

Heat Transfer

- Energy Transfer through a temperature difference
- (i) Conduction - Energy Transfer due to the direct interaction of particles

Simple Conductor: $\dot{Q}_{cond} = KA_c \frac{T_1 - T_2}{L}$

↳ From Fourier's Law: $\dot{Q}_{cond} = -KA_c \frac{dT}{dx}$

- Gases & Solids due to crystallinity

- (ii) Convection: Enhanced conduction through bulk fluid motion

↳ moving fluid has large temperature gradient

Described by: $\dot{Q}_{conv} = h \cdot A_s (T_s - T_\infty)$

↳ From Newton's Law of Cooling

Heat Transfer Coefficient: $h \uparrow$ as $U_\infty \uparrow$

- (iii) Radiation: Energy transfer due to electromagnetic waves

- Does not require a medium, and all matter at a non-zero temperature emit radiation.

Described by: $\dot{Q}_{emitted} = \epsilon \sigma A_s T^4$

↳ Stefan-Boltzmann Eq: $\dot{Q}_{emitted} = \epsilon \sigma A_s T^4$
where $\sigma = \text{Stefan-Boltzmann constant} = 5.67 \times 10^{-8} \frac{\text{W}}{\text{m}^2 \text{K}^4}$

Conduction in Radial Systems

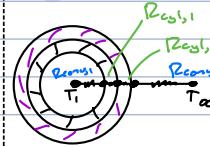
- (i) Cylindrical Coordinates

Still Fourier's Law:

$$\dot{Q}_c = \frac{2\pi L K}{\ln(\frac{r_2}{r_1})} (T_1 - T_2)$$

$$R_{cyl} = \frac{\ln(\frac{r_2}{r_1})}{2\pi L K}$$

Radial can also be in series!!



- (ii) Spherical Coordinates ($A_c = 4\pi r^2$) ← Still Fourier's

$$\dot{Q} = \frac{4\pi r_1 r_2 K}{r_2 - r_1} (T_1 - T_2) \quad R_{sph} = \frac{r_2 - r_1}{4\pi r_1 r_2 K}$$

Finned Heat Transfer

↳ Still uses Fourier's Law to Represent $\dot{Q}_{conduction}$

$$\dot{Q}_{cond} = -KA_c \frac{dT}{dx}$$

$$\Rightarrow Q = \frac{d}{dx} (KA_c \frac{dT}{dx}) - hP \sum T(x) - T_\infty$$

Two Different Boundary Condition Assumptions

- (i) Assume infinitely long fin tip

$$a) T(x \rightarrow \infty) \rightarrow T_\infty \Rightarrow \theta(x \rightarrow \infty) \rightarrow 0 \quad (\theta = T(x) - T_\infty)$$

$$b) T(x=0) \rightarrow T_b \Rightarrow \theta(x=0) \rightarrow \theta_b$$

$$\frac{T(x) - T_\infty}{T_b - T_\infty} = e^{-mx} \quad \text{for } m = \sqrt{\frac{hP}{KA_c}}$$

$$T(x) = (e^{-mx})(T_b - T_\infty) + T_\infty$$

$$\dot{Q} = \sqrt{h P K A_c} (T_b - T_\infty)$$

- (ii) Assume Adiabatic Fin Tip, Finite length

$$a) \dot{Q}|_{x=L} = 0 = -KA_c \frac{dT}{dx}|_{x=L} \Rightarrow \frac{d\theta}{dx}|_{x=L} = 0$$

$$b) \theta(0) = \theta_b$$

$$T(x) - T_\infty = \frac{\cosh[m(L-x)]}{\cosh(mL)} \quad \text{Hyperbolic cosine and tangent}$$

$$\dot{Q} = \dot{Q}|_{x=0} = \sqrt{h P K A_c} (T_b - T_\infty) + \tanh(mL)$$

$$T(x) = \frac{\cosh[m(L-x)]}{\cosh(mL)} (T_b - T_\infty) + T_\infty \quad m = \sqrt{\frac{hP}{KA_c}}$$

where...

$$\cosh(x) = \frac{1 + e^{-2x}}{2e^{-x}} \quad \text{and} \quad \tanh(x) = \frac{e^{2x} - 1}{e^{2x} + 1}$$

Thermal Resistance Network

- (i) Conduction

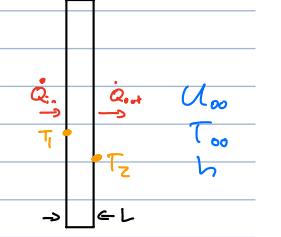
$$\dot{Q}_{cond} = KA_c \frac{(T_1 - T_2)}{L} = KA_c \frac{\Delta T}{L}$$

$$Q_{cond, \text{wall}} = \frac{\Delta T}{R_{cond}} \Rightarrow R_{cond} = \frac{L}{KA_c}$$

- (ii) Convection

$$\dot{Q}_{conv} = h A_s (T_2 - T_\infty) = h A_s \Delta T$$

$$Q_{conv} = \frac{\Delta T}{R_{conv}} \Rightarrow R_{conv} = \frac{1}{h A_s}$$



Series Resistances

$$\bullet \dot{Q} \text{ is constant} \Rightarrow \Delta T = \dot{Q} R_{tot}$$

$$\bullet R_{tot} = R_{conv,1} + R_{cond,1} + R_{cond,2} + R_{cond,3} + R_{conv,2}$$

$$\dot{Q} = \frac{\Delta T}{R_{tot}} = \frac{T_{0oo} - T_1}{R_{conv,1} + R_{cond,1} + R_{cond,2} + R_{cond,3}}$$

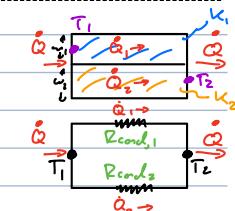
↳ Resist areas must correspond with temperatures

Resistors in Parallel

$$\bullet \Delta T \text{ is constant}, \dot{Q}_i \text{ added individually} \quad \dot{Q}_{tot} = \sum \dot{Q}_i$$

$$\Rightarrow \dot{Q} = \frac{\Delta T}{R_{tot}}$$

- ↳ For some problems like collinear parallel resistances to get $R_{parallel}$ add to the ones in series!!



Transient Heat Transfer

Lumped System Analysis: Assume temp varies only with time $T = T(t)$

$$\frac{T(t) - T_\infty}{T_1 - T_\infty} = e^{-bt} \quad \text{where } b = \frac{h A_s}{\rho V c} = \frac{h}{\rho L c} \quad \text{Assumption holds for short time scale}$$

$$Q(t) = h A_s [T_\infty - T(t)] \quad \text{decreases with time}$$

$$Q_{tot} = mc \sum [T(t_{final}) - T_i] \quad \text{total heat transfer from } t=0 \text{ to } t=f$$

$$Q_{max} = mc [T_\infty - T_i] \quad \text{maximum possible energy change}$$

- Assumption holds for $Biot \# \gg 0.1$ no completely uniform temperature and can use Lumped System Analysis (LSA)

$$Biot \# = Bi = \frac{\text{Convection } Q_{surface}}{\text{Conduction within}} = \frac{h L_c}{K} \quad \text{where } L_c = \sqrt{\frac{V}{A_s}} \quad \text{volume surface area}$$

↳ If only given thickness (t) of object, assume $t \ll \text{width } (w)$

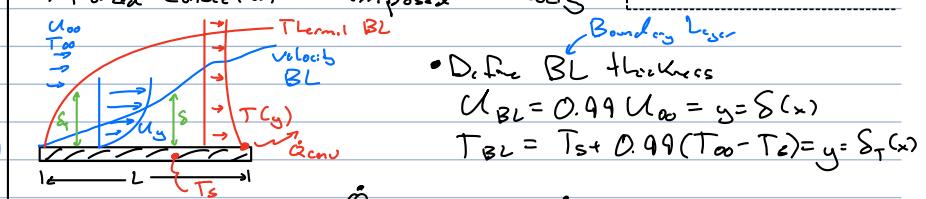
$$= Bi = \frac{V}{A_s} = \frac{L w t}{Z L w + Z t} = \frac{w t}{Z(w+t)} \approx \frac{t}{Z} \quad \text{can use 'b' in 'Bi'}$$

Convection and External Flow

- 1) Natural / Free Convection

- Flow is induced by temp gradients

- 2) Forced Convection → Imposed Velocity



$$\text{Flux: } q''_{conv} = h \Delta T = \frac{\dot{Q}_{conv}}{A_s} = \frac{\dot{Q}_{conv}}{\Delta T} = \frac{q''_{cond}}{A_s} = \frac{q''_{cond} \cdot A_s}{\Delta T}$$

$$Pr = \frac{v \cdot \text{viscosity}}{k \cdot \text{thermal diffusivity}} = \frac{v}{k} \cdot \frac{\text{viscosity}}{\text{thermal diffusivity}} \quad \text{from tables!}$$

Prandtl Number

$$Nusselt Number: Nu_L = \frac{q''_{conv} \cdot L}{q''_{cond} \cdot k} = \frac{q''_{conv} \cdot L}{q''_{cond} \cdot \frac{L}{h}} = \frac{h L}{q''_{cond}}$$

$$(i) \text{ Local: } \text{Laminar: } Nu_x = 0.322 \cdot Re_x^{1/2} \cdot Pr^{1/3}$$

$$\text{b) Turbulent: } Nu_x = 0.0296 \cdot Re_x^{4/5} \cdot Pr^{1/3}$$

$$(ii) \text{ Average: } \text{Note: unless stated take } Re_x^{4/5} \cdot Pr^{1/3}$$

$$\text{a) Laminar: } Nu_L = 0.667 \cdot Re_L^{1/2} \cdot Pr^{1/3}$$

$$\text{b) Turbulent: } Nu_L = 0.027 \cdot Re_L^{4/5} \cdot Pr^{1/3}$$

$$(c) L \& T: Nu_L = (0.037 Re_L^{4/5} - 8.71) \cdot Pr^{1/3}$$

Typical Values

- (i) $Pr \ll 1$ ⇒ liquids

- (ii) $Pr \approx 1$ ⇒ gases (air)

- (iii) $Pr \gg 1$ ⇒ oils, water

Conduction & Lateral Flow

Energy Contained in the Fluid as:

(i) Internal Energy

(ii) Kinetic (KE) & Potential (PE) Energy

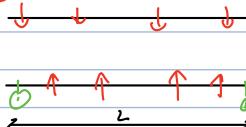
(iii) Flow Work Energy

For true liquids: $\partial h = c_p \Delta T$

Consider Two Possible Cases

(i) q_s^e constant (along x) \rightarrow $q_s^e = \text{const}$ + \dot{Q}_c , i.e. solar radiation

\hookrightarrow constant surface heat flux



$$\dot{Q}_{in} = q_s^e A_s = \dot{m} c_p (T_e - T_i)$$

$$T_e = T_i + \frac{\dot{q}_s^e A_s}{\dot{m} c_p} = T_i + \frac{\dot{q}_s^e P_L}{\dot{m} c_p}$$

$$T_m(x) = \frac{\dot{q}_s^e P_x}{\dot{m} c_p}$$

What about T_s

Newton's Law of Cooling

$$q_s^e = h \sum [T_s(x) - T_m(x)]$$

constant \Rightarrow fully developed

$$\Rightarrow T_s(x) = T_m(x) + \frac{q_s^e}{h}$$

Determining h

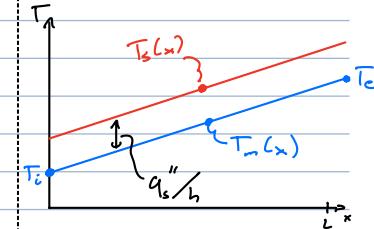
• For Laminar Pipe

• Circular P: pc

• q_s^e constant

$$Nu_p = \frac{h D}{k_{fluid}} = \frac{4.8}{\pi} \approx 4.36$$

$$\Rightarrow h = \frac{4.8}{\pi} \cdot \frac{k_{fluid}}{D}$$



(ii) T_s constant (along x) \rightarrow i.e. ducts, submerged pipes

\hookrightarrow constant surface temperature

$$\dot{Q}_{conv} = h A_s [T_s - T_m(x)]$$

$$\dot{Q}_{in} = h A_s \Delta T_{lm} = \dot{m} c_p (T_e - T_i)$$

$$\hookrightarrow T_s - T_m(x) = \Delta T_{avg} \quad \Delta T_{lm} = \frac{\Delta T_e - \Delta T_i}{\ln(\frac{\Delta T_e}{\Delta T_i})}$$

use log mean diff bc Tm changes

with position x .

$$T_e = T_s - (T_s - T_i) \cdot \left(\exp \left[-\frac{h P_L}{\dot{m} c_p} \right] \right)$$

$$T_m(x) = T_s - (T_s - T_i) \cdot \left(\exp \left[-\frac{h P_x}{\dot{m} c_p} \right] \right)$$

h for constant T_s ,

Laminar flow,

circular p: pc

$$Nu_p = 3.66$$

$$\Rightarrow h = 3.66 \cdot \frac{k_{fluid}}{D}$$

$T_s(x)$

$T_m(x)$

T_e

ΔT_e

Laminar Flow for other cross-sections

TABLE 10-3 Nusselt number and friction factor for fully developed laminar flow in tubes of various cross sections. $D_H = \frac{4A_c}{P}$			
Tube Geometry	D_H or R	$f_f = \text{Const.}$	Nusselt Number
Circle	—	3.66	4.36
Rectangle	0.5	2.98	3.61
	1	3.39	4.12
	2	4.00	4.79
	4	4.44	5.33
	8	5.00	6.00
	16	5.60	6.49
Ellipse	0.5	3.66	4.36
	1	3.74	4.56
	4	4.88	5.53
	8	5.72	6.20
	16	6.65	7.18
Isosceles Triangle	10°	1.61	2.45
	30°	2.26	2.91
	60°	2.67	3.11
	90°	2.34	2.98
	120°	2.60	3.68

Turbulent Flow ($Re_p > 2300$)

Dittus-Boelter Equations

$$Nu_p = 0.023 \cdot Re_p^{4/5} \cdot Pr^n$$

where: $n=0.4$ if fluid is heated

$n=0.3$ if fluid is cooled

Blackbody Radiation

• Perfect emitter and absorber

a) Emits maximum radiation at a given T

b) Absorbs all incident radiation

c) Uniform/Non-uniform

Examples: Sun @ 5800K, Snow & white paint @ 18

Can be temperature or wavelength specific

Blackbody Emissive Power

$$E_b(T) = \sigma T^4$$

$\sigma = 5.67 \times 10^{-8} \frac{\text{W}}{\text{m}^2 \text{K}^4}$

Stefan-Boltzmann constant

Wavelength Dependence (+temperature)

• Different amounts of radiation given off at different wavelengths

\Rightarrow Spectral blackbody emissive power

$$E_{b,\lambda}(T) = \frac{C_1}{\lambda} \left[\exp \left(\frac{C_2}{\lambda T} \right) - 1 \right]$$

$$\text{where } C_1 = 3.74 \times 10^8 \frac{\text{W}}{\text{m}^2 \text{nm}^2}$$

$$C_2 = 1.44 \times 10^4 \frac{\text{nm} \cdot \text{K}}{\text{W}}$$

For wavelength in microns (μm)

Wien's Displacement Law

$$(LT)_{max} = 2897.8 \frac{\text{nm} \cdot \text{K}}{\text{W}}$$

Total Emission (usually most important)

$$E_b(T) = \int_0^{\infty} E_{b,\lambda}(T) d\lambda = \sigma T^4$$

Thermal Radiation

• Radiation requires no medium

• All objects at non-zero (absolute) temperature emit radiation

• Amount of radiation proportional to temperature

• Thermal Radiation is a subset of the full EM spectrum ($0.1 \mu\text{m} \text{ to } 100 \mu\text{m}$)

• Effects that affect radiative transfer

(i) Temperature (ii) Wavelength

(i) Geometry, (iii) Material Properties

Two Primary Types of Materials

(i) Opaque Materials: All radiation absorbed over first few microns of surface ex: wood, bark, metals

(ii) Semi-transparent Materials: Transparency can be wavelength dependent

↳ Greenhouse effect: Visible light through glass gets absorbed and IR released that cannot pass through glass, temperature ↑

Radiative Properties

(i) Emissivity: Real bodies emit a fraction of the blackbody emissive power

$$E(T) = \frac{E_b(T)}{E_b(T)_{ideal}} \sim \text{Ideal Blackbody Emissivity}$$

\Rightarrow Emissivity: Emitted = $E \sigma T^4$

• Properties for incident radiation (irradiance)

(i) Absorptivity (α)

$$\alpha = \frac{G_{absorbed}}{G_{incident}}$$

↳ Incident Radiation

↳ Transmitted Radiation

(ii) Reflectivity (ρ)

$$\rho = \frac{G_{reflected}}{G_{incident}}$$

Energy Balance

$$G = \alpha G + \rho G + \tau G$$

$$\Rightarrow 1 = \alpha + \rho + \tau$$

if opaque ($\tau=0$)

$$\alpha + \rho = 1$$

$$1 = \frac{\log(G_2/G_1)}{m \cdot \epsilon^2} = \frac{J}{m^2}$$