PyFANT Documentation

Release 17.12.21

Julio Trevisan

CONTENTS

1	Introduction	3
2	Installation	5
3	Coding using the API	7
4	Index of applications (scripts)	19

Welcome!

PyFANT is a Python interface to the PFANT stellar spectral synthesis code, with a varied set of extra tools.

CONTENTS 1

2 CONTENTS

CHAPTER

ONE

INTRODUCTION

PyFANT is a Python interface to PFANT, a stellar spectral synthesis code for Astronomy originally written in Fortran. PyFANT provides different ways to run spectral synthesis: command-line or graphical interface; single or batch mode. It also includes many additional tools to open, edit, visualize, convert and manipulate data files and spectra, and an API (application programming interfce) to create your own spectral synthesis applications in Python.

This documentation addresses coding with the PyFANT API, and provides a reference to the scripts included in the pyfant package.

For a tutorial on spectral synthesis using PFANT and PyFANT, please visit the PFANT project website: https://trevisanj.github.io/PFANT.

1.1 Acknowledgement

The project started in 2015 at IAG-USP (Institute of Astronomy, Geophysics and Atmospheric Sciences at University of São Paulo, Brazil).

Funded by FAPESP - Research Support Foundation of the State of São Paulo, Brazil (2015-2017).

1.2 Contact

For bugs reports, questions, suggestions, etc., please open an issue at the project site on GitHub: http://github.com/trevisanj/pyfant.

CHAPTER

TWO

INSTALLATION

If you have both Python 3 and PFANT installed, then simply type:

pip install pyfant

2.1 Pre-requisites

2.1.1 **PFANT**

The PFANT spectral synthesis software installation instructions can be found at http://trevisanj.github.io/PFANT/install.html.

2.1.2 Python 3

If you need to set up your Python 3 environment, one option is to visit project F311 installation instructions at http://trevisanj.github.io/f311/install.html. That page also provides a troubleshooting section that applies.

2.2 Installing PyFANT in developer mode

This is an alternative to the "pip" at the beginning of this section. Use this option if you would like to download and modify the Python source code.

First clone the "pyfant" GitHub repository:

```
git clone ssh://git@github.com/trevisanj/pyfant.git
```

or

git clone http://github.com/trevisanj/pyfant

Then, install PyFANT in **developer** mode:

```
cd pyfant
python setup.py develop
```

2.3 Upgrade pyfant

Package pyfant can be upgraded to a new version by typing:

pip install pyfant --upgrade

CODING USING THE API

This section contains a series of examples on how to use the PFANT Fortran executables from a Python script. These "bindings" to the Fortran binaries, together with the ability to load, manipulate and save PFANT data files allow for complex batch operations.

3.1 Spectral synthesis

The following code generates Figure 3.1.

```
"""Runs synthesis over short wavelength range, then plots normalized and convolved spectrum"""
import f311.pyfant as pf
import f311.explorer as ex
import matplotlib.pyplot as plt
if __name__ == "__main__":
   # Copies files main.dat and abonds.dat to local directory (for given star)
   pf.copy_star(starname="sun-grevesse-1996")
   # Creates symbolic links to all non-star-specific files, such as atomic & molecular lines,
   # partition functions, etc.
   pf.link_to_data()
   # # First run
   # Creates object that will run the four Fortran executables (innewmarcs, hydro2, pfant, nulbad)
   obj = pf.Combo()
   # synthesis interval start (angstrom)
   obj.conf.opt.llzero = 6530
   # synthesis interval end (angstrom)
   obj.conf.opt.llfin = 6535
   # Runs Fortrans and hangs until done
   obj.run()
   # Loads result files into memory. obj.result is a dictionary containing elements ...
   obj.load_result()
   print("obj.result = {}".format(obj.result))
   res = obj.result
   plt.figure()
   ex.draw_spectra_overlapped([res["norm"], res["convolved"]])
    plt.savefig("norm-convolved.png")
    plt.show()
```

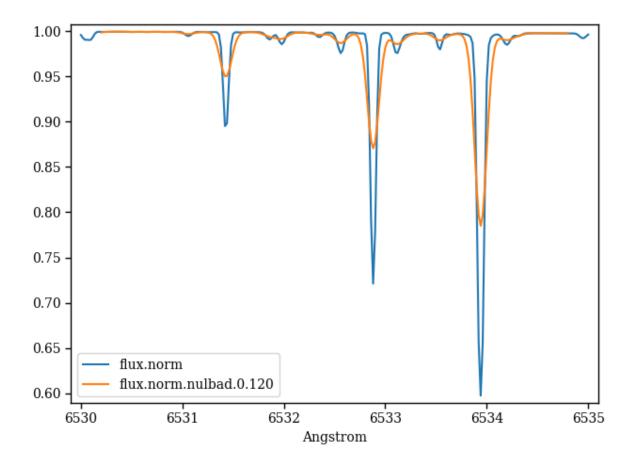


Figure 3.1: – pfant (file "flux.norm") and nulbad outputs.

3.2 Spectral synthesis - convolutions

The following example convolves the synthetic spectrum (file "flux.norm") with Gaussian profiles of different FWHMs (Figure 3.2).

```
#!/usr/bin/env python
"""Runs synthesis over short wavelength range, then plots normalized and convolved spectrum"""
import f311.pyfant as pf
import f311.explorer as ex
import matplotlib.pyplot as plt
import a99
# FWHM (full width at half of maximum) of Gaussian profiles in angstrom
FWHMS = [0.03, 0.06, 0.09, 0.12, 0.15, 0.20, 0.25, 0.3, 0.5]
if __name__ == "__main__":
    # Copies files main.dat and abonds.dat to local directory (for given star)
   pf.copy_star(starname="sun-grevesse-1996")
   # Creates symbolic links to all non-star-specific files
   pf.link_to_data()
   # # 1) Spectral synthesis
   # Creates object that will run the four Fortran executables (innewmarcs, hydro2, pfant, nulbad)
   ecombo = pf.Combo()
   # synthesis interval start (angstrom)
   ecombo.conf.opt.llzero = 6530
   # synthesis interval end (angstrom)
   ecombo.conf.opt.llfin = 6535
   # Runs Fortrans and hangs until done
   ecombo.run()
   ecombo.load_result()
   # Retains un-convolved spectrum for comparison
   spectra = [ecombo.result["norm"]]
   # # 2) Convolutions
    for fwhm in FWHMS:
        enulbad = pf.Nulbad()
        enulbad.conf.opt.fwhm = fwhm
       enulbad.run()
        enulbad.load_result()
        # Appends convolved spectrum for comparison
        spectra.append(enulbad.result["convolved"])
   # # 3) Plots
   plt.figure()
   ex.draw_spectra_overlapped(spectra)
   K = 1.1
   a99.set_figure_size(plt.gcf(), 1000*K, 500*K)
   plt.tight_layout()
   plt.savefig("many-convs.png")
    plt.show()
```

3.3 Spectral synthesis - Continuum

The following code generates Figure 3.3.

```
"""Runs synthesis over large wavelength range, then plots continuum"""
```

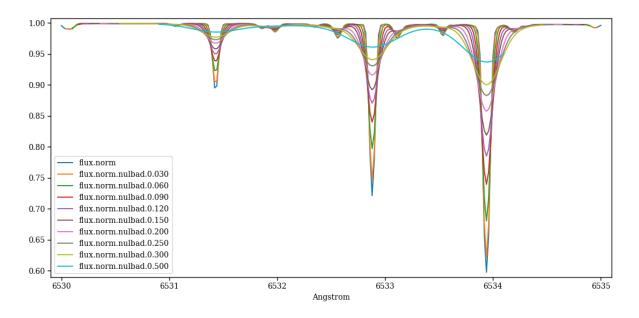


Figure 3.2: – single pfant output and several nulbad outputs.

```
import f311.pyfant as pf
import f311.explorer as ex
import matplotlib.pyplot as plt
import a99
if __name__ == "__main__":
    # Copies files main.dat and abonds.dat to local directory (for given star)
   pf.copy_star(starname="sun-grevesse-1996")
   # Creates symbolic links to all non-star-specific files, such as atomic & molecular lines,
    # partition functions, etc.
   pf.link_to_data()
   # Creates object that will run the four Fortran executables (innewmarcs, hydro2, pfant, nulbad)
   obj = pf.Combo()
   oo = obj.conf.opt
   # synthesis interval start (angstrom)
   oo.llzero = 2500
   # synthesis interval end (angstrom)
   oo.11fin = 30000
   # savelength step (angstrom)
   oo.pas = 1.
   # Turns off hydrogen lines
   oo.no_h = True
    # Turns off atomic lines
   oo.no_atoms = True
    # Turns off molecular lines
   oo.no_molecules = True
   obj.run()
   obj.load_result()
   print("obj.result = {}".format(obj.result))
   res = obj.result
   ex.draw_spectra_stacked([res["cont"]], setup=ex.PlotSpectrumSetup(fmt_ylabel=None))
   K = .75
   a99.set_figure_size(plt.gcf(), 1300*K, 450*K)
   plt.tight_layout()
    plt.savefig("continuum.png")
   plt.show()
```

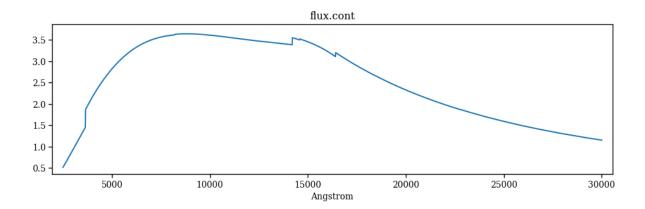


Figure 3.3: – continuum.

3.4 Spectral synthesis - Separate atomic species

PFANT atomic lines files contains wavelength, log_gf and other tabulated information for several (element, ionization level) atomic species.

The following code calculates isolated atomic spectra for a list of arbitrarily chosen atomic species (Figure 3.4).

```
"""Runs synthesis for specified atomic species separately. No molecules or hydrogen lines."""
import f311.pyfant as pf
import f311.explorer as ex
import matplotlib.pyplot as plt
import f311.filetypes as ft
import a99
\# ["<element name><ionization level>", ...] for which to draw panels
MY_SPECIES = ["Fe1", "Fe2", "Ca1", "Ca2", "Na1", "Si1"]
if __name__ == "__main__":
    pf.copy_star(starname="sun-grevesse-1996")
   pf.link_to_data()
    # Loads full atomic lines file
    fatoms = ft.FileAtoms()
   fatoms.load()
    runnables = []
    for elem_ioni in MY_SPECIES:
        atom = fatoms.find_atom(elem_ioni)
        # Creates atomic lines file object containing only one atom
        fatoms2 = ft.FileAtoms()
        fatoms2.atoms = [atom]
        ecombo = pf.Combo()
        # Overrides file "atoms.dat" with in-memory object
        ecombo.conf.file_atoms = fatoms2
        ecombo.conf.flag_output_to_dir = True
        oo = ecombo.conf.opt
        # Assigns synthesis range to match atomic lines range
        oo.llzero, oo.llfin = fatoms2.llzero, fatoms2.llfin
        # Turns off hydrogen lines
        oo.no_h = True
```

```
# Turns off molecular lines
       oo.no_molecules = True
       runnables.append(ecombo)
   pf.run_parallel(runnables)
   # Draws figure
   f = plt.figure()
   a99.format_BLB()
   for i, (title, ecombo) in enumerate(zip(MY_SPECIES, runnables)):
       ecombo.load_result()
       plt.subplot(2, 3, i+1)
       ex.draw_spectra_overlapped([ecombo.result["spec"]],
                                   setup=ex.PlotSpectrumSetup(flag_xlabel=i/3 >= 1, flag_
→legend=False))
       plt.title(title)
   K = 1.
   a99.set_figure_size(plt.gcf(), 1300*K, 740*K)
   plt.tight_layout()
   plt.savefig("synthesis-atoms.png")
   plt.show()
```

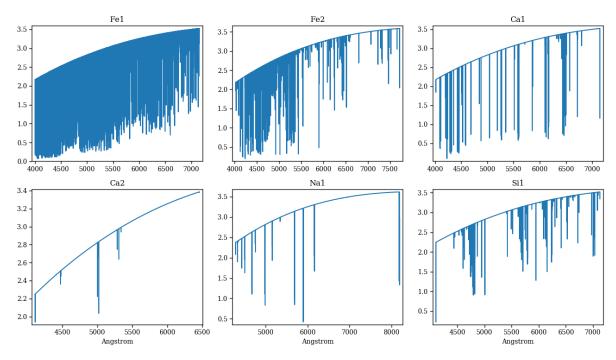


Figure 3.4: – atomic lines synthesized separately for each species.

3.5 Spectral synthesis - Separate molecules

The following code generates Figure 3.5, Figure 3.6, and additional plots not shown here.

```
"""Runs synthesis for molecular species separately. No atomic nor hydrogen lines."""

import f311.pyfant as pf
import f311.explorer as ex
import matplotlib.pyplot as plt
```

```
import f311.filetypes as ft
import a99
SUBPLOT_NUM_ROWS = 2
SUBPLOT_NUM_COLS = 2
if __name__ == "__main__":
   pf.copy_star(starname="sun-grevesse-1996")
   pf.link_to_data()
   # Loads full molecular lines file
   fmol = ft.FileMolecules()
   fmol.load()
   runnables = []
   for molecule in fmol:
        fmol2 = ft.FileMolecules()
        fmol2.molecules = [molecule]
        ecombo = pf.Combo()
        # Overrides file "molecules.dat" with in-memory object
        ecombo.conf.file_molecules = fmol2
        ecombo.conf.flag_output_to_dir = True
        oo = ecombo.conf.opt
        # Assigns synthesis range to match atomic lines range
        oo.llzero, oo.llfin = fmol2.llzero, fmol2.llfin
        # Turns off hydrogen lines
        oo.no_h = True
        # Turns off atomic lines
        oo.no_atoms = True
        # Adjusts the wavelength step according to the calculation interval
        oo.pas = max(1, round(oo.llfin*1./20000/2.5)*2.5)
        oo.aint = max(50., oo.pas)
        runnables.append(ecombo)
   pf.run_parallel(runnables)
   num_panels = SUBPLOT_NUM_COLS*SUBPLOT_NUM_ROWS
   num_molecules = len(runnables)
   ifigure = 0
   a99.format_BLB()
    for i in range(num_molecules+1):
        not\_first = i > 0
        first_panel_of_figure = (i / num_panels - int(i / num_panels)) < 0.01</pre>
        is_panel = i < num_molecules</pre>
        if not_first and (not is_panel or first_panel_of_figure):
            plt.tight_layout()
            K = 1.
            a99.set_figure_size(plt.gcf(), 1500 * K, 740 * K)
            plt.tight_layout()
            filename_fig ="synthesis-molecules-{}.png".format(ifigure)
            print("Saving figure '{}'...".format(filename_fig))
            plt.savefig(filename_fig)
            plt.close()
            ifigure += 1
        if first_panel_of_figure and is_panel:
            plt.figure()
        if is_panel:
```

```
ecombo = runnables[i]
ecombo.load_result()

isubplot = i % num_panels + 1
plt.subplot(SUBPLOT_NUM_ROWS, SUBPLOT_NUM_COLS, isubplot)
ex.draw_spectra_overlapped([ecombo.result["spec"]],
    setup=ex.PlotSpectrumSetup(flag_xlabel=i/3 >= 1, flag_legend=False))

_title = fmol[i].description
if "]" in _title:
    title = _title[:_title.index("]")+1]
else:
    title = _title[:20]
plt.title(title)
```

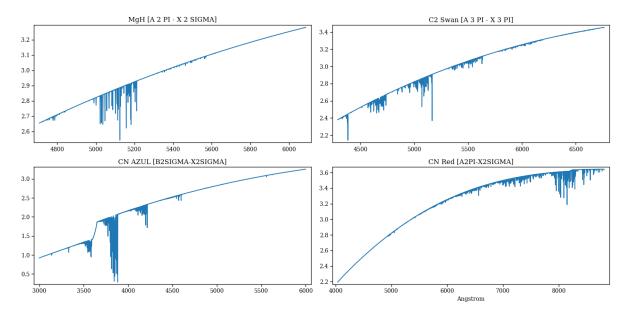


Figure 3.5: - molecular lines synthesized separately for each species

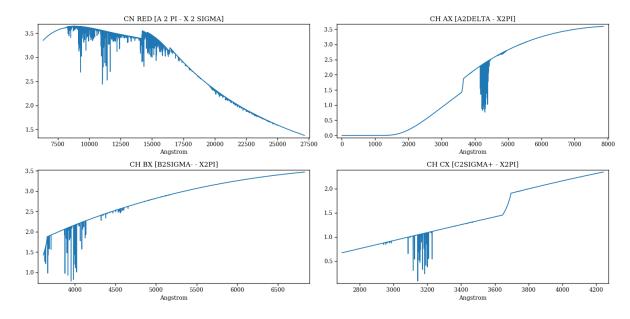


Figure 3.6: – molecular lines synthesized separately for each species.

3.6 Gaussian profiles as nulbad outputs

nulbad is one of the Fortran executables of the PFANT package. It is the one that convolves the synthetic spectrum calculated by pfant with a Gaussian profile specified by a "fwhm" parameter (Figure 3.7).

```
Nulbad's "impulse response"
Saves "impulse" spectrum (just a spike at lambda=5000 angstrom) as "flux.norm",
then runs 'nulbad' repeatedly to get a range of Gaussian profiles.
import f311.pyfant as pf
import f311.explorer as ex
import matplotlib.pyplot as plt
import a99
import f311.filetypes as ft
import numpy as np
# FWHM (full width at half of maximum) of Gaussian profiles in angstrom
FWHMS = [0.03, 0.06, 0.09, 0.12, 0.15, 0.20, 0.25, 0.3]
if __name__ == "__main__":
   # Copies files main.dat and abonds.dat to local directory (for given star)
   pf.copy_star(starname="sun-grevesse-1996")
   # Creates symbolic links to all non-star-specific files
   pf.link_to_data()
   # # 1) Creates "impulse" spectrum
   fsp = ft.FileSpectrumPfant()
   sp = ft.Spectrum()
   N = 2001
   sp.x = (np.arange(0, N, dtype=float)-(N-1)/2)*0.001+5000
   sp.y = np.zeros((N,), dtype=float)
   sp.y[int((N-1)/2)] = 1.
   fsp.spectrum = sp
   fsp.save_as("flux.norm")
   # # 2) Convolutions
    spectra = []
    for fwhm in FWHMS:
       enulbad = pf.Nulbad()
       enulbad.conf.opt.fwhm = fwhm
       enulbad.run()
        enulbad.load_result()
        enulbad.clean()
        # Appends convolved spectrum for comparison
        spectra.append(enulbad.result["convolved"])
   # # 3) Plots
   f = plt.figure()
   ex.draw_spectra_overlapped(spectra)
   K = 0.7
   a99.set_figure_size(plt.gcf(), 1300*K, 500*K)
    plt.tight_layout()
    plt.savefig("gaussian-profiles.png")
    plt.show()
```

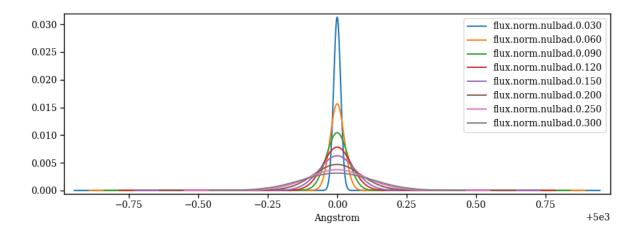


Figure 3.7: - Gaussian profiles illustrated for different FWHMs.

3.7 Plot hydrogen profiles

The following code generates Figure 3.8.

```
Calculates hydrogen lines profiles, then plots them in several 3D subplots
import pyfant
import a99
import os
import shutil
import matplotlib.pyplot as plt
from mpl_toolkits.mplot3d import Axes3D # yes, required (see below)
def mylog(*args):
   print("^^ {}".format(", ".join(args)))
def main(flag_cleanup=True):
   tmpdir = a99.new_filename("hydrogen-profiles")
    # Saves current directory
   pwd = os.getcwd()
   mylog("Creating directory '{}'...".format(tmpdir))
   os.mkdir(tmpdir)
   try:
        pyfant.link_to_data()
        _main()
    finally:
        # Restores current directory
        os.chdir(pwd)
        # Removes temporary directory
        if flag_cleanup:
            mylog("Removing directory '{}'...".format(tmpdir))
            shutil.rmtree(tmpdir)
        else:
            mylog("Not cleaning up.")
def _main():
   fm = pyfant.FileMain()
```

```
fm.init_default()
   fm.llzero, fm.llfin = 1000., 200000. # spectral synthesis range in Angstrom
   ei = pyfant.Innewmarcs()
   ei.conf.file_main = fm
   ei.run()
   ei.clean()
   eh = pyfant.Hydro2()
   eh.conf.file_main = fm
   eh.run()
   eh.load_result()
   eh.clean()
   _plot_profiles(eh.result["profiles"])
def _plot_profiles(profiles):
   fig = plt.figure()
   for filename, ftoh in profiles.items():
        if ftoh is not None:
            mylog("Drawing '{}'...".format(filename))
            \# ax = plt.subplot(2, 3, i+1)
            ax = fig.add_subplot(2, 3, i+1, projection='3d')
            ax.set_title(filename)
            pyfant.draw_toh(ftoh, ax)
            i += 1
   plt.tight_layout()
   plt.savefig("hydrogen-profiles.png")
   plt.show()
if __name__ == "__main__":
   main(flag_cleanup=True)
```

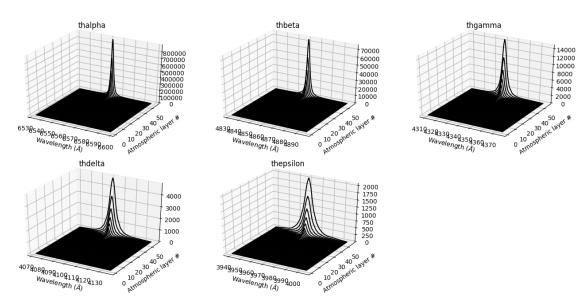


Figure 3.8: - hydrogen lines profiles as calculated by hydro2.

3.8 Import Kurucz' molecular linelist file

```
This example loads file "c2dabrookek.asc" and prints a memory representation of its first line.

This file can be obtained at http://kurucz.harvard.edu/molecules/c2/. First lines of file:

287.7558-14.533 23.0 2354.082 24.0 -37095.578 6063a00e1 3d10e3 12 677 34741.495
287.7564-14.955 22.0 2282.704 23.0 -37024.124 6063a00f1 3d10f3 12 677 34741.419
287.7582-14.490 21.0 2214.696 22.0 -36955.900 6063a00e1 3d10e3 12 677 34741.205
287.7613-15.004 24.0 2428.453 25.0 -37169.280 6063a00f1 3d10f3 12 677 34740.828
287.7650-14.899 20.0 2149.765 21.0 -36890.147 6063a00f1 3d10f3 12 677 34740.382

"""

import f311.filetypes as ft

f = ft.load_any_file("c2dabrookek.asc")

print(repr(f[0]).replace(", ", ",\n "))
```

This code should print the following:

```
KuruczMolLine(lambda_=2877.558,
              loggf=-14.533,
              J21=23.0,
              E21=2354.082,
              J1=24.0,
              E1=37095.578,
              atomn0=6,
              atomn1=6,
              state21='a',
              v21=0,
              lambda_doubling2l='e',
              spin2l=1,
              statel='d',
              vl=10,
              lambda_doublingl='e',
              spinl=3,
              iso=12)
```

INDEX OF APPLICATIONS (SCRIPTS)

This chapter is a reference to all scripts in project PyFANT

4.1 Script copy-star.py

```
usage: copy-star.py [-h] [-l] [-p] [directory]
Copies stellar data files (such as main.dat, abonds.dat, dissoc.dat) to local directory
examples of usage:
 > copy-star.py
 (displays menu)
 > copy-star.py arcturus
 ("arcturus" is the name of a subdirectory of PFANT/data)
 > copy-star.py -p /home/user/pfant-common-data
  (use option "-p" to specify path)
 > copy-star.py -l
 (lists subdirectories of PFANT/data , doesn't copy anything)
positional arguments:
             name of directory (either a subdirectory of PFANT/data or the
              path to a valid system directory (see modes of operation)
              (default: None)
optional arguments:
 -h, --help show this help message and exit
 -1, --list lists subdirectories of
              /home/j/Documents/projects/astro/github/PFANT/code/data\\
              (default: False)
 -p, --path system path mode (default: False)
```

This script belongs to package *pyfant*

4.2 Script create-grid.py

```
Working modes (option "-m"):
"opa" (default mode): looks for MARCS[1] ".mod" and ".opa" text file pairs and
                       creates a *big* binary file containing *all* model
                       information including opacities.
                       Output will be in ".moo" format.
 \hbox{\tt "modtxt": looks for MARCS ".mod" text files only. Resulting grid will not contain}
           opacity information.
           Output will be in binary ".mod" format.
 "modbin": looks for binary-format ".mod" files. Resulting grid will not contain
           opacity information.
           Output will be in binary ".mod" format.
References:
 [1] http://marcs.astro.uu.se/
positional arguments:
 fn_output
                        output file name (default: "grid.moo" or "grid.mod",
                        depending on mode)
optional arguments:
 -h, --help
                        show this help message and exit
  --pattern [PATTERN]
                        file name pattern (with wildcards) (default: *.mod)
 -m [{opa,modtxt,modbin}], --mode [{opa,modtxt,modbin}]
                        working mode (see description above) (default: opa)
```

4.3 Script cut-atoms.py

This script belongs to package pyfant

4.4 Script cut-molecules.py

```
usage: cut-molecules.py [-h] llzero llfin fn_input fn_output

Cuts molecular lines file to wavelength interval specified
```

4.5 Script hitran-scraper.py

```
usage: hitran-scraper.py [-h] [-t T] [M] [I] [llzero] [llfin]
Retrieves molecular lines from the HITRAN database [Gordon2016]
This script uses web scraping and the HAPI to save locally molecular lines from the HITRAN database.
While the HAPI provides the downloading facility, web scraping is used to get the lists of molecules
and isotopologues from the HITRAN webpages and get the IDs required to run the HAPI query.
The script is typically invoked several times, each time with an additional argument.
References:
[Gordon2016] I.E. Gordon, L.S. Rothman, C. Hill, R.V. Kochanov, Y. Tan, P.F. Bernath, M. Birk,
   V. Boudon, A. Campargue, K.V. Chance, B.J. Drouin, J.-M. Flaud, R.R. Gamache, J.T. Hodges,
   D. Jacquemart, V.I. Perevalov, A. Perrin, K.P. Shine, M.-A.H. Smith, J. Tennyson, G.C. Toon,
   H. Tran, V.G. Tyuterev, A. Barbe, A.G. Császár, V.M. Devi, T. Furtenbacher, J.J. Harrison,
   J.-M. Hartmann, A. Jolly, T.J. Johnson, T. Karman, I. Kleiner, A.A. Kyuberis, J. Loos,
   O.M. Lyulin, S.T. Massie, S.N. Mikhailenko, N. Moazzen-Ahmadi, H.S.P. Müller, O.V. Naumenko,
   A.V. Nikitin, O.L. Polyansky, M. Rey, M. Rotger, S.W. Sharpe, K. Sung, E. Starikova,
   S.A. Tashkun, J. Vander Auwera, G. Wagner, J. Wilzewski, P. Wcisło, S. Yu, E.J. Zak,
   The HITRAN2016 Molecular Spectroscopic Database, J. Quant. Spectrosc. Radiat. Transf. (2017).
   doi:10.1016/j.jqsrt.2017.06.038.
positional arguments:
 М
             HITRAN molecule number (default: (lists molecules))
 Ι
             HITRAN isotopologue number (not unique, starts over at each
             molecule) (default: (lists isotopologues))
 llzero
             Initial wavelength (Angstrom) (default: None)
 llfin
             Final wavelength (Angstrom) (default: None)
optional arguments:
 -h, --help show this help message and exit
  -t T
             Table Name (default: (molecular formula))
```

This script belongs to package pyfant

4.5.1 Usage examples

```
$ hitran-scraper.py
List of all HITRAN molecules
```

=======================================						
ID	Formula	Name				
1	H20	Water				
2	C02	Carbon Dioxide				
3	03	Ozone				
4	N20	Nitrous Oxide				
5	CO	Carbon Monoxide				
6	CH4	Methane				
7	02	Molecular Oxygen				
8	NO	Nitric Oxide				
9	S02	Sulfur Dioxide				
10	NO2	Nitrogen Dioxide				
11	NH3	Ammonia				
12	HNO3	Nitric Acid				
13	OH	Hydroxyl Radical				
14	HF	Hydrogen Fluoride				
15	HC1	Hydrogen Chloride				
16	HBr	Hydrogen Bromide				
17	HI	Hydrogen Iodide				
18	C10	Chlorine Monoxide				
19	OCS	Carbonyl Sulfide				
20	H2CO	Formaldehyde				
21	HOC1	Hypochlorous Acid				
22	N2	Molecular Nitrogen				
23	HCN Hydrogen Cyanide					
24						
25	H202	Hydrogen Peroxide				
26	C2H2	Acetylene				
27	C2H6					
28	PH3	Phosphine				
29	COF2	Carbonyl Fluoride				
31	H2S	Hydrogen Sulfide				
32	HC00H	Formic Acid				
33	H02	Hydroperoxyl Radical				
34	0	Oxygen Atom				
36	NO+					
37	HOBr	Hypobromous Acid				
38	C2H4	Ethylene				
39 40	CH30H	Methanol Mathyl Brazida				
40	CH3Br	Methyl Gyanida				
41	CH3CN	Methyl Cyanide				
43	C4H2 HC3N	Diacetylene Cyanoacetylene				
44 45	HC3N H2	Cyanoacetylene Molecular Hydrogen				
45 46	CS	Carbon Monosulfide				
47	S03	Sulfur trioxide				
		otopologues for a given molecule, please type:				
h	itran-scrap	per.py <molecule id=""></molecule>				
wnere	where <molecule id=""> is one of the IDs listed above.</molecule>					

Now suppose we want the molecule OH molecule:

```
OH 1 13 (16)OH 61 0.997473
OH 2 13 (18)OH 81 0.002
OH 3 13 (16)OD 62 1.553710 × 10-4

Now, to download lines, please type:
hitran-scraper.py 13 <isotopologue ID> <llzero> <llfin>
where <isotopologue ID> is one the numbers from the 'ID' column above,
and [<llzero>, <llfin>] defines the wavelength interval in Angstrom.
```

Now selecting the first isotopologue and specifying the visible wavelength range:

```
$ hitran-scraper.py 13 1 3000 7000
Isotopologue selected:
Value
Field name
m_formula OH
ID_molecule 13
           (16)0H
Formula
AFGL_Code
             61
Abundance
            0.997473
Wavelength interval (air): [3000.0, 7000.0] Angstrom
Wavenumber interval (vacuum): [14289.61969369552, 33342.42546386186] cm**-1
Table name: '(16)0H'
Fetching data...
=== BEGIN messages from HITRAN API ===
BEGIN DOWNLOAD: (16)OH
 65536 bytes written to ./(16)0H.data
 65536 bytes written to ./(16)OH.data
 65536 bytes written to ./(16)0H.data
 65536 bytes written to ./(16)OH.data
 65536 bytes written to ./(16)0H.data
 65536 bytes written to ./(16)0H.data
Header written to ./(16)0H.header
END DOWNLOAD
                    Lines parsed: 3855
PROCESSED
=== END messages from HITRAN API ===
...done
Please check files '(16)OH.header', '(16)OH.data'
```

4.5.2 Quick note on the HITRAN API

The files created ('(16)OH.header', '(16)OH.data') can be opened using the HAPI. They are also accessed by the application convmol.py.

The HAPI can be downloaded, but one version is also included with the f311 package. The following is an example of how the HITRAN data could be accessed from the Python console:

To work properly with these data in your code, you may have a look at the HAPI source code and manual, as this library is superbly documented.

Within f311, the code in f311.convmol.conv_hitran.hitran_to_sols() contains a usage example of HITRAN data.

4.6 Script link.py

```
usage: link.py [-h] [-l] [-p] [-y] [directory]
Creates symbolic links to PFANT data files as an alternative to copying these (sometimes large)_
→files into local directory
A star is specified by three data files whose typical names are:
main.dat, abonds.dat, and dissoc.dat .
The other data files (atomic/molecular lines, partition function, etc.)
are star-independent, and this script is a proposed solution to keep you from
copying these files for every new case.
How it works: link.py will look inside a given directory and create
symbolic links to files *.dat and *.mod.
The following files will be skipped:
 - main files, e.g. "main.dat"
 - dissoc files, e.g., "dissoc.dat"
 - abonds files, e.g., "abonds.dat"
  - .mod files with a single model inside, e.g., "modeles.mod"
 - hydrogen lines files, e.g., "thalpha", "thbeta"
This script works in two different modes:
a) default mode: looks for files in a subdirectory of PFANT/data
  > link.py common
   (will create links to filess inside PFANT/data/common)
b) "-1" option: lists subdirectories of PFANT/data
c) "-p" option: looks for files in a directory specified.
  Examples:
```

```
> link.py -p /home/user/pfant-common-data
  > link.py -p ../../pfant-common-data
Note: in Windows, this script must be run as administrator.
positional arguments:
             name of directory (either a subdirectory of PFANT/data or the
 directory
             path to a valid system directory (see modes of operation)
              (default: common)
optional arguments:
 -h, --help show this help message and exit
 -1, --list lists subdirectories of
             /home/j/Documents/projects/astro/github/PFANT/code/data
             (default: False)
 -p, --path system path mode (default: False)
             Automatically answers 'yes' to eventual question (default:
 -y, --yes
             False)
```

4.7 Script merge-molecules.py

This script belongs to package pyfant

4.8 Script nist-scraper.py

```
usage: nist-scraper.py [-h] [-u] formula

Retrieves and prints a table of molecular constants from the NIST Chemistry Web Book [NISTRef]

To do so, it uses web scraping to navigate through several pages and parse the desired information from the book web pages.

It does not provide a way to list the molecules yet, but will give an error if the molecule is not found in the NIST web book.

Example:

print-nist.py OH

**Note** This script was designed to work with **diatomic molecules** and may not work with other molecules.

**Warning** The source material online was known to contain mistakes (such as an underscore instead)
```

4.8.1 Usage examples

Usage examples:

```
nist-scraper.py TiO
```

will print

*** titanium oxide ***								
		•	omega_ex_e nu_00 A	omega_ey_e	B_e	alpha_e	gamma_e	u
D	31920.0	1040						
\hookrightarrow			31940					
e 1Sigma+	a + 26598.1	845.2	4.2		0.4892	0.0023		4.7e-
⇔ 07	1.695	e <-> d R	24297.5					
f 1Delta	(a + 19132)	890			0.50221			6.4e-
	1.67292							
	a + 17890.2				0.523	0.00313		3.9e-
	1.6393							
_			4.76	0.047	0.48989	0.00306	-3e-05	6.7e-
	1.69383							
	16331.3		5		0.50617			6.
→86e-07 1.66636 B <-> X R 16066.7 b 1Pi a+11322.0_3 911.2 3.72 0.51337 0.00291					C 1.			
					0.51337	0.00291		6.1e-
	1.65464 14431.0				0 50720	0.00315	-1a-0E	6.
_	2e-09 1.664				0.30739	0.00313	-1e-05	0.
	12025.0							
L 31 1	12023.0							<u>.</u>
					0.54922	0.00337		6e-
d 1Sigma+ a + 2215.6 1014.6 4.64 0.54922 0.00337 6e− →07 1.59972						-		
a 1Delta	а	1009.3	3.93		0.5376	0.00298		5.9e-
	1.61692							
X 3Delta_r	197.5	1009.02	4.498	-0.0107	0.53541	0.00301	-1.1e-05	6.
→ 03e-07	→03e-07 3e-09 1.62022							

4.9 Script run-multi.py

```
usage: run-multi.py [-h] [--abs ABS] [--absoru ABSORU] [--aint AINT]
                    [--allow ALLOW] [--amores AMORES] [--convol CONVOL]
                    [--explain EXPLAIN] [--flam FLAM] [--flprefix FLPREFIX]
                    [--fn_abonds FN_ABONDS] [--fn_absoru2 FN_ABSORU2]
                    [--fn_atoms FN_ATOMS] [--fn_cv FN_CV]
                    [--fn_dissoc FN_DISSOC] [--fn_flux FN_FLUX]
                    [--fn_hmap FN_HMAP] [--fn_lines FN_LINES]
                    [--fn_log FN_LOG] [--fn_logging FN_LOGGING]
                    [--fn_main FN_MAIN] [--fn_modeles FN_MODELES]
                    [--fn_modgrid FN_MODGRID] [--fn_molecules FN_MOLECULES]
                    [--fn_moo FN_MOO] [--fn_opa FN_OPA]
                    [--fn_partit FN_PARTIT] [--fn_progress FN_PROGRESS]
                    [--fwhm FWHM] [--interp INTERP] [--kik KIK] [--kq KQ]
                    [--llfin LLFIN] [--llzero LLZERO]
                    [--logging_console LOGGING_CONSOLE]
                    [--logging_file LOGGING_FILE]
                    [--logging_level LOGGING_LEVEL] [--no_atoms NO_ATOMS]
                    [--no_h NO_H] [--no_molecules NO_MOLECULES] [--norm NORM]
                    [--opa OPA] [--pas PAS] [--pat PAT] [--play PLAY]
                    [--sca SCA] [--zinf ZINF] [--zph ZPH] [-f FN_ABXFWHM]
                    [-s CUSTOM_SESSION_ID]
Runs pfant and nulbad in "multi mode" (equivalent to Tab 4 in ``x.py``) (several abundances X FWHM
→ 's)
optional arguments:
 -h, --help
                        show this help message and exit
  --abs ABS
  --absoru ABSORU
  --aint AINT
  --allow ALLOW
  --amores AMORES
  --convol CONVOL
 --explain EXPLAIN
  --flam FLAM
  --flprefix FLPREFIX
  --fn_abonds FN_ABONDS
  --fn_absoru2 FN_ABSORU2
  --fn_atoms FN_ATOMS
  --fn_cv FN_CV
  --fn_dissoc FN_DISSOC
  --fn_flux FN_FLUX
  --fn_hmap FN_HMAP
  --fn_lines FN_LINES
  --fn_log FN_LOG
  --fn_logging FN_LOGGING
  --fn_main FN_MAIN
  --fn_modeles FN_MODELES
  --fn_modgrid FN_MODGRID
  --fn_molecules FN_MOLECULES
  --fn_moo FN_MOO
  --fn_opa FN_OPA
  --fn_partit FN_PARTIT
  --fn_progress FN_PROGRESS
  --fwhm FWHM
  --interp INTERP
 --kik KIK
 --kq KQ
  --llfin LLFIN
  --llzero LLZERO
```

```
--logging_console LOGGING_CONSOLE
--logging_file LOGGING_FILE
--logging_level LOGGING_LEVEL
--no_atoms NO_ATOMS
--no_h NO_H
--no molecules NO MOLECULES
--norm NORM
--opa OPA
--pas PAS
--pat PAT
--play PLAY
--sca SCA
--zinf ZINF
--zph ZPH
-f FN_ABXFWHM, --fn_abxfwhm FN_ABXFWHM
                      Name of file specifying different abundances and
                      FWHM's (default: abxfwhm.py)
-s CUSTOM_SESSION_ID, --custom_session_id CUSTOM_SESSION_ID
                      Name of directory where output files will be saved
                      (default: multi-session-<i>)
```

4.10 Script run4.py

```
usage: run4.py [-h] [--abs ABS] [--absoru ABSORU] [--aint AINT]
               [--allow ALLOW] [--amores AMORES] [--convol CONVOL]
               [--explain EXPLAIN] [--flam FLAM] [--flprefix FLPREFIX]
               [--fn_abonds FN_ABONDS] [--fn_absoru2 FN_ABSORU2]
               [--fn_atoms FN_ATOMS] [--fn_cv FN_CV] [--fn_dissoc FN_DISSOC]
               [--fn_flux FN_FLUX] [--fn_hmap FN_HMAP] [--fn_lines FN_LINES]
               [--fn_log FN_LOG] [--fn_logging FN_LOGGING] [--fn_main FN_MAIN]
               [--fn_modeles FN_MODELES] [--fn_modgrid FN_MODGRID]
               [--fn_molecules FN_MOLECULES] [--fn_moo FN_MOO]
               [--fn_opa FN_OPA] [--fn_partit FN_PARTIT]
               [--fn_progress FN_PROGRESS] [--fwhm FWHM] [--interp INTERP]
               [--kik KIK] [--kq KQ] [--llfin LLFIN] [--llzero LLZERO]
               [--logging_console LOGGING_CONSOLE]
               [--logging_file LOGGING_FILE] [--logging_level LOGGING_LEVEL]
               [--no_atoms NO_ATOMS] [--no_h NO_H]
               [--no_molecules NO_MOLECULES] [--norm NORM] [--opa OPA]
               [--pas PAS] [--pat PAT] [--play PLAY] [--sca SCA] [--zinf ZINF]
               [--zph ZPH]
Runs the four Fortran binaries in sequence: 'innewmarcs', 'hydro2', 'pfant', 'nulbad'
Check session directory "session-<number>" for log files.
optional arguments:
 -h, --help
                        show this help message and exit
  --abs ABS
  --absoru ABSORU
  --aint AINT
  --allow ALLOW
 --amores AMORES
 --convol CONVOL
 --explain EXPLAIN
 --flam FLAM
  --flprefix FLPREFIX
  --fn_abonds FN_ABONDS
```

```
--fn_absoru2 FN_ABSORU2
--fn_atoms FN_ATOMS
--fn_cv FN_CV
--fn_dissoc FN_DISSOC
--fn_flux FN_FLUX
--fn_hmap FN_HMAP
--fn_lines FN_LINES
--fn_log FN_LOG
--fn_logging FN_LOGGING
--fn_main FN_MAIN
--fn_modeles FN_MODELES
--fn_modgrid FN_MODGRID
--fn_molecules FN_MOLECULES
--fn_moo FN_MOO
--fn_opa FN_OPA
--fn_partit FN_PARTIT
--fn_progress FN_PROGRESS
--fwhm FWHM
--interp INTERP
--kik KIK
--kg KQ
--llfin LLFIN
--11zero LLZERO
--logging_console LOGGING_CONSOLE
--logging_file LOGGING_FILE
--logging_level LOGGING_LEVEL
--no_atoms NO_ATOMS
--no_h NO_H
--no_molecules NO_MOLECULES
--norm NORM
--opa OPA
--pas PAS
--pat PAT
--play PLAY
--sca SCA
--zinf ZINF
--zph ZPH
```

4.11 Script vald3-to-atoms.py

4.12 Script abed.py

This script belongs to package pyfant

4.13 Script ated.py

This script belongs to package pyfant

4.14 Script convmol.py

```
--fn_config [FN_CONFIG]

File name for Configuration file for molecular lines

conversion GUI (Python code) (default:

configconvmol.py)
```

4.15 Script mained.py

This script belongs to package pyfant

4.16 Script mced.py

This script belongs to package pyfant

4.17 Script mled.py

This script belongs to package pyfant

4.18 Script moldbed.py

This script belongs to package pyfant

4.19 Script optionsed.py

This script belongs to package pyfant

4.20 Script tune-zinf.py

```
usage: tune-zinf.py [-h] [--min [MIN]] [--max [MAX]] [--inflate [INFLATE]]
                    [--ge_current] [--no_clean]
                    fn_input [fn_output]
Tunes the "zinf" parameter for each atomic line in atomic lines file
The "zinf" parameter is a distance in angstrom from the centre of an atomic
line. It specifies the calculation range for the line:
[centre-zinf, centre+zinf].
This script runs pfant for each atomic line to determine the width of each
atomic line and thus zinf.
Note: pfant is run using most of its default settings and will require the
following files to exist in the current directory:
 - main.dat
 - dissoc.dat
 - abonds.dat
 - modeles.mod
 - partit.dat
 - absoru2.dat
Note: the precision in the zinf found depends on the calculation step ("pas")
specified in main.dat. A higher "pas" means lower precision and a tendency to
get higher zinf's. This is really not critical. pas=0.02 or pas=0.04 should do.
```

positional arguments:

fn_input input file name

fn_output output file name (default: <made-up filename>)

optional arguments:

-h, --help show this help message and exit

--min [MIN] minimum zinf. If zinf found for a particular line is

smaller than this value, this value will be used

instead (default: 0.1)

--max [MAX] maximum zinf. If zinf found for a particular line is

 $\ensuremath{\mathsf{greater}}$ than this value, this value will be used

instead (default: 50.0)

--inflate [INFLATE] Multiplicative constant to apply a "safety margin".

Each zinf found will be multiplied by this value. For example a value of INFLATE=1.1 means that all the zinf's saved will be 10 percent larger than those

calculated (default: 1.1)

--ge_current "Greater or Equal to current": If this option is set,

the current zinf in the atomic lines file is used as a

lower boundary. (default: False)

(default: False)

This script belongs to package *pyfant*

4.21 Script x.py

```
usage: x.py [-h]
PFANT Launcher -- Graphical Interface for Spectral Synthesis
Single and multi modes.
Multi mode
Runs pfant for different abundances for each element, then run nulbad for each
pfant result for different FWHMs.
The configuration is read from a .py file.
The user must specify a list of FWHM values for nulbad convolutions, and
a dictionary containing element symbols and respective list containing n_abdif
differential abundances to be used for each element.
pfant will be run n_abdif times, each time adding to each element in ab the i-th
value in the vector for the corresponding element.
nulbad will run n_abdif*n_fwhms times, where n_fwhms is the number of different
FWHMs specified.
The result will be
- several spectra saved as "<star name><pfant name or counter>.sp"
- several "spectra list" files saved as "cv_<FWHM>.spl". As the file indicates,
 each ".spl" file will have the names of the spectrum files for a specific FWHM.
 .spl files are subject to input for lineplot.py by E.Cantelli
optional arguments:
```

4.21. Script x.py 33

-h, --help show this help message and exit

This script belongs to package *pyfant*