

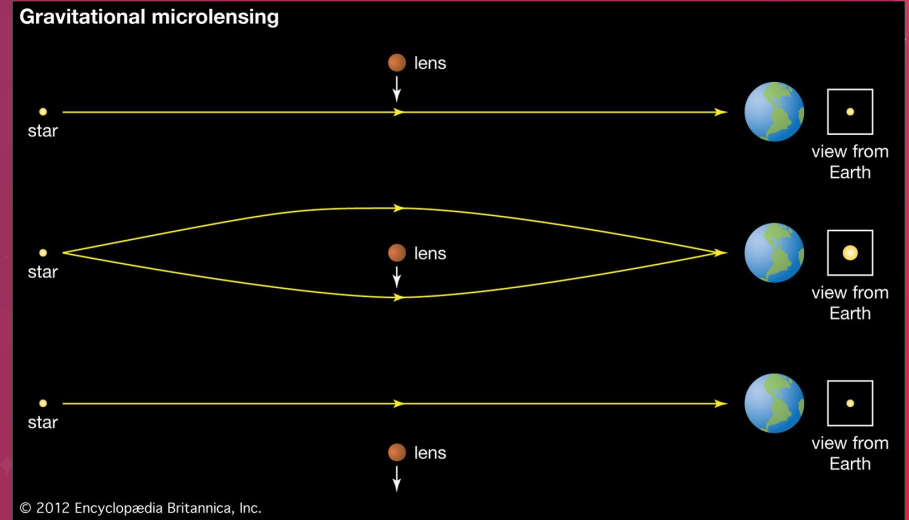
The background is a vibrant, stylized space scene. It features large, flowing, organic shapes in shades of red, purple, and blue. Scattered throughout are various celestial bodies: a large red planet with orange and yellow patterns in the top right, a yellow and orange striped planet in the bottom left, and several smaller blue and purple planets. Numerous small white stars and sparkles are also visible.

# Gravitational Microlensing

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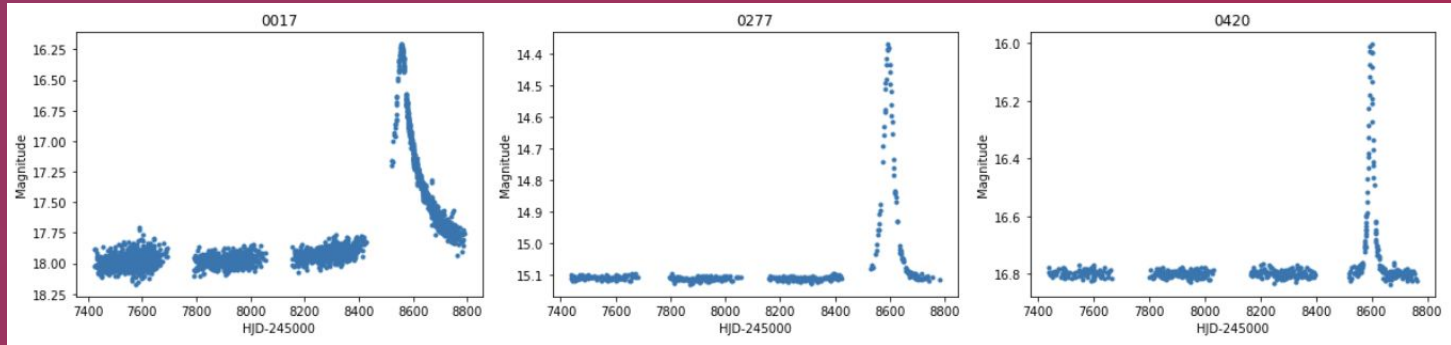
# Background

- Gravitational microlensing
  - The focusing of light as mass moves past star
- Took data from Optical Gravitational Lensing Experiment (OGLE)
  - Polish based
  - Long running observations of trying to find microlensing events
- Our Goal:
  - Fit and analyze Paczynski Light Curve Model to OGLE's data



# Data

- Collected from OGLE-IV project
  - Database of microlensing events from 2019
  - One of the largest sky variability surveys
- X-axis measures time, the Y-axis measures the magnitude, and the data points themselves represent a measure of brightness as a function of time
- Picked three microlensing events



# Methods

- Paczynski microlensing model

$$u(t) = \sqrt{u_{min}^2 + \left(\frac{t - t_0}{t_E}\right)^2}$$

Non-blended model

$$A(t) = \frac{u^2 + 2}{u\sqrt{u^2 + 4}}$$

Blended model

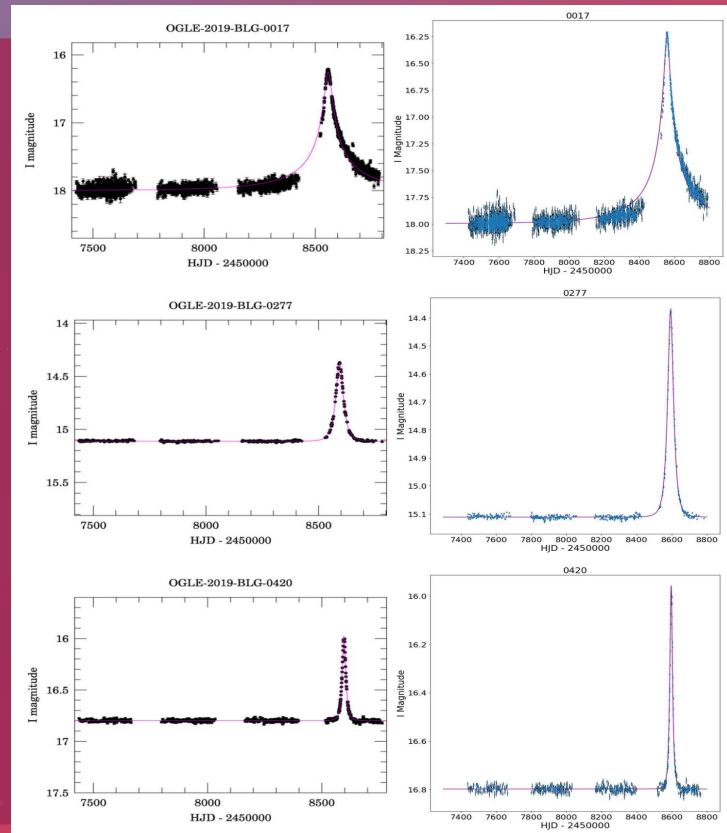
$$A(t) = f_{bl} \left( \frac{u^2 + 2}{u\sqrt{u^2 + 4}} - 1 \right) + 1$$

$$m_{mod}(t) = m_{bl} - 2.5 \log_{10}(A(t))$$

- $u(t)$ : angular separation of source and lens
- $u_{min}$ : minimum separation
- $A(t)$ : Magnification of light from lensing
- $m_{bl}$ : intensity of light with  $A = 1$
- $f_{bl}$ : fraction of incoming light that is from the source
- $m_{mod}(t)$ : intensity of light predicted by our model
- Found parameters by optimizing a  $\chi^2$  function for both models
- Used Markov Chain Monte Carlo (MCMC) simulation to analyze error in parameters

# Results

- Both model's lined up well with the data
- We were able to mimic OGLE's results
  - With aesthetic differences
- Blended vs Unblended parameter fitting

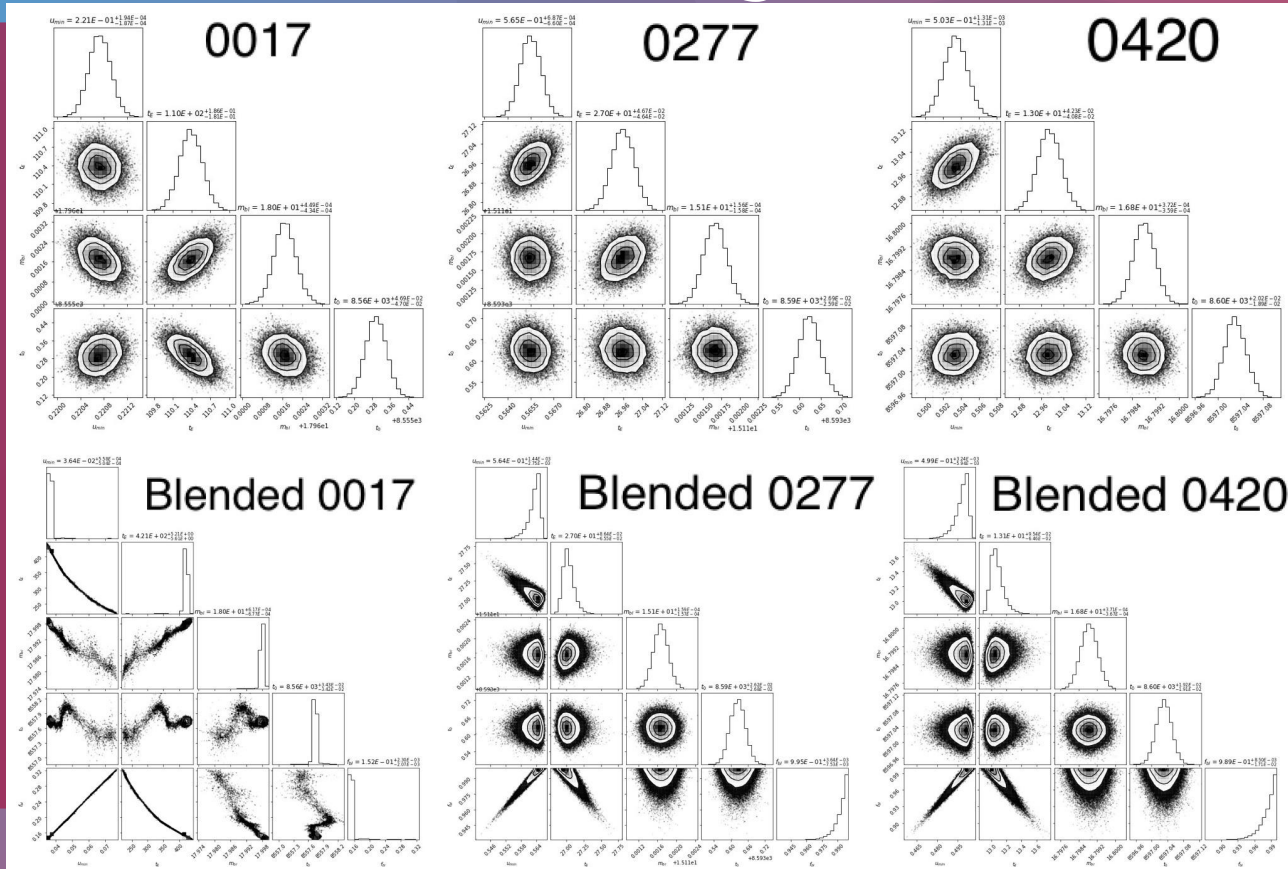


# Results

| Parameters↓ Event→    | 0017     | 0277     | 0420     | OGLE 0017 | OGLE 0277 | OGLE 0420 |
|-----------------------|----------|----------|----------|-----------|-----------|-----------|
| $u_{min}$             | 0.221    | 0.565    | 0.504    | 0.36      | 0.565     | 0.503     |
| $t_E$ [days]          | 110.381  | 26.958   | 12.989   | 421.250   | 26.595    | 12.989    |
| $t_0$ [HJD - 2450000] | 8555.296 | 8593.624 | 8597.027 | 85557.734 | 8593.624  | 8597.030  |
| $m_{bl}$              | 17.962   | 15.112   | 16.795   | 17.998    | 15.112    | 16.799    |

| Parameters[Blended]↓ Event→ | 0017     | 0277     | 0420     | OGLE 0017 | OGLE 0277 | OGLE 0420 |
|-----------------------------|----------|----------|----------|-----------|-----------|-----------|
| $u_{min}$                   | 0.036    | 0.565    | 0.503    | 0.036     | 0.565     | 0.503     |
| $t_E$ [days]                | 421.242  | 26.959   | 12.989   | 421.250   | 26.595    | 12.989    |
| $t_0$ [HJD - 2450000]       | 8557.734 | 8593.624 | 8597.030 | 85557.734 | 8593.624  | 8597.030  |
| $m_{bl}$                    | 17.998   | 15.112   | 16.799   | 17.998    | 15.112    | 16.799    |
| $f_{bl}$                    | 0.152    | 1.0      | 1.0      | 0.152     | 1.0       | 1.0       |

# Error Analysis





# Conclusion

- Limitations:
  - Cannot determine mass of lens
    - Need distances (us to lens, us to source, and lens to source)

$$M = \frac{c^2}{4G} \frac{D_l D_s}{D_{ls}} |\theta_+ \theta_-|$$

- Limitations → Further research, observation, and experimentation



# Questions?

