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*Mathématiques et Sciences et Technologies
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Spécialité : *Mathématiques et Interactions*

Par

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Méthodes d'analyse asymptotique et d'approximation numérique

Problèmes d'évolution multi-échelles de type oscillatoires ou dissipatifs

Thèse présentée et soutenue à « Lieu », le « date »

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I would like to thank. my parents..

J'adresse également toute ma reconnaissance à

....

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INTRODUCTION

Le développement de modèles mathématiques en sciences naturelles bénéficie d'avancées mathématiques qui permettent de vérifier le caractère bien posé des équations, ou le bon comportement des solutions. Une classe de modèles très prisés depuis quelques dizaines d'années sont les modèles multi-échelles, dont l'étude concerne principalement sur les modèles double-échelle. Dans ces modèles, on distingue deux dynamiques : une d'échelle caractéristique « rapide » ε et l'autre d'échelle 1. Dans ce manuscrit, on s'intéresse essentiellement à une sous-classe de ces modèles : ceux dont la dynamique rapide est une relaxation. Pour nos résultats, on s'inspire de méthodes développées pour les problèmes dont la dynamique rapide est oscillatoire.

Ces systèmes à relaxation rapide apparaissent en physique dans un cadre fonctionnel de modèles cinétiques [BGK54; LM08] ou dans certains systèmes dérivés de problèmes non-linéaires [JX95]. On les observe également en dynamique des populations, e.g. dans [GHM94; AP96; SAAP00; CCS18]. Les systèmes hautement oscillants sont également fréquents en physique. Certains exemples sont présentés dans l'ouvrage de référence [HLW06, Chap. I], comme le modèle de Hénon-Heiles [HH64] dans un contexte de mouvements célestes, ou le problème de Fermi-Pasta-Ulam-Tsingou [For92] en théorie du chaos. On peut aussi étudier certains phénomènes de dynamique quantique non-linéaires, tels que le modèle de Klein-Gordon [BD12], l'équation de Schrödinger [GV11] ou le modèle de Wigner en milieu périodique [CJL17; MS11].

La simulation de tels systèmes présente des défis particuliers, qui peuvent se ramener aux concepts de base des méthodes numériques de *stabilité* et de *convergence*. La stabilité consiste à déterminer une condition pour que la solution numérique soit bien définie, qu'elle ne diverge pas. La convergence trace un lien direct entre le coût de calcul et la précision de l'approximation numérique. Dans ce chapitre d'introduction, on introduit mathématiquement les modèles qui nous concernent, et on fait une brève description de leur comportement. À cet égard, on introduit deux exemples de systèmes « jouet » que nous suivrons tout au long du chapitre. Ensuite, on illustre les limitations des méthodes numériques de l'état de l'art, en introduisant certains concepts de convergence liés à la présence du paramètre ε . Enfin, on présente brièvement la contribution de ce travail de

thèse, et on annonce le plan pour la suite du manuscrit.

Introduction mathématique

Ce manuscrit se concentre sur des problèmes de Cauchy de la forme

$$\partial_t u^\varepsilon = -\frac{1}{\varepsilon} A u^\varepsilon + f(u^\varepsilon), \quad u^\varepsilon(0) = u_0 \quad (1)$$

où t évolue dans l'intervalle $[0, 1]$. On considère ce problème dans un Banach $(E, |\cdot|)$ de dimension finie $d > 0$, avec $A : E \rightarrow E$ un opérateur linéaire et $f : E \rightarrow E$ un champ de vecteurs régulier. On se concentre sur le cas où A est diagonal avec des valeurs propres positives entières. Souvent, on séparera le problème entre le noyau et l'image de A , pour obtenir un problème de la forme

$$\begin{cases} \partial_t x^\varepsilon = a(x^\varepsilon, z^\varepsilon), & x^\varepsilon(0) = x_0, \\ \partial_t z^\varepsilon = -\frac{1}{\varepsilon} \Lambda z^\varepsilon + b(x^\varepsilon, z^\varepsilon), & z^\varepsilon(0) = z_0 \end{cases} \quad (2a)$$

$$(2b)$$

avec $u = \begin{pmatrix} x \\ z \end{pmatrix}$, $A = \begin{pmatrix} 0 & 0 \\ 0 & \Lambda \end{pmatrix}$ et $f = \begin{pmatrix} a \\ b \end{pmatrix}$. En général, on omettra l'exposant ε . Le lecteur peut supposer que, sauf mention contraire, toutes les variables dépendent de ε . D'ailleurs, on peut supposer que le champ de vecteurs f évolue de manière régulière en fonction de ε sans impacter les résultats.

Dans cette section on décrit et illustre le comportement de la solution $(x^\varepsilon, z^\varepsilon)$ à travers deux exemples. En particulier, on énonce le théorème de variété centrale, qui décrit le comportement de la solution en temps long, et on présente rapidement une méthode pour calculer cette variété centrale. On introduit ensuite quelques hypothèses qui permettront de citer des résultats d'estimation numérique rigoureux dans la prochaine section.

Quelques observations

Pour démarrer, considérons le problème jouet suivant

$$\partial_t z(t) = -\frac{1}{\varepsilon} z(t) + \sin(t), \quad z(0) = 1, \quad (3)$$

qui peut être transformé en un problème de la forme (2) en posant $x(t) = t$, soit $\partial_t x = 1$, $x(0) = 0$. Ce problème peut être obtenu à partir de

$$\partial_t y(t) = -\frac{1}{\varepsilon} (y(t) - \cos(t)), \quad y(0) = 0,$$

en posant $z(t) = y(t) - \cos(t)$. C'est un problème de référence pour l'introduction aux systèmes raides : c'est le premier exemple présenté dans [HW96]. La solution exacte se calcule sans difficulté en intégrant $\partial_t [e^{t/\varepsilon} z(t)]$, ce qui donne

$$z(t) = e^{-t/\varepsilon} \left(1 + \frac{\varepsilon^2}{1 + \varepsilon^2} \right) + \frac{\varepsilon}{1 + \varepsilon^2} (\sin(t) - \varepsilon \cos(t)). \quad (4)$$

On observe que la solution comporte deux parties de natures différentes, la phase transitoire (en $e^{-t/\varepsilon}$) et la variété centrale (en t) de taille ε . Ces deux phases apparaissent clairement sur la figure ci-dessous où on a tracé la solution et la variété centrale associée pour trois valeurs de ε . En effet, le temps d'atteinte de la variété semble proportionnel à ε pour des petites valeurs.

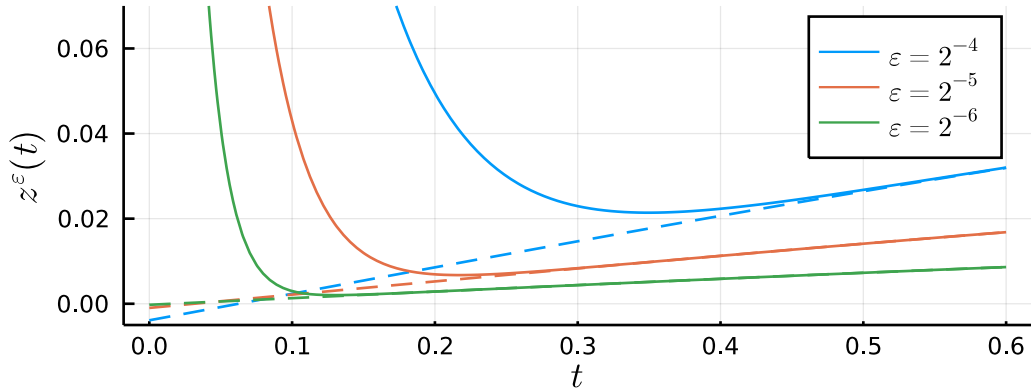


FIGURE 1 – Comportement de la solution (4) pour différentes valeurs de ε , avec chaque variété centrale associée en pointillés.

En fait, dans le cas d'un système de la forme (2), la dynamique en temps long est déterminé entièrement par la variable x . Il y a une réduction de dimension, qui est traduite dans le théorème de variété centrale.

Théorème (Variété centrale, [Car82]). *En supposant que le champ f est dans $C^\infty(E)$ et que la donnée initiale $u_0 \in E$ du problème (1) est bornée, il existe un temps final $T > 0$ et*

un paramètre limite ε_0 tels que le problème (1) est bien posé sur $[0, T]$. En outre, il existe une application régulière $x \mapsto z = \varepsilon h^\varepsilon(x)$, un taux $\mu > 0$ et une constante $c > 0$ tels que pour tout $t \in [0, T]$,

$$|z(t) - \varepsilon h^\varepsilon(x(t))| \leq ce^{-\mu t/\varepsilon}.$$

En outre, quitte à modifier c , il existe une donnée initiale « de l'ombre » x_0^ε telle que

$$|x(t) - \varphi_t^\varepsilon(x_0^\varepsilon)| \leq c\varepsilon e^{-\mu t/\varepsilon},$$

où φ_t^ε est le t -flot associé au champ de vecteurs $x \mapsto a(x, \varepsilon h^\varepsilon(x))$. On appelle l'ensemble des $(x, \varepsilon h^\varepsilon(x))$ la variété centrale, qui est stable et « attire » la solution.

Par exemple dans le cas (3), le morphisme de variété centrale $\varepsilon h^\varepsilon$ est donné par

$$h^\varepsilon(x) = \frac{1}{1 + \varepsilon^2} (\sin(x) - \varepsilon \cos(x)). \quad (5)$$

Il est impossible d'espérer calculer le morphisme de variété centrale explicitement en général. On peut néanmoins appliquer une méthode de point fixe en remarquant qu'en temps long, $z \approx \varepsilon h^\varepsilon(x)$, d'où

$$\partial_t z \approx \varepsilon \partial_x h^\varepsilon(x) \cdot \partial_t x \approx \varepsilon \partial_x h^\varepsilon(x) \cdot a(x, \varepsilon h^\varepsilon(x)).$$

Ainsi on peut poser $h^{[0]} = 0$ et itérer de manière explicite

$$\varepsilon \partial_x h^{[n]}(x) \cdot a(x, \varepsilon h^{[n]}(x)) = -h^{[n+1]}(x) + b(x, \varepsilon h^{[n]}(x)). \quad (6)$$

Le calcul de la donnée initiale de l'ombre x_0^ε n'est en revanche pas si simple. La calculer demande de faire des développements sur l'intégralité du système et demande un certain investissement. Dans [CCS16], les auteurs font appel à des B-séries¹ pour obtenir un modèle asymptotique sur l'intégralité du problème (1), à partir duquel ils trouvent en particulier cette donnée initiale modifiée, mais les calculs sont bien plus compliqués que (6).

Remarque. *Le phénomène de donnée initiale de l'ombre n'est cependant pas visible sur*

1. Les B-séries sont des séries formelles qui décrivent les solutions d'EDO. Cet outil, souvent utilisé pour développer des schémas numériques, est présenté en détails dans [HLW06, Chap. III] ou de manière plus concise dans [CHV10]. Malgré une apparence encombrante, les B-séries possèdent une structure algébrique élégante basée sur des opérations sur les arbres.

cet exemple, puisque la variable x ne dépend pas de la dynamique sur z (pour rappel, on a $x(t) = t$). L'interdépendance entre x et z apparaîtra dans le cas test de la section suivante.

On remarque deux caractéristiques importantes de la méthode d'approximation de h^ε : elle nécessite de calculer une dérivée, ce qui demande du calcul symbolique potentiellement coûteux, et la convergence de la méthode n'est pas assurée. Même dans l'exemple jouet (3), cette méthode génère

$$h^{[n]}(x) = R_n(\varepsilon) \sin(x) + \varepsilon R_{n-1}(\varepsilon) \cos(x)$$

avec $R_n(\varepsilon) = \sum_{k=0}^{\lfloor n/2 \rfloor} \varepsilon^{2k}$ et par convention $R_{-1} = 0$. En d'autres termes, on construit le développement en série entière en ε de la partie lente dans (4), i.e. $h^{[1]}(x) = \sin(x)$, $h^{[2]}(x) = \sin(x) + \varepsilon \cos(x)$ etc. Ainsi, le développement n'est convergent que pour $\varepsilon < 1$, alors que la solution $z(t)$ converge vers la variété $\varepsilon h^\varepsilon(t)$ définie par (5) pour tout $\varepsilon > 0$. Cette limitation apparaîtra couramment au cours de ce manuscrit sous le format plus contraignant

$$\varepsilon \leq \frac{\varepsilon_0}{n+1},$$

qu'on peut remarquer par exemple avec $b(x, z) = (1+x)^{-1}$ dont la taille des dérivées évolue de manière factorielle avec l'ordre de dérivation.

Définitions et hypothèses

On considère le problème (1) en dimension finie d , et on suppose que la donnée initiale u_0 appartient à un ensemble \mathcal{U}_0 bien choisi de sorte qu'en tout temps, la solution appartient à un ensemble borné \mathcal{K} de la forme

$$\mathcal{K} := \left\{ \begin{pmatrix} x \\ z \end{pmatrix} \in E, \text{ s.t. } x \in \mathcal{K}_x, |z| \leq \rho \right\}$$

avec \mathcal{K}_x borné et $\rho > 0$. Grâce au théorème de variété centrale, ces hypothèses peuvent être considérées valides pour tout $\varepsilon \in (0, \varepsilon_0]$ pour un certain $\varepsilon_0 > 0$ fixé.

En outre, on suppose que le champ de vecteurs $u \mapsto f(u)$ est de classe C^∞ sur \mathcal{K}_R pour un certain rayon $R > 0$, où

$$\mathcal{K}_R = \{u \in \mathbb{R}^d, d(u, \mathcal{K}) \leq R\}$$

avec $d(u, \mathcal{K}) = \inf_{v \in \mathcal{K}} |u - v|$ la distance de u à \mathcal{K} . Cette hypothèse permet d'assurer que les méthodes numériques se comportent bien au voisinage de la solution, sans nécessiter d'être exact pour autant.

Paradigme numérique

En général, on suppose $\varepsilon \ll 1$, et donc le système (1) comporte une dynamique *rapide* par rapport au temps d'étude. À cet égard, des méthodes d'*analyse asymptotique* ont été développées, c'est-à-dire des méthodes qui permettent de caractériser le système dans cette limite ε « petit », en général en découplant ces deux dynamiques. Pour les problèmes hautement oscillants, trois exemples particulièrement célèbres sont les méthodes d'homogénéisation [GM03], de moyennisation [Per69 ; SVM07 ; LM88] et de formes normales [Mur06 ; Bam03]. Pour les problèmes à relaxation rapide, la littérature est moins fournie. Qu'il s'agisse de calculer la variété centrale comme précédemment ou d'un développement de Chapman-Enskog [SBD86 ; Deg04 ; CCLM15], la phase transitoire n'est pas calculée.

Plus récemment dans [CCS16], les auteurs capturent aussi la phase transitoire, mais la méthode n'est valide que dans la limite $\varepsilon \rightarrow 0$. Dans cette section, on étudie l'application de méthodes numériques « standards » de l'état de l'art, et on observe le comportement de l'erreur numérique non seulement en fonction du pas de temps Δt , mais aussi en fonction du paramètre ε .

On commence par décrire ce qu'on entend par « méthode numérique » et le contexte dans lequel on va les étudier. Ensuite, on présente trois méthodes d'ordre 2 de l'état de l'art : le splitting de Strang, un schéma IMEX-BDF et une méthode de Runge-Kutta exponentielle.

Mise en place de la résolution

Pour étudier le comportement des schémas numériques sur les problèmes de la forme (1), on considère l'exemple jouet suivant

$$\begin{cases} \partial_t v_1 = v_2, & v_1(0) = 1, \\ \partial_t v_2 = -\frac{1}{\varepsilon}(v_1 + v_2), & v_2(0) = 0. \end{cases} \quad (7a)$$

$$\quad (7b)$$

Cet exemple ressemble à certains problèmes hyperboliques avec relaxation, et sa linéarité le rend simple à étudier. Il prend facilement la forme (2) en posant par exemple $x = v_1$ et $z = v_1 + v_2$, ce qui donne

$$\begin{cases} \partial_t x = -x + z, & x(0) = 1, \\ \partial_t z = -\frac{1}{\varepsilon}z - x + z, & z(0) = 1. \end{cases} \quad (8a)$$

$$\quad (8b)$$

Ce problème est linéaire et se diagonalise sans difficulté pour $\varepsilon < 1/4$, ce qui génère

$$\tilde{u} = \underbrace{\begin{pmatrix} -1 & 1 - r_\varepsilon \\ \varepsilon & 1 - \varepsilon - \varepsilon r_\varepsilon \end{pmatrix}}_P \begin{pmatrix} x \\ z \end{pmatrix}, \quad \text{tel que} \quad \partial_t \tilde{u} = \begin{pmatrix} -r_\varepsilon & 0 \\ 0 & -\frac{1}{\varepsilon} + r_\varepsilon \end{pmatrix} \tilde{u}$$

avec $r_\varepsilon = \frac{1}{2\varepsilon}(1 - \sqrt{1 - 4\varepsilon})$. On obtient ainsi une expression explicite pour $u = \begin{pmatrix} x \\ z \end{pmatrix}$ de la forme

$$u(t) = P^{-1} \begin{pmatrix} e^{-tr_\varepsilon} & 0 \\ 0 & e^{-t/\varepsilon} e^{tr_\varepsilon} \end{pmatrix} Pu(0)$$

$$\text{où } P^{-1} = \frac{1}{\sqrt{1-4\varepsilon}} \begin{pmatrix} -1 + \varepsilon + \varepsilon r_\varepsilon & 1 - r_\varepsilon \\ \varepsilon & 1 \end{pmatrix}.$$

Remarque. On voit bien dans la définition de r_ε que le problème change de nature entre $\varepsilon \leq 1/4$ et $\varepsilon > 1/4$. En effet, dans le premier cas le système est purement dissipatif, alors que dans le second, des oscillations apparaissent. Cette singularité apparaît également dans la matrice de changement de variable, dont le déterminant vaut $-\sqrt{1 - 4\varepsilon}$.

En pratique, on ne sait pas résoudre tous les systèmes de la forme (1) de manière exacte. On va donc appliquer des méthodes d'*approximation numérique* pour calculer une solution approchée. Plus spécifiquement, on va implémenter certains *schémas numériques* et observer la qualité d'approximation sur l'exemple (8). Définissons ce qu'on entend par « schéma numérique ».

On démarre par considérer une discrétisation de l'intervalle temporel $[0, T]$, c'est-à-dire qu'au lieu de considérer cet objet comme continu, on le considère comme une suite de $N + 1$ points $(t_n)_{0 \leq n \leq N}$ avec $N \geq 1$. On choisit de se restreindre à une discrétisation *uniforme*, c'est-à-dire que l'intervalle de temps $[0, T]$ est divisé en N intervalles de taille égale notée Δt .

$$0 \quad \Delta t \quad 2\Delta t \quad \dots \quad T - \Delta t \quad T$$

De manière équivalente, les points de séparation $(t_n)_{0 \leq n \leq N}$ sont donnés par

$$t_{n+1} = t_n + \Delta t \quad \text{avec} \quad t_0 = 0 \quad \text{et} \quad \Delta t = \frac{T}{N},$$

ou encore $t_n = \frac{n}{N}T$. À cette discrétisation, on peut associer une *approximation* $(u_n)_{0 \leq n \leq N}$ telle que $u_n \approx u(t_n)$. On peut voir un exemple d'une telle approximation en Figure 2, où les points carrés sont une approximation² des points ronds.

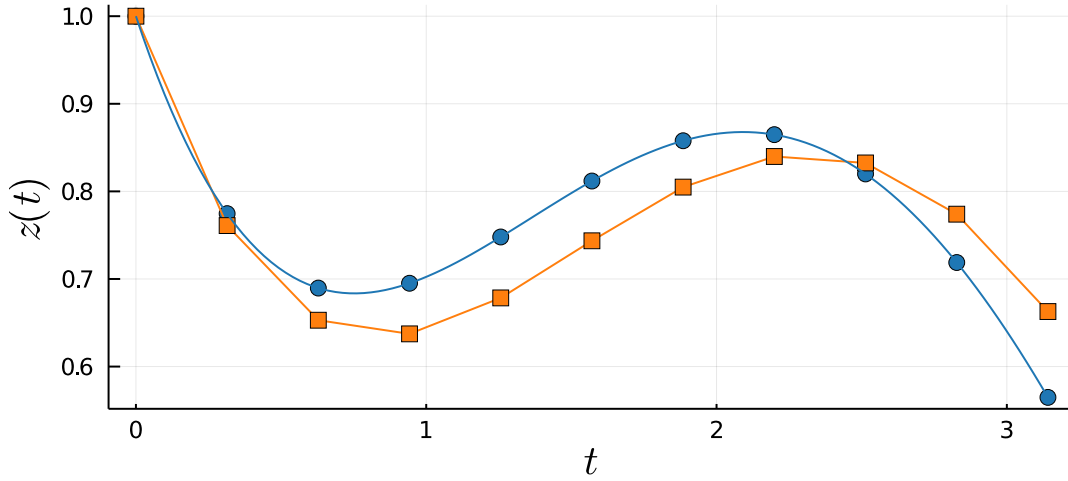


FIGURE 2 – Tracé de la solution exacte (en bleu, marqueurs ronds) au problème (3) avec $\varepsilon = 1$ et d'une approximation (orange, marqueurs carrés) sur une discrétisation uniforme de $[0, \pi]$ à 11 points.

Définition. On définit l'erreur d'une approximation comme l'erreur maximale sur les points d'interpolation, i.e. étant donnée une discrétisation à $N + 1$ points $(t_n)_{0 \leq n \leq N}$ et une approximation (u_n) d'une fonction $t \mapsto u(t)$, l'erreur est définie par la formule

$$\text{err} = \max_{0 \leq n \leq N} |u_n - u(t_n)|. \quad (9)$$

Si en outre la solution et l'approximation dépendent du paramètre $\varepsilon \in (0, \varepsilon_0]$, on écrit

2. Cette approximation est obtenue par itération en posant $z_{n+1} = z_n - \frac{\Delta t}{\varepsilon} z_{n+1} + \Delta t \sin(t_n)$ avec $z_0 = z(0) = 1$, ce qui correspond à une méthode appelée IMEX-BDF1.

$\text{err}(\varepsilon)$ l'erreur d'approximation, et on définit l'erreur uniforme $\overline{\text{err}}$ par

$$\overline{\text{err}} = \sup_{\varepsilon \in (0, \varepsilon_0]} \text{err}(\varepsilon). \quad (10)$$

Ces deux erreurs peuvent avoir des comportements différents.

Parfois, l'erreur est définie comme l'erreur au temps final, ce qui peut paraître moins contraignant mais a peu d'influence en pratique car les estimations d'erreur sont croissantes avec l'indice n . Ainsi ces deux définitions coïncident pour les résultats théoriques.

Remarque. À partir d'une approximation (u_n) , on peut obtenir une approximation sur une discrétisation plus fine par interpolation (typiquement avec des splines cubiques ou de manière linéaire comme en Figure 2). Il semble alors naturel de se demander s'il est possible de réduire le coût de calcul en calculant une approximation sur une discrétisation grossière pour l'interpoler ensuite vers une discrétisation plus fine. Néanmoins, si l'erreur telle que définie en (9) est mauvaise, on ne peut pas espérer l'améliorer par interpolation. C'est pourquoi ce manuscrit se concentre sur cette approche « simple » de l'erreur.

Lorsqu'il s'agit de trouver une approximation à une solution d'équation différentielle, il semble naturel de procéder de manière itérative : les seules données accessibles sont la condition initiale $u(0)$ et le champ de vecteurs suivi par la solution, donc il faut trouver un moyen de combiner ces deux informations pour obtenir une approximation $u_1 \approx u(\Delta t)$. Une fois cette information obtenue, on peut l'utiliser avec les autres pour calculer $u_2 \approx u(2\Delta t)$, etc. Dans les faits, on se limite à un nombre fixe $s \geq 1$ de points pour extrapoler le suivant. On parle alors de méthode multipoints. Les méthodes à un point sont appelées méthodes de Runge-Kutta.

La méthode utilisée pour calculer un terme à partir des précédents est appelé un schéma numérique $\Phi_{\Delta t}^\varepsilon$, et elle peut s'écrire

$$u_{n+s} = u_{n+s-1} + \Delta t \Phi_{\Delta t}^\varepsilon(u_{n+s-1}, \dots, u_n).$$

Cette notation peut être pratique pour étudier le schéma, mais il faut garder à l'esprit qu'elle peut camoufler de nombreuses difficultés. Par exemple, le schéma d'Euler implicite s'écrit

$$u_{n+1} = u_n + \Delta t \left(-\frac{1}{\varepsilon} A u_{n+1} + f(u_{n+1}) \right).$$

Pour trouver $\Phi_{\Delta t}^\varepsilon$, il faut inverser cette relation ; on parle de schéma *implicite*, par oppo-

sition aux schémas *explicites*. Cette inversion peut s'avérer particulièrement coûteuse si f présente des non-linéarités et si le système est grand. Ainsi par la suite on se limite à des schémas *explicites* en f , et on fait l'hypothèse suivante :

Hypothèse. *On sait calculer $t \mapsto e^{-tA}$ et $t \mapsto (\text{id} + tA)^{-1}$ de manière exacte.*

Calculer le semi-groupe $t \mapsto e^{-tA}$ peut sembler contraignant, mais rappelons nous qu'on se restreint ici au cas où A est une matrice diagonale.

Définition. *On dit qu'un schéma numérique $\Phi_{\Delta t}$ est d'ordre q s'il existe une constante $C > 0$ et un pas de temps maximal $\Delta t_0 > 0$ tel que pour toute subdivision de pas de temps $\Delta t \leq \Delta t_0$, l'erreur de schéma est bornée par $C\Delta t^q$.*

Dans le contexte d'un schéma qui dépend du paramètre ε , la constante d'erreur C et le pas de temps maximal Δt_0 dépendent généralement de ε , ainsi il faut distinguer l'ordre du schéma (calculé pour $\Delta t \ll \varepsilon$) de l'ordre de « convergence uniforme » du schéma, qui fait disparaître la dépendance en ε . Cette distinction sera clarifiée par les exemples à venir.

Méthodes numériques

On présente les résultats associés à trois méthodes d'ordre 2, qui traitent la partie raide différemment de la partie non-raide. Les méthodes sont bien définies dans la limite $\varepsilon \rightarrow 0$, et on s'intéresse au comportement de l'erreur en fonction de Δt et de ε .

On ne considèrera pas de méthodes complètement implicites parce qu'elles sont très coûteuses notamment dans un contexte d'EDP. Néanmoins la convergence de ces méthodes est souvent excellente et des implémentations très efficaces existent. La référence sur le sujet est [HW96], et toutes les bonnes boîtes à outils de résolution d'EDO contiennent la méthode RadauII.³ On ne considère pas non plus de méthodes purement explicites demandant $\Delta t < \varepsilon$, ce qui est beaucoup trop coûteux en pratique.

Il est important de noter que cette introduction ne présente pas une étude des schémas présentés. Il s'agit d'une compilation non-exhaustive de résultats et d'observations sur les propriétés de ces méthodes appliquées à (1) afin de contextualiser les contributions du manuscrit par la suite. Néanmoins, les schémas sont présentés plus en détails dans l'Annexe B.

3. Attention cependant, la plupart de ces boîtes à outils masquent la difficulté associée aux méthodes implicites, à savoir la résolution d'un système et l'erreur associée. En outre, il est parfois difficile de désactiver le pas de temps adaptatif, ce qui est problématique pour une étude de convergence.

Splitting de Strang

Une approche courante consiste à séparer le problème (1) en deux parties, l'une raide et l'autre non-raide. La manière naturelle de procéder fournit

$$\begin{cases} \partial_t u^{(1)} = -\frac{1}{\varepsilon} A u^{(1)}, \\ \partial_t u^{(2)} = f(u^{(2)}). \end{cases}$$

On note φ_t , $\varphi_t^{(1)}$ et $\varphi_t^{(2)}$ les t -flots associés aux problèmes en u , $u^{(1)}$ et $u^{(2)}$ respectivement. On remarque qu'il est simple de calculer $\varphi^{(1)}$ de manière exacte, et simple de calculer $\varphi^{(2)}$ de manière numérique. Cependant, ces deux dynamiques sont mélangées dans φ , ce qui rend le flot du problème d'origine difficile à calculer. Ainsi, on est en droit de se poser la question : est-il possible d'obtenir φ à partir de $\varphi^{(1)}$ et de $\varphi^{(2)}$?

La réponse est négative en général, mais on peut *approcher* φ à partir de $\varphi^{(1)}$ et de $\varphi^{(2)}$ à l'aide de compositions successives. C'est cette approche qu'on appelle *splitting*. Le plus couramment utilisé est le splitting de Strang, qui s'écrit

$$\varphi_t = \varphi_{t/2}^{(1)} \circ \varphi_t^{(2)} \circ \varphi_{t/2}^{(1)} + \mathcal{O}(t^3)$$

avec la constante d'erreur dans $\mathcal{O}(t^3)$ qui dépend du paramètre ε de manière raide. Pour la plupart des équations, l'ordre des opérations n'a pas d'importance, mais lorsque le système présente une partie de relaxation raide comme ici, il a été remarqué dans [Spo00 ; DM04] qu'il vaut mieux « terminer » par la relaxation.

Notons que le splitting de Strang peut être obtenu par « symétrisation » du splitting de Lie $\varphi_t^{(2)} \circ \varphi_t^{(1)}$, d'ordre 1. Le splitting est exact si et seulement si les champs A et f commutent, c'est-à-dire si on vérifie l'identité

$$Af - \partial_u f \cdot A = [A, f] = 0.$$

Dans ce cas, le splitting de Lie génère un flot qui coïncide avec φ . Évidemment, ce n'est pas le cas en général. En particulier dans le cas test (8), on a $[A, f] = \begin{pmatrix} 0 & -1 \\ -1 & 0 \end{pmatrix}$, donc on s'attend à avoir une erreur qui décroît comme Δt^2 lors de la simulation. Cependant, on observe un résultat différent en Figure 3.

Dans cette figure, on observe que le comportement de la solution est celui attendu

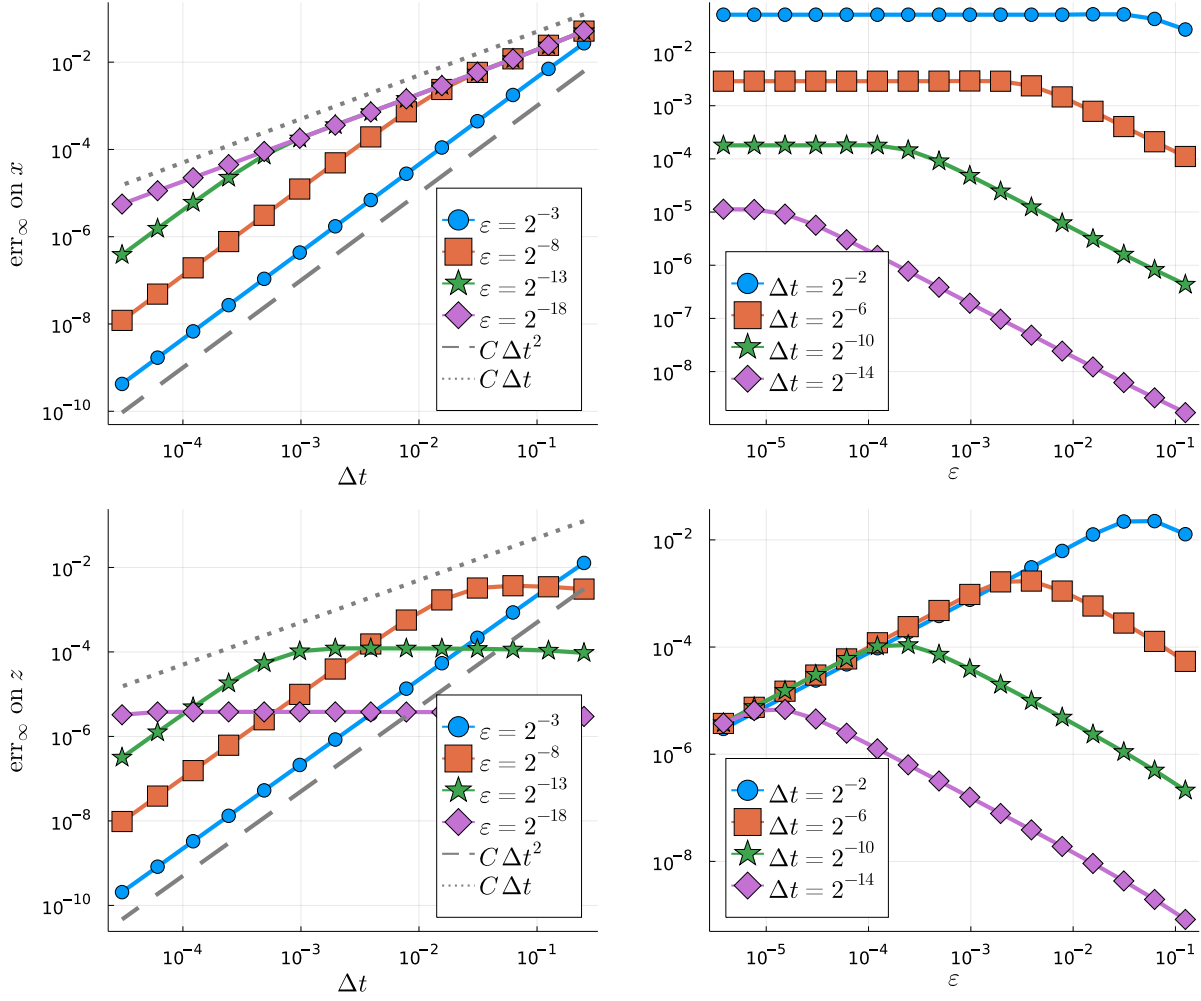


FIGURE 3 – Erreurs sur x (en haut) et z (en bas) en fonction de Δt (à gauche) et de ε (à droite) avec la méthode de Strang. Sur les tracés de l'erreur en fonction de Δt , les convergences théorique et uniforme sont tracées.

pour $\Delta t \ll \varepsilon$. Néanmoins, lorsqu'on trace l'erreur en fonction de ε , on voit qu'à Δt fixé, il y a toujours un seuil à partir duquel une réduction de ε entraîne une augmentation de l'erreur. Cette augmentation entraîne une *réduction d'ordre*, c'est-à-dire qu'on ne peut pas avoir

$$\sup_{0 < \varepsilon \leq \varepsilon_0} \text{err} \leq C \Delta t^q,$$

avec $q = 2$, ce n'est possible qu'avec $q = 1$. Cette réduction d'ordre a été notée dans différents contextes [Spo00; FOS15], et différentes méthodes ont été développées pour dépasser cette limite [EO15; CR17; BV20], même si la méthode reste au plus d'ordre 2.

Méthodes exponentielles Runge-Kutta (expRK)

Ces méthodes exponentielles⁴ proviennent de la formulation intégrale du problème (1),

$$u(t) = e^{-\frac{t}{\varepsilon}A}u_0 + \int_0^t e^{-\frac{t-\tau}{\varepsilon}A}f(u(s))ds.$$

La partie du semi-groupe est gardée exacte tandis que la partie (possiblement) non-linéaire est approchée, ce qui conduit au schéma d'ordre 1 suivant,

$$u_{n+1} = e^{-\Delta t A/\varepsilon}u_n + \left(\int_0^{\Delta t} e^{(\tau-\Delta t)A/\varepsilon}d\tau \right) f(u_n). \quad (11)$$

Pour passer à l'ordre supérieur, les parties raide et non-raide sont liées, donc l'approche demande plus de subtilité, mais on peut obtenir des méthodes exponentielles Runge-Kutta (i.e. des méthodes à un pas, pour obtenir $u_{n+1} \approx u(t_{n+1})$ à partir de seulement $u_n \approx u(t_n)$) d'ordre arbitraire. Une grande classe de schémas de ce type est compilée dans [HO05], et dans un autre article les mêmes auteurs obtiennent une convergence théorique.

Erreur de convergence (Hochbruck, Ostermann - [HO04])

Avec un schéma expRK d'ordre $q \geq 1$, l'erreur vérifie la borne à une constante multiplicative près

$$\left| \left(I + \frac{1}{\varepsilon}A \right) (u_n - u(t_n)) \right| \leq \Delta t^q \left(\sup_{0 \leq t \leq t_n} |\partial_t^{q-1} \mathcal{U}(t)| + \int_0^{t_n} |\partial_t^q \mathcal{U}(t)| dt \right)$$

où $\mathcal{U} = \partial_t u + \frac{1}{\varepsilon}Au$.

4. À ne pas confondre avec les méthodes de Lawson (voir [Law67; HLO20]) qui procèdent en appliquant des méthodes de Runge-Kutta standards sur la variable filtrée $v(t) = e^{tA/\varepsilon}u(t)$, puis en multipliant le résultat par $e^{-tA/\varepsilon}$. Les résultats théoriques sur la convergence de ces dernières sont encore très récents.

Ce résultat est généralement évoqué dans un contexte fonctionnel de problème parabolique, mais cette version simplifiée est suffisante ici. Un intérêt remarquable de ces méthodes est que dans l'erreur, la composante z est renormalisée par ε . Dans notre cas, cela permet d'obtenir une sorte d'erreur relative puisque $z(t) = \varepsilon h^\varepsilon(x(t)) + \mathcal{O}(e^{-t/\varepsilon})$. D'ailleurs, le schéma (11) (parfois appelé « Euler exponentiel ») avait été proposé dans [VS98] pour obtenir une meilleure convergence que le splitting de Strang en erreur relative.

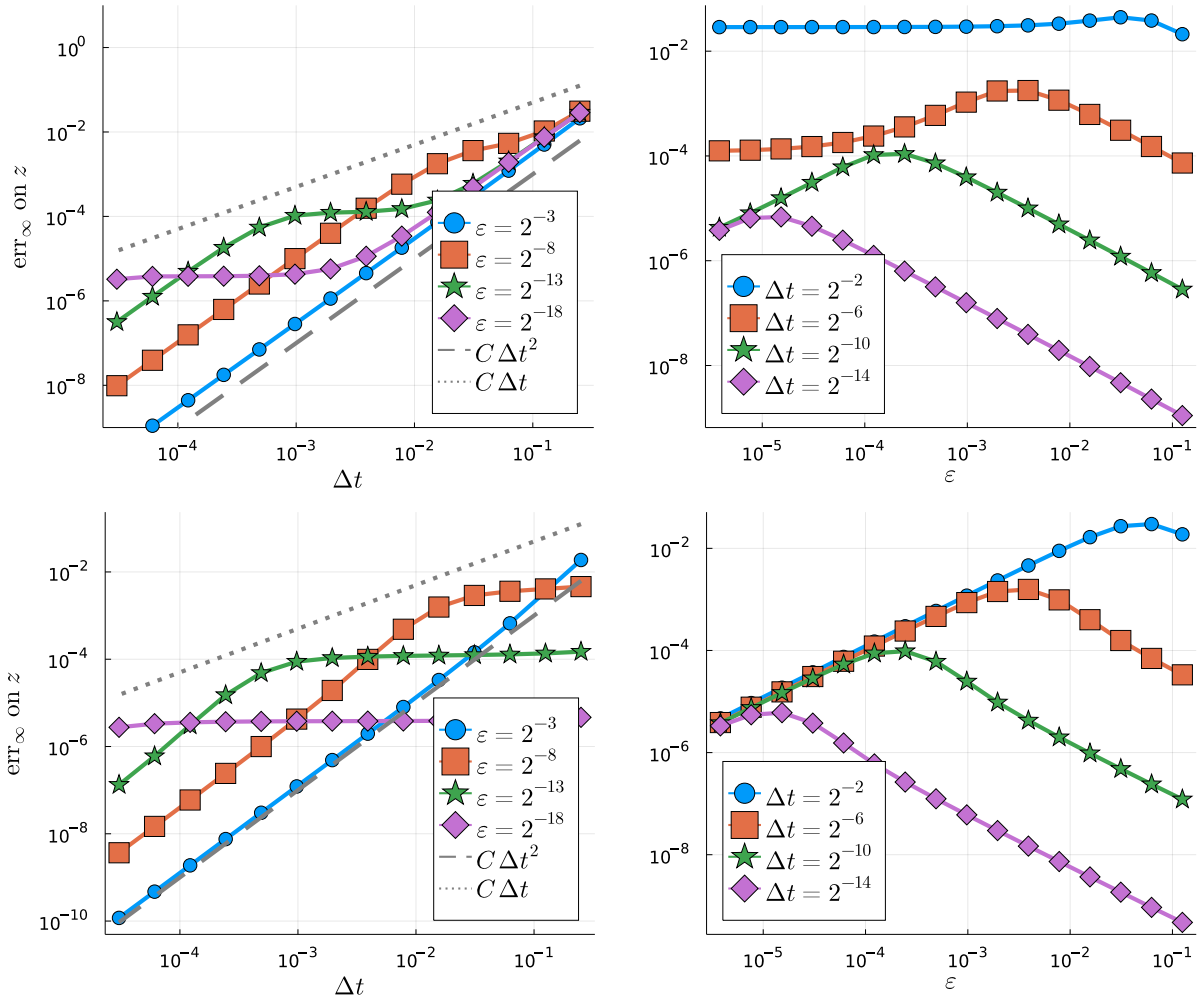


FIGURE 4 – Erreurs sur x (en haut) et z (en bas) en fonction de Δt (à gauche) et de ε (à droite) avec le schéma exponentiel RK2. Sur les tracés de l'erreur en fonction de Δt , les convergences théorique et uniforme sont tracées.

Cette renormalisation par ε de la composante z se voit nettement en Figure 4 sur l'erreur en fonction de ε est bornée par une fonction de la forme $C\varepsilon$. Sur x , à ε fixé, on observe trois phases dans les variations d'erreur en fonction de Δt :

- une décroissance en Δt^2 ;
- un plateau à partir de $\Delta t^2 \approx \varepsilon$ jusqu'à $\Delta t \approx \varepsilon$;
- de nouveau une décroissance en Δt^2 .

Cette deuxième phase engendre une réduction d'ordre, où l'ordre « uniforme » est 1, et la troisième phase présente l'ordre « naturel » du schéma. Pour la composante z , il n'y a pas de première phase, mais la réduction d'ordre est la même. Malgré cela, la convergence est meilleure que pour le splitting de Strang : dans le paradigme $\varepsilon \rightarrow 0$, seule la première phase est observée et l'erreur décroît comme Δt^2 . On dit que le schéma « préserve l'asymptote ». Cette description ne permet pas de décrire le comportement de l'erreur dans les phases 1 et 3, mais pas dans la phase 2.

Méthode IMEX-BDF

L'inconvénient des méthodes exponentielles est qu'elles demandent une intégration très précise du semi-groupe $t \mapsto e^{-tA/\varepsilon}$. On peut considérer des méthodes moins coûteuses, qui demandent seulement d'inverser un système linéaire. C'est le cas par exemple des méthodes implicite-explicites (IMEX), où la partie raide est implicite et la partie non-linéaire est explicite. Ainsi la méthode d'Euler implicite-explicite appliquée à (1) s'écrit

$$\frac{u_{n+1} - u_n}{\Delta t} = -\frac{1}{\varepsilon} A u_{n+1} + f(u_n).$$

On se restreint ici aux méthodes multipas dénommées IMEX-BDF (*backwards differentiation formula*), initialement développées dans [Cro80], puis dans [ARW95 ; ACM99 ; HR07 ; DP17]. On a là aussi une erreur théorique.

Résultat sur l'erreur (Crouzeix - [Cro80])

Avec une méthode IMEX-BDF d'ordre q à s pas, si on suppose qu'on a pu obtenir les $(s-1)$ -ièmes premiers pas avec une manière quelconque, l'erreur sur les pas suivants peut être bornée par

$$\left| u_n - u(t_n) \right| \leq \sum_{i=0}^{s-1} \left| u_i - u(t_i) \right| + \Delta t^q \int_0^{t_n} \left(\left| \partial_t^{q+1} u(t) \right| + \frac{1}{\varepsilon} \left| A \partial_t^q u(t) \right| \right) dt$$

à une constante multiplicative près.

La différence principale de cette erreur avec celle des méthodes expRK est que la composante z n'est pas « normalisée ». On voit ainsi en Figure 5 que l'erreur uniforme

sur z dégénère à l'ordre *zéro*. Il est même possible d'augmenter l'erreur en diminuant le pas de temps Δt .

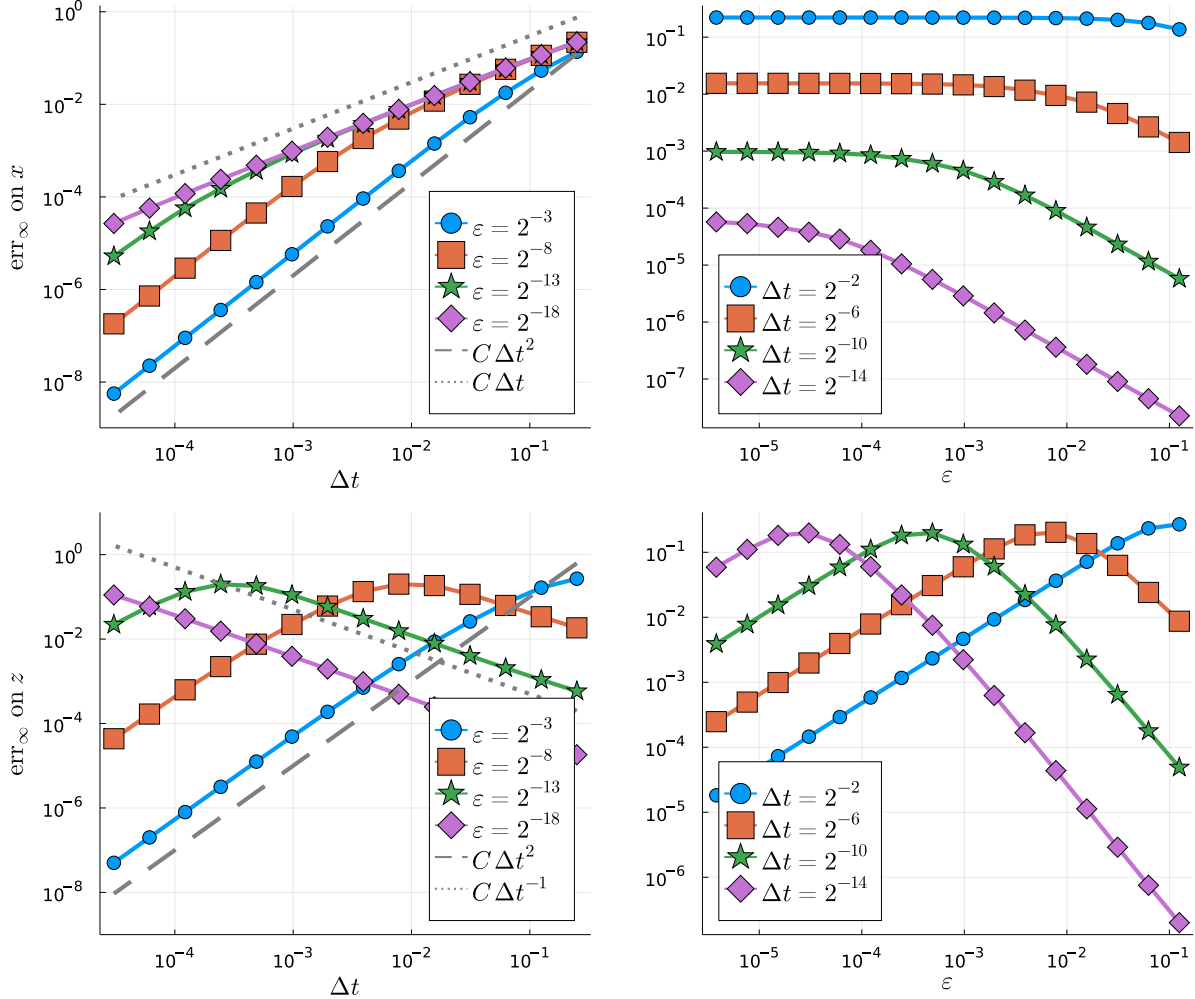


FIGURE 5 – Erreurs sur x (en haut) et z (en bas) en fonction de Δt (à gauche) et de ε (à droite) avec le schéma IMEX-BDF2. Sur les tracés de l'erreur en fonction de Δt , la convergence théorique est tracée, ainsi que pour x la convergence uniforme et pour z l'ordre négatif observé sur une portion de valeurs de Δt .

Ces méthodes sont néanmoins très utilisées, notamment dans le contexte de modèles cinétiques. Par exemple les méthodes IMEX-LM développées dans [LM08 ; BPR17 ; ADP20] sur un système

$$\partial_t \rho + \partial_x j = 0, \quad \partial_t j + \frac{1}{\varepsilon} \partial_x \rho = -\frac{1}{\varepsilon} j$$

traitent de manière la partie en j en gardant explicite la partie en ρ . Cela permet d'avoir des schémas qui se comportent bien dans la limite $\varepsilon \rightarrow 0$ malgré la raideur sur le transport

en ρ .

En outre, si la donnée initiale se situe proche de l'équilibre $z = \varepsilon h^\varepsilon(x)$ de sorte que $\partial_t^{q+1}u$ reste bornée dans la limite $\varepsilon \rightarrow 0$, alors on peut obtenir une convergence uniforme. On peut voir une interprétation de ce résultat en Figure 6 où l'erreur sur z est améliorée en prenant une donnée initiale nulle.⁵ Ainsi, il est fréquent de choisir une donnée initiale bien préparée et d'annoncer que ce type de schéma est « uniformément précis », par exemple dans [JPT00 ; HS21]. La même confusion est faite pour les schémas IMEX-RK dans [BR09 ; BPR17], par exemple.

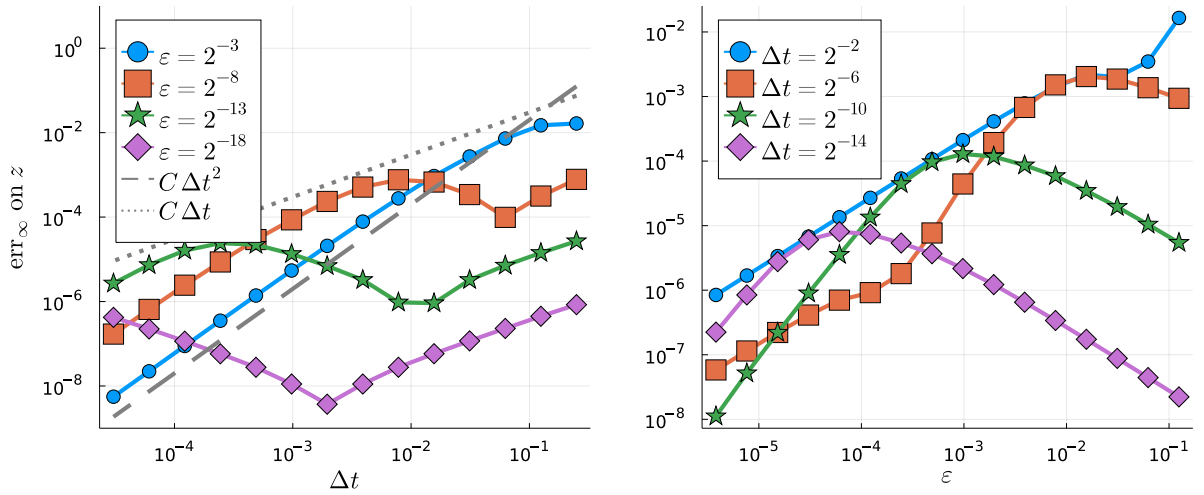


FIGURE 6 – Erreur sur z en fonction de Δt (à gauche) et de ε (à droite) avec le schéma IMEX-BDF2, avec une donnée initiale $z(0) = 0$. Sur les tracés de l'erreur en fonction de Δt , les convergences théorique et uniforme sont tracées.

D'autres notions de convergence

On voit que toutes ces méthodes subissent une réduction d'ordre : on passe d'une convergence d'ordre 2 à une convergence d'ordre 1, ou pire. Pourtant, les schémas sont bien *d'ordre 2* au sens asymptotique $\Delta t \rightarrow 0$, comme on l'observe pour $\varepsilon = 2^{-3}$ en Figures 3, 4 et 5. Il apparaît donc nécessaire d'introduire des concepts de convergence qui diffèrent des définitions usuelles.

Définition. *Considérons la solution $t \mapsto u^\varepsilon(t)$ du problème (1) avec une donnée initiale $u(0) = u_0$ indépendante de ε . On construit une solution approchée (u_n^ε) en appli-*

5. Le même phénomène est observé pour le schéma expRK.

quant un schéma $\Phi_{\Delta t}^\varepsilon$ d'ordre $q \geq 1$, et on suppose que $\Phi_{\Delta t}^\varepsilon$ admet une limite $\varepsilon \rightarrow 0$, qu'on note $\Phi_{\Delta t}^0$.

On dit que le schéma $\Phi_{\Delta t}^\varepsilon$ est asymptotic preserving (AP) ou qu'il préserve l'asymptote si le schéma limite $\Phi_{\Delta t}^0$ existe et est du même ordre q que le schéma d'origine. On dit en outre que le schéma est uniformly accurate (UA) ou uniformément précis si l'erreur uniforme présente le même ordre de convergence que l'erreur « standard » du schéma.

On peut résumer ces propriétés sur le diagramme de commutation suivant :

$$\begin{array}{ccc}
 u^{\varepsilon_0}(t) & \xrightarrow{\mathcal{O}(\Delta t^q)} & (u_n^{\varepsilon_0}) \\
 \downarrow \varepsilon \rightarrow 0 & \searrow & \downarrow \\
 u^0(t) & \xrightarrow{\mathcal{O}(\Delta t^q)} & (u_n^0)
 \end{array}$$

Les flèches verticales représentent le passage à la limite $\varepsilon \rightarrow 0$ tandis que les flèches horizontales représentent un calcul de solution numérique avec un ordre q . Un schéma UA permet d'emprunter la flèche en pointillés sans perte de précision (c'est-à-dire avec n'importe quelle valeur de ε), tandis qu'un schéma AP ne présente une bonne convergence que le long des flèches solides.

Comme mentionné précédemment, certains articles annoncent que des schémas sont UA mais en rajoutant une hypothèse de donnée initiale « bien préparée », i.e. proche de l'équilibre. Cette donnée initiale se traduit généralement en l'identité $z(0) = \varepsilon h^\varepsilon(x(0)) + \mathcal{O}(\varepsilon^q)$ grâce au théorème de variété centrale. On parlera alors de schéma UA « à l'équilibre ». Parmi les méthodes précédentes, on a les propriétés suivantes

	AP	UA éq.	UA
Strang			
IMEX-BDF2		✓	
expRK2	✓	✓	

La colonne UA étant vide, on cherche à développer de telles méthodes.

Contribution personnelle

Suite aux résultats de précision uniforme obtenus pour les problèmes hautement oscillants [CCLM15 ; CJL17 ; CLMV20], ce travail de thèse a cherché à développer des méthodes à précision uniforme dans le cadre des problèmes à relaxation rapide de type (1). Il existe deux grandes stratégies pour obtenir une convergence uniforme. La première est le développement double échelle, où on écrit la solution $t \mapsto u(t)$ comme une évaluation particulière d'une fonction à deux variable $(t, \theta) \mapsto U(t, \theta)$ en posant

$$u(t) = U(t, \theta)|_{\theta=t/\varepsilon}.$$

L'apparition de cette seconde variable permet de choisir une donnée initiale $U(0, \theta)$ qui réduit la raideur dans la direction t . La seconde stratégie est la décomposition micro-macro. L'idée est similaire à celle du double-échelle ; il s'agit de séparer la dynamique rapide oscillante (en $e^{it/\varepsilon}$) et la dynamique de *dérive* (en t). Ainsi on écrit

$$u(t) = \Omega_{t/\varepsilon}^\varepsilon \circ \Gamma_t^\varepsilon \circ \left(\Omega_0^\varepsilon\right)^{-1}(u_0)$$

où $\theta \mapsto \Omega_\theta^\varepsilon$ est périodique et Γ_t^ε est le t -flot d'un champ de vecteurs non-raide F^ε . C'est cette seconde approche que nous avons privilégié, puisqu'elle permet de ne pas avoir à considérer de seconde variable θ lors de la résolution numérique et d'utiliser des schémas standards. Néanmoins, on fournit en Annexe A des pistes pour adapter un développement double-échelle au cadre dissipatif.

L'idée de la méthode micro-macro est de chercher des approximations de Ω^ε et de F^ε à un certain ordre ε^{n+1} près, dans la veine des méthodes de moyennisation [Per69 ; LM88 ; SVM07 ; CMS10 ; CMS12a ; CCMM15]. La nouveauté ici est de s'intéresser en outre au reste de ce développement asymptotique, ainsi on obtient une décomposition exacte

$$u(t) = \Omega_{t/\varepsilon}^{[n]}(v(t)) + w(t)$$

avec $w(t) = \mathcal{O}(\varepsilon^{n+1})$ et $\partial_t v = F^{[n]}(v)$. Les applications $\Omega^{[n]}$ et $F^{[n]}$ sont des approximations de Ω^ε et F^ε respectivement. Dans [CLMV20], le problème sur (v, w) est moins raide et peut être résolu avec un schéma numérique d'ordre uniforme n .

L'utilisation de ces méthodes de moyennisation a permis, au cours de la troisième année de thèse, de dériver les résultats préexistants en moyennisation avec certaines preuves originales, notamment sur le caractère géométrique de la moyennisation dite *strobosco-*

pique. Précédemment, les démonstrations demandaient de construire le morphisme Ω^ε et le champ de vecteurs moyen F^ε , pour ensuite raisonner par récurrence ou avec des arbres. Ces nouvelles preuves ne font appel qu'à l'équation homologique (i.e. l'équation algébrique vérifiée par Ω^ε et F^ε) et sont plus élégantes. En outre, on met en évidence certains liens forts entre moyennisation et formes normales, qui sont déjà connus (voir [SVM07]) mais peu référencés. Cette synthèse fait l'objet du Chapitre 1.

L'essentiel de ce travail de thèse consisté à construire des méthodes micro-macro adaptées aux problèmes à relaxation rapide de type (1). Formellement, au lieu de voir le changement de variable $\Omega_\theta^\varepsilon$ comme une série de Fourier en $\theta \in \mathbb{R}/\mathbb{Z}$, il est vu ici comme une série exponentielle $\Omega_\tau^\varepsilon = \sum_{k \geq 0} \omega_k e^{-k\tau}$ avec $\tau \in \mathbb{R}_+$. La nouvelle difficulté est alors de calculer l'équivalent formel de la « moyenne » de cette série exponentielle. Pour cette raison, on considère $\Omega_{i\tau}^\varepsilon$ et on adapte les résultats du cas périodique. En effectuant des développements à l'ordre n , on obtient ainsi un problème micro-macro en (v, w) qu'on peut résoudre avec une précision uniforme d'ordre $n + 1$ (le caractère de relaxation permet un gain d'ordre par rapport aux problèmes hautement oscillants) en utilisant des schémas expRK. On peut aussi obtenir une précision uniforme d'ordre n avec des schémas IMEX-BDF. Il est intéressant de noter que ce résultat est étendu partiellement à des EDP bien connues : les problèmes hyperboliques relaxés

$$\begin{cases} \partial_t v_1 + \partial_x v_2 = 0, \\ \partial_t v_2 + \partial_x v_1 = \frac{1}{\varepsilon} (g(v_1) - v_2), \end{cases}$$

et l'équation de télégraphe (i.e. BGK à vitesses discrètes)

$$\begin{cases} \partial_t \rho + \partial_x j = 0, \\ \partial_t j + \frac{1}{\varepsilon} \partial_x \rho = -\frac{1}{\varepsilon} j. \end{cases}$$

De nombreuses méthodes AP ou UA à l'équilibre existent pour ces problèmes –on peut citer [Jin99 ; LM08 ; DP11 ; DP17 ; BPR17 ; ADP20]– mais le développement de méthodes UA est encore un sujet de recherche actif. On peut voir ce travail de thèse comme une étape préliminaire au développement de méthodes UA pour cette catégorie de problèmes. Ces résultats ont fait l'objet d'une publication,

Philippe CHARTIER, Mohammed LEMOU et Léopold TRÉMANT,
« A uniformly accurate numerical method for a class of dissipative
systems », à paraître dans *Mathematics of Computation* (2021),

qui est présentée en Chapitre 2.

En Chapitre 3, on discute d'extensions directes des résultats de notre article, qui servent à rendre le résultat plus robuste, et on fournit quelques pistes pour traiter l'équation de télégraphe de manière complète.

LES PROPRIÉTÉS REVISITÉES DE LA MOYENNISATION

Ce chapitre est un petit article qui offre une nouvelle approche sur des résultats connus de moyennisation. Dans la littérature les preuves de ces résultats sont laborieuses, faisant appel soit à des récurrences, soit à des propriétés sur les arbres. En outre, ces preuves requièrent de construire les morphismes de moyennisation, ce qui réduit les raisonnements à la construction en question. Dans cet article, on présente rapidement le cadre formel qui décrit ce qu'on entend par « moyennisation », puis on prouve quelques propriétés en supposant que les morphismes existent. En particulier, on s'intéresse aux propriétés géométrique de commutation, de conservation de volume et de structure hamiltonienne ou de Poisson. Dans une dernière partie, on adapte ces propriétés au cas de morphismes *approchés*.

1.1 Introduction

This paper compiles results pertaining to *high-order averaging*, that is to say the problem of separating the slow and fast dynamics in a highly-oscillatory setting. The type of problem we consider may arise in many realistic physical models, such as molecular dynamics [GSS98] or charged-particle dynamics under a strong magnetic field [CCLMZ20; FSS09; FS00]. It may also arise in functional spaces; two examples are the nonlinear Klein-Gordon equation in the nonrelativistic limit regime [BCZ14; BZ19; CLMV20] and the oscillatory nonlinear Schrödinger equation [CCLM15; CCMM15].

Mathematically speaking, we consider problems with forced oscillations of the form

$$\partial_t u(t) = f_{t/\varepsilon}(u(t)), \quad u(0) = u_0 \in X, \quad t \in [0, 1] \quad (1.1)$$

where X is a Banach space of norm $|\cdot|$, the non-autonomous vector field $(\theta, u) \in \mathbb{T} \times X \mapsto$

$f_\theta(u)$ is 1-periodic w.r.t. θ on the torus $\mathbb{T} := \mathbb{R}/\mathbb{Z}$. As mentioned, the space X may be simply \mathbb{R}^d , in which case the problem is a simple ordinary differential equation in finite dimension, or it may be a functional space, such as the space of square-integrable function $L^2(\mathbb{R})$. Note that this type of equation can result from the *filtering* of an autonomous equation

$$\dot{v}^\varepsilon = \frac{1}{\varepsilon}G(v) + K(v), \quad v^\varepsilon(0) = v_0 \in X \quad (1.2)$$

if G generates a 1-periodic flow $(\theta, u) \mapsto \chi_\theta(u)$. It links to 1.1 using the filtered variable $u^\varepsilon(t) = \chi_{-t/\varepsilon}(v^\varepsilon(t))$ which follows an equation of the form (1.1) with $f_\theta(u) = (\partial_u \chi_{-\theta} \cdot K) \circ \chi_\theta(u)$.

The approach of averaging can be summarized as the decomposition of the solution $u(t)$ into a *near-identity, rapidly oscillating* change of variable $\Phi_{t/\varepsilon}^\varepsilon$ and the dynamics of an *average* autonomous vector field F^ε . This can be written

$$u(t) = \Phi_{t/\varepsilon}^\varepsilon \circ \Psi_t^\varepsilon \circ (\Phi_0^\varepsilon)^{-1}(u_0), \quad (1.3)$$

where $(\theta, u) \mapsto \Phi_\theta^\varepsilon(u)$ is 1-periodic w.r.t. θ and $(t, u) \mapsto \Psi_t^\varepsilon(u)$ is the t -flow associated to F^ε , i.e. for $(t, u) \in [0, 1] \times X$,

$$\frac{d}{dt}\Psi_t^\varepsilon(u) = F^\varepsilon(\Psi_t^\varepsilon(u)), \quad \Psi_0^\varepsilon = \text{id}. \quad (1.4)$$

We refer to Lochak-Meunier [LM88] and Sanders-Verhulst-Murdock [SVM07] for textbooks on these issues. Since the goal is to separate the fast periodic part in θ and the slow drift in t , averaging is can be seen as analogous to the two-scale expansion $u^\varepsilon(t) = U^\varepsilon(t, \theta)|_{\theta=t/\varepsilon}$ often found in the context of high-frequency PDEs. It is also similar to WKB expansions [Wen26; Kra26; Bri26], since in some sense Φ^ε captures the rapid phase dynamics and Ψ^ε the slow amplitude changes. Admittedly, these dynamics are usually more intertwined in averaging than in WKB expansions, but this allows the preservation of geometric structures such as volume invariance or symplecticity. This is the subject of Section 1.4.

In this work, we shall not discuss specific methods to compute the periodic change of variable or the averaged vector fields, the traditional approach dating back to [Per69] consists in assuming the maps are power series in ε and injecting the ansatz $\Phi_\theta^\varepsilon = \text{id} + \sum_{n \geq 1} \Phi_\theta^{[n]}$ in (1.3) and identifying like terms in ε . This formal series approach has been revisited using B-series or the Magnus expansion in [CMS10; CMS12a; CCM19].

Another approach is that of “successive substitution” dating back to [Nei84] (albeit in a slightly different context), and more recently in [CCMM15; CLMV20]. This circumvents the ansatz and seems to yield better convergence properties. Both approaches coincide formally. Our goal in this paper is to present known results in a new light, and offer original proofs without having to invoke any ansatz, formal series or construction process.

A particularly well-studied case is that of the autonomous problems with linearly-generated oscillations (i.e. linear G), for which the problem of averaging can often be reduced to finding some θ -independent change of variable $(\Phi_0^\varepsilon)^{-1}$, or some equivalent. It is then possible to consider the problem on this new variable $(\Phi_0^\varepsilon)^{-1}(u(t))$. As such, a link can be made with normal forms, and specifically Birkhoff’s forms technique have been considered in this context by Bambusi [Bam03; BB05; Bam06; Bam08], Bourgain [Bou96], Colliander [CKSTT10; CKO12] and Grébert [Bam06; GV11; GT12], to mention only a few. We offer some insight on this approach in Section 1.3.3. Note that many of these works consider the setting of multiple non-resonant frequencies, which is akin to considering f as a function of multiple phases $\theta_1, \theta_2, \dots$ in (1.1). This setting has also been studied with averaging using diophantine approximations in [CMTZ17] and with B -series in [CMS12b].

In Section 1.2, we present some general properties of averaging, detailing the differences between standard and stroboscopic averaging. In Section 1.3, we present some remarkable properties of averaging in the autonomous case. In Section 1.4, we restrain ourselves to stroboscopic averaging, and present some of its geometric properties. Finally, in Section 1.5, we discuss what becomes of the previous results in the case of a bounded domain.

1.2 A brief presentation of averaging

Differentiating (1.3) w.r.t. t generates

$$f_{t/\varepsilon} \circ \Phi_{t/\varepsilon}^\varepsilon(v(t)) = \frac{1}{\varepsilon} \partial_\theta \Phi_{t/\varepsilon}^\varepsilon(v(t)) + \partial_u \Phi_{t/\varepsilon}^\varepsilon(v(t)) \cdot F^\varepsilon(v(t))$$

with $v(t) = \Psi_t^\varepsilon \circ (\Phi_0^\varepsilon)^{-1}(u_0)$ the average dynamics. By separating the rapid oscillations in t/ε and the slow drift in t , one obtains the homological equation, which is for $(\theta, u) \in$

$\mathbb{T} \times X$,

$$\partial_\theta \Phi_\theta^\varepsilon(u) = \varepsilon (f_\theta \circ \Phi_\theta^\varepsilon(u) - \partial_u \Phi_\theta^\varepsilon(u) F^\varepsilon(u)). \quad (1.5)$$

Now taking the average, it appears that the change of variable Φ^ε alone stores the information of the averaged vector field. Indeed, for u in X , $F^\varepsilon(u)$ is given by

$$F^\varepsilon(u) = \left(\partial_u \langle \Phi^\varepsilon \rangle(u) \right)^{-1} \langle f \circ \Phi \rangle(u), \quad (1.6)$$

where $\langle \cdot \rangle$ denotes the average, defined for a periodic map $(\theta, u) \in \mathbb{T} \times X \mapsto \varphi_\theta(u)$ by

$$\langle \varphi \rangle(u) = \int_0^1 \varphi_\theta(u) d\theta. \quad (1.7)$$

Up to a change of variable, Φ^ε is assumed to be near identity, i.e.

$$\Phi^\varepsilon = \text{id} + \mathcal{O}(\varepsilon). \quad (1.8)$$

It is known that equation (1.5) generally has no rigorous solution, only solutions as a formal series in ε . An example where this divergence is observed can be found in [CMS10]. However the series converges in the case where f_θ is a linear and bounded operator, for ε small enough.

Perhaps the most straightforward approach to solve the homological equation is a fixed point method separating the right-hand side of the equation (of size ε) and the left (of size 1). It immediately appears that a closure condition on Φ^ε is needed to properly invert ∂_θ . Two choices are often considered.

Standard averaging : $\langle \Phi^\varepsilon \rangle = \text{id}$,

also called the Chapmann-Enskog method in the context of kinetic theory (see [CCLM20]). This choice circumvents the computation of an inverse, as then $F^\varepsilon = \langle f \circ \Phi^\varepsilon \rangle$, therefore computations are not too costly. As highlighted in [CLMZ20], in numerical contexts the $\partial_u \Phi^\varepsilon \cdot F^\varepsilon$ -term can be replaced by a finite-differences approximation up to some order in ε , thereby removing the need to compute an exact derivative and making automatic computations much simpler.

Stroboscopic averaging : $\Phi_0^\varepsilon = \text{id}$,

for which the solution $u(t)$ coincides with the average $\Psi_t^\varepsilon(u_0)$ at “stroboscopic” times $t \in \varepsilon\mathbb{N}$. This produces more complex computations but renders fairly straightforward the conservation of geometric properties, such as energy preservation or symplectic structure.

We shall focus on the properties of stroboscopic averaging in the upcoming section, but it is important to keep in mind that these choices are conjugate. Indeed, the latter can be obtained from the former by setting

$$\Phi^{strob} = \Phi^{std} \circ (\Phi_0^{std})^{-1} \quad \text{and} \quad \Psi^{strob} = \Phi_0^{std} \circ \Psi^{std} \circ (\Phi_0^{std})^{-1},$$

i.e. $F^{strob} = (\partial_u \Phi_0^{std} \cdot F^{std}) \circ (\Phi_0^{std})^{-1}$. Conversely, standard averaging can be obtained from stroboscopic averaging with the relations

$$\Phi^{std} = \Phi^{strob} \circ \langle \Phi^{strob} \rangle^{-1} \quad \text{and} \quad \Psi^{std} = \langle \Phi^{strob} \rangle \circ \Psi^{strob} \circ \langle \Phi^{strob} \rangle^{-1}. \quad (1.9)$$

Thus some properties of standard averaging will also be discussed.

1.3 Commutation of flows in the autonomous case

In this section we restrict ourselves to the case of an autonomous equation of the form

$$\dot{v}^\varepsilon = \frac{1}{\varepsilon} G(v^\varepsilon) + K(v^\varepsilon), \quad v^\varepsilon(0) = v_0 \in X \quad (1.10)$$

where G and K are smooth function from a Banach space X into itself and where G generates a 1-periodic flow $(\theta, u) \mapsto \chi_\theta(u)$. The approach is the same as for the non-autonomous problem, which is to say we search a solution under the form

$$v^\varepsilon(t) = \Omega_{t/\varepsilon}^\varepsilon \circ \Psi_t^\varepsilon \circ (\Omega_0^\varepsilon)^{-1}(v_0) \quad (1.11)$$

where $\theta \mapsto \Omega_\theta^\varepsilon$ is assumed to be 1-periodic and Ψ_t^ε is the t -flow associated to the averaged vector flow K^ε . The reasons why the notation of the change of variable changed but not that of the average flow will be made clear as this section progresses. The homological equation is now

$$\partial_\theta \Omega_\theta^\varepsilon(u) - G \circ \Omega_\theta^\varepsilon(u) = \varepsilon \left(K \circ \Omega_\theta^\varepsilon(u) - \partial_u \Omega_\theta^\varepsilon(u) K^\varepsilon(u) \right). \quad (1.12)$$

It appears that the closure condition of standard averaging must be reconsidered. Indeed, in the limit $\varepsilon \rightarrow 0$, the change of variable $\Omega_\theta^\varepsilon$ approaches $\chi_{\theta+\theta_0}$ for some initial phase θ_0 . Consider for instance the case $G(u) = 2\pi \begin{pmatrix} -u_2 \\ u_1 \end{pmatrix} = 2\pi Ju$, then clearly choosing the standard closure condition $\langle \Omega^\varepsilon \rangle = \text{id}$ cannot hold, as $\langle \chi \rangle = 0$. Rather than discarding standard averaging altogether, we may *filter* the equation, which is to say transform it into a forcibly-oscillating problem by left-multiplying it by $\partial_u \chi_{-\theta+\theta_1}(\Omega_\theta^\varepsilon)$ for some arbitrary phase θ_1 . Define the filtered change of variable $\Phi_{\theta,\theta_1}^\varepsilon = \chi_{-\theta+\theta_1} \circ \Omega_\theta^\varepsilon$, it satisfies¹

$$\partial_\theta \Phi_{\theta,\theta_1}(u) = \varepsilon \left(f_{\theta,\theta_1} \circ \Phi_{\theta,\theta_1}(u) - \partial_u \Phi_{\theta,\theta_1}(u) K^\varepsilon(u) \right) \quad (1.13)$$

with $f_{\theta,\theta_1}(u) = (\partial_u \chi_{-\theta+\theta_1} \cdot K) \circ \chi_{\theta-\theta_1}(u)$. Note that we exploited the identity $\partial_\theta \chi_\theta = G \circ \chi_\theta = \partial_u \chi_\theta G$. Take now the average on θ of (1.13),

$$0 = \varepsilon \left(\langle f_{\cdot,\theta_1} \circ \Phi_{\cdot,\theta_1} \rangle(u) - \partial_u \langle \Phi_{\cdot,\theta_1} \rangle(u) K^\varepsilon(u) \right). \quad (1.14)$$

The standard choice of closure condition therefore seems to be $\langle \Phi_{\cdot,\theta_1} \rangle = \text{id}$, i.e. $\Omega_\theta^\varepsilon$ close to $\chi_{\theta-\theta_1}$. Remember however that the phase shift θ_1 is arbitrary, therefore there are an infinite number of standard closure conditions, the canonical one being $\langle \chi_{-\theta} \circ \Omega_\theta^\varepsilon \rangle = \text{id}$.

Whatever the closure condition, it is possible to obtain K^ε from (1.14), since $\partial_u \langle \Phi_{\cdot,\theta_1}^\varepsilon \rangle$ is invertible. Indeed, assuming that $\Omega_\theta^\varepsilon$ is close to $\chi_{\theta+\theta_0}$, all filtered changes of variable satisfy

$$\Phi_{\theta,\theta_1}^\varepsilon = \chi_{-\theta+\theta_1} \circ (\chi_{\theta+\theta_0} + \mathcal{O}(\varepsilon)) = \chi_{\theta_1+\theta_0} + \mathcal{O}(\varepsilon).$$

This generates the identity

$$K^\varepsilon(u) = \left(\partial_u \langle \chi_{-\theta+\theta_1} \circ \Omega_\theta^\varepsilon \rangle(u) \right)^{-1} \left\langle (\partial_u \chi_{-\theta+\theta_1} \cdot K) \circ \Omega_\theta^\varepsilon \right\rangle(u). \quad (1.15)$$

Defining an operator extracting the average behaviour

$$\mathcal{A}^{\theta_1}[\varphi] := \left(\partial_u \langle \chi_{-\theta+\theta_1} \circ \varphi_\theta \rangle \right)^{-1} \left\langle (\partial_u \chi_{-\theta+\theta_1} \cdot K) \circ \varphi_\theta \right\rangle, \quad (1.16)$$

the change of variable Ω^ε may be defined as the unique solution to the homological

1. This homological equation can also be obtained directly by considering the filtered problem of form (1.1) satisfied by $u_{\theta_1}^\varepsilon(t) = \chi_{-t/\varepsilon+\theta_1}(v^\varepsilon(t))$, which is $\partial_t u_{\theta_1}^\varepsilon(t) = f_{t/\varepsilon,\theta_1}(u_{\theta_1}^\varepsilon(t))$.

equation

$$\partial_\theta \Omega_\theta^\varepsilon - G \circ \Omega_\theta^\varepsilon = \varepsilon \left(K \circ \Omega_\theta^\varepsilon - \partial_u \Omega_\theta^\varepsilon \cdot \mathcal{A}^{\theta_1}[\Omega^\varepsilon] \right) \quad (1.17)$$

that is 1-periodic and satisfies some closure condition. Note that the above equation is considered with fixed θ_1 , but modifying this phase has no impact on the definition of Ω^ε . Linking with (1.12), this may be restated as

$$\forall \theta_1 \in \mathbb{T}, \quad K^\varepsilon = \mathcal{A}^{\theta_1}[\Omega^\varepsilon] = \mathcal{A}^0[\Omega^\varepsilon].$$

Thanks to this invariance, a group relation may be found in the case of stroboscopic averaging, summarized by the following proposition.

Proposition 1.3.1. *When considering stroboscopic averaging, for all θ and all θ_0 , the following group relation is satisfied*

$$\Omega_\theta^\varepsilon \circ \Omega_{\theta_0}^\varepsilon = \Omega_{\theta+\theta_0}^\varepsilon.$$

Equivalently, there exists a vector field G^ε such that

$$\forall \theta, \forall u, \quad \frac{d}{d\theta} \Omega_\theta^\varepsilon(u) = G^\varepsilon \circ \Omega_\theta^\varepsilon(u).$$

Démonstration. Consider the θ -map

$$\widetilde{\Omega}_\theta^\varepsilon = \Omega_{\theta+\theta_0}^\varepsilon \circ (\Omega_{\theta_0}^\varepsilon)^{-1}.$$

Writing equation (1.12) with θ replaced by $\theta + \theta_0$ and $(\Omega_{\theta_0}^\varepsilon)^{-1}(u)$ in lieu of u , we obtain with all maps evaluated in u ,

$$\partial_\theta \widetilde{\Omega}_\theta^\varepsilon - G \circ \widetilde{\Omega}_\theta^\varepsilon = \varepsilon \left(K \circ \widetilde{\Omega}_\theta^\varepsilon - \partial_u \widetilde{\Omega}_\theta^\varepsilon \cdot \left(\partial_u (\Omega_{\theta_0}^\varepsilon)^{-1} \right)^{-1} \cdot K^\varepsilon \circ (\Omega_{\theta_0}^\varepsilon)^{-1} \right). \quad (1.18)$$

The new averaged vector field $\widetilde{K}^\varepsilon = \left(\partial_u (\Omega_{\theta_0}^\varepsilon)^{-1} \right)^{-1} \cdot K^\varepsilon \circ (\Omega_{\theta_0}^\varepsilon)^{-1}$ can be written

$$\widetilde{K}^\varepsilon = \left(\left(\partial_u \langle \chi_{-\theta+\theta_0} \circ \Omega_\theta^\varepsilon \rangle \right) \circ (\Omega_{\theta_0}^\varepsilon)^{-1} \cdot \partial_u (\Omega_{\theta_0}^\varepsilon)^{-1} \right)^{-1} \left\langle \partial_u \chi_{-\theta+\theta_0} \cdot K \right\rangle \circ \Omega_\theta^\varepsilon \circ (\Omega_{\theta_0}^\varepsilon)^{-1},$$

exploiting (1.15) with $\theta_1 = \theta_0$. The derivatives can be concatenated into $\partial_u \langle \chi_{-\theta+\theta_0} \circ \Omega_\theta^\varepsilon \circ (\Omega_{\theta_0}^\varepsilon)^{-1} \rangle$. Exploiting then the phase invariance of the average, i.e. $\langle \varphi_\theta \rangle = \langle \varphi_{\theta+\theta_0} \rangle$, the

identity appears

$$\widetilde{K}^\varepsilon = \left(\partial_u \langle \chi_{-\theta} \circ \widetilde{\Omega}_\theta^\varepsilon \rangle \right)^{-1} \left\langle (\partial_u \chi_{-\theta} \cdot K) \circ \widetilde{\Omega}_\theta^\varepsilon \right\rangle = \mathcal{A}^0[\widetilde{\Omega}^\varepsilon].$$

Injecting this into (1.18), we find that $\widetilde{\Omega}^\varepsilon$ is a 1-periodic map which satisfies an equation of the form (1.17). As we only consider stroboscopic averaging, $\widetilde{\Omega}^\varepsilon$ also satisfies the same closure condition as Ω^ε , which is to say $\widetilde{\Omega}_0^\varepsilon = \Omega_0^\varepsilon = \text{id}$. Therefore, the two maps coincide and the proof is over. \square

Proposition 1.3.2. *The flows $\theta \mapsto \Omega_\theta^\varepsilon$ and $t \mapsto \Psi_t^\varepsilon$ commute with each other, i.e.*

$$\forall \theta, \quad \forall t, \quad \Omega_\theta^\varepsilon \circ \Psi_t^\varepsilon = \Psi_t^\varepsilon \circ \Omega_\theta^\varepsilon.$$

Equivalently, the vector fields G^ε and K^ε commute with each other, i.e.

$$[G^\varepsilon, K^\varepsilon] = 0$$

where $[\cdot, \cdot]$ is the usual Lie-bracket.

Démonstration. The group law for $t \mapsto \Omega_{t/\varepsilon}^\varepsilon \circ \Psi_t^\varepsilon$ (recall that equation (1.10) is autonomous) gives for all s and t

$$\left(\Omega_{s/\varepsilon}^\varepsilon \circ \Psi_s^\varepsilon \right) \circ \left(\Omega_{t/\varepsilon}^\varepsilon \circ \Psi_t^\varepsilon \right) = \Omega_{(s+t)/\varepsilon}^\varepsilon \circ \Psi_{s+t}^\varepsilon. \quad (1.19)$$

The t -flow Ψ_t^ε satisfies a group-law by construction and owing to Proposition 1.3.1, this is also the case for Ω_τ^ε . Hence, we can compose equation (1.19) from the left by $\Omega_{-s/\varepsilon}^\varepsilon$ and from the right by Ψ_{-t}^ε and obtain

$$\Psi_s^\varepsilon \circ \Omega_{t/\varepsilon}^\varepsilon = \Omega_{t/\varepsilon}^\varepsilon \circ \Psi_s^\varepsilon.$$

The commutation of the vector fields then follows in a standard way. \square

Note that this result can also be obtained from the proof of Proposition 1.3.1, since there we find

$$K^\varepsilon = \widetilde{K}^\varepsilon = \left(\partial_u (\Omega_{\theta_0}^\varepsilon)^{-1} \right)^{-1} \cdot K^\varepsilon \circ (\Omega_{\theta_0}^\varepsilon)^{-1},$$

i.e. K^ε is invariant when conjugated by $\Omega_{\theta_0}^\varepsilon$.

Remark 1.3.3. If G is linear, then differentiating $\Phi_{\theta, \theta_0}^\varepsilon$ w.r.t. θ and taking the average generates

$$G^\varepsilon = \left\langle \partial_u \Phi_{\theta, \theta_0}^\varepsilon \right\rangle^{-1} G \left\langle \Phi_{\theta, \theta_0}^\varepsilon \right\rangle = \left(\partial_u \Omega_0^{std} \cdot G \right) \circ \left(\Omega_0^{std} \right)^{-1}$$

if Ω^{std} is such that $\langle e^{-(\theta - \theta_0)G} \Omega^{std} \rangle = \text{id}$ owing to (1.9). Furthermore the average vector field K^{std} commutes with G , thanks to the identity

$$[G, K^{std}] = [\mathbb{S}(G^\varepsilon), \mathbb{S}(K^\varepsilon)] = \mathbb{S}([G^\varepsilon, K^\varepsilon]) = 0$$

with $\mathbb{S}(F) = \left(\partial_u \Phi_0^{std} \cdot F \right) \circ \left(\Phi_0^{std} \right)^{-1}$. In other words, the change of variable $\left(\Omega_0^{std} \right)^{-1}$ transforms the perturbed vector field $G + \varepsilon K$ into $G + \varepsilon K^{std}$, where G and K^{std} commute. This links to the vision of normal forms as presented in [SVM07, Chap. IX].

1.4 Stroboscopic averaging and geometry

We start by introducing some geometric properties, then prove they are preserved by stroboscopic averaging.

1.4.1 Definitions of geometric properties

Definition 1.4.1. Define the matrix $J \in \mathcal{M}(\mathbb{R}^{2n})$ as the block matrix

$$J = \begin{pmatrix} 0 & I_n \\ -I_n & 0 \end{pmatrix}.$$

A vector field function $f : \mathbb{R}^{2n} \mapsto \mathbb{R}^{2n}$ is said to be canonically Hamiltonian if there exists a scalar smooth function $H : \mathbb{R}^{2n} \mapsto \mathbb{R}$ such that

$$\forall u \in \mathbb{R}^{2n}, \quad f(u) = J^{-1} \nabla_u H(u).$$

A smooth map $(\tau, u) \in \mathbb{R} \times \mathbb{R}^{2n} \mapsto S_\tau(u) \in \mathbb{R}^{2n}$ is said to be symplectic iff

$$\forall u \in \mathbb{R}^{2n}, \quad (\partial_u S_\tau(u))^T J (\partial_u S_\tau(u)) = J, \quad (1.20)$$

or equivalently

$$\forall u \in \mathbb{R}^{2n}, \quad (\partial_u S_\tau(u)) J^{-1} (\partial_u S_\tau(u))^T = J^{-1}. \quad (1.21)$$

Remark 1.4.2. *It is known that the τ -flow of a canonically Hamiltonian system is symplectic and that the reverse is also true, at least on connected sets. This is proved by derivation and use of the integrability Lemma, which asserts that a vector function derives from a gradient iff its jacobian is symmetric (on a connected set at least).*

Definition 1.4.3. *A vector field function $f : \mathbb{R}^n \mapsto \mathbb{R}^n$ is said to be divergence free*

$$\forall u \in \mathbb{R}^{2n}, \quad \sum_{i=1}^n \partial_i f_i(u) = \text{tr}(\partial_u f) = 0.$$

A smooth function $(\tau, u) \in \mathbb{R} \times \mathbb{R}^n \mapsto S_\tau(u) \in \mathbb{R}^n$ is said to be volume-preserving iff

$$\forall u \in \mathbb{R}^n, \quad \det(\partial_u S_\tau(u)) = 1.$$

Remark 1.4.4. *By differentiation of the determinant, it is straightforward that the τ -flow of a divergence-free vector field is volume preserving. The converse is true as well.*

Definition 1.4.5. *A matrix $B(u) \in \mathcal{M}(\mathbb{R}^n)$ is said to be a Poisson matrix if it is skew-symmetric and satisfies the Jacobi relation*

$$\forall i, j, k \in \{1, \dots, n\}, \quad \sum_{l=1}^n (\partial_l b_{ij}) b_{lk} + (\partial_l b_{jk}) b_{li} + (\partial_l b_{ki}) b_{lj} = 0.$$

A vector field function $f : \mathbb{R}^n \mapsto \mathbb{R}^n$ is said to be Poisson if there exists a scalar smooth function $H : \mathbb{R}^n \mapsto \mathbb{R}$ and a Poisson matrix $B(u)$ such that

$$\forall u \in \mathbb{R}^n, \quad f(u) = B(u) \nabla_u H(u).$$

A smooth function $(\tau, u) \in \mathbb{R} \times \mathbb{R}^n \mapsto S_\tau(u) \in \mathbb{R}^n$ is said to be a Poisson map iff

$$\forall u \in \mathbb{R}^n, \quad (\partial_u S_\tau(u)) B(u) (\partial_u S_\tau(u))^T = B(S_\tau(u)).$$

Remark 1.4.6. *The τ -flow of a Poisson system is a Poisson map symplectic, and the converse is locally true if in addition the Casimirs (spanning the null space of B) are preserved by the flow. This result is found for instance in [HLW06, Chap. VII, Thm. 4.5].*

The proof consists in separating the null space and the invertible space of B with a change of variable. The invertible space can be treated as in the Hamiltonian case, and the null space is preserved by definition of the Casimirs.

1.4.2 The geometry of stroboscopic averaging

If both G and K satisfy a geometric property (divergence-free, an invariant or B -Poisson) in the autonomous case, then this property is also satisfied by f_θ for all $\theta \in \mathbb{T}$ in the associated forced problem. Indeed, since G and K satisfy the same property, they belong to the same Lie algebra, and the flow of G is in the generating group of the algebra. Because f_θ is obtained by conjugating K by the flow of G , i.e. by conjugating an element of the algebra by an element of the group, it still belongs to the algebra. As the averaged vector field of the autonomous case coincides with the averaged vector field of this forced case, any geometric reasoning conducted on the forced case can be extended to the autonomous case. For this reason, we only state the following result in the forced case.

Theorem 1.4.7. *If $\varepsilon \partial_u F^\varepsilon$ is bounded, then stroboscopic averaging is a geometric procedure. More precisely, for ε small enough, if for all $\theta \in \mathbb{T}$,*

- (i) *f_θ is a divergence-free vector field and X is of dimension $d < \infty$, then F^ε is also divergence-free.*
- (ii) *the functional I is preserved by the flow of f_θ , then it is an invariant of F^ε .*
- (iii) *f_θ is a Hamiltonian vector field, then F^ε is Hamiltonian.*
- (iv) *f_θ is a B -Poisson vector field, then F^ε is B -Poisson.*

Démonstration. For the sake of the upcoming proofs, we shall denote for any map $(t, u) \mapsto \varphi_t(u) \in E$ with $(E, |\cdot|)$ a Banach space, at fixed t ,

$$\|\varphi_t\| = \sup_{u \in X} |\varphi_t(u)| \quad \text{and} \quad \|\varphi\|_\varepsilon = \sup_{t \leq \varepsilon} \|\varphi_t\|.$$

We also set $C = \sup_\varepsilon \varepsilon \|\partial_u F^\varepsilon\|$.

- (i) For $(t, u) \in [0, \varepsilon] \times X$, write $R_t(u)$ the deviation from volume preservation,

$$R_t(u) = \det \left(\partial_u \Psi_t^\varepsilon(u) \right) - 1.$$

Setting $L_t(u) = S_t(u) = \text{tr}(\partial_u F^\varepsilon) \circ \Psi_t^\varepsilon(u)$, it satisfies

$$\frac{d}{dt} R_t = L_t R_t + S_t, \quad \text{i.e.} \quad R_t = \int_0^t L_\tau R_\tau d\tau + \int_0^t S_\tau d\tau. \quad (1.22)$$

Taylor's theorem with integral remainder generates the identity

$$R_\varepsilon = R_0 + \varepsilon S_0 + \int_0^\varepsilon (\varepsilon - t) \dot{R}_t dt,$$

which can be simplified thanks the periodicity of Φ^ε , as then $R_\varepsilon = R_0 = 0$. Hence the bound

$$\|S_0\| \leq \frac{\varepsilon}{2} \|\dot{R}\|_\varepsilon \leq \frac{\varepsilon}{2} (\|L\|_\varepsilon \|R\|_\varepsilon + \|S\|_\varepsilon).$$

Now applying Gronwall's lemma to the integral form of R yields $\|R_t\| \leq t \|S\|_\varepsilon e^{t\|L\|_\varepsilon}$, which injects into the previous bound

$$\|S_0\| \leq \varepsilon \left(\varepsilon \|L\|_\varepsilon e^{\varepsilon\|L\|_\varepsilon} + 1 \right) \|S\|_\varepsilon.$$

Since $S_t = S_0 \circ \Psi_t^\varepsilon$, it appears by one-to-one property of Ψ_t^ε that $\|S_0\| = \|S\|_\varepsilon$. The same can be said for L , thus $\varepsilon \|L\|_\varepsilon \leq dC$. We finally obtain

$$\|S\|_\varepsilon \leq \varepsilon \left(1 + dC e^{dC} \right) \|S\|_\varepsilon$$

therefore for $\varepsilon < \left(1 + dC e^{dC} \right)^{-1}$, the source term in (1.22) is zero, and the t -flow Ψ_t^ε is volume-preserving. Equivalently, the averaged vector field F^ε is divergence-free.

(ii) From the identity

$$\frac{d}{dt} [I \circ \Psi_t^\varepsilon] = (\partial_u I \cdot F^\varepsilon) \circ \Psi_t^\varepsilon,$$

Taylor's theorem generates

$$I \circ \Psi_\varepsilon^\varepsilon = I + \varepsilon \partial_u I \cdot F^\varepsilon + \int_0^\varepsilon (\varepsilon - t) (\partial_u I \cdot F^\varepsilon) \circ \Psi_t^\varepsilon dt.$$

By periodicity of Φ^ε , $I \circ \Psi_\varepsilon^\varepsilon = I$, therefore $\|\partial_u I \cdot F^\varepsilon\| \leq \frac{\varepsilon}{2} \|(\partial_u I \cdot F^\varepsilon) \circ \Psi^\varepsilon\|_\varepsilon$. By one-to-one property of Ψ_t^ε at all t , both norms are equal, therefore for $\varepsilon < 2$,

$$\|\partial_u I \cdot F^\varepsilon\| = 0,$$

i.e. I is an invariant of F^ε .

(iii) Writing R_t the deviation from simplicity,

$$R_t = \partial_u \Psi_t^\varepsilon J^{-1} (\partial_u \Psi_t^\varepsilon)^T - J^{-1},$$

it satisfies an equation of the form (1.22) with $L_t M = \partial_u F^\varepsilon(\Psi_t^\varepsilon) M + M (\partial_u F^\varepsilon(\Psi_t^\varepsilon))^T$ and $S_t = L_t J^{-1}$. Exactly the same reasoning can be made as in (i), therefore the t -flow Ψ_t^ε is symplectic, i.e. the averaged vector field F^ε is Hamiltonian.

(iv) This follows from (ii) and (iii) thanks to Remark 1.4.6.

□

Remark 1.4.8. *It may be of interest to note that in the linear autonomous case $\partial_t u = \frac{1}{\varepsilon} Gu + Ku$, property (i) of volume-preservation does not involve the dimension. Indeed differentiating the filtered change of variable $\Phi_\theta = e^{-\theta G} e^{\theta G^\varepsilon}$ and taking the average yields*

$$G^\varepsilon = \langle \Phi \rangle^{-1} G \langle \Phi \rangle.$$

In the homological equation $\partial_\theta \Omega_\theta = \varepsilon(K\Omega_\theta - \Omega_\theta K^\varepsilon)$, we obtain

$$K^\varepsilon = \langle \Phi \rangle^{-1} K \langle \Phi \rangle.$$

The involvement of the dimension in our proof actually seems purely technical, since the averaged vector field F^ε can be expressed as a power series in ε which converges for ε small enough. Our result shows that every term of the series must be divergence-free, but the radius of convergence of the series ε_0 may be independent of the dimension of the space.

1.5 Approximations on bounded domains

In this section we discuss what becomes of the previous results in actual applications, which is to say when the maps Φ^ε and F^ε of averaging are not known exactly, but only up to error terms of size ε^{n+1} for some order $n \in \mathbb{N}$. We also get rid of the assumption that the vector field $(\theta, u) \mapsto f_\theta(u)$ is uniformly bounded on the entire space X , and conduct our study on a possibly-bounded open subset $\mathcal{K} \subset X$.

1.5.1 Assumptions

For technical purposes, define \mathcal{K}_ρ this subset extended by a radius of $\rho \geq 0$, i.e.

$$\mathcal{K}_\rho = \{u \in X \quad \text{s.t.} \quad \exists v \in \mathcal{K}, |u - v| \leq \rho\}.$$

We also define, given a map φ from \mathcal{K}_ρ to some Banach space $(E, |\cdot|)$, the norm

$$\|\varphi\|_\rho = \sup_{u \in \mathcal{K}_\rho} |\varphi(u)|.$$

In particular for the vector fields and morphisms $E = X$, and for their derivatives $E = \mathcal{L}(E, E)$.

Assumption 1.5.1. *The vector field $(\theta, u) \mapsto f_\theta(u)$ and its derivative are bounded (uniformly w.r.t. θ) on K_{3R} for some radius $R > 0$. There exist positive constants ε_0 and C such that for any rank $n \in \mathbb{N}$, there is a continuous near-identity 1-periodic change of variable $\Phi^{[n]}$ and a near-averaged vector field $F^{[n]}$, both well-defined on \mathcal{K}_{3R} for $\varepsilon \leq \varepsilon_n := \varepsilon_0/(n+1)$. Precisely,*

$$\sup_{\theta \in \mathbb{T}} \|\Phi_\theta^{[n]} - \text{id}\|_{3R} \leq \frac{\varepsilon}{\varepsilon_n} R \quad \text{and} \quad \|F^{[n]}\|_{3R} \leq C.$$

Furthermore, the error of approximation is of size ε^{n+1} , i.e. writing $\Psi_t^{[n]}$ the t -flow of $F^{[n]}$,

$$u(t) = \Phi_{t/\varepsilon}^{[n]} \circ \Psi_t^{[n]} \circ (\Phi_0^{[n]})^{-1}(u_0) + \mathcal{O}(\varepsilon^{n+1}) \quad (1.23)$$

until some time $T_R > 0$.

The error of approximation is characterised by the defect $\delta^{[n]}$ defined by

$$\delta_\theta^{[n]} = \frac{1}{\varepsilon} \partial_\theta \Phi_\theta^{[n]} - f_\theta \circ \Phi_\theta^{[n]} + \partial_u \Phi_\theta^{[n]} \cdot F^{[n]},$$

which corresponds to the error in the homological equation (1.5). The previous assumptions corresponds to the situation

$$\sup_{\theta \in \mathbb{T}} \|\delta^{[n]}\|_{3R} = \mathcal{O}(\varepsilon^n) \quad \text{and} \quad \langle \delta^{[n]} \rangle = \mathcal{O}(\varepsilon^{n+1}). \quad (1.24)$$

This assumption matches the behaviour generally observed with averaging, found for instance in [CCMM15] when assuming $(\theta, u) \mapsto f_\theta(u)$ analytic w.r.t. u . As noted in [CMS15],

this is enough to ensure the historical optimal “exponential” error bound of [Nei84], which can be stated as such : There is a positive constant c such that for all $\varepsilon > 0$ there is an integer n such that for all t ,

$$\left| u(t) - \Phi_{t/\varepsilon}^{[n]} \circ \Psi_t^{[n]} \circ \left(\Phi_0^{[n]} \right)^{-1}(u_0) \right| \leq ce^{-c/\varepsilon}.$$

This reflects the fact that the maps Φ^ε and F^ε can only be obtained as diverging power series in ε , therefore the error is *formal*, up to a flat function. Indeed, in order to increase the order of the approximation, ε must get smaller and smaller, such that an error $\mathcal{O}(\varepsilon^\infty)$ is impossible with $\varepsilon \neq 0$.

Note furthermore that this assumption is enough to ensure that $\Phi_0^{[n]}$ and $\langle \Phi^{[n]} \rangle$ are invertible from \mathcal{K}_ρ to $\mathcal{K}_{\rho+R}$ for any $\rho \in [0, 3R]$. Indeed for $u \in \mathcal{K}_\rho$, the map $\varphi(v) = u + v - \Phi_0^{[n]}(u + v)$ maps the closed ball of radius R onto itself,² thus admits a fixed point by Brouwer’s fixed point theorem. Therefore there exists $u^* = u + v^* \in \mathcal{K}_{\rho+R}$ such that $u = \Phi_0^{[n]}(u^*)$. The same reasoning holds for $\langle \Phi^{[n]} \rangle$.

1.5.2 Autonomous case

Consider the autonomous problem (1.10) of Section 1.3,

$$\dot{v}^\varepsilon = \frac{1}{\varepsilon} G(v^\varepsilon) + K(v^\varepsilon), \quad v^\varepsilon(0) = v_0.$$

The flow of G , denoted $(\theta, u) \mapsto \chi_\theta(u)$, is assumed 1-periodic w.r.t. θ , and we assume that for every radius ρ , the set \mathcal{K}_ρ is invariant by the flow of G . Performing averaging on this problem is equivalent to performing it on the filtered problem

$$\dot{u}^\varepsilon(t) = \left(\partial_u \chi_{-t/\varepsilon} \cdot K \right) \circ \chi_{t/\varepsilon}(u^\varepsilon(t)), \quad u^\varepsilon(0) = v_0.$$

The unfiltered variable is obtained as $v^\varepsilon(t) = \chi_{t/\varepsilon}(u^\varepsilon(t))$. Given an approximation $v^\varepsilon(t) = \Omega_{t/\varepsilon}^{[n]} \circ \Psi_t^{[n]} \circ \left(\Omega_0^{[n]} \right)^{-1} + \mathcal{O}(\varepsilon^{n+1})$, an approximation on u^ε of the form (1.23) is obtained by setting $\Phi_\theta^{[n]} = \chi_{-\theta} \circ \Omega_\theta^{[n]}$. Conversely, it is also possible to obtain $\Omega^{[n]}$ from working on the filtered problem, and in the case where $u \mapsto G(u)$ is non-linear, this latter approach is generally more straightforward. The defect associated to averaging on the autonomous

2. If $\varepsilon \leq \alpha \varepsilon_n$ for $\alpha \in (0, 1]$, then this radius becomes αR , therefore $\Phi_0^{[n]}$ injects \mathcal{K}_ρ into $\mathcal{K}_{\rho+\alpha R}$.

problem is

$$\eta_\theta^{[n]} := \frac{1}{\varepsilon} \left(\partial_\theta \Omega_\theta^{[n]} - G \circ \Omega_\theta^{[n]} \right) - K \circ \Omega_\theta^{[n]} + \partial_u \Omega_\theta^{[n]} K^{[n]} \quad (1.25)$$

and the link is made with the filtered averaging with the formula

$$\eta_\theta^{[n]} = \partial_u \chi_\theta \left(\Phi_\theta^{[n]} \right) \cdot \delta_\theta^{[n]}.$$

Theorem 1.5.2 (Adaptation of Propositions 1.3.1 and 1.3.2).

Given averaging maps $\Phi^{[n]}$ and $K^{[n]}$ which satisfy Assumption 1.5.1 (with $F^{[n]}$ replaced by $K^{[n]}$) and such that the associated defect $\delta^{[n]}$ satisfies (1.24), define the change of variable $(\theta, u) \mapsto \Omega_\theta^{[n]}(u) = \chi_{-\theta} \circ \Phi_\theta^{[n]}(u)$ for autonomous averaging. With this definition, $\Omega_\theta^{[n]}$ is the θ -flow of a vector field $G^{[n]}$ defined on \mathcal{K}_R up to $\mathcal{O}(\varepsilon^{n+2})$. Furthermore, $G^{[n]}$ and $K^{[n]}$ commute up to $\mathcal{O}(\varepsilon^{n+2})$ on \mathcal{K}_R .

D  monstration. The first step of the proof is to show

$$K^{[n]} = \mathcal{A}^{\theta_1} [\Omega^{[n]}] + \mathcal{O}(\varepsilon^{n+1})$$

for all phases $\theta_1 \in \mathbb{T}$, with \mathcal{A}^{θ_1} the operator defined in (1.16). This result stems from the identity on $\widetilde{\Phi}_\theta^{[n]} = \chi_{-\theta-\theta_1} \circ \Omega_\theta^{[n]}$,

$$\partial_\theta \widetilde{\Phi}_\theta^{[n]} = \varepsilon \left(f_{\theta+\theta_1} \circ \widetilde{\Phi}_\theta^{[n]} - \partial_u \widetilde{\Phi}_\theta^{[n]} \cdot K^{[n]} \right) - \varepsilon \partial_u \chi_{-\theta-\theta_1} (\Omega_\theta^{[n]}) \cdot \eta_\theta^{[n]}.$$

Before taking the average, compute

$$\begin{aligned} \partial_u \chi_{-\theta-\theta_1} (\Omega_\theta^{[n]}) \cdot \eta_\theta^{[n]} &= \partial_u \chi_{-\theta-\theta_1} \left(\chi_\theta \Phi_\theta^{[n]} \right) \partial_u \chi_\theta (\Phi_\theta^{[n]}) \delta_\theta^{[n]} \\ &= \partial_u (\chi_{-\theta-\theta_1} \circ \chi_\theta \circ \Phi_\theta^{[n]}) \left(\partial_u \Phi_\theta^{[n]} \right)^{-1} \delta_\theta^{[n]} \\ &= \partial_u \chi_{-\theta_1} (\Phi_\theta^{[n]}) \delta_\theta^{[n]} \end{aligned}$$

Hence this term can be written as $\partial_u \chi_{-\theta_1} \left(\text{id} + \mathcal{O}(\varepsilon) \right) \delta_\theta^{[n]}$, and its average is of size $\mathcal{O}(\varepsilon^{n+1})$ thanks to the assumption on $\delta^{[n]}$. Taking the average of the previous identity, we finally obtain

$$K^{[n]} = \mathcal{A}^{\theta_1} [\Omega^{[n]}] + \mathcal{O}(\varepsilon^{n+1}).$$

We then proceed in the same manner as for the proof of Proposition 1.3.1. For some phase $\theta_0 \in \mathbb{T}$, consider the map $\widetilde{\Omega}_\theta^{[n]} = \Omega_{\theta+\theta_0}^{[n]} \circ \left(\Omega_{\theta_0}^{[n]} \right)^{-1}$ defined on \mathcal{K}_R . By definition of

the defect, this new map satisfies the equation,

$$\partial_\theta \widetilde{\Omega}_\theta^{[n]} - G \circ \widetilde{\Omega}_\theta^{[n]} = \varepsilon \left(K \circ \widetilde{\Omega}_\theta^{[n]} - \partial_u \widetilde{\Omega}_\theta^{[n]} \cdot \widetilde{K}^{[n]} \right) - \varepsilon \widetilde{\eta}_\theta^{[n]}.$$

with $\widetilde{K}^{[n]} = \left(\partial_u (\Omega_{\theta_0}^{[n]})^{-1} \right)^{-1} \cdot K^{[n]} \circ (\Omega_{\theta_0}^{[n]})^{-1}$ and $\widetilde{\eta}_\theta^{[n]} = \eta_{\theta+\theta_0}^{[n]} \circ (\Omega_{\theta_0}^{[n]})^{-1}$. From (1.25), it appears in particular that $K^{[n]} = \mathcal{A}^{\theta_0}[\Omega^{[n]}] + \mathcal{O}(\varepsilon^{n+1})$. Injected into $\widetilde{K}^{[n]}$, this generates

$$\widetilde{K}^{[n]} = \left(\partial_u \langle \chi_{-\theta} \circ \widetilde{\Omega}_\theta^\varepsilon \rangle \right)^{-1} \left\langle (\partial_u \chi_{-\theta} \cdot K) \circ \widetilde{\Omega}_\theta^\varepsilon \right\rangle + \mathcal{O}(\varepsilon^{n+1}) = \mathcal{A}^0[\widetilde{\Omega}^\varepsilon] + \mathcal{O}(\varepsilon^{n+1}).$$

Hence $\Omega^{[n]}$ and $\widetilde{\Omega}^{[n]}$ satisfy the same equation up to a modification of the defect while still respecting (1.24). In other words, we can replace $\Omega^{[n]}$ by $\widetilde{\Omega}^{[n]}$ in the following equation

$$\partial_\theta \Omega_\theta^{[n]} - G \circ \Omega_\theta^{[n]} = \varepsilon \left(K \circ \Omega_\theta^{[n]} - \partial_u \Omega_\theta^{[n]} \cdot \mathcal{A}^0[\Omega^{[n]}] \right) + \mathcal{O}(\varepsilon^{n+1})$$

without impacting the result. Since these two maps satisfy the same closure condition $\Omega_0^{[n]} = \text{id} + \mathcal{O}(\varepsilon^{n+1})$, they differ by only $\mathcal{O}(\varepsilon^{n+1})$ at any phase $\theta \in \mathbb{T}$. We can finally define

$$G^{[n]} = \partial_\theta \Omega_\theta^{[n]} \Big|_{\theta=0}.$$

The second part of the theorem stems from the identity $K^{[n]} = \widetilde{K}^{[n]} + \mathcal{O}(\varepsilon^{n+1})$ which becomes

$$K^{[n]} \circ \Omega_{\theta_0}^{[n]} = \partial_u \Omega_{\theta_0}^{[n]} \cdot K^{[n]} + \mathcal{O}(\varepsilon^{n+1}).$$

□

Note that the exact flow of $G^{[n]}$ may not be 1-periodic depending on its definition. Think for instance of the one-dimensional converging example $G^{[n]} = i(1 - \varepsilon) \sum_{k=0}^n \varepsilon^k = i(1 - \varepsilon^{n+1})$.

1.5.3 Geometric properties

Here is what the preservation of geometric properties presented in Section 1.4 becomes.

Theorem 1.5.3 (Adaptation of Theorem 1.4.7).

Consider Assumption 1.5.1 met and denote φ_t^ε the t -flow associated to Problem (1.1). Up to a reduction of ε_0 , the following properties are satisfied up to an error of size $\mathcal{O}(\varepsilon^{n+1})$: if for all $t \in [0, T_R]$,

- (i) $u \mapsto \varphi_t^\varepsilon(u)$ is volume-preserving on \mathcal{R} , then $\Psi_t^{[n]}$ is volume-preserving on \mathcal{K}_R ;
- (ii) the functional I is preserved by φ_t^ε on \mathcal{K}_{2R} , then it is preserved by $\Psi_t^{[n]}$ on \mathcal{K}_R ;
- (iii) φ_t^ε is symplectic on \mathcal{K}_{2R} , then $\Psi_t^{[n]}$ is symplectic on \mathcal{K}_R ;
- (iv) φ_t^ε is B -symplectic and preserves Casimirs on \mathcal{K}_{2R} , then $\Psi_t^{[n]}$ also does on \mathcal{K}_R .

Note that since $\Phi_\theta^{[n]} = \varphi_{\varepsilon\theta}^\varepsilon \circ \Psi_\theta^{[n]} + \mathcal{O}(\varepsilon^{n+1})$, these properties are also true for $\Phi_\theta^{[n]}$. It is therefore possible to modify $\Phi^{[n]}$ and $F^{[n]}$ and have these properties met exactly, although the impact of this process on the well-posedness of the maps is unclear.

Démonstration. As can be seen in the proof of Theorem 1.4.7, every property can be proven in the same way. Therefore we will only describe how to prove (iii), as it is probably the most interesting property for the majority of readers. We refer to the other proof for the adaptation to other properties.

Set $(t, u) \mapsto \Delta_t(u)$ the deviation from symplecticity,

$$\Delta_t = \left(\partial_u \Psi_t^{[n]} \right) J^{-1} \left(\partial_u \Psi_t^{[n]} \right)^T - J^{-1},$$

defined and bounded for $u \in \mathcal{K}_{R_n}$. Thanks the periodicity of $\Phi^{[n]}$, $t \mapsto \Delta_t$ is almost zero at stroboscopic times, meaning that for all $k \in \mathbb{N}$ such that $\varepsilon k \leq T_R$, since $\Psi_{\varepsilon k}^{[n]} = \varphi_{\varepsilon k}^\varepsilon + \mathcal{O}(\varepsilon^{n+1})$,

$$\Delta_{\varepsilon k} = \left(\partial_u \varphi_{\varepsilon k}^\varepsilon \right) J^{-1} \left(\partial_u \varphi_{\varepsilon k}^\varepsilon \right)^T - J^{-1} + \mathcal{O}(\varepsilon^{n+1}) = \mathcal{O}(\varepsilon^{n+1}).$$

Setting $L_t M = \partial_u F^{[n]}(\Psi_t^{[n]})M + M \left(\partial_u F^{[n]}(\Psi_t^{[n]}) \right)^T$ and $S_t = L_t J^{-1}$, it satisfies

$$\partial_u \Delta_t = L_t \Delta_t + S_t, \quad \text{i.e.} \quad \Delta_t = \Delta_0 + \int_0^t L_\tau \Delta_\tau d\tau + \int_0^t S_\tau d\tau. \quad (1.26)$$

We want to prove $\sup_{0 \leq t \leq \varepsilon} \|\Delta_t\|_R = \mathcal{O}(\varepsilon^{n+1})$. To that effect, introduce the norm $\|\cdot\|_{\varepsilon, \rho}$ and the radii R_k ,

$$\|g\|_{\varepsilon, \rho} = \sup_{0 \leq t \leq \varepsilon} \|g_t\|_\rho \quad \text{and} \quad R_k = R + \frac{k}{n+1} R,$$

and set $\alpha > 0$ such that $\|\Delta_0\|_{2R}, \|\Delta_\varepsilon\|_{2R} \leq \alpha \varepsilon^{n+1}$. Gronwall's lemma in the integral form of Δ_t yields

$$\|\Delta\|_{\varepsilon, R} \leq \left(\alpha \varepsilon^{n+1} + \varepsilon \|S\|_{\varepsilon, R} \right) e^{\varepsilon \|L\|_{\varepsilon, R}}, \quad (1.27)$$

therefore we want to show $\|S\|_{\varepsilon,R} = \mathcal{O}(\varepsilon^{n+1})$. Because S is transported by $F^{[n]}$, i.e. $S_t = S_0 \circ \Psi_t^{[n]}$, it is possible to bound S_t on some space \mathcal{K}_ρ by the norm of S_0 on a larger space. In particular, assuming $\varepsilon_0 \leq R/C$,

$$\|S\|_{\varepsilon,R_k} \leq \|S_0\|_{R_{k+1}} \quad (1.28)$$

since $\|\Psi_t^{[n]} - \text{id}\|_{R_n} \leq tC$. This bound is fairly useful, as Taylor's theorem with integral remainder generates the identity

$$\Delta_\varepsilon = \Delta_0 + \varepsilon S + \int_0^\varepsilon (\varepsilon - t) \partial_t \Delta_t dt,$$

from which a Cauchy inequality yields

$$\|S_0\|_{R_1} \leq \frac{\varepsilon}{2} \|\partial_t \Delta\|_{\varepsilon,R_1} + 2\alpha\varepsilon^n \leq \frac{\varepsilon}{2} (\|L\|_{\varepsilon,R_1} \|\Delta\|_{\varepsilon,R_1} + \|S\|_{\varepsilon,R_1}) + 2\alpha\varepsilon^n.$$

Injecting (1.27) and (1.28) into the right-hand term, we obtain

$$\|S_0\|_{R_1} \leq \frac{\varepsilon}{2} (1 + C_L) \|S_0\|_{R_2} + (2 + \varepsilon C_L) \alpha \varepsilon^n.$$

where $C_L = \varepsilon \|L_0\|_{2R} e^{\varepsilon \|L_0\|_{2R}}$ (exploiting the fact that L is transported by $F^{[n]}$). We set $\kappa = \frac{1}{2}(1 + C_L)$ and $q = \alpha(2 + \varepsilon C_L)$ for brevity, and successive applications of this reasoning on $\|S_0\|_{R_k}$ generate

$$\|S_0\|_{R_1} \leq (\varepsilon\kappa)^n \|S_0\|_{2R} + \sum_{k=0}^{n-1} (\varepsilon\kappa) q \varepsilon^n \leq \left(\frac{\varepsilon}{2\varepsilon_0}\right)^n \|S_0\|_{2R} + 2q\varepsilon^n$$

assuming $\varepsilon_0 \leq 1/(2\kappa)$. Finally, $\|S_0\|_{R_1} = \mathcal{O}(\varepsilon^{n+1})$ thus $\|\Delta\|_{\varepsilon,R} = \mathcal{O}(\varepsilon^{n+1})$. This reasoning can be conducted on any time interval of the form $[\varepsilon k, \varepsilon(k+1)]$, proving that R is of size (ε^{n+1}) at all times.

□

CONVERGENCE UNIFORME POUR UN PROBLÈME DISSIPATIF

Ce chapitre reprend un article à paraître dans *Mathematics of Computation*, intitulé

A uniformly accurate numerical method for a class of dissipative systems,

co-écrit avec mes directeurs, Philippe CHARTIER et Mohammed LEMOU. Dans cet article, on construit un problème micro-macro pour une classe de problèmes à relaxation rapide, ce qui permet de résoudre le problème avec une précision uniforme d'ordre arbitraire. Le caractère bien posé de ce problème est prouvé en dressant un lien original avec les problèmes hautement oscillant pour lesquels des résultats existaient déjà.

2.1 Introduction

We are interested in problems of the form, for $x^\varepsilon(t) \in \mathbb{R}^{d_x}$ and $z^\varepsilon(t) \in \mathbb{R}^{d_z}$,

$$\begin{cases} \dot{x}^\varepsilon = a(x^\varepsilon, z^\varepsilon), & x^\varepsilon(0) = x_0, \\ \dot{z}^\varepsilon = -\frac{1}{\varepsilon}Az^\varepsilon + b(x^\varepsilon, z^\varepsilon), & z^\varepsilon(0) = z_0, \end{cases} \quad (2.1)$$

with $\varepsilon \in (0, 1]$ a small parameter, A a diagonal positive matrix with integer coefficients, and where a, b are respectively the x -component and the z -component of an analytic map f which smoothly depends on ε . We look for a solution $x^\varepsilon(t), z^\varepsilon(t)$, defined for $t \in [0, 1]$, irrespectively of the value of ε . The exact value of the right bound of the interval of definition of the solution, here 1, is somehow arbitrary, as it can be rescaled by changing the value of $\frac{1}{\varepsilon}\Lambda$. In the limit when ε goes to zero, the problem becomes stiff on the considered interval : in other words, the problem resorts to long-time integration as 1 becomes large compared to ε . In the sequel we shall more often write the equations in

compact form as

$$\dot{u}^\varepsilon = -\frac{1}{\varepsilon}\Lambda u^\varepsilon + f(u^\varepsilon), \quad u^\varepsilon(0) = u_0, \quad (2.2)$$

where $u = \begin{pmatrix} x \\ z \end{pmatrix}$, $\Lambda = \begin{pmatrix} 0 & 0 \\ 0 & A \end{pmatrix}$ and $f(u) = \begin{pmatrix} a(x, z) \\ b(x, z) \end{pmatrix}$. We set $d = d_x + d_z$ the dimension of u such that $u \in \mathbb{R}^d$. Note that x^ε may be zero-dimensional without impacting our results, or that it may include a component $\tilde{x}(t) = t$ such that f depends on t in a “hidden” manner. In contrast, it should be emphasized that we do not address the case where the map $u \mapsto f(u)$ is a differential operator and u lies in a functional space : the theory required for that situation is outside the scope of our theorems. Nonetheless, two of our examples are discretized hyperbolic partial differential equations (PDEs) for which the method is successfully applied, even though an additional specific treatment is required.

Problems of the form (2.2) recurrently appear in population dynamics (see [GHM94; AP96; SAAP00; CCS18]), where A accounts for migration (in space and/or age) and a and b account for both the demographic and inter-population dynamics. In this context, the factor $1/\varepsilon$ accounts for the fact that the migration dynamics is quantifiably faster than other dynamics involved.

When solving this kind of system numerically, problems arise due to the large range of values that ε can take. To be more specific, the error for standard methods of order $q > 1$ behave like

$$E_\varepsilon(\Delta t) \leq \min \left(C_q \frac{\Delta t^q}{\varepsilon^r}, C_s \Delta t^s \right),$$

for some positive constants C_q and C_s independent of ε and integers $s \leq q$ and $r \geq 0$. This forces very small values of Δt in order to achieve some accuracy and causes the computational cost of the simulation to increase greatly, often prohibitively so. Additionally, the order is reduced to s in the sense that ¹

$$\sup_{\varepsilon \in (0,1]} E_\varepsilon(\Delta t) \leq C \Delta t^s. \quad (2.3)$$

This behaviour is documented for instance in [HW96, Section IV.15] or in [HR07]. In order to ensure a given error bound, one must either accept this order reduction (if $s > 0$), as is done for asymptotic-preserving (AP) schemes [Jin99] by taking a modified time-step $\tilde{\Delta t} = \Delta t^{q/s}$, or use an ε -dependent time-step $\Delta t = \mathcal{O}(\varepsilon^{r/q})$.

1. In particular, the scheme cannot be any usual explicit scheme since it would require a stability condition of the form $\Delta t/\varepsilon < C$ with C independent of ε .

A common approach to circumvent this difficulty is to invoke the *center manifold theorem* (see [Vas63 ; Car82 ; Sak90]), which dictates the long-time behaviour of the system and presents useful characteristics for numerical simulations : the dimension of the system is reduced and the dynamics on the manifold is non-stiff. However, this approach does not allow to capture the *transient phase* of the solution, i.e. the solution in short time before it reaches the stable manifold. Insofar as one wishes to describe the system out of equilibrium, this is clearly unsatisfactory. Furthermore, even if the solution is exponentially (w.r.t. time) close to the manifold, the center manifold approximation is accurate up to a certain error $\mathcal{O}(\varepsilon^n)$, rendering it useless if ε is of the order of 1.

The strategy developed in this paper is based on a *micro-macro* decomposition of the problem in combination with the use of standard q^{th} -order *exponential Runge-Kutta* methods. It aims at deriving an overall scheme with an error $E_\varepsilon(\Delta t)$ that can be bounded from above independently of ε , that is to say

$$E_\varepsilon(\Delta t) \leq C \Delta t^q$$

for some positive constant C independent of ε . In order to construct the appropriate transformation of the original system, we first provide a systematic way to compute asymptotic models at any order in ε approaching the solution over the *whole interval of time*. We then use the defect of this approximation to compute the solution with usual explicit numerical schemes and *uniform* accuracy (i.e. the cost and error of the scheme must be independent of ε). This approach automatically overcomes the challenges posed by both extremes $\varepsilon \ll 1$ and $\varepsilon \sim 1$.

The aforementioned micro-macro decomposition is obtained by writing the solution u^ε of (2.2) as the following composition of maps

$$u^\varepsilon(t) = \Omega_{t/\varepsilon}^\varepsilon \circ \Gamma_t^\varepsilon \circ (\Omega_0^\varepsilon)^{-1}(u_0) \quad (2.4)$$

where $(\tau, u) \in \mathbb{R}_+ \times \mathbb{R}^d \mapsto \Omega_\tau^\varepsilon(u) \in \mathbb{R}^d$ is a change of variable ε -close to the map $(\tau, u) \mapsto e^{-\tau\Lambda}u$ and where $(t, u) \in [0, T] \times \mathbb{R}^d \mapsto \Gamma_t^\varepsilon(u)$ is the flow associated to a *non-stiff* autonomous vector field $u \mapsto F^\varepsilon(u)$, yet to be defined. The formal maps Ω^ε and F^ε are approached at an arbitrary order $n \in \mathbb{N}$ by $\Omega^{[n]}$ and $F^{[n]}$ respectively such that the equality

$$u^\varepsilon(t) = \Omega_{t/\varepsilon}^{[n]}(v^{[n]}(t)) + w^{[n]}(t) \quad (2.5)$$

holds true, where $v^{[n]}(t) = \Gamma_t^{[n]} \circ (\Omega_0^{[n]})^{-1}(u_0)$ and $w^{[n]}$ are respectively called the *macro* component and the *micro* component. A crucial feature of this decomposition is that $w^{[n]}$ remains of size $\mathcal{O}(\varepsilon^{n+1})$.

Now, the main contribution of this work is to prove that, using explicit exponential Runge-Kutta (ERK) schemes of order $n + 1$ (which can be found for instance in [HO05]), it is possible to approximate u^ε with *uniform accuracy* and at *uniform computational cost* with respect to ε . In other words, we prove that formula (2.3) holds with $s = q = n + 1$ and $r = 0$. More precisely, if $(t_i)_{0 \leq i \leq N}$ is a time-step grid of mesh-size Δt , and if (v_i) and (w_i) are computed numerically by applying the ERK method to the micro-macro decomposition, then there exists C independent of ε such that ($|\cdot|$ stands for the usual Euclidian norm)

$$\max_{0 \leq i \leq N} \left\{ |x^\varepsilon(t_i) - x_i| + \frac{1}{\varepsilon} |z^\varepsilon(t_i) - z_i| \right\} \leq C \Delta t^{n+1} \quad \text{with} \quad \begin{pmatrix} x_i \\ z_i \end{pmatrix} = \Omega_{t_i/\varepsilon}^{[n]}(v_i) + w_i.$$

We emphasize here the expected occurrence of the scaling factor $1/\varepsilon$ accounts for the fact that z becomes of size $\mathcal{O}(\varepsilon)$ after a time $\mathcal{O}(\varepsilon \log(1/\varepsilon))$. IMEX methods such as CNLF and SBDF (see [ARW95; ACM99; HS21]), which mix implicit and explicit parts are not the focus of the article, but their use is briefly discussed in Remark 2.2.9.

The present work is related to the recent paper [CCS16], where asymptotic expansions of the solution of (2.1) are constructed for the special case where A is the identity matrix. The theory developed therein is however of no relevance for the construction of micro-macro decompositions as it relies heavily on trees and associated elementary differentials which can hardly be computed in practice. Our approach actually shares more similarities with the one introduced for highly-oscillatory problems in [CLMV20] and later modified to become amenable for actual computations at any order [CLMZ20]. As a matter of fact, the technical arguments that sustain decomposition (2.4) are essentially adapted from [CCMM15] in a way that will be fully explained in Section 2.3.

The rest of the paper is organized as follows. In Section 2.2, we show our method to construct a micro-macro problem up to any order, and state our main result, i.e that solving this micro-macro problem with ERK schemes generates uniform accuracy on u^ε . In Section 2.3, we give proofs of all the results from Section 2.2. In Section 2.4, we present some techniques to adapt our method to discretized hyperbolic PDEs. Namely, we study a

relaxed conservation law and the telegraph equation, which can be respectively found for instance in [JX95] and [LM08]. In Section 2.5, we verify our theoretical result of uniform accuracy by successfully obtaining uniform convergence when numerically solving micro-macro problems obtained from a toy ODE and from the two aforementioned discretized PDEs.

2.2 Uniform accuracy from a decomposition

We start by considering the solution u of

$$\partial_t u^\varepsilon = -\frac{1}{\varepsilon} \Lambda u^\varepsilon + f(u^\varepsilon), \quad u^\varepsilon(0) = u_0 \in \mathbb{R}^d, \quad (2.6)$$

and write it as the composition of a *non-stiff* flow $(t, u) \mapsto \Gamma_t^\varepsilon(u)$ with a change of variable $(\tau, u) \mapsto \Omega_\tau^\varepsilon(u)$ with $\tau \in \mathbb{R}_+$,

$$u^\varepsilon(t) = \Omega_{t/\varepsilon}^\varepsilon \circ \Gamma_t^\varepsilon \circ (\Omega_0^\varepsilon)^{-1}(u_0). \quad (2.7)$$

In order for our approach to be rigorous, we start by introducing some definitions and assumptions in Subsection 2.2.1. We then present a way to approach these maps at any rank $n \in \mathbb{N}$ by $\Gamma^{[n]}$ and $\Omega^{[n]}$ in Subsection 2.2.2. This approximation is such that the error in (2.7) is of size $\mathcal{O}(\varepsilon^{n+1})$. In Subsection 2.2.3, we use this approximation to construct a micro-macro problem which can be solved numerically using standard IMEX schemes. This leads to our main result : reconstructing the solution u^ε of (2.6) from the numerical solution of the micro-macro problem yields an error *independent of ε* on u^ε . All proofs are delayed until Section 2.3.

2.2.1 Definitions and assumptions

Before proceeding, we must first state the assumptions on the vector field $u \mapsto f(u)$ and the operator Λ .

Assumption 2.2.1. *The matrix Λ is diagonal with nonnegative integer eigenvalues, and these values are nondecreasing when following the diagonal. In other words, $\Lambda = \text{Diag}(\lambda_1, \dots, \lambda_d)$ with $(\lambda_i)_{1 \leq i \leq d} \in \mathbb{N}^d$ and $\lambda_1 \leq \dots \leq \lambda_d$.*

Thanks to this assumption, we write $u = \begin{pmatrix} x \\ z \end{pmatrix}$, with (x, z) such that $\Lambda u = \begin{pmatrix} 0 \\ Az \end{pmatrix}$

for some A positive definite. The dimension of z may be zero without making our results invalid.

Assumption 2.2.2. *Let us set d_x and d_z the dimensions of x and z respectively. There exists a compact set $X_1 \subset \mathbb{R}^{d_x}$ and a radius $\check{\rho} > 0$ such that for every x in X_1 , the map $u \in \mathbb{R}^d \mapsto f(u) \in \mathbb{R}^d$ can be developed as a Taylor series around $\begin{pmatrix} x \\ 0 \end{pmatrix}$, and the series converges with a radius not smaller than $\check{\rho}$.*

It is therefore possible to naturally extend f to compact subsets of \mathbb{C}^d defined by

$$\mathcal{U}_\rho := \left\{ u \in \mathbb{C}^d ; \exists x \in X_1, \left| u - \begin{pmatrix} x \\ 0_{d_z} \end{pmatrix} \right| \leq \rho \right\},$$

for all $0 \leq \rho < \check{\rho}$ as it is represented by a Taylor series in $u \in \mathbb{C}^d$ on these sets. Here $|\cdot|$ is the natural extension of the Euclidian norm on \mathbb{R}^d to \mathbb{C}^d .

It may seem particularly restrictive to assume that the z -component of the solution u^ε of (2.2) stays in a neighborhood of 0, however this is somewhat ensured by the *center manifold theorem*. This theorem states that there exists a map $x \in \mathbb{R}^{d_x} \mapsto \varepsilon h^\varepsilon(x) \in \mathbb{R}^{d_x}$ smooth in ε and x , such that the manifold \mathcal{M} defined by

$$\mathcal{M} = \left\{ (x, z) \in \mathbb{R}^{d_x} \times \mathbb{R}^{d_z} : z = \varepsilon h^\varepsilon(x) \right\}$$

is a stable invariant for (2.1). It also states that all solutions $(x^\varepsilon, z^\varepsilon)$ of (2.1) converge towards it exponentially quickly, i.e. there exists $\mu > 0$ independent of ε such that

$$|z^\varepsilon(t) - \varepsilon h^\varepsilon(x^\varepsilon(t))| \leq C e^{-\mu t/\varepsilon}. \quad (2.8)$$

This means that the growth of z^ε is bounded by that of x^ε , and that after a time $t \geq \varepsilon \log(1/\varepsilon)$, $z^\varepsilon(t)$ is of size $\mathcal{O}(\varepsilon)$. Therefore it is credible to assume that z^ε stays somewhat close to 0. This is translated into a final assumption.

Assumption 2.2.3. *There exist two radii $0 < \rho_0 \leq \rho_1 < \check{\rho}$ and a closed subset $X_0 \subset X_1 \subset \mathbb{R}^{d_x}$ such that the initial condition $u_0 \in \mathbb{C}^d$ satisfies*

$$\min_{x \in X_0} \left| u_0 - \begin{pmatrix} x \\ 0_{d_z} \end{pmatrix} \right| \leq \rho_0,$$

and for all $\varepsilon \in (0, 1]$, Problem (2.6) is well-posed on $[0, 1]$ with its solution u^ε in \mathcal{U}_{ρ_1} .

Note that this is different to assuming that the initial data (x_0, z_0) is close to the center manifold. Indeed, the size of the initial condition is supposed independent of ε , therefore the distance from $z(0)$ to the center manifold is always $\mathcal{O}(1)$.

For $\rho \in [0, \check{\rho} - \rho_1)$, we define the sets

$$\mathcal{K}_\rho := \mathcal{U}_{\rho_1 + \rho} = \left\{ u \in \mathbb{C}^d; \exists x \in X_1, \left| u - \begin{pmatrix} x \\ 0 \end{pmatrix} \right| \leq \rho_1 + \rho \right\} \quad (2.9)$$

which help quantify the distance to the solution u^ε . By Assumption 2.2.3, the solution of (2.2) is in \mathcal{K}_0 at all time.

Definition 2.2.4. *We introduce some technical constants :*

(i) A radius $0 < R < \frac{1}{2}(\check{\rho} - \rho_1)$

(ii) An arbitrary rank p and a positive constant M such that for all $0 \leq \alpha, \beta \leq p + 2$ and all $\sigma \in [0, 6\|\Lambda\|]$,

$$\frac{\sigma^\beta}{\beta!} \left\| (\rho_1 + 2R)^\alpha \partial_u^\alpha f \right\| \leq M$$

Given a radius $0 \leq \rho \leq 2R$ and a map $(\tau, u) \in \mathbb{R}_+ \times \mathcal{K}_\rho \mapsto \psi_\tau(u)$, we define the norm,

$$\|\psi\|_\rho := \sup_{(\tau, u) \in \mathbb{R}_+ \times \mathcal{K}_\rho} |\psi_\tau(u)|. \quad (2.10)$$

If the map is furthermore p -times continuously differentiable w.r.t. τ , then we define

$$\|\psi\|_{\rho, p} := \max_{0 \leq \nu \leq p} \|\partial_\tau^\nu \psi\|_\rho. \quad (2.11)$$

2.2.2 Constructing the micro-macro problem

We assume that the vector field in (2.7) follows an autonomous vector field F^ε , i.e.

$$\frac{d}{dt} \Gamma_t^\varepsilon(u) = F^\varepsilon(\Gamma_t^\varepsilon(u)). \quad (2.12)$$

Injecting this and (2.7) into (2.6) and writing $v_0 = (\Omega_0^\varepsilon)^{-1}(u_0)$

$$(\partial_\tau + \Lambda) \Omega_{t/\varepsilon}^\varepsilon(\Gamma_t^\varepsilon(v_0)) = \varepsilon \left(f \circ \Omega_{t/\varepsilon}^\varepsilon(\Gamma_t^\varepsilon(v_0)) - \partial_u \Omega_{t/\varepsilon}^\varepsilon(\Gamma_t^\varepsilon(v_0)) \cdot F^\varepsilon(\Gamma_t^\varepsilon(v_0)) \right)$$

which by separation of scales t and t/ε generates the homological equation on Ω^ε , for all $(\tau, u) \in \mathbb{R}_+ \times K_\rho$,

$$(\partial_\tau + \Lambda)\Omega_\tau^\varepsilon(u) = \varepsilon(f \circ \Omega_\tau^\varepsilon(u) - \partial_u \Omega_\tau^\varepsilon(u) \cdot F^\varepsilon(u)). \quad (2.13)$$

It is furthermore possible to extract the vector field F^ε from this equation to get

$$F^\varepsilon = \langle \partial_u \Omega^\varepsilon \rangle^{-1} \langle f \circ \Omega^\varepsilon \rangle \quad (2.14)$$

where $\langle \cdot \rangle$ is defined by the following formula

$$\langle \psi \rangle := \frac{1}{2\pi} \int_0^{2\pi} e^{i\theta\Lambda} \psi_{i\theta} \, d\theta, \quad (2.15)$$

with the canonical definition $\psi_{i\theta} = \sum_{k \geq 0} e^{-ik\theta} \hat{\psi}_k$. To see this, we first observe that for an exponential series $\tau \in \mathbb{R}_+ \mapsto \psi_\tau$ which converges absolutely for $\tau = 0$, i.e. $\psi_\tau = \sum_{k \geq 0} e^{-k\tau} \hat{\psi}_k$ with $\sum_k \hat{\psi}_k$ absolutely converging, we can extract the coefficient $\hat{\psi}_k$ as the Fourier coefficient of $\psi_{i\theta}$ according to

$$\hat{\psi}_k = \frac{1}{2\pi} \int_0^{2\pi} e^{ik\theta} \psi_{i\theta} \, d\theta. \quad (2.16)$$

Therefore, we write equation (2.13) as follows

$$\partial_\tau(e^{\tau\Lambda}\Omega_\tau^\varepsilon)(u) = \varepsilon(e^{\tau\Lambda}f \circ \Omega_\tau^\varepsilon(u) - e^{\tau\Lambda}\partial_u \Omega_\tau^\varepsilon(u) \cdot F^\varepsilon(u)), \quad (2.17)$$

and apply the Fourier operator (2.16) to get

$$\widehat{\partial_\tau(e^{\tau\Lambda}\Omega_\tau^\varepsilon)(u)}_k = \varepsilon \left(\widehat{(e^{\tau\Lambda}f \circ \Omega_\tau^\varepsilon(u))}_k - \widehat{(\partial_u \Omega_\tau^\varepsilon(u) \cdot F^\varepsilon(u))}_k \right).$$

Taking now $k = 0$ and using definition (2.15) we get the expression (2.14). This framework of exponential series comes naturally thanks to Assumption 2.2.1.

The homological equation (2.13) has no unique solution in general, however we can approximate a solution as a *formal* solution as a power series in ε . This is generally the idea behind *normal forms*, where different methods have been developed (see [Mur06] for instance). Here we only consider a basic method to compute approximations $\Omega^{[n]}$ and $F^{[n]}$

of Ω^ε and F^ε at any rank $n \in \mathbb{N}$ by setting

$$(\partial_\tau + \Lambda)\Omega_\tau^{[n+1]} = \varepsilon \left(f \circ \Omega_\tau^{[n]} - \partial_u \Omega_\tau^{[n]} \cdot F^{[n]} \right). \quad (2.18)$$

with initial condition $\Omega_\tau^{[0]} = e^{-\tau\Lambda}$. Because we want $\Omega^{[n+1]}$ to be an exponential series, it appears that necessarily,

$$F^{[n]} = \langle \partial_u \Omega^{[n]} \rangle^{-1} \langle f \circ \Omega^{[n]} \rangle. \quad (2.19)$$

However these equations alone are not enough to obtain $\Omega^{[n]}$ at any order. Indeed, from (2.18), one gets

$$\Omega_\tau^{[n+1]} = e^{-\tau\Lambda} \Omega_0^{[n+1]} + \varepsilon \int_0^\tau e^{(\sigma-\tau)\Lambda} \left(f \circ \Omega_\sigma^{[n]} - \partial_u \Omega_\sigma^{[n]} \cdot F^{[n]} \right) d\sigma \quad (2.20)$$

meaning a choice of initial data $\Omega_0^{[n+1]}$ is needed. One could think that choosing $\Omega_0^{[n+1]} = \text{id}$ is the easiest choice, but computing (2.19) requires an inversion of $\langle \partial_u \Omega^\varepsilon \rangle$. Therefore we choose $\Omega_0^{[n+1]}$ such that $\langle \Omega^{[n+1]} \rangle = \text{id}$, i.e. for all $n \in \mathbb{N}$,

$$\Omega_0^{[n+1]} = \text{id} - \varepsilon \left\langle \int_0^\cdot e^{(\sigma-\cdot)\Lambda} \left(f \circ \Omega_\sigma^{[n]} - \partial_u \Omega_\sigma^{[n]} \cdot F^{[n]} \right) d\sigma \right\rangle \quad \text{thus} \quad F^{[n]} = \langle f \circ \Omega^{[n]} \rangle. \quad (2.21)$$

Now that we have a way to compute an approximate solution of (2.13), we introduce the error of approximation

$$\eta_\tau^{[n]} = \frac{1}{\varepsilon} (\partial_\tau + \Lambda) \Omega_\tau^{[n]} + \partial_u \Omega_\tau^{[n]} \cdot F^{[n]} - f \circ \Omega_\tau^{[n]}. \quad (2.22)$$

With these definitions, the maps $(\tau, u) \mapsto \Omega_\tau^{[n]}(u)$, $u \mapsto F^{[n]}(u)$ and $(\tau, u) \mapsto \eta_\tau^{[n]}$ have the following properties.

Theorem 2.2.5. *For n in \mathbb{N} , let us denote $r_n = R/(n+1)$ and $\varepsilon_n := r_n/16M$ with R and M from Definition 2.2.4. For all $\varepsilon > 0$ such that $\varepsilon \leq \varepsilon_n$, the maps $(\tau, u) \mapsto \Omega_\tau^{[n]}(u)$, $u \mapsto F^{[n]}(u)$ and $(\tau, u) \mapsto \eta_\tau^{[n]}(u)$ given by (2.20) and (2.21) are well-defined on $\mathbb{R}_+ \times \mathcal{K}_R$ and are analytic w.r.t. u . The change of variable $\Omega^{[n]}$ and the residue $\eta^{[n]}$ are both $p+1$ -times continuously differentiable w.r.t. τ . Moreover, with $\|\cdot\|_R$ and $\|\cdot\|_{R,p+1}$ given by (2.10)*

and (2.11), the following bounds are satisfied for all $0 \leq \nu \leq p + 1$,

$$\begin{aligned} (i) \quad & \left\| \Omega^{[n]} - e^{-\tau\Lambda} \right\|_R \leq 4\varepsilon M, & (ii) \quad & \left\| \partial_\theta^\nu \left[\Omega^{[n]} - e^{-\tau\Lambda} \right] \right\|_R \leq 8 \left(1 + \|\Lambda\| \right)^\nu \varepsilon M \nu! \\ (iii) \quad & \|F^{[n]}\|_R \leq 2M & (iv) \quad & \|\eta_\tau^{[n]}(u)\|_{R,p} \leq 2M \left(1 + \|\Lambda\| \right)^p \left(2\mathcal{Q}_p \frac{\varepsilon}{\varepsilon_n} \right)^n \end{aligned}$$

where $\|\cdot\|$ is the induced norm from \mathbb{R}^d to \mathbb{R}^d , and \mathcal{Q}_p is a p -dependent constant.

The proof will be treated in Subsection 2.3.1, and this results remains valid with the choice $\Omega_0^{[n]} = \text{id}$.

2.2.3 A result of uniform accuracy

Given a rank $n \in \mathbb{N}$, we now denote $v^{[n]}(t) := \Gamma_t^{[n]} \circ \left(\Omega_0^{[n]} \right)^{-1}(u_0)$ and inject the decomposition

$$u^\varepsilon(t) = \Omega_{t/\varepsilon}^{[n]}(v^{[n]}(t)) + w^{[n]}(t) \quad (2.23)$$

into Problem (2.6) in order to find an equation on $w^{[n]}$. The main interests of this decomposition can be roughly summarized as follows. First, the change of variable $\Omega_{t/\varepsilon}^{[n]}$ is known explicitly and the macro solution $v^{[n]}$ is smooth in ε , in the sense that time derivatives of $v^{[n]}$ at any order are uniformly bounded with respect to $\varepsilon \in (0, 1]$. Second, the micro part $w^{[n]}$ is less stiff than the original solution u^ε in the sense that its time derivatives, up to order $n + 1$, are uniformly bounded in ε . These important properties naturally allow the construction of numerical schemes on $v^{[n]}$ and $w^{[n]}$ that enjoy the *uniform accuracy*, i.e. in which the order of the numerical methods is independent of ε and is not degraded by the stiffness generated by the possibly small values of ε .

From decomposition (2.23) we obtain the following system

$$\begin{cases} \partial_t v^{[n]}(t) = F^{[n]}(v^{[n]}), \\ \partial_t w^{[n]}(t) = -\frac{1}{\varepsilon} \Lambda \left(\Omega_{t/\varepsilon}^{[n]}(v^{[n]}) + w^{[n]} \right) + f \left(\Omega_{t/\varepsilon}^{[n]}(v^{[n]}) + w^{[n]} \right) - \frac{d}{dt} \Omega_{t/\varepsilon}^{[n]}(v^{[n]}), \end{cases}$$

with initial conditions $v^{[n]}(0) = \left(\Omega_0^{[n]} \right)^{-1}(u_0)$ and $w^{[n]}(0) = 0$. By definition of $v^{[n]}$ and

using (2.22),

$$\begin{aligned}\frac{d}{dt}\Omega_{t/\varepsilon}^{[n]}(v^{[n]}(t)) &= \frac{1}{\varepsilon}\partial_\tau\Omega_{t/\varepsilon}^{[n]}(v^{[n]}) + \partial_u\Omega_{t/\varepsilon}^{[n]}(v^{[n]}) \cdot F^{[n]}(v^{[n]}) \\ &= -\frac{1}{\varepsilon}\Lambda\Omega_{t/\varepsilon}^{[n]}(v^{[n]}) + \eta_{t/\varepsilon}^{[n]}(v^{[n]}) + f(\Omega_{t/\varepsilon}^{[n]}(v^{[n]})).\end{aligned}$$

We get the micro-macro problem

$$\begin{cases} \partial_t v^{[n]}(t) = F^{[n]}(v^{[n]}), \end{cases} \quad (2.24a)$$

$$\begin{cases} \partial_t w^{[n]}(t) = -\frac{1}{\varepsilon}\Lambda w^{[n]} + f(\Omega_{t/\varepsilon}^{[n]}(v^{[n]}) + w^{[n]}) - f(\Omega_{t/\varepsilon}^{[n]}(v^{[n]})) - \eta_{t/\varepsilon}^{[n]}(v^{[n]}). \end{cases} \quad (2.24b)$$

with initial conditions $v^{[n]}(0) = (\Omega_0^{[n]})^{-1}(u_0)$, $w^{[n]}(0) = 0$. The properties of this micro-macro problem can be summed up as followed.

Theorem 2.2.6. *For all $n \in \mathbb{N}^*$, let us define $r_n = R/n$ and $\varepsilon_n := r_n/16M$, with R and M from Definition 2.2.4. For all $\varepsilon \leq \varepsilon_n$, Problem (2.24) is well-posed until some final time T_n independent of ε , and the following bounds are satisfied for all $t \in [0, T_n]$ and $0 \leq \nu \leq \min(n, p)$,*

$$\begin{aligned} (i) \quad & v^{[n]}(t) \in \mathcal{K}_R & (ii) \quad & |w^{[n]}(t)| \leq \frac{R}{4} \left(\frac{\varepsilon}{\varepsilon_n}\right)^{n+1} \\ (iii) \quad & |\partial_t^\nu E^{[n]}(t)| = \mathcal{O}(\varepsilon^{n-\nu}) & (iv) \quad & \|\partial_t^{\nu+1} E^{[n]}\|_{L^1} = \mathcal{O}(\varepsilon^{n-\nu}) \end{aligned}$$

where $E^{[n]} = \partial_t w^{[n]} + \frac{1}{\varepsilon}\Lambda w^{[n]}$.

Remark 2.2.7. *The attentive reader may notice that, while we made the computation of $F^{[n]}$ easy with (2.21), the initial condition of the macro part, $v^{[n]}(0) = (\Omega_0^{[n]})^{-1}(u_0)$, is not explicit. However, this system must be solved only once, while $F^{[n]}$ is used at every time-step. Furthermore, it is possible to compute an approximation of $v^{[n]}(0)$ explicitly up to $\mathcal{O}(\varepsilon^{n+1})$ using²*

$$v^{[n+1]}(0) = u_0 - \left(\Omega_0^{[n+1]} - \text{id}\right)(v^{[n]}(0)) + \mathcal{O}(\varepsilon^{n+2}) \quad (2.25)$$

with initialization $v^{[0]}(0) = u_0$. Because $\Omega_0^{[n+1]}$ is near-identity (up to $\mathcal{O}(\varepsilon)$), an error of

2. The above formula is a consequence of the behaviour of the error, $\Omega^{[n+1]} = \Omega^{[n]} + \mathcal{O}(\varepsilon^{n+1})$ (see [CLMV20]), therefore $v^{[n+1]}(0) = v^{[n]}(0) + \mathcal{O}(\varepsilon^{n+1})$. Injecting this last approximation in $v^{[n+1]}(0) = u_0 - (\Omega^{[n+1]} - \text{id})(v^{[n+1]}(0))$ generates the formula.

size ε^{n+1} on $v^{[n]}(0)$ will only translate in an error of size ε^{n+2} on $v^{[n+1]}(0)$.

We can now define approached initial conditions for the micro-macro problem iterating (2.25) at each rank n and truncating the $\mathcal{O}(\varepsilon^{n+2})$ term. The initial condition of the micro part becomes

$$w^{[n]}(0) = u_0 - \Omega_0^{[n]}(v_n) \quad (2.26)$$

which ensures $w^{[n]}(0) = \mathcal{O}(\varepsilon^{n+1})$, meaning our results are not jeopardised.

Using a standard explicit scheme to solve Problem (2.24) cannot work due to the term $\frac{1}{\varepsilon}\Lambda w^{[n]}$. This is why we focus on exponential schemes, which render this term non-problematic in terms of stability (see [MZ09]). Of course, the only use of these exponential schemes does not solve the problem of non-uniform order of accuracy however, as these schemes all reduce to order 1 when taking the supremum of the error for $\varepsilon \in (0, \varepsilon^*]$. This is where our micr-macro formulation plays a crucial role since it allows standard numerical schemes (like exponential Runge-Kutta schemes for instance) to *keep their order uniformly* in $\varepsilon \in (0, 1]$. It should be noted that exponential schemes are well-established and the formulas to implement them can be found for example in [HO05] up to the fourth-order.

The first-order Euler method applied to (2.2) would yield

$$u_{i+1} = e^{-\frac{\Delta t}{\varepsilon}\Lambda} u_i + \Delta t \, \varphi\left(-\frac{\Delta t}{\varepsilon}\Lambda\right) f(u_i)$$

with $\varphi(-h\Lambda) = \frac{1}{h} \int_0^h e^{-s\Lambda} ds$. Because Λ is diagonal, this type of integral is easy to compute. There is no computational drawback to exponential schemes in this case. Furthermore, for these schemes the error bound involves the “modified” norm

$$|u|_\varepsilon = \left| u + \frac{1}{\varepsilon}\Lambda u \right|. \quad (2.27)$$

This norm is interesting because after a short time $t \geq \varepsilon \log(1/\varepsilon)$, the z -component of the solution u^ε of (2.2) is of size ε , as evidenced by the center manifold theorem in (2.8). Using the norm $|\cdot|_\varepsilon$ somewhat rescales z^ε (but not x^ε) by ε^{-1} such that studying the error in this norm can be seen as a sort of “relative” error.

The following result asserts that, indeed, our micro-macro reformulation of the problem allows any numerical scheme of order p , namely exponential schemes, to enjoy the uniform accuracy property, with the same order p . A detailed presentation of exponential Runge-Kutta schemes can be found for instance in [HO05; HO04].

Theorem 2.2.8. *Under the assumptions of Theorem 2.2.6 and denoting $T_n \leq T$ a final time such that Problem (2.24) is well-posed on $[0, T_n]$. Given $(t_i)_{i \in \llbracket 0, N \rrbracket}$ a discretisation of $[0, T_n]$ of time-step $\Delta t := \max_i |t_{i+1} - t_i|$. computing an approximate solution (v_i, w_i) of (2.24) using an exponential Runge-Kutta scheme of order $q := \min(n, p) + 1$ yields a uniform error of order q , i.e.*

$$\max_{0 \leq i \leq N} |u^\varepsilon(t_i) - \Omega_{t_i/\varepsilon}^{[n]}(v_i) - w_i|_\varepsilon \leq C \Delta t^q \quad (2.28)$$

where C is independent of ε .

The left-hand side of this inequality involves $|\cdot|_\varepsilon$ and shall be called the modified error. It dominates the absolute error which uses $|\cdot|$.

Remark 2.2.9. *Only exponential schemes are considered here rather than for instance IMEX-BDF schemes which are sometimes preferred (as in [HS21]). The reason for this is twofold.*

First, as was mentioned already, iterations are easy to compute because of the diagonal nature of Λ . Second, the error bounds are generally better for these schemes. Indeed, an IMEX-BDF scheme of order q involves the L^1 norm of $\partial_t^{q+1} w^{[n]}$, which is worse than the L^1 norm of $\partial_t^q E^{[n]}$. The former is of size $\mathcal{O}(\varepsilon^{n-q})$ while the latter is of size $\mathcal{O}(\varepsilon^{n+1-q})$. We made the choice to prioritize methods of order $n+1$ rather than n .

2.3 Proofs of theorems from Section 2.2

2.3.1 Proof of Theorem 2.2.5 : properties of the decomposition

For some rank $n \in \mathbb{N}$, consider the change of variable $(\tau, u) \mapsto \Omega_\tau^{[n]}(u)$ given by (2.20) and (2.21). From a straightforward induction using Assumptions 2.2.1 and 2.2.2, it appears that this change of variable can be written as a *formal* exponential series,

$$\Omega_\tau^{[n]}(u) = \sum_{k \in \mathbb{N}} e^{-k\tau} \widehat{\Omega^{[n]}_k}(u).$$

This can be associated to a power series $\Xi^{[n]}(\xi; u) = \sum_{k \in \mathbb{N}} \xi^k \widehat{\Omega^{[n]}_k}(u)$, $\xi \in \mathbb{C}$, $|\xi| \leq 1$, which is entirely determined by its behaviour on the border, i.e. by the periodic map

$$\Phi_\theta^{[n]}(u) = \Xi^{[n]}(e^{i\theta}; u) = \Omega_{-i\theta}^{[n]}(u) = \sum_{k \in \mathbb{N}} e^{ik\theta} \widehat{\Omega^{[n]}_k}(u). \quad (2.29)$$

Differentiating $\Phi^{[n+1]}$ w.r.t. θ and identifying the coefficients in (2.18), we obtain a (still formal) homological equation on $\Phi^{[n]}$:

$$(\partial_\theta - i\Lambda)\Phi_\theta^{[n+1]} = -i\varepsilon \left(f \circ \Phi_\theta^{[n]} - \partial_u \Phi_\theta^{[n]} \cdot F^{[n]} \right). \quad (2.30)$$

The periodic defect $\delta_\theta^{[n]} = -i\eta_{-i\theta}^{[n]}$ satisfies

$$\delta_\theta^{[n]} = \frac{1}{\varepsilon} \left((\partial_\theta - i\Lambda)\Phi_\theta^{[n]} + i f \circ \Phi_\theta^{[n]} - i \partial_u \Phi_\theta^{[n]} \cdot F^{[n]} \right) \quad (2.31)$$

Note that these relations both use the identity

$$\sum_{k \in \mathbb{N}} \xi^k \widehat{f \circ \Omega^{[n]}}_k = f \left(\sum_{k \in \mathbb{N}} \xi^k \widehat{\Omega^{[n]}}_k \right) \quad (2.32)$$

which seems fairly evident, but requires the right-hand side of the equation to be well-defined for all $|\xi| \leq 1$.

Setting the filtered map $\widetilde{\Phi}_\theta^{[n]} = e^{-i\theta\Lambda}\Phi_\theta^{[n]}$, it satisfies

$$\partial_\theta \widetilde{\Phi}_\theta^{[n+1]} = \varepsilon \left(g_\theta \circ \widetilde{\Phi}_\theta^{[n]} - \partial_u \widetilde{\Phi}_\theta^{[n]} \cdot G^{[n]} \right) \quad (2.33)$$

with $g_\theta(u) = e^{-i\theta\Lambda} f(e^{i\theta\Lambda} u)$ and $G^{[n]} = iF^{[n]}$.

Property 2.3.1. *Assumptions 2.2.2 and 2.2.3 ensure the following properties, with R, M and p given in Definition 2.2.4 :*

- (i) *For all $\varepsilon \in (0, 1]$, the Cauchy problem $\partial_t y^\varepsilon = g_{t/\varepsilon}(y^\varepsilon)$, $y^\varepsilon(0) = u_0$ is well-posed in \mathcal{K}_0 up to some final time independent of ε .*
- (ii) *For all $\theta \in \mathbb{T}$, the function $u \mapsto g_\theta(u)$ is analytic from \mathcal{K}_{2R} to \mathbb{C}^d .*
- (iii) *For all $\sigma \in [0, 3]$,*

$$\forall 0 \leq \nu \leq p+2, \quad \frac{\sigma^\nu}{\nu!} \|\partial_\theta^\nu g\|_{\mathbb{T}, 2R} \leq M, \quad (2.34)$$

Initial condition (2.21) means that the periodic change of variable would be defined by

$$\widetilde{\Phi}_\theta^{[n+1]} = \text{id} + \varepsilon \left(T_\theta^{[n]} - \Pi(T^{[n]}) \right) \quad \text{and} \quad \Phi_\theta^{[n+1]} = e^{i\theta\Lambda} \widetilde{\Phi}_\theta^{[n+1]} \quad (2.35)$$

with Π the average³ and $T_\theta^{[n]} = \int_0^\theta \left(g_\sigma \circ \widetilde{\Phi}_\sigma^{[n]} - \partial_u \widetilde{\Phi}_\sigma^{[n]} \cdot G^{[n]} \right) d\sigma$. Because $\widetilde{\Phi}^{[n]}$ is periodic at

3. Explicitely, $\Pi(\varphi) = \frac{1}{2\pi} \int_0^{2\pi} \varphi_\sigma d\sigma$

all rank n , taking the average in (2.33) gives the vector field

$$G^{[n]} = \Pi(g \circ \widetilde{\Phi}^{[n]}). \quad (2.36)$$

This is known as *standard averaging*. We introduce norms on periodic maps akin to (2.10) and (2.11), namely for $0 \leq \rho \leq 2R$, given a periodic map $(\theta, u) \in \mathbb{T} \times \mathcal{K}_\rho \mapsto \varphi_\theta(u)$,

$$\|\varphi\|_{\mathbb{T}, \rho} := \sup_{(\theta, u) \in \mathbb{T} \times \mathcal{K}_\rho} |\varphi_\theta(u)| \quad \text{and} \quad \|\varphi\|_{\mathbb{T}, \rho, \nu} := \max_{0 \leq \alpha \leq \nu} \|\partial_\theta^\alpha \varphi_\theta(u)\|_{\mathbb{T}, \rho} \quad (2.37)$$

where the second norm assumes that φ is ν -times continuously differentiable w.r.t. θ . Then the following bounds are satisfied.

Theorem 2.3.2 (from [CLMV20] and [CCMM15]). *For $n \in \mathbb{N}$, let us denote $r_n = R/(n+1)$ and $\varepsilon_n := r_n/16M$. For all $\varepsilon > 0$ such that $\varepsilon \leq \varepsilon_n$, the maps $\Phi^{[n]}$ and $G^{[n]}$ are well-defined by (2.35) and (2.36). The change of variable $\Phi^{[n]}$ and the defect $\delta^{[n]}$ are both $(p+2)$ -times continuously differentiable w.r.t. θ , and $\Phi_0^{[n]}$ is invertible with analytic inverse on $\mathcal{K}_{R/4}$. Moreover, the following bounds are satisfied for $1 \leq \nu \leq p+1$,*

$$\begin{aligned} (i) \quad & \|\widetilde{\Phi}^{[n]} - \text{id}\|_{\mathbb{T}, R} \leq 4\varepsilon M \leq \frac{r_n}{4}, & (ii) \quad & \|\partial_\theta^\nu \widetilde{\Phi}^{[n]}\|_{\mathbb{T}, R} \leq 8\varepsilon M \nu! \\ (iii) \quad & \|G^{[n]}\|_{\mathbb{T}, R} \leq 2M & (iv) \quad & \|\widetilde{\delta}^{[n]}\|_{\mathbb{T}, R, p+1} \leq 2M \left(2\mathcal{Q}_p \frac{\varepsilon}{\varepsilon_n}\right)^n \end{aligned}$$

where $\widetilde{\Phi}_\theta^{[n]} = e^{-i\theta\Lambda} \Phi_\theta^{[n]}$ and $\widetilde{\delta}^{[n]} = e^{-i\theta\Lambda} \delta_\theta^{[n]}$ correspond to the filtered equation (2.33), and \mathcal{Q}_p is a p -dependent constant.

In order to prove Theorem 2.2.5, we show that the previous calculations of this section are rigorous rather than formal. Let us work by induction and assume that the negative modes of $\Phi^{[n]}$ vanish (this is true for $\Phi_\theta^{[0]} = e^{i\theta\Lambda}$ since Λ is positive semidefinite). Because $(\theta, u) \mapsto \Phi_\theta^{[n]}(u)$ is continuously differentiable w.r.t. θ , its Fourier series converges absolutely, thus $(\xi, u) \mapsto \Xi^{[n]}(\xi; u)$ is well-defined for all $|\xi| \leq 1$ and $u \in \mathcal{K}_R$. By maximum modulus principle,

$$\|\Omega^{[n]} - e^{-\tau\Lambda}\|_R \leq \sup_{|\xi| \leq 1, u \in \mathcal{K}_R} |\Xi^{[n]}(\xi; u) - \xi^\Lambda| \leq \|\Phi^{[n]} - e^{i\theta\Lambda}\|_{\mathbb{T}, R} \leq \|\widetilde{\Phi}^{[n]} - \text{id}\|_{\mathbb{T}, R}$$

The reasoning also stands for all derivatives $1 \leq \nu \leq p+1$,

$$\left\| \partial_\tau^\nu [\Omega^{[n]} - e^{-\tau\Lambda}] \right\| \leq \sup_{\xi, u} \left| (\xi \partial_\xi)^\nu [\Xi^{[n]}(\xi; u) - \xi^\Lambda] \right| \leq \left\| \partial_\theta^\nu [\Phi^{[n]} - e^{i\theta\Lambda}] \right\|_{\mathbb{T}, R}$$

and $\|\partial_\theta^\nu[\Phi^{[n]} - e^{i\theta\Lambda}]\|_{\mathbb{T},R} \leq (1 + \|\Lambda\|)^\nu \|\partial_\theta^\nu \widetilde{\Phi}^{[n]}\|_{\mathbb{T},R,\nu}$. Furthermore, for $u \in \mathcal{K}_R$, let $x \in X_1$ s.t. $\left|u - \begin{pmatrix} x \\ 0 \end{pmatrix}\right| \leq \rho_1 + R$. Then for all $|\xi| \leq 1$,

$$\left|\Xi^{[n]}(\xi; u) - \begin{pmatrix} x \\ 0 \end{pmatrix}\right| \leq \left|\Phi_\theta^{[n]}(u) - \begin{pmatrix} x \\ 0 \end{pmatrix}\right| \leq \left|\widetilde{\Phi}_\theta^{[n]}(u) - \begin{pmatrix} x \\ 0 \end{pmatrix}\right|,$$

since a multiplication by $e^{-i\theta\Lambda}$ has no influence on the norm, nor on $\begin{pmatrix} x \\ 0 \end{pmatrix}$. A triangle inequality yields

$$\left|\Xi^{[n]}(\xi; u) - \begin{pmatrix} x \\ 0 \end{pmatrix}\right| \leq |\widetilde{\Phi}^{[n]} - u| + \left|u - \begin{pmatrix} x \\ 0 \end{pmatrix}\right| < \rho_1 + 2R,$$

therefore $f(\Xi^{[n]}(\xi; u))$ is well-defined for all $|\xi| \leq 1$ and $u \in \mathcal{K}_R$, by expanding it into an absolutely converging series around $\begin{pmatrix} x \\ 0 \end{pmatrix}$, thereby justifying relations (2.32) and (2.31).

The maximum modulus principle can finally be applied to the couple $(\eta^{[n]}, \delta^{[n]})$ in order to obtain the last estimate of Theorem 2.2.5.

□

2.3.2 Proof of Theorem 2.2.6 : well-posedness of the micro-macro problem

This proof is in several parts : first we show that problem (2.24a) is well-posed, and use this result to show that the bound on $w^{[n]}$ is satisfied, thereby also proving that (2.24b) is well-posed. Finally we focus on the bounds on $E^{[n]}$.

Let us set $\varphi(v) = u_0 + v - \Omega_0^{[n]}(u_0 + v)$. Using Theorem 2.2.5, if $|v| \leq R/4$ then $|\varphi(v)| \leq R/4$. By Brouwer fixed-point theorem, there exists v^* such that $\varphi(v^*) = v^*$, i.e. $u^* \in \mathcal{K}_{R/4}$ such that $\Omega_0^{[n]}(u^*) = u_0$. Therefore $v^{[n]}(0) := u^* \in \mathcal{K}_{R/4}$.

Given $t > 0$ and assuming $v^{[n]}(s) \in \mathcal{K}_R$ for all $s \in [0, t]$, one can bound $v^{[n]}(t)$ using Theorem 2.2.5 :

$$|v^{[n]}(t) - v^{[n]}(0)| = \left| \int_0^t F^{[n]}(v^{[n]}(s)) \, ds \right| \leq 2Mt.$$

Setting $T_v := \frac{3R}{8M}$ ensures $|v^{[n]}(t) - v^{[n]}(0)| \leq 3R/4$, meaning that for all $t \in [0, T_v]$, $v^{[n]}(t)$ exists and is in \mathcal{K}_R . Again from Theorem 2.2.5, we deduce $\Omega_\tau^{[n]}(v^{[n]}(t)) \in \mathcal{K}_{5R/4}$.

Focusing now on $w^{[n]}$ and assuming for all $s \in [0, t]$, $|w^{[n]}(s)| \leq R/4$, the linear term $L^{[n]}(\tau, s, w^{[n]}(s))$ is bounded using a Cauchy estimate :

$$|L^{[n]}(\tau, s, w^{[n]}(s))| \leq \|\partial_u f\|_{3R/2} \leq \frac{\|f\|_{2R}}{2R - \frac{3}{2}R} \leq \frac{2M}{R}$$

using a Cauchy estimate. The integral form then gives the bounds

$$\begin{aligned} |w^{[n]}(t)| &\leq \left| \int_0^t e^{\frac{s-t}{\varepsilon}\Lambda} L^{[n]}(s/\varepsilon, s, w^{[n]}(s)) w^{[n]}(s) ds + \int_0^t e^{\frac{s-t}{\varepsilon}\Lambda} S^{[n]}(s/\varepsilon, s) ds \right| \\ &\leq \int_0^t \frac{2M}{R} |w^{[n]}(s)| ds + \left| \int_0^t e^{\frac{s-t}{\varepsilon}\Lambda} S^{[n]}(s/\varepsilon, s) ds \right| \end{aligned} \quad (2.38)$$

Using the notation of the previous subsection, $\tilde{\delta}_\theta^{[n]} = -ie^{-i\theta\Lambda}\eta_{-\theta}^{[n]}$, from which

$$\eta_\tau^{[n]}(u) = \sum_{k \in \mathbb{Z}} e^{-(k+\Lambda)\tau} c_k^{[n]}(u) \quad \text{with} \quad c_k^{[n]}(u) = \frac{1}{2\pi} \int_0^{2\pi} e^{-ik\theta} \tilde{\delta}_\theta^{[n]}(u) d\theta.$$

Since $\langle \eta^{[n]} \rangle = 0$, i.e. $c_0^{[n]} = 0$, it is possible to bound the source term in $w^{[n]}$ by

$$\begin{aligned} \left| \int_0^t e^{\frac{s-t}{\varepsilon}\Lambda} S^{[n]}(s/\varepsilon, s) ds \right| &\leq \left\| e^{-\frac{t}{\varepsilon}\Lambda} \right\| \int_0^t \sum_{k \in \mathbb{Z}^*} \left(e^{-k\frac{s}{\varepsilon}} \|c_k^{[n]}\|_{\mathbb{T}, R} \right) ds \\ &\leq \sum_{k \in \mathbb{Z}^*} \frac{\varepsilon}{k} \|c_k^{[n]}\|_{\mathbb{T}, R} \leq \varepsilon \left(\sum_{k \in \mathbb{Z}^*} \frac{1}{k^2} \right) \|\partial_\theta \tilde{\delta}^{[n]}\|_{\mathbb{T}, R} \end{aligned}$$

where $\|\cdot\|_{\mathbb{T}, R}$ is given by (2.37). Using Theorem 2.3.2, there exists a constant $M_n > 0$ such that for all $t \in [0, T_v]$,

$$\left| \int_0^t e^{\frac{s-t}{\varepsilon}\Lambda} S^{[n]}(s/\varepsilon, s) ds \right| \leq M_n \left(\frac{\varepsilon}{\varepsilon_n} \right)^{n+1}. \quad (2.39)$$

Using Gronwall's lemma in (2.38) with this inequality yields

$$|w^{[n]}(t)| \leq M_n e^{\frac{2M}{R}t} \left(\frac{\varepsilon}{\varepsilon_n} \right)^{n+1} \leq M_n e^{\frac{2M}{R}t}.$$

We now set $T_w > 0$ such that $M_n e^{\frac{2M}{R}T_w} \leq R/4$ (T_w may therefore depend on n , but does

not depend on ε) and

$$T_n = \min(T_v, T_w).$$

This ensures the well-posedness of (2.24) on $[0, T_n]$ as well as the size of $w^{[n]}$.

Finally, the results on $E^{[n]}$ are a direct consequence of the bounds on the linear term

$$\sup_{\alpha+\beta+\gamma \leq p+1} \|\partial_\tau^\alpha \partial_t^\beta \partial_u^\gamma L^{[n]}\| < +\infty$$

and on the source term

$$\sup_{0 \leq \alpha+\beta \leq p} \|\partial_\tau^\alpha \partial_t^\beta S^{[n]}\|_{L^\infty} = \mathcal{O}(\varepsilon^n), \quad \sup_{\substack{\beta \geq 1 \\ 1 \leq \alpha+\beta \leq p+1}} \|\partial_\tau^\alpha \partial_t^\beta S^{[n]}\|_{L^1} = \mathcal{O}(\varepsilon^{n+1}).$$

This stems directly from Cauchy estimates and Theorem 2.2.5.

□

2.3.3 Proof of Theorem 2.2.8 : uniform accuracy

The idea in this proof is to bound the errors on the macro part and micro part separately, using

$$\left| u^\varepsilon(t_i) - \Omega_{t_i/\varepsilon}^{[n]}(v_i) - w_i \right|_\varepsilon \leq \left| \Omega_{t_i/\varepsilon}^{[n]}(v^{[n]}(t_i)) - \Omega_{t_i/\varepsilon}^{[n]}(v_i) \right|_\varepsilon + \left| w^{[n]}(t_i) - w_i \right|_\varepsilon.$$

As the macro part $v^{[n]}$ involves no linear term, the scheme acts like any RK scheme on this part. Since $v^{[n]}$ and $F^{[n]}$ are non-stiff, the scheme is necessarily *uniformly* of order q , i.e.

$$\left| v^{[n]}(t_i) - v_i \right| \leq \Delta t^q \cdot t_i \cdot \|\partial_t^{q+1} v^{[n]}\|_{L^\infty}$$

using usual error bounds on RK schemes. The reader may notice that the absolute error involving $|\cdot|$ was used, not the modified error involving $|\cdot|_\varepsilon$. The results in [HO04] state that an exponential RK scheme of order q generates an error given by

$$\left| w^{[n]}(t_i) - w_i \right|_\varepsilon \leq C \Delta t^q \left(\|\partial_t^{q-1} E^{[n]}\|_\infty + \|\partial_t^q E^{[n]}\|_{L^1} \right). \quad (2.40)$$

The bounds on $E^{[n]} = \partial_t w^{[n]} + \frac{1}{\varepsilon} \Lambda w^{[n]}$ and its derivatives w.r.t. ε can be found in Theorem 2.2.6, rendering the computation of bounds on the error of the micro part straightforward. From Theorem 2.2.5.(i), $\Omega_\tau^{[n]}(u) = e^{-\tau \Lambda} u + \mathcal{O}(\varepsilon)$, therefore the error on $\Omega_{t/\varepsilon}^{[n]}(v^{[n]})$

is of the form

$$\Omega_{t_i/\varepsilon}^{[n]}(v^{[n]}(t_i)) - \Omega^{[n]}(v_i) = e^{-t_i\Lambda/\varepsilon}(v^{[n]}(t_i) - v_i) + \varepsilon r_i$$

where $v^{[n]}(t_i) - v_i$ and r_i are of size $t_i \cdot \Delta t^q$. The error can therefore be bounded, denoting $\|\cdot\|$ the induced norm from \mathbb{R}^d to \mathbb{R}^d ,

$$\left| \Omega_{t_i/\varepsilon}^{[n]}(v^{[n]}(t_i)) - \Omega^{[n]}(v_i) \right|_\varepsilon \leq \left(1 + \left\| \frac{t_i}{\varepsilon} \Lambda e^{-\frac{t_i}{\varepsilon} \Lambda} \right\| \right) |v^{[n]}(t_i) - v_i| + (\varepsilon + \|\Lambda\|) |r_i|.$$

From this we get the desired result on u^ε .

□

2.4 Application to some ODEs derived from discretized PDEs

In this section, we construct micro-macro problems for two *discretized* hyperbolic relaxation systems of the form

$$\begin{cases} \partial_t u + \partial_x \tilde{u} = 0 \\ \partial_t \tilde{u} + \partial_x u = \frac{1}{\varepsilon}(g(u) - \tilde{u}) \end{cases}$$

where g acts either as a differential operator on u (telegraph equation, Subsection 2.4.1), or as a scalar value function (relaxed conservation law, Subsection 2.4.2). These two problems may seem similar in theory, and the latter actually serves as a stepping stone to treat the former in [JPT98; JPT00], but we will treat them quite differently in practice. Some recent AP schemes with promising convergence have been developed for this type of problems in [BPR17; ADP20].

Let us insist that we only consider these problems *after discretization* (using either Fourier modes or an upwind scheme), yet even in a discrete framework, it will be apparent that a direct application of the method is impossible, often because of the apparition of a backwards heat equation. The goal of this section is precisely to present some possible workarounds to overcome the problems that appear. Should the reader wish to see a more detailed and direct application of our method, they can find one in Subsection 2.5.1.

2.4.1 The telegraph equation

A commonly studied equation in kinetic theory is the one-dimensional Goldstein-Taylor model, also known as the telegraph equation (see [JPT98; LM08], for instance). It can be written, for $(t, x) \in [0, T] \times \mathbb{R}/2\pi\mathbb{Z}$

$$\begin{cases} \partial_t \rho + \partial_x j = 0, \\ \partial_t j + \frac{1}{\varepsilon} \partial_x \rho = -\frac{1}{\varepsilon} j, \end{cases} \quad (2.41)$$

where ρ and j represent the mass density and the flux respectively. Using Fourier transforms in x , it is possible to represent a function $v(t, x)$ by

$$v(t, x) = \sum_{k \in \mathbb{Z}} v_k(t) e^{ikx}.$$

Considering a given frequency $k \in \mathbb{Z}$ the problem can be reduced to

$$\begin{cases} \partial_t \rho_k = -ik j_k, \\ \partial_t j_k = -\frac{1}{\varepsilon} (j_k + ik \rho_k). \end{cases}$$

Treating this problem using our method directly leads to dead-ends, therefore we will guide the reader through our reasoning navigating some of these dead-ends. This will lead to micro-macro decompositions of orders 0 and 1. These struggles can be seen as limitations of our approach, however we show that with only slight tweaks, it is possible to obtain an error of uniform order 2 using a standard exponential RK scheme. This result is summed up at the end of this subsection as Proposition 2.4.1.

In order to make a component $-\frac{1}{\varepsilon} z$ appear, it would be tempting to set $z_k = j_k + ik \rho_k$. This quantity would verify the following differential equation

$$\partial_t z_k = -\frac{1}{\varepsilon} z_k + k^2 z_k - ik^3 \rho_k.$$

Integrating this differential equation gives

$$z_k(t) = \exp\left(-\lambda \frac{t}{\varepsilon}\right) z_k(0) - ik^3 \int_0^t e^{(s-t)\lambda/\varepsilon} \rho_k(s) ds. \quad (2.42)$$

where $\lambda = 1 - \varepsilon k^2$. Because $\varepsilon \in (0, 1]$ and $k \in \mathbb{Z}$ should not be correlated, λ can take any value in $(-\infty, 1)$. For λ negative, this equation is unstable and cannot be solved

numerically using standard tools. To overcome this, we consider the stabilized change of variable instead

$$z_k = j_k + \frac{ik}{1 + \alpha \varepsilon k^2} \rho_k$$

where α is a positive constant which we shall calibrate as the study progresses. This is the same change of variable as before up to $\mathcal{O}(\varepsilon)$, but $ik\rho_k$ was regularized with an elliptic operator to help with high frequencies. The problem to solve becomes

$$\begin{cases} \partial_t \rho_k = -\frac{k^2}{1 + \alpha \varepsilon k^2} \rho_k - ik z_k, \\ \partial_t z_k = -\frac{1}{\varepsilon} z_k + \frac{k^2}{1 + \alpha \varepsilon k^2} z_k - \frac{ik^3}{1 + \alpha \varepsilon k^2} \left(\alpha + \frac{1}{1 + \alpha \varepsilon k^2} \right) \rho_k. \end{cases} \quad (2.43)$$

As in (2.42), the growth of z_k is given by $e^{-\lambda t/\varepsilon}$ if λ is defined by

$$\lambda = 1 - \frac{\varepsilon k^2}{1 + \alpha \varepsilon k^2} \in \left(1 - \frac{1}{\alpha}, 1 \right].$$

For stability reasons λ must be positive, therefore we shall choose $\alpha \geq 1$.

Let us set $u_k = (\rho_k, z_k)^T$ and $\Lambda = \text{Diag}(0, 1)$ such that $\partial_t u_k = -\frac{1}{\varepsilon} \Lambda u_k + f(u_k)$ with

$$f(u) = \begin{pmatrix} -\frac{k^2}{1 + \alpha \varepsilon k^2} u_1 - ik u_2 \\ \frac{k^2}{1 + \alpha \varepsilon k^2} u_2 - \frac{ik^3}{1 + \alpha \varepsilon k^2} \left(\alpha + \frac{1}{1 + \alpha \varepsilon k^2} \right) u_1 \end{pmatrix}. \quad (2.44)$$

In the upcoming study, we usually prefer the notation $f(\rho, z)$ rather than $f(u)$ so as to keep the distinction between both coordinates clear. Assuming $|k| \leq k_{\max}$, it is possible to bound $f(\rho_k, z_k)$ independently of k and of ε , allowing us to apply the method developed in this paper in order to approximate every ρ_k and j_k , and eventually $\rho(x, t)$ and $j(x, t)$. Note that no rigorous aspects of convergence in functional spaces are considered here – this will be treated in a forthcoming work. We omit the index k going forward for the sake of clarity.

The micro-macro method is initialized by setting the change of variable $\Omega_\tau^{[0]}(\rho, z) = (\rho, e^{-\tau} z)^T$. The vector field followed by the macro part $v^{[0]}$ is $F^{[0]}$ given by

$$F^{[0]}(\rho, z) = \hat{k}^2 \begin{pmatrix} -\rho \\ z \end{pmatrix} \quad \text{with} \quad \hat{k} = \frac{k}{\sqrt{1 + \alpha \varepsilon k^2}}. \quad (2.45)$$

This means that the macro variable $v^{[0]}(t)$ is given by

$$v^{[0]}(t) = \begin{pmatrix} e^{-\hat{k}^2 t} & 0 \\ 0 & e^{\hat{k}^2 t} \end{pmatrix} v^{[0]}(0).$$

Notice that the growth of $v_2^{[0]}(t)$ is in $e^{\hat{k}^2 t}$, which is akin to the heat equation in reverse time. This is problematic, as it is possible for \hat{k} to be quite big. For example with $k = 10, \alpha = 2$ and $\varepsilon = 10^{-2}$, one gets $e^{\hat{k}^2} \approx 3 \cdot 10^{14}$. However in order to obtain the solution of (2.41), $u_k(t) = \Omega_{t/\varepsilon}^{[0]}(v^{[0]}(t)) + w^{[0]}(t)$, we are only interested in $\Omega_{t/\varepsilon}^{[0]}(v^{[0]}(t))$ for the macro part, and $\eta_{t/\varepsilon}^{[0]}(v^{[0]}(t))$ for the micro part, which only depend on $e^{-\frac{t}{\varepsilon}\Lambda} v^{[0]}(t)$ as can be seen in the upcoming expression of $\eta^{[0]}$ and using $\Omega_\tau^{[0]}(u) = e^{-\tau\Lambda} u$. This means that the interesting quantity is

$$e^{-\frac{t}{\varepsilon}\Lambda} v^{[0]}(t) = \begin{pmatrix} e^{-\hat{k}^2 t} & 0 \\ 0 & e^{-(1-\varepsilon\hat{k}^2)\frac{t}{\varepsilon}} \end{pmatrix} v^{[0]}(0). \quad (2.46)$$

Recognizing $1 - \varepsilon\hat{k}^2 = \lambda > 0$ in this expression, it follows that $v_2^{[0]}$ is a decreasing function of time, therefore it is bounded uniformly for all t, k and ε . Because the exact computation of $e^{-\frac{t}{\varepsilon}\Lambda} v^{[0]}(t)$ is readily available, it is used during implementation, leaving only $w^{[0]}$ to be computed numerically using ERK schemes. Should the reader wish to conduct their own implementation, they should use the defect

$$\eta_\tau^{[0]}(\rho, z) = \begin{pmatrix} ike^{-\tau}z \\ \hat{k}^2 \left(\alpha + \frac{1}{1 + \alpha\varepsilon k^2} \right) ik\rho \end{pmatrix} = \eta_0^{[0]}(\rho, e^{-\tau}z).$$

By linearity of f , the micro variable $w^{[0]}$ follows the differential equation

$$\partial_t w^{[0]} = -\frac{1}{\varepsilon} \Lambda w^{[0]} + f(w^{[0]}) - \eta_0^{[0]} \left(e^{-\frac{t}{\varepsilon}\Lambda} v^{[0]}(t) \right), \quad w^{[0]}(0) = 0.$$

The rescaled macro variable $e^{-\frac{t}{\varepsilon}\Lambda} v^{[0]}(t)$ is given by relation (2.46) with initial condition $v^{[0]}(0) = u(0) = (\rho_k(0), z_k(0))^T$.

Extending our expansion to order 1 is not trivial either. Direct application of itera-

tions (2.20) yields

$$\Omega_\tau^{[1]}(\rho, z) = \begin{pmatrix} \rho + \varepsilon i k e^{-\tau} z \\ e^{-\tau} z - \varepsilon \hat{k}^2 \left(\alpha + \frac{1}{1 + \alpha \varepsilon k^2} \right) i k \rho \end{pmatrix}$$

from which the vector field for the macro part is

$$F^{[1]}(\rho, z) = \hat{k}^2 \left(1 + \varepsilon k^2 \left(\alpha + \frac{1}{1 + \alpha \varepsilon k^2} \right) \right) \begin{pmatrix} -\rho \\ z \end{pmatrix}.$$

Following the same reasoning as before, one should study the evolution of the z -component of the rescaled macro variable $e^{-\frac{t}{\varepsilon} \Lambda} v^{[1]}(t)$. This evolution is in $e^{-\tilde{\lambda} t / \varepsilon}$ where $\tilde{\lambda} = 1 - \varepsilon \hat{k}^2 \left(1 + \varepsilon k^2 \left(\alpha + \frac{1}{1 + \alpha \varepsilon k^2} \right) \right)$. Studying $\tilde{\lambda}$ as a function of εk^2 in \mathbb{R}_+ shows that it is negative for $\varepsilon k^2 > 1$, whatever the value of $\alpha \geq 1$.

To circumvent this, we replace ε by $\frac{\varepsilon}{1 + \alpha \varepsilon k^2}$ in iterations (2.20). This adds terms of order ε^2 in the definition of $\Omega^{[1]}$ that do not modify any properties of the micro-macro decomposition but it regularises the problem. Specifically, we define

$$\Omega_0^{[1]}(\rho, z) = \begin{pmatrix} \rho + \frac{\varepsilon}{1 + \alpha \varepsilon k^2} i k z \\ z - \frac{\varepsilon}{1 + \alpha \varepsilon k^2} \hat{k}^2 \left(\alpha + \frac{1}{1 + \alpha \varepsilon k^2} \right) i k \rho \end{pmatrix}, \quad (2.47)$$

from which we get the vector field

$$F^{[1]}(\rho, z) = \hat{k}^2 \left(1 + \varepsilon \hat{k}^2 \left(\alpha + \frac{1}{1 + \alpha \varepsilon k^2} \right) \right) \begin{pmatrix} -\rho \\ z \end{pmatrix}.$$

This time also, the identities $\Omega_\tau^{[1]}(u) = \Omega_0^{[1]}(e^{-\tau \Lambda} u)$ and $\eta_\tau^{[1]}(u) = \eta_0^{[1]}(e^{-\tau \Lambda} u)$ are satisfied, therefore the interesting variable is $e^{-\frac{t}{\varepsilon} \Lambda} v^{[1]}(t)$. The quantity dictating its growth is

$$\tilde{\lambda} = 1 - \varepsilon \hat{k}^2 \left(1 + \varepsilon \hat{k}^2 \left(\alpha + \frac{1}{1 + \alpha \varepsilon k^2} \right) \right)$$

which is positive for all $\varepsilon k^2 \in \mathbb{R}_+$ if and only if $\alpha \geq 2$. As with the expansion of order 0, the macro variable should be rescaled and computed exactly. The micro part $w^{[1]}$ is given

by the differential equation

$$\partial_t w^{[1]} = -\frac{1}{\varepsilon} \Lambda w^{[1]} + f(w^{[1]}) - \eta_0^{[1]} \left(e^{-\frac{t}{\varepsilon} \Lambda} v^{[1]}(t) \right), \quad w^{[1]}(0) = u_k(0) - \Omega_0^{[1]} \left(v^{[1]}(0) \right) \quad (2.48)$$

where, writing $\hat{I} = (1 + \alpha \varepsilon k^2)^{-1}$,

$$\eta_\tau^{[1]}(\rho, z) = ik \cdot \varepsilon \hat{k}^2 \left(\alpha + \hat{I} \left(2 + \varepsilon \hat{k}^2 (\alpha + \hat{I}) \right) \right) \begin{pmatrix} e^{-\tau} z \\ \hat{k}^2 (\alpha + \hat{I}) \rho \end{pmatrix} \quad (2.49)$$

$$\text{and } v^{[1]}(0) = \begin{pmatrix} \rho_k(0) - \varepsilon \hat{I} i k z_k(0) \\ z_k(0) + \varepsilon \hat{k}^2 (\alpha + \hat{I}) i k \rho_k(0) \end{pmatrix}. \quad (2.50)$$

We approached the initial condition using Remark 2.2.7, but an exact computation of the exact initial condition $(\Omega_0^{[1]})^{-1}(u_0)$ is possible, as the map $u \mapsto \Omega_0^{[1]}(u)$ is linear.

Proposition 2.4.1. *Given a maximum frequency $k_{\max} > 0$ and a scalar $\alpha \geq 2$, and assuming $|k| \leq k_{\max}$, the solution u_k of problem (2.43) can be decomposed into*

$$u_k(t) = \Omega_0^{[1]} \left(e^{-\frac{t}{\varepsilon} \Lambda} v^{[1]}(t) \right) + w^{[1]}(t)$$

where $\Omega_0^{[1]}$ is given by (2.47) and $w^{[1]}(t) = \mathcal{O}(\varepsilon^2)$. The macro component $v^{[1]}$ is given by

$$e^{-\frac{t}{\varepsilon} \Lambda} v^{[1]}(t) = \begin{pmatrix} e^{-K^{[1]}t} & 0 \\ 0 & e^{-(1-\varepsilon K^{[1]})\frac{t}{\varepsilon}} \end{pmatrix} v^{[1]}(0)$$

with $K^{[1]} = \hat{k}^2 \left(1 + \varepsilon \hat{k}^2 \left(\alpha + \frac{1}{1+\alpha \varepsilon k^2} \right) \right)$, $\hat{k} = \frac{k}{\sqrt{1+\alpha \varepsilon k^2}}$ and $v^{[1]}(0)$ is either $(\Omega_0^{[k]})^{-1}(u_k(0))$ or its approximation (2.50). The micro component $w^{[1]}$ is the solution to

$$\partial_t w^{[1]} = -\frac{1}{\varepsilon} \Lambda w^{[1]} + f(w^{[1]}) - \eta_0^{[1]} \left(e^{-\frac{t}{\varepsilon} \Lambda} v^{[1]}(t) \right), \quad w^{[1]}(0) = u_k(0) - \Omega_0^{[k]} \left(v^{[1]}(0) \right)$$

with f and $\eta_0^{[1]}$ given respectively by (2.44) and (2.49). With these definitions, $w^{[1]}$ can be computed with a uniform error of order 2, therefore u_k can be computed with a uniform error of order 2.

The reader may notice that only a finite number of modes is considered. This is required so that there exists a bound uniform w.r.t. k and ε on the micro part of the problem (2.48)

in order to apply our method. This is amenable to a CFL condition, i.e. some stiffness still exists due to the nature of the problem, but this stiffness is independent of ε . This is what we mean by uniform accuracy.

2.4.2 Relaxed conservation law

Our second test case is a hyperbolic problem for $(t, x) \in [0, T] \times \mathbb{R}/2\pi\mathbb{Z}$,

$$\begin{cases} \partial_t u + \partial_x \tilde{u} = 0, \\ \partial_t \tilde{u} + \partial_x u = \frac{1}{\varepsilon}(g(u) - \tilde{u}), \end{cases} \quad (2.51)$$

with smooth initial conditions $u(0, x)$ and $\tilde{u}(0, x)$. This is a stiffly relaxed conservation law, as presented in [JX95].

Remark 2.4.2. *Note that the assumption that Λ has integer coefficients is restrictive in this case. One may want to consider the equation on the second coordinate to be*

$$\partial_t \tilde{u} + \partial_x u = \frac{\sigma(x)}{\varepsilon}(b(x)u - \tilde{u})$$

as is done in [HS21], however this is not possible with our method. Perhaps a link can be made with the highly-oscillatory study [CJL17] where the “phase” t/ε in (2.4) is replaced by a space-dependent function $\varphi(t, x)/\varepsilon$.

Without the space-derivatives, this problem is straightforward : One can simply set $x = u$ and $z = g(u) - \tilde{u}$. Here new difficulties appear. For instance, in order to proceed, we require the following condition to be met :

$$|g'(u)| < 1 \quad (2.52)$$

This is a known stability condition when deriving asymptotic expansions for this kind of problem.

We start by discretising this system in space with $N > 0$ points. Going forward, $(x_j)_{j \in \mathbb{Z}/N\mathbb{Z}}$ denotes a fixed uniform discretisation of $\mathbb{R}/2\pi\mathbb{Z}$, of mesh size $\Delta x := 2\pi/N$. We define the vectors $U = (u_j)_j, \tilde{U} = (\tilde{u}_j)_j$ and, given a vector $V = (v_j)_j$ of size N , $g(V) = (g(v_j))_j$. For simplicity, $u_j(t)$ is the approximation of $u(t, x_j)$, and the same goes for \tilde{u} . We denote D the matrix of centered finite differences and L the standard discrete

Laplace operator, which is to say

$$DV = \left(\frac{1}{2\Delta x} (v_{j+1} - v_{j-1}) \right)_j \quad \text{and} \quad LV = \left(\frac{1}{\Delta x^2} (v_{j+1} - 2v_j + v_{j-1}) \right)_j$$

Using an upwind scheme after diagonalising problem (2.51) yields

$$\begin{cases} \partial_t U + D\tilde{U} - \frac{\Delta x}{2} LU = 0, \\ \partial_t \tilde{U} + DU - \frac{\Delta x}{2} L\tilde{U} = \frac{1}{\varepsilon} (g(U) - \tilde{U}). \end{cases} \quad (2.53)$$

Setting $U_1 = U$ and $U_2 := \tilde{U} - g(U_1)$, and neglecting the terms involving L for clarity, this problem becomes

$$\begin{cases} \partial_t U_1 = -D(U_2 + g(U_1)), \\ \partial_t U_2 = -\frac{1}{\varepsilon} U_2 + g'(U_1) D U_2 - T(U_1) \end{cases} \quad (2.54)$$

where we defined $T(U_1) := D U_1 - g'(U_1) D g(U_1)$. From this, our method can be applied, but precautions must be taken in order to avoid having to solve the heat equation in backwards time. Therefore we set

$$\Omega_\tau^{[1]}(U_1, U_2) = \begin{pmatrix} U_1 + \varepsilon(1 - 2\varepsilon D^2)^{-1} D U_2 \\ e^{-\tau} U_2 - \varepsilon T(U_1) \end{pmatrix}.$$

Similarly to the manipulations for the telegraph equation, we multiplied ε by $(I_N - 2\varepsilon D^2)^{-1}$, but this time only for the first component. Writing $\widetilde{D} = (I_N - 2\varepsilon D^2)^{-1} D$, the associated vector field is

$$F^{[1]}(U_1, U_2) = \begin{pmatrix} -Dg(U_1) + \varepsilon D T(U_1) \\ g'(U_1) D U_2 - \varepsilon T'(U_1) \widetilde{D} U_2 - \varepsilon^2 g''(U_1) (T(U_1), \widetilde{D} U_2) \end{pmatrix}.$$

It is possible to obtain $\Omega^{[0]}$ and $F^{[0]}$ by neglecting the terms of order ε and above in the

expressions above.

Remark 2.4.3. Remember that for the telegraph equation, the macro variable $v^{[1]}(t)$ needed to be rescaled by $e^{-t\Lambda/\varepsilon}$. This is not the case here : In the limit $\Delta x \rightarrow 0$, the macro variable $v^{[1]} = (\bar{u}_1, \bar{u}_2)^T$ is given by

$$\begin{cases} \partial_t \bar{u}_1 = -\partial_x \left[g(\bar{u}_1) - \varepsilon(1 - g'(\bar{u}_1)^2) \partial_x \bar{u}_1 \right], \\ \partial_t \bar{u}_2 = g'(\bar{u}_1) \partial_x \bar{u}_2 - (1 - g'(\bar{u}_1)^2) \cdot (1 - 2\varepsilon \partial_x^2)^{-1} \varepsilon \partial_x^2 \bar{u}_2 + \varepsilon \phi^\varepsilon(\bar{u}_1, \widetilde{D} \bar{u}_2) \end{cases}$$

with $\widetilde{D} = (1 - 2\varepsilon \partial_x^2)^{-1} \partial_x$ and $\phi^\varepsilon(u_1, u_2) = g''(u_1) (2g'(u_1) - \varepsilon(1 - g'(u_1)^2) \partial_x u_1) u_2$. The operator $(1 - 2\varepsilon \partial_x^2)^{-1} \varepsilon \partial_x^2$ is bounded, therefore \bar{u}_2 is well-defined. The equation on \bar{u}_1 is a well-known result. If ε was also relaxed in the U_2 -component of $\Omega^{[1]}$, there might be no need for condition (2.52) but the result would be different.

Because D^2 is sparse, it is not too costly to compute $(I_N - \varepsilon D^2)^{-1}$, however the conditioning may depend on the ratio between ε and Δx . Indeed, studying the eigenvalues of D reveals that the eigenvalues $(\mu_k)_{k \in \mathbb{Z}/N\mathbb{Z}}$ of $I_N - \varepsilon D^2$ are

$$\mu_k = 1 + \frac{\varepsilon}{\Delta x^2} \sin^2 \left(2\pi \frac{k}{N} \right) \quad (2.55)$$

meaning that for N big, the conditioning is approximately $1 + \varepsilon/\Delta x^2$. Therefore, for ε big and Δx small, this inversion can become very costly, even though the cost remains bounded independently of ε .

Obtaining the defects of order 0 and 1 from these expressions presents no difficulty. For $\eta^{[1]}$, we separate here the U_1 -component and the U_2 -component for clarity.

$$\eta_\tau^{[0]}(U_1, U_2) = \begin{pmatrix} e^{-\tau} D U_2 \\ T(U_1) \end{pmatrix},$$

$$\begin{aligned} \eta_0^{[1]}(U_1, U_2)_{U_1} &= D(g(U_1 + \varepsilon \widetilde{D} W) - g(U_1)) + (D - \widetilde{D}) U_2 \\ &\quad + \varepsilon \widetilde{D} \left(g'(U_1) D W - \varepsilon T'(U_1) \widetilde{D} W - \varepsilon^2 g''(U_1) (T(U_1), \widetilde{D} W) \right), \end{aligned} \quad (2.56a)$$

$$\begin{aligned}
 \eta_0^{[1]}(U_1, U_2)_{U_2} = & -\left(g'(U_1 + \varepsilon \widetilde{D}U_2) - g'(U_1)\right)DU_2 \\
 & + T(U_1 + \varepsilon \widetilde{D}U_2) - T(U_1) - \varepsilon T'(U_1)\widetilde{D}U_2 \\
 & + \varepsilon g'(U_1 + \varepsilon \widetilde{D}U_2)DT(U_1) - \varepsilon^2 g''(U_1)(\widetilde{D}U_2, T(U_1)) \\
 & + \varepsilon T'(U_1)(Dg(U_1) - \varepsilon T(U_1)).
 \end{aligned} \tag{2.56b}$$

The values of $\eta_\tau^{[1]}(U_1, U_2)$ can be recovered using the identity

$$\eta_\tau^{[1]}(U_1, U_2) = \eta_0^{[1]}(U_1, e^{-\tau}U_2).$$

2.5 Numerical simulations

In this section we shall demonstrate our results by confirming the theoretical convergence rates of exponential Runge-Kutta (ERK) schemes from [HO05]. We also use these schemes on the original problem (2.1), thereby exhibiting the problem of order reduction.

In Subsection 2.5.1 we study a toy model and a PDE-inspired problem with some non-linearity, for which we compute the micro-macro expansion up to order 2. In Subsection 2.5.2, we showcase the results of uniform convergence for the partial differential equations of Section 2.4. For these, the exact solution shall not take into account the error in space, i.e. it will be the solution to the discretized problem. Finally in Subsection 2.5.3, we share our thoughts on some remaining avenues of research following this paper.

2.5.1 Application to some ODEs

Slowly oscillating toy problem

We first study an “oscillating” problem presented in [CCS16] which demonstrates a possible use of the method when studying non-linear problems :

$$\begin{cases} \dot{x} = (1 - z) \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix} x \\ \dot{z} = -\frac{1}{\varepsilon}z + x_1^2 x_2^2 \end{cases} \tag{2.57}$$

with initial conditions $x_0 = (0.1, 0.7)^T$ and $z_0 = 0.05$, and final time $T = 1$. This is of the

form $\partial_t u = -\frac{1}{\varepsilon}\Lambda u + f(u)$ when setting

$$u = \begin{pmatrix} x \\ z \end{pmatrix}, \quad \Lambda = \text{Diag}(0, 0, 1) \quad \text{and} \quad f(u) = \begin{pmatrix} -(1 - u_3)u_2 \\ (1 - u_3)u_1 \\ (u_1 u_2)^2 \end{pmatrix}.$$

The macro part of our micro-macro decomposition is built by solving iterations on the homological equation

$$(\partial_\tau + \Lambda)\Omega_\tau^{[n+1]} = \varepsilon (f \circ \Omega_\tau^{[n]} - \partial_u \Omega_\tau^{[n]} F^{[n]}) \quad (2.58)$$

where $F^{[n]} = \langle f \circ \Omega^{[n]} \rangle$ with $\langle \cdot \rangle$ the projector on the $e^{-\tau\Lambda}$ -component parallel to the other components of the exponential series. We choose the initial condition $\Omega_\tau^{[0]} = e^{-\tau\Lambda}$ and closure condition $\langle \Omega^{[0]} \rangle = e^{-\tau\Lambda}$. The first iteration yields

$$\Omega_\tau^{[1]}(x, z) = \begin{pmatrix} x_1 - \varepsilon e^{-\tau} x_2 z \\ x_2 + \varepsilon e^{-\tau} x_1 z \\ e^{-\tau} z + \varepsilon (x_1 x_2)^2 \end{pmatrix} \quad \text{and} \quad F^{[1]}(x, z) = \begin{pmatrix} -(1 - \varepsilon (x_1 x_2)^2) x_2 \\ (1 - \varepsilon (x_1 x_2)^2) x_1 \\ 2\varepsilon x_1 x_2 z (x_1^2 - x_2^2) \end{pmatrix}.$$

In order to compute the second order decomposition, one must compute the difference $T^{[1]} = f \circ \Omega^{[1]} - \partial_u \Omega^{[1]} F^{[1]}$, which is also used to compute the defect $\delta^{[1]} = \frac{1}{\varepsilon}(\partial_\tau + \Lambda)\Omega^{[1]} - T^{[1]}$. From a direct calculation this writes,

$$T_\tau^{[1]}(x, z) = \begin{pmatrix} e^{-\tau} z (x_2 + \varepsilon e^{-\tau} x_1 z + 2\varepsilon^2 x_1 x_2^2 (x_1^2 - x_2^2)) \\ -e^{-\tau} z (x_1 - \varepsilon e^{-\tau} x_2 z - 2\varepsilon^2 x_1^2 u_2 (x_1^2 - x_2^2)) \\ Z_0 + \varepsilon Z_1 + \varepsilon^2 Z_2 \end{pmatrix}$$

where for clarity we defined

$$Z_0 = \left(x_1^2 + \varepsilon^2 e^{-2\tau} (x_2 z)^2 \right) \left(x_2 + \varepsilon^2 e^{-2\tau} (x_1 z)^2 \right),$$

$$Z_1 = -2x_1 x_2 (x_1^2 - x_2^2) \left(1 - \varepsilon (x_1 x_2)^2 + \varepsilon e^{-3\tau} z^3 \right) \quad \text{and} \quad Z_2 = -e^{-2\tau} (2u_1 u_2 u_3)^2.$$

To compute the expansion of order 2, we truncate terms of order ε^2 and above in $T^{[1]}$

(which will not impact results of uniform accuracy) and solve (2.58). This yields ⁴

$$\Omega_\tau^{[2]}(x, z) = \begin{pmatrix} x_1 - \varepsilon e^{-\tau} x_2 z - \frac{1}{2} \varepsilon^2 e^{-2\tau} z^2 x_1 \\ x_2 + \varepsilon e^{-\tau} x_1 z - \frac{1}{2} \varepsilon^2 e^{-2\tau} z^2 x_2 \\ z + \varepsilon (x_1 x_2)^2 - 2\varepsilon^2 x_1 x_2 (x_1^2 - x_2^2) \end{pmatrix},$$

$$F^{[2]}(x, z) = \begin{pmatrix} x_2(-1 + \varepsilon (x_1 x_2)^2 - 2\varepsilon^2 x_1 x_2 (x_1^2 - x_2^2)) \\ x_1(1 - \varepsilon (x_1 x_2)^2 + 2\varepsilon^2 x_1 x_2 (x_1^2 - x_2^2)) \\ 2\varepsilon z x_1 x_2 (x_1^2 - x_2^2) \end{pmatrix}.$$

The defect $\eta^{[2]}$ is obtained using relation (2.22) or by computing $\delta^{[2]}$ and identifying the Fourier coefficients.

Remark 2.5.1. *It is possible to find an approximation of the center manifold $x \mapsto \varepsilon h^\varepsilon(x)$ by taking the limit $\tau \rightarrow \infty$ of the z -component of $\Omega^{[k]}$. For example here*

$$\varepsilon h^\varepsilon(x) = \varepsilon (x_1 x_2)^2 - 2\varepsilon^2 x_1 x_2 (x_1^2 - x_2^2) + \mathcal{O}(\varepsilon^3).$$

This coincides with the results in [CCS16].

We remind the reader that the problem that is solved at times $(t_i)_{0 \leq i \leq N}$ is

$$\begin{cases} \partial_t v^{[k]}(t) = F^{[k]}(v^{[k]}), \\ \partial_t w^{[k]}(t) = -\frac{1}{\varepsilon} \Lambda w^{[k]} + f\left(\Omega_{t/\varepsilon}^{[k]}(v^{[k]}) + w^{[k]}\right) - f\left(\Omega_{t/\varepsilon}^{[k]}(v^{[k]})\right) - \eta_{t/\varepsilon}^{[k]}(v^{[k]}), \end{cases}$$

with $k = 1, 2$. This yields vectors $(v_i) \approx (v^{[k]}(t_i))$ and $(w_i) \approx (w^{[k]}(t_i))$, from which an approximation $u_i \approx u^\varepsilon(t_i)$ is then obtained by setting $u_i = \Omega_{t_i/\varepsilon}^{[k]}(v_i) + w_i$. Initial conditions $v^{[k]}(0)$ and $w^{[k]}(0)$ are computed using Remark 2.2.7.

4. It has been pointed out to the authors that the same result is obtained using nonlinear coordinate transforms described in [Rob14]. Some normal form methods compiled in [Mur06] also yield this result.

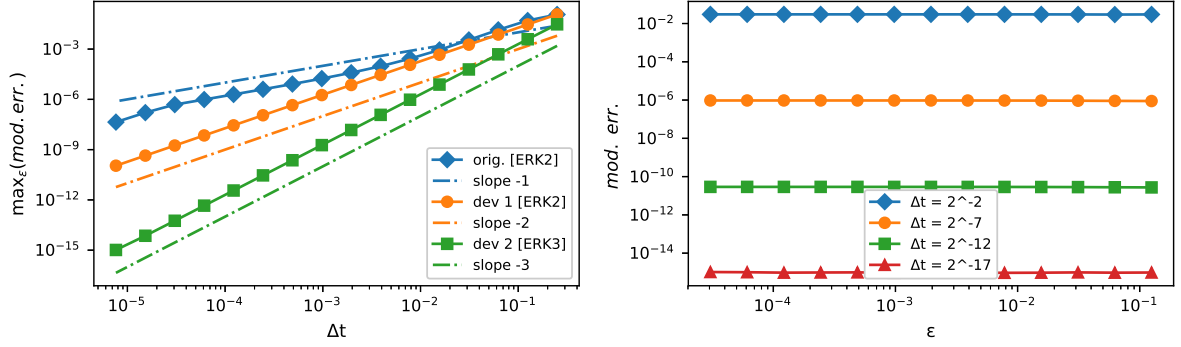


FIGURE 2.1 – Oscillating case : On the left, maximum error on ε (for $\varepsilon = 2^{-k}$ with k spanning $\{3, \dots, 15\}$) as a function of Δt when using exponential RK schemes (abbr. ERK) of different orders. On the right, the error as a function of ε when solving the micro-macro problem of order 2 using ERK3.

The difference $f\left(\Omega_{t/\varepsilon}^{[2]}(v^{[2]}) + w^{[2]}\right) - f\left(\Omega_{t/\varepsilon}^{[2]}(v^{[2]})\right)$ is computed using

$$f(x + \tilde{x}, z + \tilde{z}) - f(x, z) = \begin{pmatrix} -(1-z)\tilde{x}_2 + (x_2 + \tilde{x}_2)\tilde{z} \\ (1-z)\tilde{x}_1 - (x_1 + \tilde{x}_1)\tilde{z} \\ \left(x_1x_2 + (x_1 + \tilde{x}_1)(x_2 + \tilde{x}_2)\right)(x_1\tilde{x}_2 + \tilde{x}_1x_2 + \tilde{x}_1\tilde{x}_2) \end{pmatrix}$$

in order to avoid rounding errors due to the size difference between u and \tilde{u} .

A PDE-inspired problem

Consider a problem similar to a relaxed conservation law (as in the next subsection) but without transport, written

$$\begin{cases} \dot{u} = \tilde{u}, & u(0) \in \mathbb{R}^d, \\ \dot{\tilde{u}} = \frac{1}{\varepsilon}(g(u) - \tilde{u}), & \tilde{u}(0) \in \mathbb{R}^d \end{cases}$$

for some smooth map $g : \mathbb{R}^d \mapsto \mathbb{R}^d$. This can be transformed in a system of the form (2.1) by setting $x = u$ and $z = g(u) - \tilde{u}$, yielding the problem

$$\begin{cases} \dot{x} = g(x) - z, & x(0) = u(0), \\ \dot{z} = -\frac{1}{\varepsilon}z + g'(x)(g(x) - z), & z(0) = g(u(0)) - \tilde{u}(0). \end{cases} \quad (2.59)$$

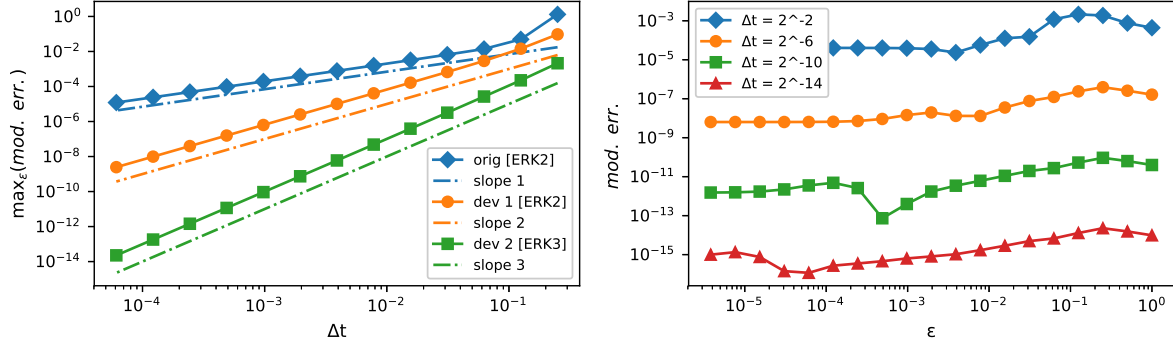


FIGURE 2.2 – PDE-inspired problem : On the left, maximum error on ε (for $\varepsilon = 2^{-k}$ with k spanning $\{1, \dots, 18\}$) as a function of Δt when using exponential RK schemes (abbr. ERK) of different orders. On the right, the error as a function of ε when solving the micro-macro problem of order 2 using ERK3.

The change of variable and vector field can be computed by hand up to order 1,

$$\Omega_{\tau}^{[1]}(x, z) = \begin{pmatrix} x + \varepsilon e^{-\tau} z \\ e^{-\tau} z + \varepsilon g'(x)g(x) \end{pmatrix},$$

$$F^{[1]}(x, z) = \begin{pmatrix} g(x) - \varepsilon g'(x)g(x) \\ -g'(x)z + \varepsilon (g'(x)^2 + g''(x)g(x) - \varepsilon g''(x)g'(x)g(x))z \end{pmatrix}.$$

Going to a higher order requires specific computations, as the expression of $\frac{1}{\varepsilon}(\partial_t + \Lambda)\Omega_{\tau}^{[2]}(x, z)$ is verbose and involves for instance $g(x + \varepsilon e^{-\tau} z) - g(x)$. It can be checked by hand that this expression involves no $e^{-\tau\Lambda}$ -term with the above expressions of $\Omega^{[1]}$ and $F^{[1]}$. For numerical testing, we chose $g(x) = -x^3/3$, $u(0) = 1$ and $\tilde{u}(0) = 0$. The micro-macro problem was computed up to order 2.

Results

Figures 2.1 and 2.2 showcase the phenomenon of order reduction when solving the original problem (2.57) : Despite using a scheme of order 2, the error depends of ε in such a way that there exists no constant C such that the error is bounded by $C\Delta t^2$ for all ε . However there exists C such that the error is bounded by $C\Delta t$. In that case, we cannot say that the error is of *uniform* order 2, as this would require the error to be independent of ε .

This order reduction disappears when solving the micro-macro problem, as can be seen on the right-hand side of the figures for a decomposition of order 2. Furthermore, the theoretical orders of convergence from Theorem 2.2.8 are confirmed. Indeed, using a scheme of order 2 (resp. 3) on the micro-macro problem of order 1 (resp. 2) generates a

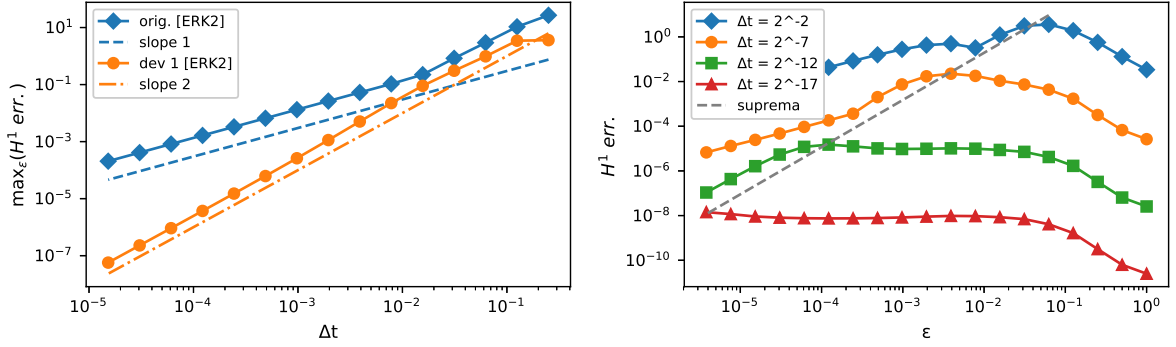


FIGURE 2.3 – Telegraph equation : Absolute H^1 error on the solution of (2.41) computed by an ERK3 scheme. Supremum on ε as a function of Δt (left) and evolution of this error as a function of ε for the 1st-order decomposition (right).

uniform error of the expected order of convergence, with no order reduction.

2.5.2 Discretized hyperbolic partial differential equations

The telegraph equation

Using a spectral decomposition, we solve the problem, for $(t, x) \in [0, T] \times \mathbb{R}/2\pi\mathbb{Z}$,

$$\begin{cases} \partial_t \rho + \partial_x j = 0, \\ \partial_t j + \frac{1}{\varepsilon} \partial_x \rho = -\frac{1}{\varepsilon} j, \end{cases}$$

by setting $z = j + (1 - \alpha\varepsilon\Delta)^{-1}\partial_x z$, yielding problem (2.43). The micro-macro decomposition of order 1 is summarized in Property 2.4.1, and its construction is detailed in Subsection 2.4.1. Implementations are conducted using $\alpha = 2$, space frequencies are bounded by $k_{\max} := 12$, and initial data is $\rho(0, x) = e^{\cos(x)}$, $j(0, x) = \frac{1}{2} \cos^3(x)$.

Results can be seen in Figure 2.3 when using a scheme of order 2. When solving the original problem, the uniform order degenerates from 2 to 1. When considering the micro-macro problem, the order of convergence is not reduced and stays of order 2. Although it varies with ε when considering a fixed Δt , when considering the supremum on ε , there is no order reduction. The dashed slope on the right plot interpolates the position of the supremum of the error for each fixed Δt . While the error seems to improve for $\varepsilon \ll \Delta t$, this does not cause any order reduction. This is stronger than the property of preservation of asymptotes (which ERK schemes have, see [DP11]), since AP schemes only need to be

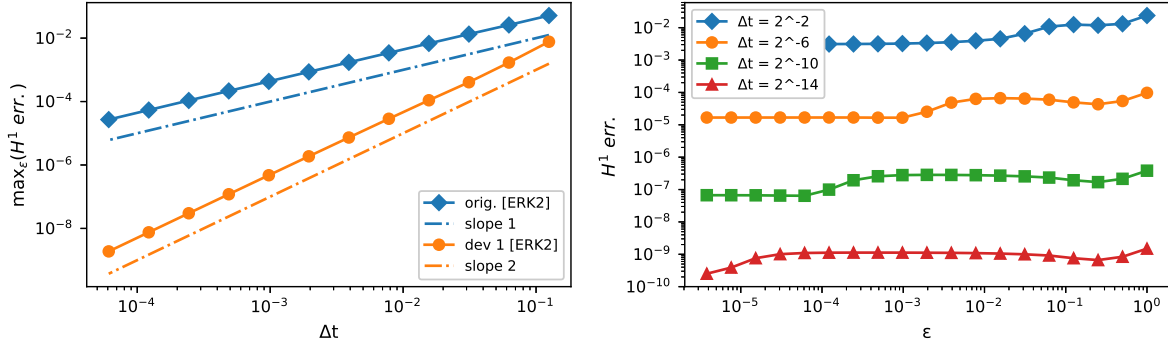


FIGURE 2.4 – Relaxed Burgers-type problem : Maximum modified H^1 error (for ε spanning 1 to 2^{-18} using an ERK3 scheme as a function of Δt (left), and H^1 error as a function of ε for the micro-macro problem of order 1 (right).

well-defined in the limit $\varepsilon \rightarrow 0$. For them, this supremum does not need to be bounded. It appears that the relationship between the error bound and the stiffness of the linear operator is rather complex when using exponential RK schemes (again, see [HO05] for details).

Relaxed conservation law

Our second test case is a hyperbolic problem for $(t, x) \in [0, T] \times \mathbb{R}/2\pi\mathbb{Z}$,

$$\begin{cases} \partial_t u + \partial_x \tilde{u} = 0, \\ \partial_t \tilde{u} + \partial_x u = \frac{1}{\varepsilon}(g(u) - \tilde{u}), \end{cases}$$

discretized with finite volumes and written in the form of (2.1) by setting $u_1 = u$ and $u_2 = \tilde{u} - g(u)$ the x^ε - and the z^ε -component respectively. The micro-macro expansion is computed to order 1 using the strategy detailed in Subsection 2.4.2.

For our tests, following [HS21], we consider $g(u) = bu^2$ with $b = 0.2$. Simulations run to a final time $T = 0.25$ and the mesh size is fixed : $N = 16$. Initial data is $u(0, x) = \frac{1}{2}e^{\sin(x)}$ and $\tilde{u}(0, x) = \cos(x)$. The reference solution was computed up to a precision 10^{-12} using an ERK2 scheme. Convergence results are presented in Figure 2.4, confirming theoretical results once more.

It should be said again that our approach does not study the error in space, only in time. For instance, the relationship between the error bound and the grid size is not considered. Further studies will be conducted, especially considering CFL conditions, L^2 and H^1 norms, and computational costs.

2.5.3 Perspectives

Computing cost

Note that when using a given scheme, solving a single step is much more costly for the micro-macro problem than for the direct problem : Not only is the system size doubled, but the functions implicated require more computing power to obtain a single value (especially the defect, see (2.56) for instance). It is therefore plausible to think that our method is best for computing values during the transient phase, after which it is possible to solve the original problem with uniform accuracy.

The regularized derivation $(I_N - 2\varepsilon D^2)^{-1}D$ which appears in the micro-macro problem of the relaxed hyperbolic system may be prohibitively costly to compute for some schemes such as WENO, for which the derivation operator is non-linear. However we may be able to work around this, as the goal of the relaxation term is only to dampen high-frequencies, and as such inverting any discrete Laplace operator should suffice, independently of the scheme used to discretize the transport. Clearly, the subject of utilizing such regularizations for numerical purposes is complex and beyond the scope of this paper.

Near-equilibrium convergence

If one chooses an initial condition $z^\varepsilon(0) = 0$ in (2.1), then it is close to the center manifold up to $\mathcal{O}(\varepsilon)$, and Problem (2.2) can be solved with uniform accuracy of order 2 but only when considering the absolute error $|\cdot|$, not the modified error $|\cdot|_\varepsilon$ from (2.27). The same behaviour is observed for the telegraph equation when setting $j(0, x) = -\partial_x \rho(0, x)$, meaning $z = \mathcal{O}(\varepsilon)$. This would theoretically mean that we need to push the micro-macro decompositions up to order 2 if we want to improve the order of convergence. However, this is not the case : uniform accuracy of order 3 is obtained from an expansion of order 1 for all test cases. This “order gain” also propagates to our micro-macro decomposition of order 2 for the oscillating toy problem. These results can be seen in Figure 2.5 and will be studied in future works.

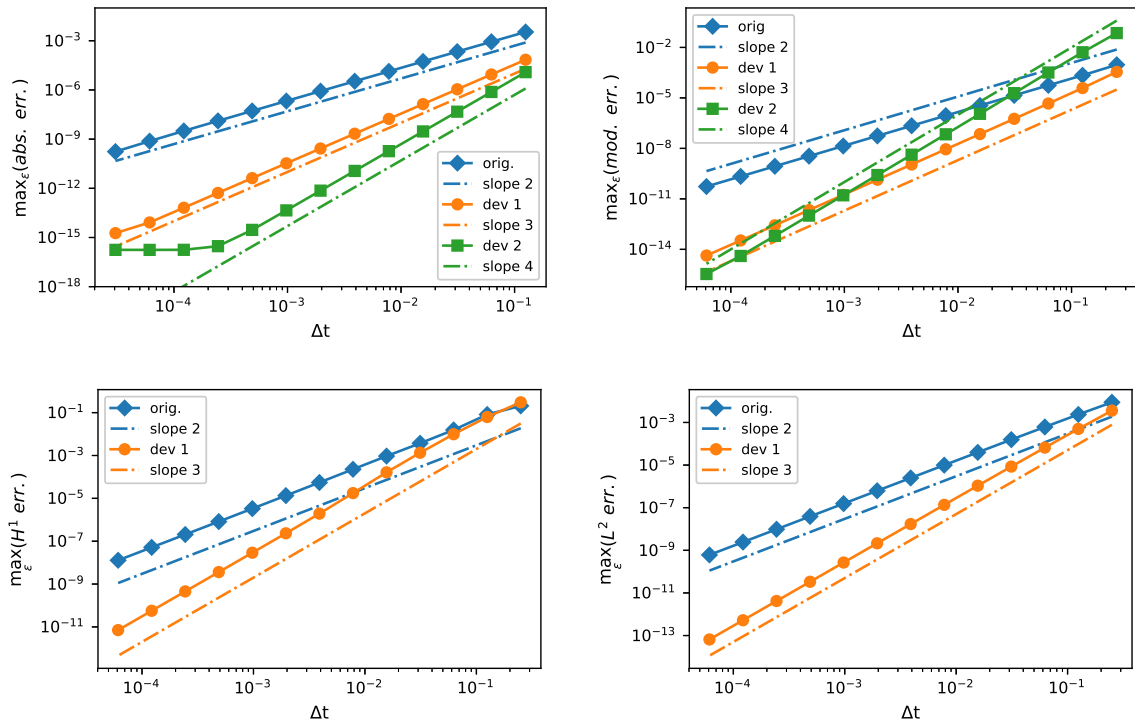


FIGURE 2.5 – In reading order, errors when solving the oscillating toy problem, the PDE-inspired problem, the telegraph equation and the relaxed conservation law. All systems start near equilibrium and are solved with exponential Runge-Kutta schemes of the observed order of convergence.

DISCUSSION D'EXTENSION DES RÉSULTATS

Closure. I keep hearing that word. It's the theater of the absurd. Everybody knows that on television they'll see the end of the story in the last 15 minutes of the thing. It's like a drug. To me, that's the beauty of 'Twin Peaks.' We throw in some curve balls. **As soon as a show has a sense of closure, it gives you an excuse to forget you've seen the damn thing.**

David Lynch, 1990

<https://web.archive.org/web/20200805032143/https://www.latimes.com/archives/la-xpm-1990-02-18-ca-1500-story.html>

3.1 Coût de calcul, erreurs d'arrondis, derivative-free, pullback

Coût de calcul avec les non-linéarités

Difficulté du calcul de la relaxation pour certains schémas

Gain d'ordre avec $z(0) = 0$ (figure avec err sup sur ε)

Donner une clé pour le gain d'ordre : $v_z(0) = \mathcal{O}(\varepsilon)$

3.2 Autour de l'équation de télégraphe

UN DÉVELOPPEMENT DOUBLE-ÉCHELLE

Résultats du stage et du début de thèse

PRÉSENTATION DE SCHÉMAS NUMÉRIQUES

Méthodes composées

Tableaux de Butcher

Splitting

Stormer-Verlett est un cas particulier de Strang mélangé à Euler explicite. En effet avec $\dot{q} = v$ et $\dot{v} = F(q)$,

$$v_{n+1/2} = v_n + \frac{\Delta t}{2} F(q_n)$$

$$q_{n+1} = q_n + \Delta t v_{n+1/2}$$

$$v_{n+1} = v_{n+1/2} + \frac{\Delta t}{2} F(q_{n+1}).$$

Ce qui revient à séparer le système en $\partial_t(q, v) = (0, F(q))$ et $\partial_t(q, v) = (v, 0)$.

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Titre : titre (en français).....

Mot clés : de 3 à 6 mots clefs

Résumé : Eius populus ab incunabulis primis ad usque pueritiae tempus extremum, quod annis circumcluditur fere trecentis, circummurana pertulit bella, deinde aetatem ingressus adultam post multiplices bellorum aerumnas Alpes transcendit et fretum, in iuvenem erectus et virum ex omni plaga quam orbis ambit inensus, reportavit laureas et triumphos, iamque vergens in senium et nomine solo aliquotiens vincens ad tranquilliora vitae discessit. Hoc immaturo interitu ipse quoque sui pertaesus excessit e vita aetatis nono anno atque vicensimo cum quadriennio imperasset. natus apud Tuscos in Massa Vaternensi, patre Constantio Constantini fratre imperatoris, matreque Galla. Thalassius vero

ea tempestate praefectus praetorio praesens ipse quoque adrogantis ingenii, considerans incitationem eius ad multorum augeri discrimina, non maturitate vel consiliis mitigabat, ut aliquotiens celsae potestates iras principum molliverunt, sed adversando iurgandoque cum parum congrueret, eum ad rabiem potius evibrabat, Augustum actus eius exaggerando creberrime docens, idque, incertum qua mente, ne lateret adfectans. quibus mox Caesar acrius efferatus, velut contumaciae quoddam vexillum altius erigens, sine respectu salutis alienae vel suae ad vertenda opposita instar rapidi fluminis irrevocabili impetu ferebatur. Hae duae provinciae bello quondam piratico catervis mixtae praedonum.

Title: titre (en anglais).....

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Abstract: Eius populus ab incunabulis primis ad usque pueritiae tempus extremum, quod annis circumcluditur fere trecentis, circummurana pertulit bella, deinde aetatem ingressus adultam post multiplices bellorum aerumnas Alpes transcendit et fretum, in iuvenem erectus et virum ex omni plaga quam orbis ambit inensus, reportavit laureas et triumphos, iamque vergens in senium et nomine solo aliquotiens vincens ad tranquilliora vitae discessit. Hoc immaturo interitu ipse quoque sui pertaesus excessit e vita aetatis nono anno atque vicensimo cum quadriennio imperasset. natus apud Tuscos in Massa Vaternensi, patre Constantio Constantini fratre imperatoris, matreque Galla. Thalassius vero

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