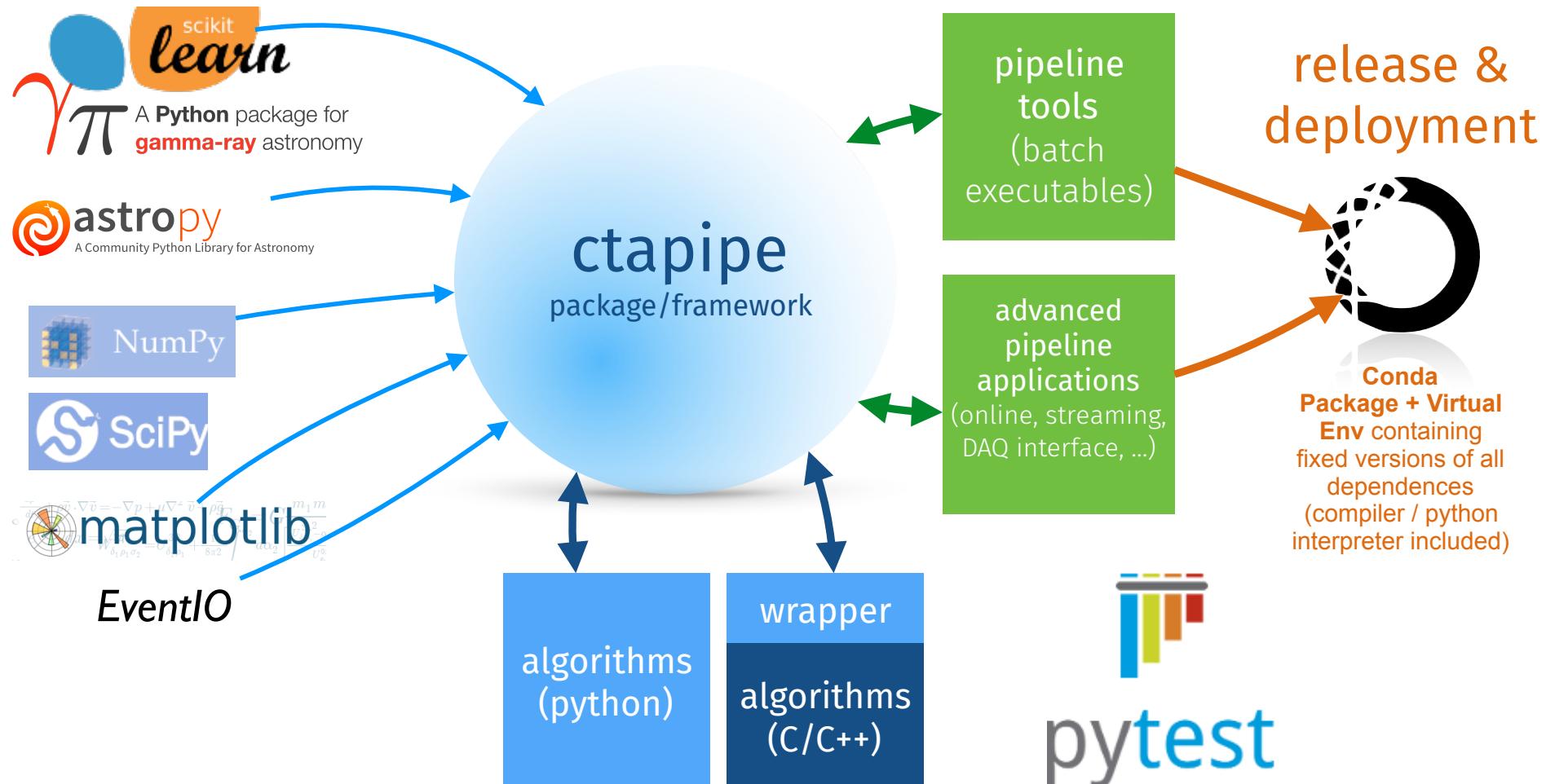
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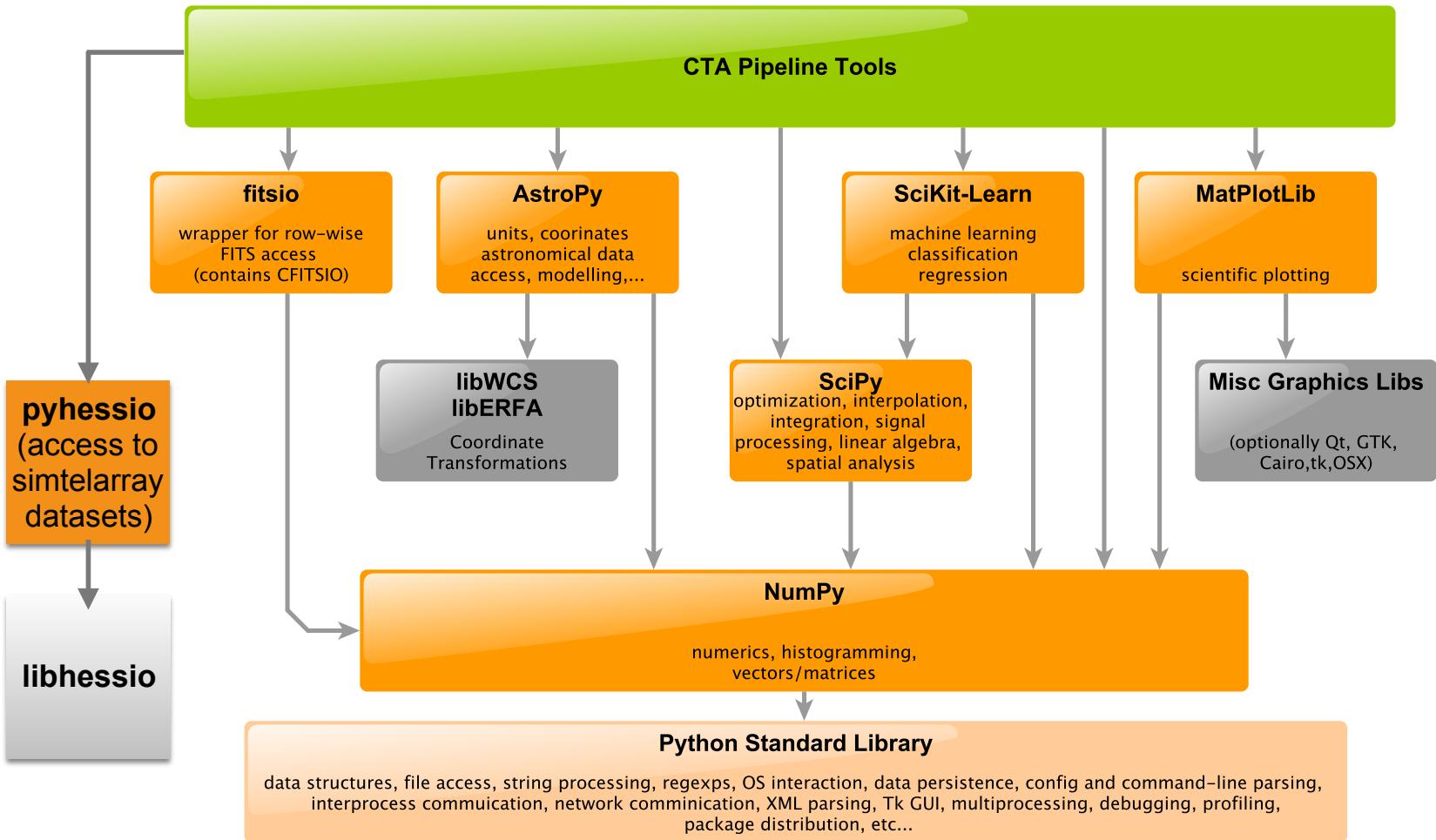




Design

- ctapipe is the “CTA Low-level Data Processing Pipeline Framework Prototype”
 - it means that it is a library for building pipelines, **not a pipeline**
- Libraries to perform the DL0 → DL3 analysis
- ctapipe is built upon the Scientific Python Stack, core dependencies are:
 - numpy
 - scipy
 - matplotlib
 - astropy
 - ...

Dependencies





Development

- ctapipe is developed as Open Source Software (Currently under MIT License)
 - <https://github.com/cta-observatory/ctapipe>
- We use the "Github-Workflow":
 - Few people (core developers) have write access to the main repository
 - Contributors fork the main repository and work on branches
 - Pull Requests are merged after Code Review and automatic execution of the test suite
- Early development stage ⇒ backwards-incompatible API changes might and will happen



What's there

- Reading simtel simulation files
- Simple calibration, cleaning and feature extraction functions
- Camera and Array plotting
- Coordinate frames and transformations
- Stereo-reconstruction using line intersections

For more info... Read the docs



ctapipe
0.1.dev1+g0fb3bcc

Search docs

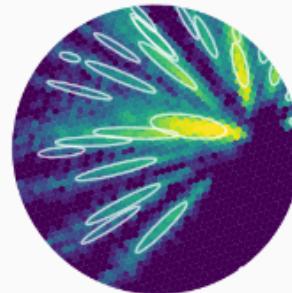
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Data Models
API Docs
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Prototype CTA Pipeline Framework (`ctapipe`)

version: 0.1.dev1+g0fb3bcc



ctapipe

What is cta pipe?

`ctapipe` is a framework for prototyping the low-level data processing algorithms for the Cherenkov Telescope Array.

For more info... Read the docs



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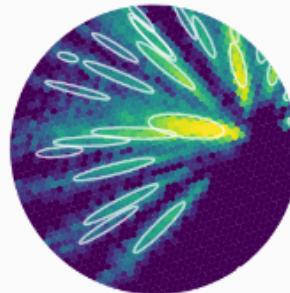
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Search docs

- Getting Started For Developers
 - Get the ctapipe software
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» Getting Started For Developers

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Getting Started For Developers

Warning

the following guide is used only if you want to *develop* the `ctapipe` package, if you just want to write code that uses it externally, you can install `ctapipe` as a conda package with
`conda install -c cta-observatory ctapipe`.

This guide assumes you are using the *Anaconda* python distribution, installed locally (*miniconda* should also work).

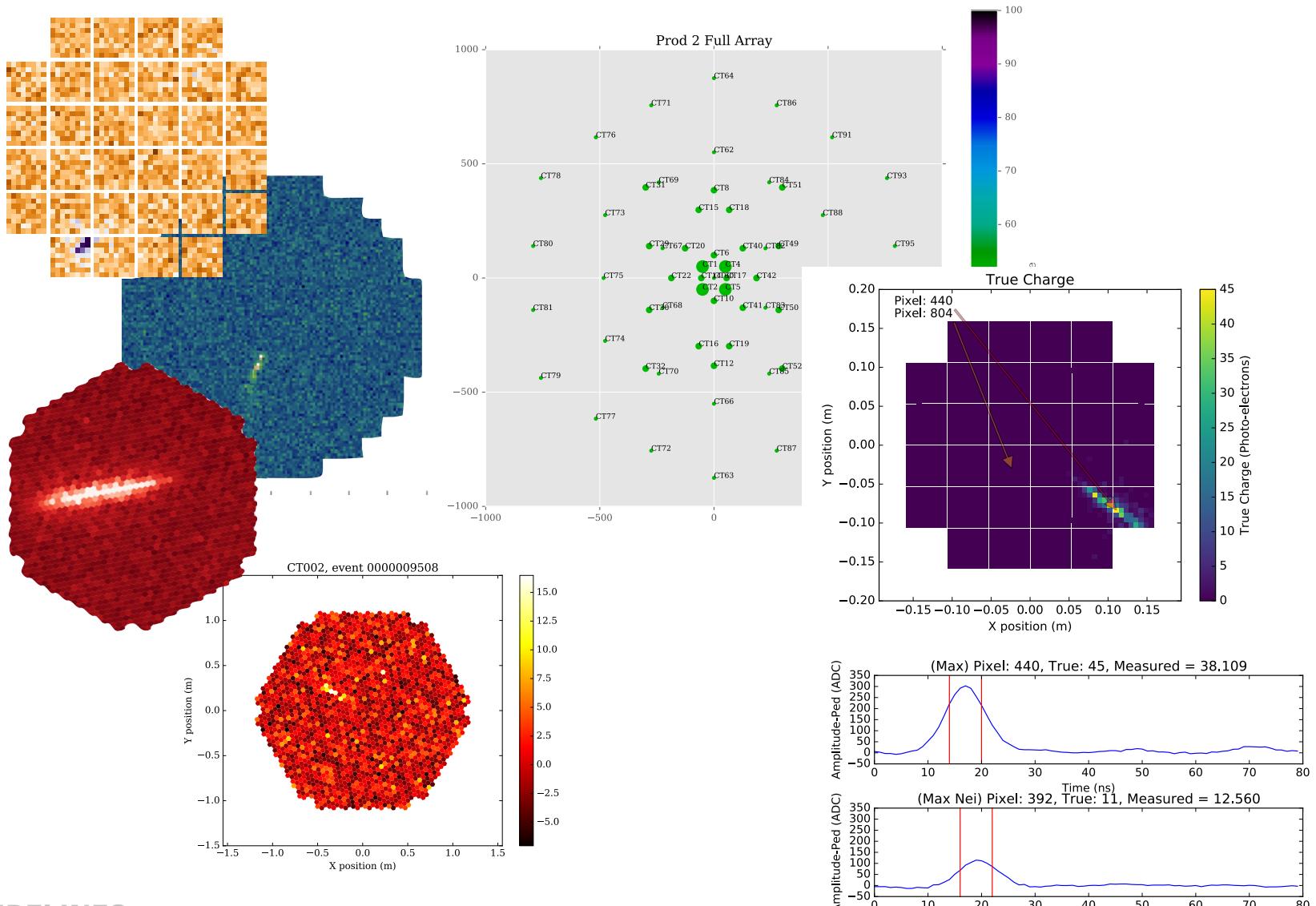
You can use *python 3.5* or above (we currently test on 3.5 and 3.6)

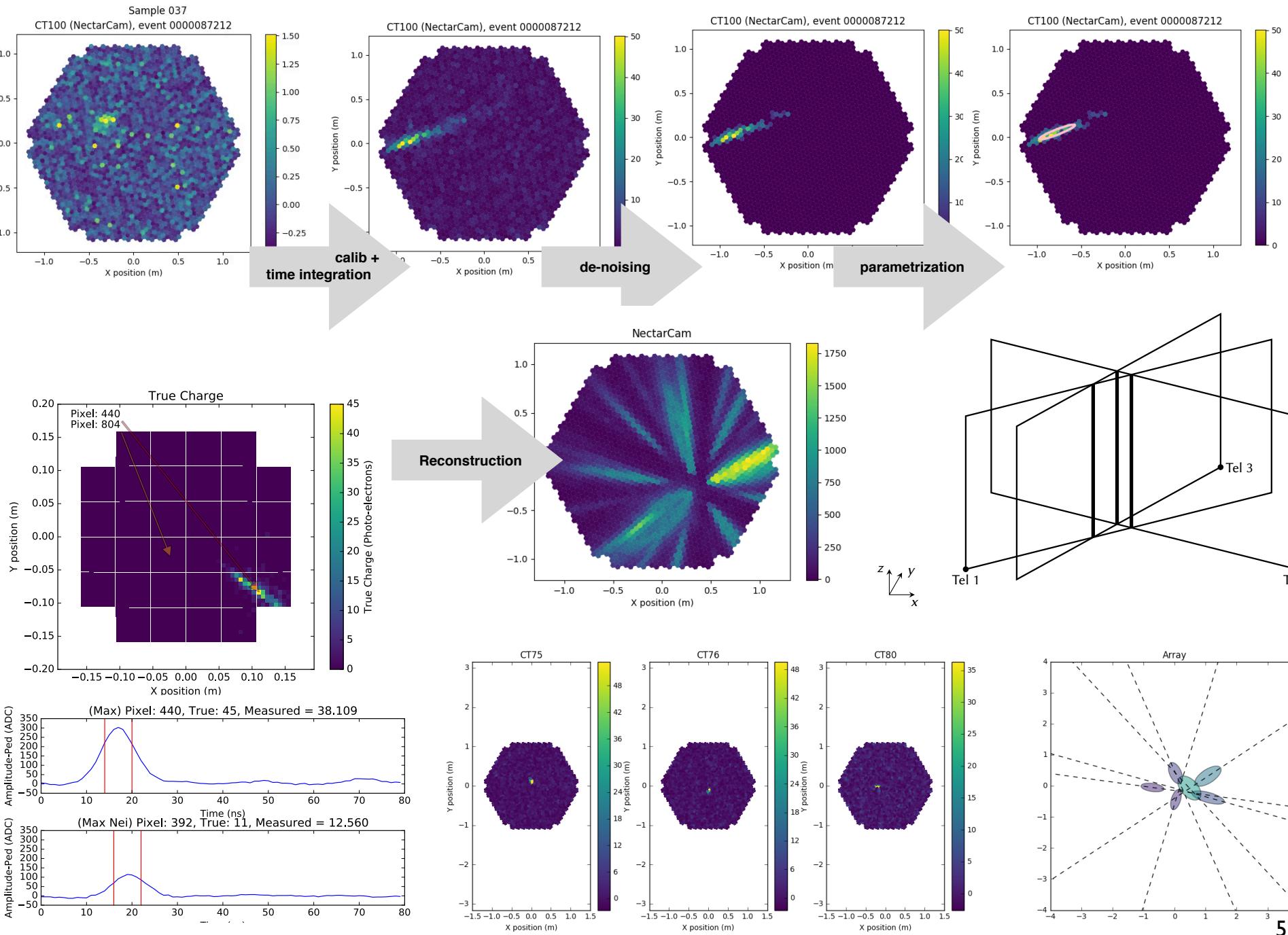
Get the ctapipe software

In order to checkout the software in such a way that you can read *and commit* changes, you need to [Fork and Clone](#) the main ctapipe repository (`cta-observatory/ctapipe`).

First, it's useful to make a directory where you have can check out cta GIT repos (this is optional - you can put it anywhere)

Visualization





CameraDisplay



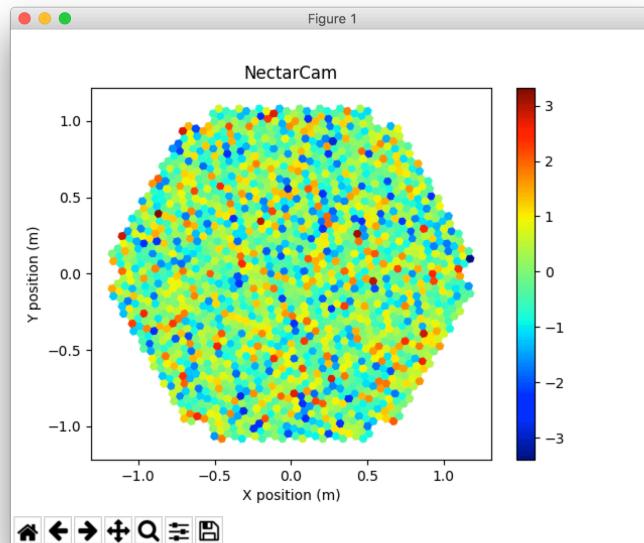
```
from ctapipe.visualization import CameraDisplay
from ctapipe.instrument import CameraGeometry

camera = CameraGeometry.from_name("NectarCam")
disp = CameraDisplay(camera)

disp.image = np.random.normal(size=camera.pix_x.shape)

disp.cmap = 'jet'
disp.add_colorbar()
```

- updating the image is fast, but constructing the display is slow
(reuse the same display when possible)



CameraDisplay
autoscale : bool
autoupdate : bool
axes
cmap
cmap : str
colorbar : NoneType
geom
image
image : NoneType
norm
norm : str
pixel_highlighting
pixels : NoneType, PatchCollection
add_colorbar()
add_ellipse()
clear_overlays()
enable_pixel_picker()
highlight_pixels()
on_pixel_clicked()
overlay_moments()
set_limits_minmax()
set_limits_percent()
show()
update()

EventSources

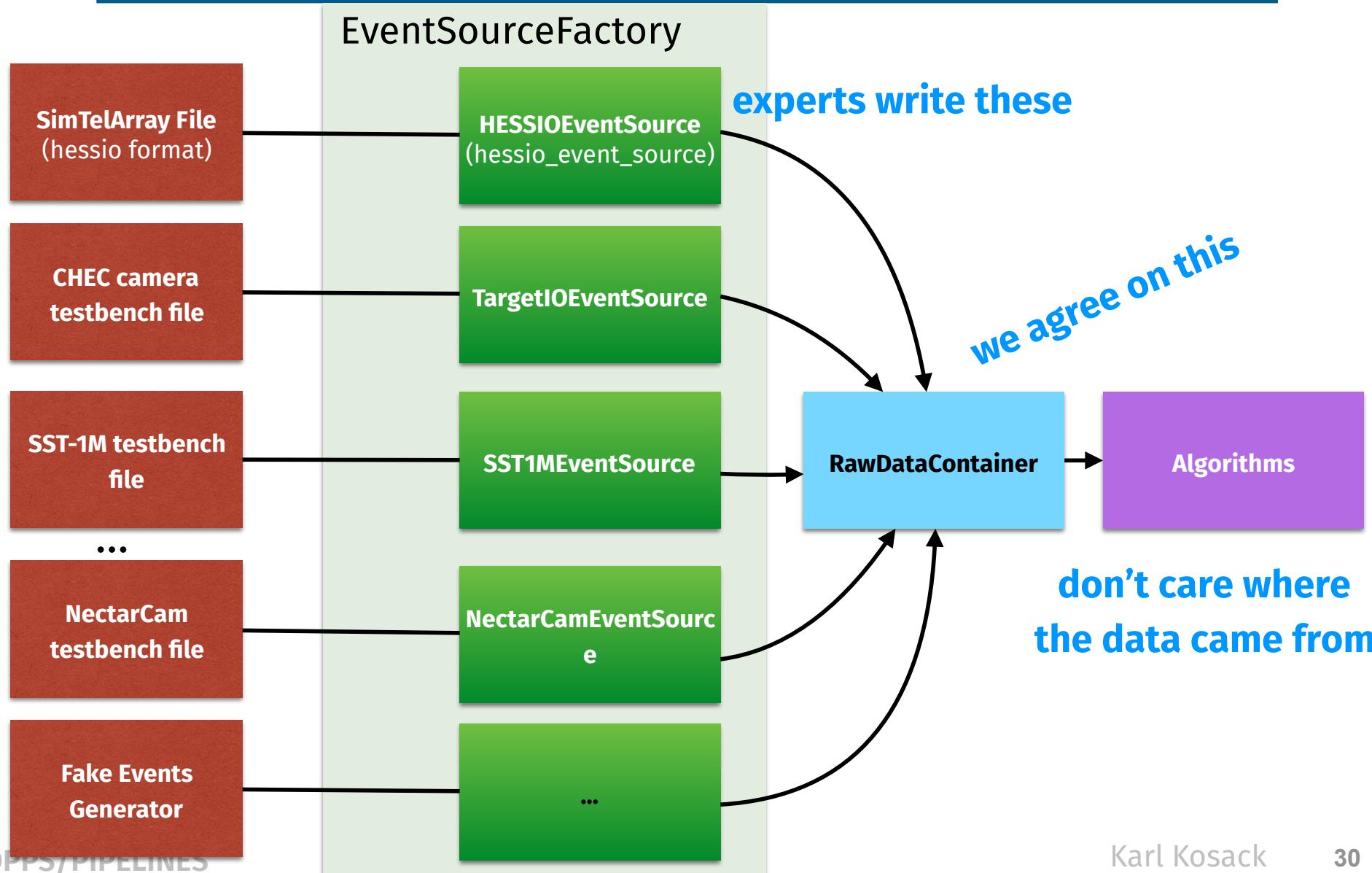
Interface:

- ▶ An *iterable* object, that emits DataContainers
- ▶ You can now loop more than once over the same EventSource (it will start from the beginning each time)
- ▶ **IMPORTANT:** by default the container returned is not a new container, but the same one with new information (no copy is made)
 - This is to prevent more than one event in memory at a time, for efficiency
 - if you want to store/cache several Containers, you must *deepcopy* them.

Useful options:

- ▶ **allowed_tels = [1,2,3,4]:**
 - specify a list of telescopes that are to be read, others will be skipped
 - *evt.r0.tels_with_data* will only show telescopes in this list
- ▶ **max_events=n:**
 - stop iteration after n events

EventSources





Hands on

- Run notebooks in the **notebooks/ctapipe** folder
 - [ctapipe_hands-on.ipynb](#)

main ▾ [intro-iact-analysis](#) / [notebooks](#) / [ctapipe](#) /

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	rlopezcoto	ctapipe notebook	62c6839 7 days ago
..			
	ctapipe_hands-on.ipynb	ctapipe notebook	7 days ago