

Roots of the Zeta Function, Random Matrices and Quantum Mechanics

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ABSTRACT. In this paper, we explore three different algorithms to numerically calculate the non-trivial zeros of the Riemann Zeta function, $\zeta(t)$, namely, the Euler-Maclaurin method, alternating series method and the Riemann-Siegel formula. We then introduce random matrices and investigate any possible connection between their eigenvalues and the non-trivial zeros of $\zeta(t)$. We introduce the Hilbert-Polya conjecture as an alternate approach to proving the Riemann hypothesis. We investigate some of the properties of the eigenvalues of the Berry-Keating Hamiltonian.

1. The Riemann Zeta Function

The Riemann Zeta function is a function of a complex variable s , that analytically continues the sum of the infinite series,

$$(1.1) \quad \zeta(s) = \sum_{n=1}^{\infty} \frac{1}{n^s} .$$

$\zeta(s)$ converges for $\text{Re}(s) > 1$, so we will restrict ourself to that domain for this paper. The Zeta function was first studied by Euler in 1740 but his study was limited to real, positive, integer values of s only. Riemann later extended the function to complex valued s in his 1859 paper “On the Number of Primes Less Than a Given Magnitude”. There are multiple ways of defining $\zeta(t)$. It can, for example, also be defined as the integral,

$$(1.2) \quad \zeta(s) = \frac{1}{\Gamma(s)} \int_0^{\infty} \frac{x^{s-1}}{e^x - 1} dx ,$$

where $\Gamma(s) = \int_0^{\infty} x^{s-1} e^{-x} dx$. A particularly important representation of the zeta is using the Euler product formula,

$$(1.3) \quad \zeta(s) = \prod_{p \text{ prime}} \frac{1}{1 - p^{-s}}$$

This form was proved by Euler in 1737 and is important, for example, because it can be used to prove that there are infinitely many primes. Consider the case when

$s = 1$. Then,

$$\begin{aligned} 1 &= \zeta(1) \left(1 - \frac{1}{2}\right) \left(1 - \frac{1}{3}\right) \left(1 - \frac{1}{5}\right) \left(1 - \frac{1}{7}\right) \dots \\ 1 &= \zeta(1) \frac{1}{2} \cdot \frac{2}{3} \cdot \frac{4}{5} \cdot \frac{6}{7} \dots \\ 1 &= \zeta(1) \frac{1 \cdot 2 \cdot 4 \cdot 6 \dots}{2 \cdot 3 \cdot 5 \cdot 7 \dots} \end{aligned}$$

But, $\zeta(1) = 1 + \frac{1}{2} + \frac{1}{3} + \frac{1}{4} + \frac{1}{5} + \frac{1}{7} + \dots$, which gives,

$$1 + \frac{1}{2} + \frac{1}{3} + \frac{1}{4} + \frac{1}{5} + \frac{1}{7} + \dots = \frac{2 \cdot 3 \cdot 5 \cdot 7 \dots}{1 \cdot 2 \cdot 4 \cdot 6 \dots}$$

But the left hand side of this equation diverges, which means the numerator in the right hand side diverges too, which means there must be infinitely many primes.

Another important form of the zeta function is the functional equation, written as,

$$(1.4) \quad \zeta(s) = 2^s \pi^{s-1} \sin\left(\frac{\pi s}{2}\right) \Gamma(1-s) \zeta(1-s).$$

2. Numerical Determination of the Roots of the Zeta Function

2.1. Euler-Maclaurin Formula. The Euler-Maclaurin formula is a very popular tool used in numerical analysis as a connection between integrals and sums. It can be used to reformulate an integral as a sum or vice versa and is very popular in analytical number theory where it is used to compute sums in terms of integrals. The formula was developed in the early eighteenth century and came as a result of trying to extend the sum $\sum_k n^k$ for k 's other than 1. It was developed independently by Leonhard Euler and Colin Maclaurin around 1735- Euler to extend the sum just mentuoned and Maclaurin to compute integrals. The formula can be written as,

$$(2.1) \quad \sum_{n=M}^N f(n) = \int_M^N f(x) dx + \frac{1}{2} [f(N) - f(M)] +$$

$$(2.2) \quad \sum_{j=1}^{\nu} \frac{B_{2j}}{(2j)!} \left[f^{(2j-1)}(N) - f^{(2j-1)}(M) \right] + R$$

The remainder R is given by,

$$(2.3) \quad R_{2\nu} = -\frac{s(s+1)\dots(s+2\nu)}{(2\nu+1)!} \int_N^\infty B_{2\nu+1}(x) x^{-s-2\nu-1} dx.$$

It can be shown that R satisfies the inequality [],

$$(2.4) \quad |R_{2\nu-2}| \leq \left| \frac{s+2\nu-1}{\operatorname{Re}(s)+2\nu-1} \right| \cdot |B_{2\nu}|.$$

This means, if $B_{2\nu}$ is the first term we ignore in our sum, the remainder will be at most $\left| \frac{s+2\nu-1}{\sigma+2\nu-1} \right| \cdot |B_{2\nu}|$.

If we apply the Euler-Maclaurin formula to $\zeta(s) = \sum_1^\infty n^{-s}$, R will not be small enough to be ignored []. So, we instead apply the formula to the sum, $\sum_N^\infty n^{-s}$.

Notice that,

$$\begin{aligned}
 \sum_{n=1}^{\infty} n^{-s} - \sum_{n=1}^{N-1} n^{-s} &= \sum_N^{\infty} n^{-s} \\
 \zeta(s) - \sum_{n=1}^{N-1} n^{-s} &= \sum_N^{\infty} n^{-s} \\
 (2.5) \quad \zeta(s) &= \sum_{n=1}^{N-1} n^{-s} + \sum_N^{\infty} n^{-s} = \sum_{n=1}^{N-1} n^{-s} + \chi(s)
 \end{aligned}$$

The first term is easy to compute and we can use the Euler-Maclaurin formula to compute $\chi(s)$. Applying the formula,

$$\begin{aligned}
 \chi(s) &= \int_N^{\infty} x^{-s} dx + \frac{1}{2}N^{-s} + \sum_{j=1}^{\nu} \frac{B_{2j}}{(2j)!} \left[-(n^{-s})^{(2j-1)}(N) \right] + R \\
 &\approx \frac{1}{s-1}N^{1-s} + \frac{1}{2}N^{-s} + \frac{B_2}{2}sN^{-s-1} + \dots + \left[\frac{B_{2\nu}}{(2\nu)!}s(s+1)\dots(s+2\nu-2)N^{(-s-2\nu+1)} \right]
 \end{aligned}$$

The sum we need to calculate can be divided into four different parts, defined as follows:

$$\begin{aligned}
 A &= \sum_{n=1}^{N-1} n^{-s} \\
 B &= \frac{1}{s-1}N^{1-s} \\
 C &= \frac{1}{2}N^{-s} \\
 D &= \frac{B_2}{2}sN^{-s-1} + \dots + \left[\frac{B_{2\nu}}{(2\nu)!}s(s+1)\dots(s+2\nu-2)N^{(-s-2\nu+1)} \right]
 \end{aligned}$$

Then,

$$\zeta(s) \approx A + B + C + D,$$

with the error at most $\left| \frac{s+2\nu-1}{\text{Re}(s)+2\nu-1} \right| \cdot |B_{2\nu}|$.

This formula can be now used to compute $\zeta(s)$ and is very easy to implement. A sample implementation can be found in Section ??.

2.2. Alternating Series (Dirichlet Eta Function). Consider the following sum:

$$(2.6) \quad \eta(s) = \sum_{n=1}^{\infty} \frac{(-1)^{n-1}}{n^s} = \frac{1}{1^s} - \frac{1}{2^s} + \frac{1}{3^s} - \frac{1}{4^s} + \dots$$

Comparing this to,

$$\zeta(s) = \sum_{n=1}^{\infty} n^{-s} = \frac{1}{1^s} + \frac{1}{2^s} + \frac{1}{3^s} + \frac{1}{4^s} + \dots$$

it's clear that,

$$(2.7) \quad \eta(s) = (1 - 2^{1-s}) \zeta(s).$$

$\eta(s)$ is called the Dirichlet eta function, sometimes also called the alternating zeta function as it's the same sum as zeta function except the sign of the terms alternates between + and -.

Equation 2.7 means that if we can find a way to evaluate $\eta(s)$, we can evaluate $\zeta(s)$ using that.

Consider using Euler's transformation of alternating series to the series defined in Equation 2.6 to accelerate its convergence. Euler's transform of alternating series is defined as,

$$(2.8) \quad \sum_{n=0}^{\infty} (-1)^n a_n = \sum_{n=0}^{\infty} (-1)^n a_0 \frac{1}{2^{n+1}} \Delta^n,$$

where Δ^n is the n^{th} -order forward difference of the function $f(x)$ given by,

$$\Delta^n[f](x) = \sum_{k=0}^n \binom{n}{k} (-1)^{n-k} f(x+k)$$

Comparing 2.8 to $\eta(s) = \sum_{n=1}^{\infty} \frac{(-1)^{n-1}}{n^s} = \sum_{n=0}^{\infty} \frac{(-1)^n}{(n+1)^s}$, it is clear that,

$$\begin{aligned} \eta(s) &= \sum_{n=0}^{\infty} (-1)^n 1^{-s} \frac{1}{2^{n+1}} \Delta^n \\ &= \sum_{n=0}^{\infty} (-1)^n 1^{-s} \frac{1}{2^{n+1}} \sum_{k=0}^n \binom{n}{k} \frac{(-1)^n}{(-1)^k} \frac{1}{(k+1)^s} \\ &= \sum_{n=0}^{\infty} \frac{1}{2^{n+1}} \sum_{k=0}^n \binom{n}{k} \frac{(-1)^k}{(k+1)^s} \end{aligned}$$

This sum converges much faster than the original sum as it is exponential in n and so, is much more efficient. A Python implementation that calculates this sum and uses it to calculate the Zeta function can be found in Section. ??.

2.3. Riemann-Siegel Formula. We won't consider the derivation of the Riemann-Siegel formula here but the formula can be expressed as,

$$(2.9) \quad Z(t) = 2 \sum_{n^2 < (t/2\pi)} n^{-1/2} \cos[\theta(t) - t \log n] + R,$$

with

$$(2.10) \quad \theta(t) = \Gamma \left[\frac{2it+1}{4} \right] - t \frac{\log \pi}{2}$$

$$(2.11) \quad R = \frac{e^{-i\nu(t)} e^{-t\pi/2}}{(2\pi)^{1/2} (2\pi)^{it} e^{-it/4} (1 - ie^{-t\pi})} \int_{C_N} \frac{(-x)^{-(1/2)+it} e^{-Nx}}{e^x - 1} dx,$$

where C_n is a contour which descends the real axis from ∞ to $(2N+1)\pi$, circles the boundary of the disk $|s| \leq (2N+1)\pi$ once and goes back to ∞ ??.

The remainder R can be simplified and expressed in the following form. The details are in Edwards book.

$$(2.12) \quad R \sim (-1)^{N-1} \left(\frac{t}{2\pi} \right)^{-1/4} \left[C_0 + C_1 \left(\frac{t}{2\pi} \right)^{-1/2} + C_3 \left(\frac{t}{2\pi} \right)^{-3/2} + C_4 \left(\frac{t}{2\pi} \right)^{-4/2} \right]$$

$$\begin{aligned}
C_0 &= \Psi(p) = \frac{\cos(p^2 - p - 1/16)}{\cos 2\pi p} \\
C_1 &= -\frac{1}{2^5 \cdot 3 \cdot \pi^2} \Psi^{(3)}(p) \\
C_2 &= \frac{1}{2^{11} \cdot 3^4 \cdot \pi^4} \Psi^{(6)}(p) + \frac{1}{2^6 \cdot \pi^2} \Psi^{(2)}(p) \\
C_3 &= -\frac{1}{2^{16} \cdot 3^4 \cdot \pi^6} \Psi^{(9)}(p) - \frac{1}{2^8 \cdot 3 \cdot 5 \cdot \pi^4} \Psi^{(5)}(p) - \frac{1}{2^6 \cdot \pi^2} \Psi^{(1)}(p) \\
C_4 &= -\frac{1}{2^{23} \cdot 3^5 \cdot \pi^8} \Psi^{(12)}(p) - \frac{11}{2^{17} \cdot 3^2 \cdot 5 \cdot \pi^6} \Psi^{(8)}(p) + \frac{19}{2^{13} \cdot 3 \cdot \pi^4} \Psi^{(4)}(p) + \frac{1}{2^7 \cdot \pi^2} \Psi(p)
\end{aligned}$$

where N is the integer part of $(t/2\pi)^{1/2}$ and p the fractional part.

The coefficients C 's are not hard to calculate. However, a more efficient approach would be to compute the Taylor series expansion of $\Psi(p)$ and then express all the C 's in terms of those coefficients. Consider the series expansion of $\Psi(p) = \sum_i a_i p^i$.

$$\begin{aligned}
\Psi(p) &= a_0 + a_1 p^1 + a_2 p^2 + a_3 p^3 + a_4 + p^4 + a_5 p^5 + \dots \\
\Psi'(p) &= a_1 + 2a_2 p + 3a_3 p^2 + 4a_4 p^3 + 5a_5 p^4 + \dots \\
\Psi''(p) &= 2a_2 + 3 \cdot 2a_3 p + 4 \cdot 3a_4 p^2 + 5 \cdot 4 \cdot a_5 p^3 + \dots \\
\Psi^{(3)}(p) &= 3 \cdot 2a_3 + 4 \cdot 3 \cdot 2a_4 p^1 + 5 \cdot 4 \cdot 3 \cdot a_5 p^2 + \dots
\end{aligned}$$

So, in general,

$$\Psi^{(m)}(p) = m! \cdot a_m + (1+m)! \cdot a_{m+1} + \frac{(2+m)!}{2!} \cdot a_{m+2} + \frac{(3+m)!}{2!} a_{m+3} + \dots$$

So, the if the i -th coefficient of series expansion of a function $\Psi(x)$ is a_i , the coefficient of $\Psi^{(m)}(p)$ is given by,

$$a_i^{(m)} = a_{i+m} \frac{(i+m)!}{i!}$$

This means, once we calculate the seires coefficients of C_0 , calculating the coefficients of other C 's does not take long. The code that does this can be found in Section. ???. The code that implements the Riemann Siegel formula gives a choice between using exact coefficients or using a series coefficients.

3. Root Finding

Now that we have three different ways to numerically compute $\zeta(t)$, it's very straightfoward to compute the zeros. We use a simple, binary-search algorithm to compute the roots, with a choice of any of the three algorithms discussed above to calculate $\zeta(t)$. We select an interval and keep bisecting it until we find an interval inside which the real part of $\zeta(t)$ changes sign, and that's narrower than whatever precision we're calculating the roots to. The first 100 roots are listed in Table 3. These were computed using the alternating series method. While the Riemann-Siegel formula is faster as it sums a smaller number of terms, for small t 's, it's not as accurate as the Euler-Maclaurin or alternating series method, particularly because we don't consider the full remainder R .

If we take the roots of $\zeta(t)$ and compute the successive differences and plot them, we get the plot in Figure 1. The successive differences have been normalized so that the mean is 1.

N	Root	N	Root	N	Root	N	Root
1	14.1347251417	26	92.4918992706	51	147.4227653426	76	202.4935945141
2	21.0220396388	27	94.6513440405	52	153.0246938112	77	204.1896718031
3	25.0108575801	28	95.8706342282	53	156.1129092942	78	205.3946972022
4	30.4248761259	29	98.8311942182	54	157.5975918176	79	207.9062588878
5	32.9350615877	30	101.3178510057	55	158.8499881714	80	209.5765097169
6	37.5861781588	31	103.7255380405	56	161.1889641376	81	211.6908625954
7	40.9187190121	32	105.4466230523	57	163.0307096872	82	213.3479193597
8	43.3270732809	33	107.1686111843	58	165.5370691879	83	214.5470447835
9	48.0051508812	34	111.0295355432	59	167.1844399782	84	216.1695385083
10	49.7738324777	35	114.3202209155	60	173.4115365196	85	219.0675963490
11	52.9703214777	36	116.2266803209	61	174.7541915234	86	220.7149188393
12	56.4462476971	37	118.7907828660	62	176.4414342977	87	221.4307055547
13	59.3470440026	38	121.3701250024	63	178.3774077761	88	227.4214442797
14	60.8317785246	39	122.9468292935	64	179.9164840203	89	229.3374133055
15	65.1125440481	40	124.2568185543	65	182.2070784844	90	233.6934041789
16	67.0798105295	41	127.5166838796	66	184.8744678484	91	236.5242296658
17	69.5464017112	42	129.5787042000	67	185.5987836777	92	237.7698204809
18	72.0671576745	43	131.0876885309	68	187.2289225835	93	239.5554775733
19	75.7046906991	44	133.4977372030	69	189.4161586560	94	241.0491577962
20	77.1448400689	45	134.7565097534	70	192.0266563607	95	242.8232719342
21	79.3373750203	46	138.1160420545	71	193.0797266038	96	244.0708984971
22	82.9103808541	47	139.7362089521	72	195.2653966795	97	247.1369900749
23	84.7354929805	48	141.1237074040	73	196.8764818410	98	248.1019900602
24	87.4252746131	49	143.1118458076	74	198.0153096763	99	249.5736896447
25	88.8091112076	50	146.0009824868	75	201.2647519437	100	251.0149477950

4. Random Matrices

A random matrix is a matrix whose elements are random variables. Two most commonly studied types of random matrices are Gaussian Orthogonal Ensemble (GOE) and Gaussian Unitary Ensemble (GUE).

4.1. Gaussian Orthogonal Ensemble (GOE). A Gaussian Orthogonal Ensemble is a set of $n \times n$ real-valued symmetric matrices H_n whose Gaussian measure has the density,

$$\frac{1}{Z_{\text{GOE}}} e^{-\frac{n}{4} \text{tr}(H^2)}.$$

GOEs model Hamiltonians with time-reversal symmetry.

Consider the ordered sequence of eigenvalues of H_n , $\lambda_1 < \lambda_2 < \dots < \lambda_{n-1} < \lambda_n$. Define $s = \frac{\lambda_{i+1} - \lambda_i}{(s)}$. Then, the probability distribution on s is given by,

$$p_{\text{GOE}}(s) = \frac{\pi}{2} s e^{-\frac{\pi}{4} s^2}$$

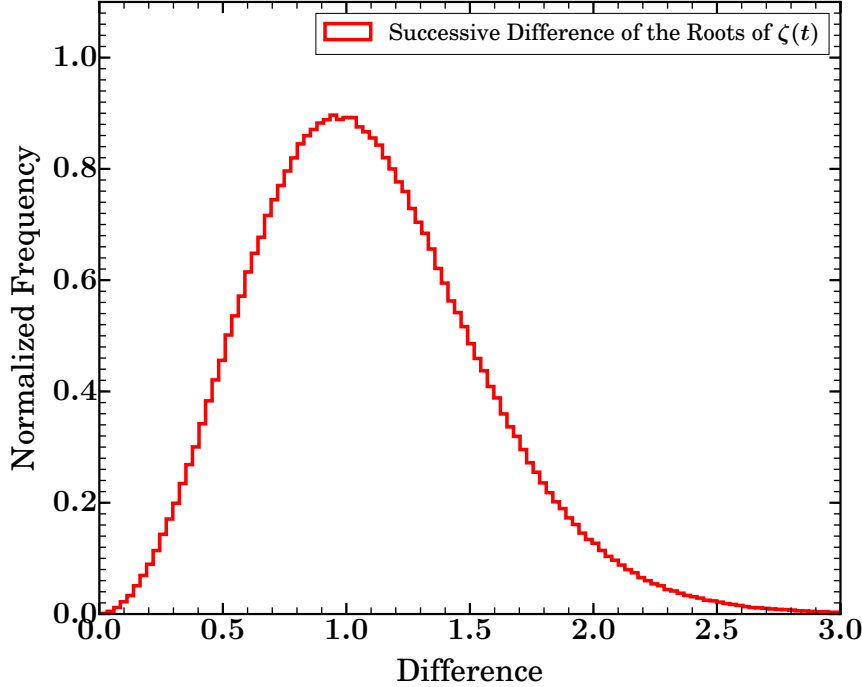


FIGURE 1. Distribution of successive differences of the first 2 million non-trivial zeros of $\zeta(t)$. The differences have been normalized so that the mean is 1.

4.2. Gaussian Unitary Ensemble (GUE). A Gaussian Unitary Ensemble is a set of $n \times n$ Hermitian matrices H_n whose Gaussian measure has the density,

$$\frac{1}{Z_{\text{GUE}}} e^{-\frac{n}{2} \text{tr}(H^2)},$$

with $Z_{\text{GUE}} = 2^{\pi/2} \pi^{n^2/2}$. GUEs model Hamiltonians lacking time-reversal symmetry.

For GUEs, the probability distribution on s , is given by,

$$p_{\text{GUE}}(s) = \frac{32}{\pi^2} s^2 e^{-\frac{4}{\pi} s^2}$$

5. Riemann Hypothesis

The Riemann hypothesis conjectures that all non-trivial roots of the zeta function have real-part $1/2$. It was proposed by Riemann in 1859 in the same paper where he introduced the zeta function, “On the Number of Primes Less Than a Given Magnitude”.

The Riemann hypothesis is arguably one of the most important unsolved problems in mathematics. Several approaches to proving it has been attempted but none has succeeded so far. Numerically, it has been tested to be true for the first

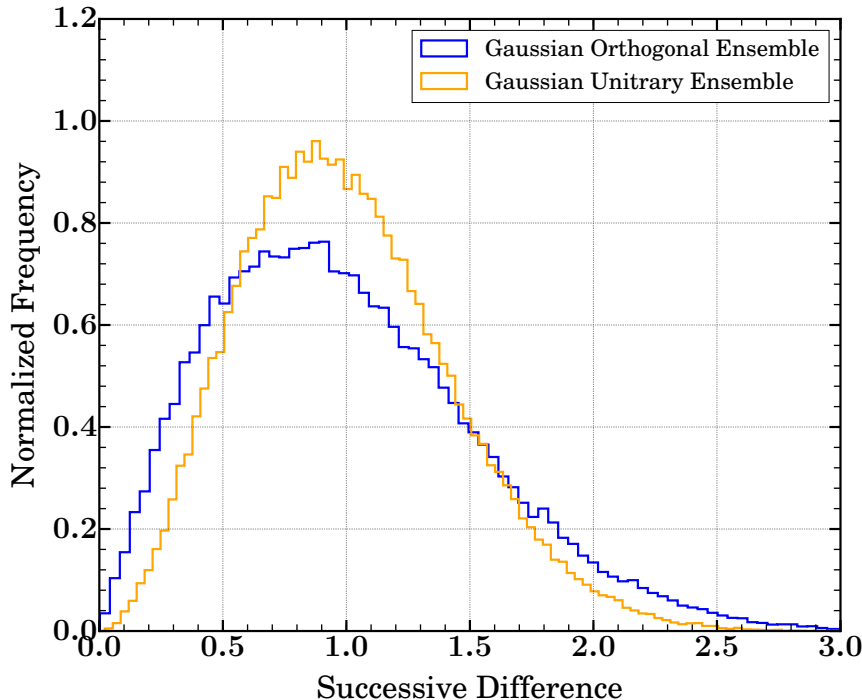


FIGURE 2. Successive eigenvalue differences for Gaussian Orthogonal and Gaussian Unitary Ensembles of 100 K 25×25 matrices. Each difference corresponds to $\lambda_{\lceil N \rceil + 1} - \lambda_{\lceil N \rceil}$.

10^{13} roots by X. Gourdon and Patrick Demichel [1]. However, numerical verification, no matter to how large values of t , can not be considered to be a compelling argument for a mathematical hypothesis to be true. One notable example is related to Skewes' number where a related conjecture breaks down for the first time at a number around 10^{316} .

6. Hilbert-Polya Conjecture

The Hilbert-Polya conjecture suggests that the Riemann hypothesis would be true if the imaginary parts of the zeros of the zeta function corresponds to the eigenvalues of some unbounded Hermitian operator. This came out of letter to Andrew Odlyzko, in 1982, in which Polya said that he was asked by Edmund Landau for a physical reason why the Riemann hypothesis must be true.

The eigenvalues of a physical quantum operator are all real values and Hermitian operators are guaranteed to have real eigenvalues. This means, a physical way of proving the Riemann hypothesis would be to try to find a quantum mechanical operator whose eigenvalues (which will be all real) coincide with the imaginary parts of the non-trivial zeros of the Riemann zeta function. Finding any such quantum mechanical operator would be a proof of the Hilbert-Polya conjecture and the Riemann hypothesis both. So in a sense, the Hilbert-Polya conjecture is just an

alternate formulation of the Riemann hypothesis that makes it possible to try to find a physical proof of the Riemann hypothesis.

It's important to note that there are infinitely many non-trivial zeros the $\zeta(t)$?? and so, if such a Hermitian matrix does exist, as conjectured by the Hilbert-Polya conjecture, that matrix would have to be infinite-dimensional. Of course, that's not a problem since most quantum mechanical operators, even the simplest ones like say the Hamiltonian corresponding the a simple harmonic oscillator, have matrix forms where the dimension is infinite-dimensional. But it's something that one needs to be aware of since, for example, because no numerical calculations can go until infinity, numerical methods of finding / testing such quantum systems must be careful, and certain properties, the most notable example being the commutation relationship, won't be valid when we arrest the matrix representation to a finite size for numerical computations. The first one of this will be particularly important in Section. ?? where we discuss one such possible quantum system.

Very broadly speaking, there are now two approaches to prove the Riemann hypothesis. Through the Hilber-Polya conjecture, we've reduced the problem to finding some quantum mechanical system with a specific property. Because the operator (most likely the Hamiltonian) needs to be Hermitian, it's reasonable to expect, though this certainly does not have to be true, would be some function of the most common physical quantum mechanical operators. Finiding such an operator and testing whether or not its eigenvalues corresponds to the imaginary part of the non-trivial zeros of $\zeta(t)$ could be done numerically or we could try to find an analytical expression for those eigenvalues of our operator, in which case the analytical form would probably reduce to essentially one of the already known forms for $\zeta(t)$ equated to 0. Of course, it's not unconceivable that this would lead to an alternate expression for $\zeta(t)$ but nevertheless, the Hilbert-Polya conjecture opens up the door for attempts to prove the Riemann hypothesis in an alternate way.

7. $\zeta(t)$ and Random Matrices

If we're to prove the Riemann hypothesis, a careful study of the distribution of the non-trivial zeros of $\zeta(t)$ is crucial. In 1973, Montgomery conjectured that there might be some possible connection between those non-trivial zeros and random matrices. Of particular interest is to study the distribution of successive differences of eigenvalues of the Gaussian Unitary Ensemble (GUE) and compare it to the successive differences of the non-trivial zeros of the zeta function.

It's interesting to note that the Gaussian Unitary Ensemble is a realization of every possible Hermitian matrices (assuming the ensemble size goes to infinity). But, through the Hilbert-Polya conjecture, there seems to be at least some connection between the eigenvalues of some such Hermitian matrix and the non-trivial zeros of the zeta function.

We need to choose some variable that is well-defined for both the GUEs and the distribution of the zeros of $\zeta(t)$ and then we can compare that variable for the two cases. One such possible variable is to consider the 2-point correlation function $R_2(E_1, E_2)$ which gives the probability density function of finding an eigenvalue of H at each of the values E_1, E_2 . Montogomery, in 1972, showed that the 2-point

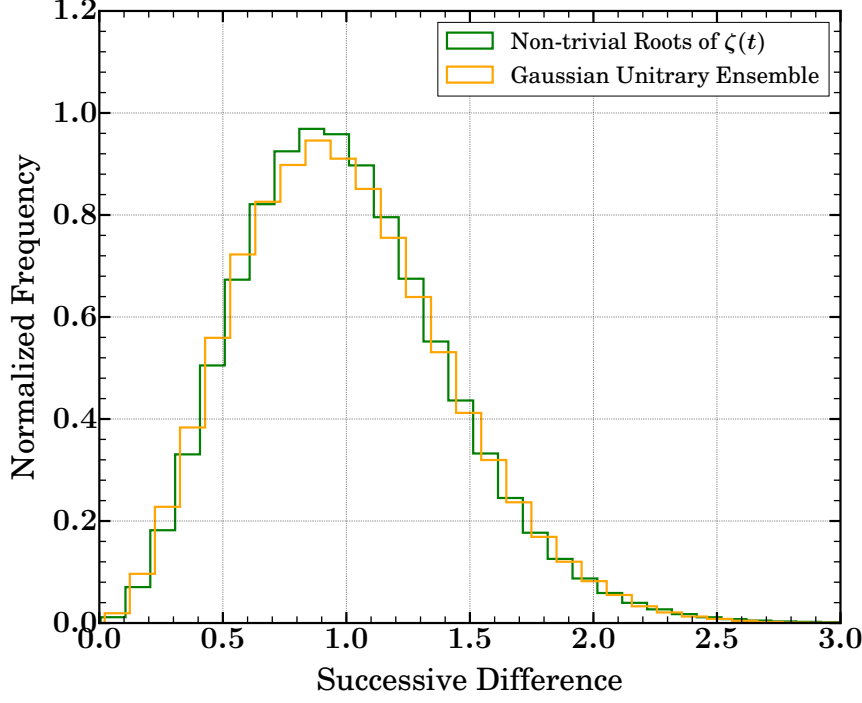


FIGURE 3. Successive eigenvalue differences for Gaussian Unitary Ensembles of 100K 25×25 matrices and the first 2M non-trivial zeros of $\zeta(t)$. For GUE, each difference corresponds to $\lambda_{[N]+1} - \lambda_{[N]}$. Both distributions have been normalized so that their mean is 1.

correlation function for GUEs is given by,

$$R_2(r) = 1 - \left[\frac{\sin \pi r}{\pi r} \right]^2,$$

and the 2-point correlation function for the Riemann zeros is,

$$R_{\zeta,2}(r) = 1 - \left[\frac{\sin \pi r}{\pi r} \right]^2.$$

This is remarkable since this implies that **(What does this imply? –at)**

Another interesting variable to compare is the level-spacing of the zeros and the eigenvalues. What this means is, we could compute the successive differences (after sorting the eigenvalues / zeros) and compare them, subject to appropriate normalization. This is exactly the probability distribution we defined in ?? as $p_{\text{ensemble}}(s)$. We normalize both the distributions so that the mean is 1, as that makes it easy to compare them.

8. Quantum Mechanics

Michael Berry and Jonathan Keating have conjectured that the Hilbert-Polya Hamiltonian H is some quantization of the Hamiltonian xp . We'll consider one such possible quantization and discuss it at length in this section. Let's define:

$$(8.1) \quad H = \frac{1}{2} (xp + px) = -i \left(x \frac{\partial}{\partial x} + \frac{1}{2} \right)$$

For a simple harmonic oscillator, x and p can be written in terms of raising and lowering operators a and a^\dagger as,

$$(8.2) \quad x = \sqrt{\frac{\hbar}{2m\omega}}(a + a^\dagger); p = i\sqrt{\frac{m\omega\hbar}{2}}(a^\dagger - a)$$

Let's ignore the factors in the front and just consider the linear combination of a and a^\dagger along with the imaginary unit. Using this expansion, it's fairly straightforward to see that,

$$(8.3) \quad H = \frac{1}{2} (xp + px) = i [(a^\dagger)^2 - (a)^2]$$

a^\dagger and a has infinite-dimensional matrix representations as follows:

$$a^\dagger = \begin{bmatrix} 0 & 0 & 0 & 0 & \dots \\ \sqrt{1} & 0 & 0 & 0 & \dots \\ 0 & \sqrt{2} & 0 & 0 & \dots \\ 0 & 0 & \sqrt{3} & 0 & \dots \\ 0 & 0 & 0 & \sqrt{4} & \dots \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ \dots & \dots & \dots & \dots & \dots \end{bmatrix}, a = \begin{bmatrix} 0 & \sqrt{1} & 0 & 0 & \dots \\ 0 & 0 & \sqrt{2} & 0 & \dots \\ 0 & 0 & 0 & \sqrt{3} & \dots \\ 0 & 0 & 0 & 0 & \dots \\ 0 & 0 & 0 & 0 & \dots \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ \dots & \dots & \dots & \dots & \dots \end{bmatrix}$$

For our computation, we will consider N dimensional matrices instead of infinite dimensional.

Let's study what the eigenvalues of H look like. Because a and a^\dagger are sparse matrices, computing their eigenvalues, even for large N , is not too bad.

We again have to make a choice of what properties of the operator H (and its eigenvalues) we want to study. The obvious choice is to study the successive difference of eigenvalues like we did for random matrices and the non-trivial zeros of $\zeta(t)$. Figure 4 shows a plot. It's interesting to note that **(what? -at)**

An interesting feature of the plot in Figure. 4 is that, while the distributions for the non-trivial roots and GUE seems to go down all the way to 0, there seems to be a finite cut for the eigenvalue difference for Berry's Hamiltonian H . The existence of this finite cut isn't a remarkable property by itself since this just means that there is no degeneracy for the Hamiltonian we're considering. What's interesting though is that, this cut seems to depend on the size of the matrix representation we choose N .

Figure 5 shows a plot of how the minimum successive eigenvalue difference (minimum level spacing) varies with the size of the matrix we choose to represent our Hamiltonian H . Two immediately interesting things can be noticed here. First, there seems to be a very well-defined, very specific way in which the minimum difference (which corresponds to the left-cutoff in Figure 4) varies with N . Second,

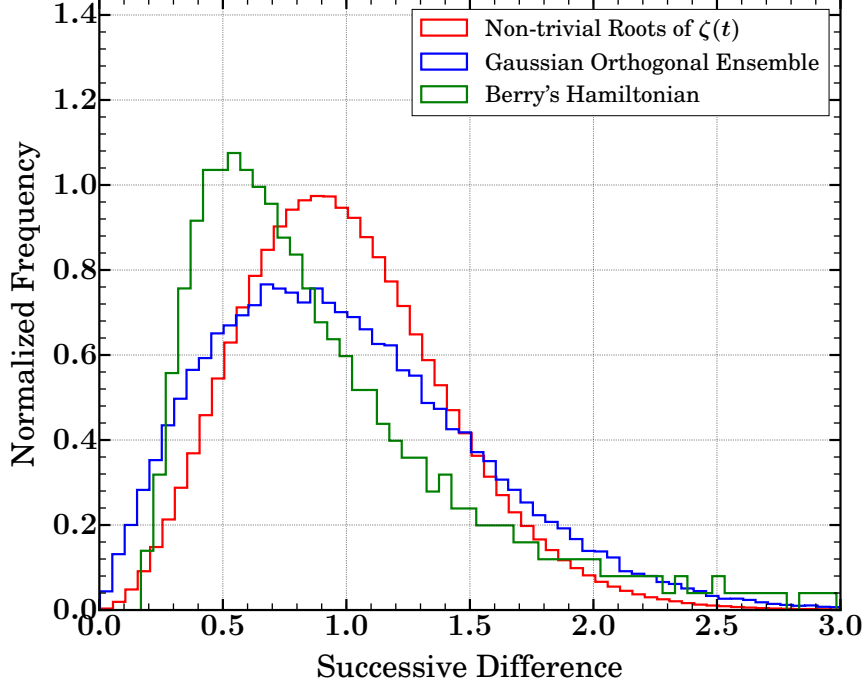


FIGURE 4. Successive eigenvalue differences for Gaussian Orthogonal Ensembles of 100K 25×25 matrices, the first 2M non-trivial zeros of $\zeta(t)$ and the Berry Hamiltonian, H . For H , we consider a 1000×1000 matrix representation. For GOE, each difference corresponds to $\lambda_{\lceil N \rceil + 1} - \lambda_{\lceil N \rceil}$. All distributions have been normalized so that their mean is 1.

that variation seems to be dependent on the parity of N . In particular, for odd values of N , the minimum difference seems to be larger than for what would otherwise be defined by whatever larger-scale model is applicable.

Figure 6 shows a histogram of eigenvalues for a given $N \times N$ representation of H . The property we're particularly interested in is that the eigenvalues seem to be clustered around the origin. So, we conjecture, without proof, that the minimum eigenvalue difference is always going to be the difference of the two eigenvalues closest to the origin.

However, there is a slight technicality here. For odd values of N , the matrix representation of H is not invertible. In particular, it's rank is always going to be $N - 1$, which means it will have at most $N - 1$ non-zero eigenvalues. Said in other words, there will be one or more 0 eigenvalues for odd N . However, if there were more than one 0 eigenvalues, the minimum eigenvalue difference would be 0, which doesn't seem to be the case for finite N according to Figure 5. This means, for odd N 's, there will always be one zero eigenvalue.

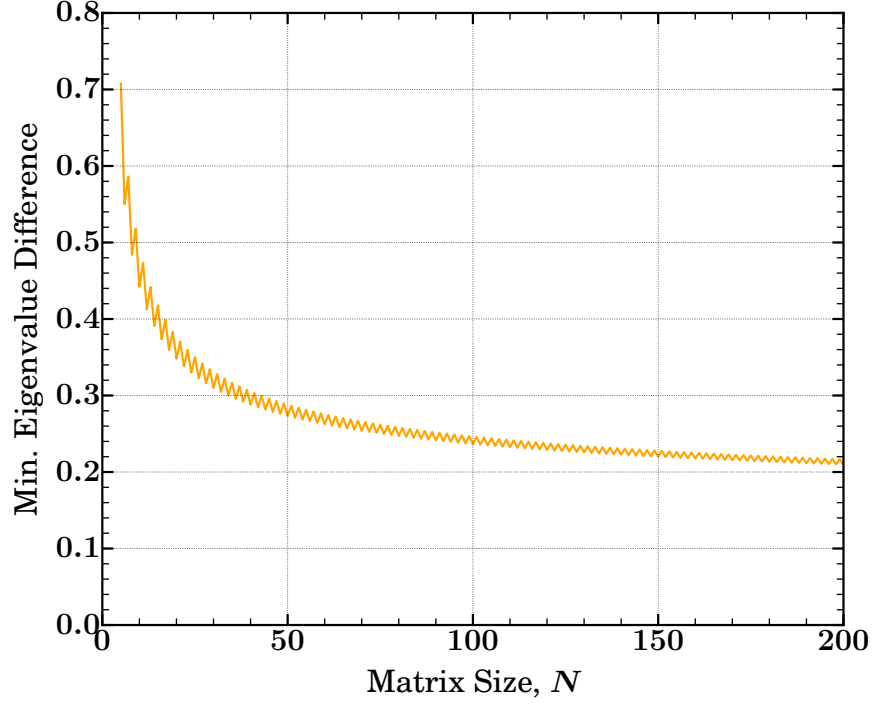


FIGURE 5. A plot of how the minimum successive eigenvalue difference (minimum level spacing) varies with the size of the matrix we choose to represent our Hamiltonian H .

Because our matrix representation of H only contains real entries, all its eigenvalues will occur as complex-conjugate pairs. These get divided by the i factor outside to give real eigenvalues but this means that the eigenvalues will be symmetric about the origin, the only difference between odd and even N being whether or not 0 is one of those eigenvalues.

For odd N 's, since we always have a 0 eigenvalue, the min. difference would be just $|\lambda_{+1}^{\text{odd}N} - \lambda_0^{\text{odd}N}| = |\lambda_{+1}^{\text{odd}N} - 0| = \lambda_{+1}^{\text{odd}N}$. However, for even values of N , since there is no zero eigenvalue, the minimum difference will be $|\lambda_{+1}^{\text{even}N} - \lambda_{-1}^{\text{even}N}| = |2\lambda_{+1}^{\text{even}N}| = 2\lambda_{+1}^{\text{even}N}$. Naively, it'd seem that therefore the minimum difference for even N 's would be larger than for odd N 's. However, as suggested by Figure 5, this is not true. It's important to note that $\lambda_{+1}^{\text{even}N} \neq \lambda_{+1}^{\text{odd}N}$, and so our argument breaks down. In fact, because the minimum difference for odd N 's seems to always be larger than for even N 's, we can say that $2\lambda_{+1}^{\text{even}N} < \lambda_{+1}^{\text{odd}N}$. Of course, this relation is not exactly well defined since N will either be odd or will be even so there isn't a proper way to compare λ_{+1} since the matrix size changes. So an implicit assumption in the relation is, "what would normally be defined by the broader large-scale model". Figure 7 demonstrates our claim.

Another important property of the plot in Figure 5 we're interested in is the asymptotic behavior of the difference as $N \rightarrow \infty$. Ultimately, our aim is to find

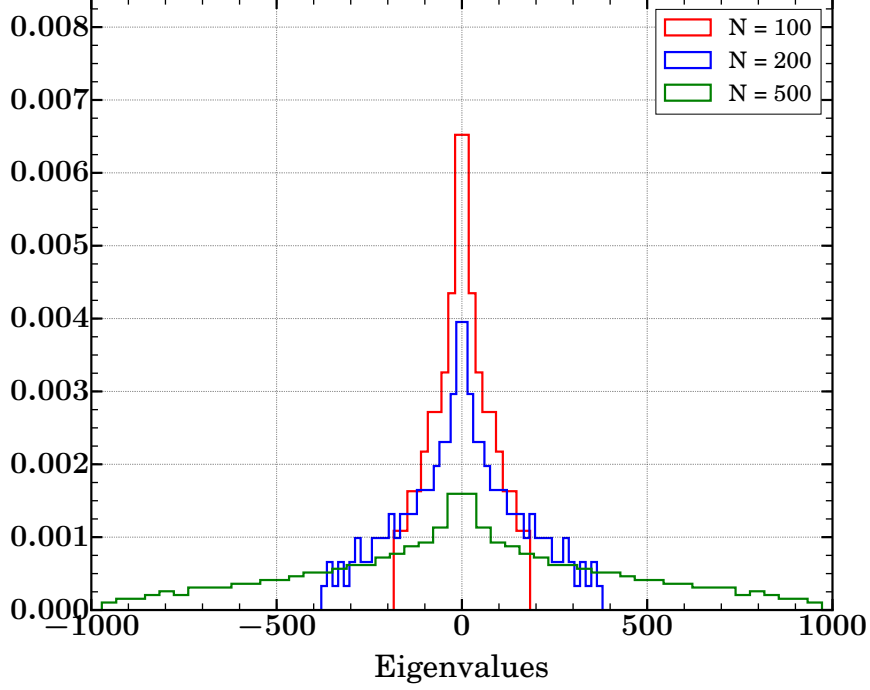


FIGURE 6. A plot of how the minimum successive eigenvalue difference (minimum level spacing) varies with the size of the matrix we choose to represent our Hamiltonian H .

a Hilbert-Polya Hamiltonian and that would have an infinite-dimensional matrix representation (with infinitely many real eigenvalues which correspond to the imaginary parts of the infinitely many non-trivial roots of $\zeta(t)$) so, it's important to study the behavior of the matrix representation of H as $N \rightarrow \infty$.

From Figure 5, it's not quite clear whether the difference approaches 0 or some finite value (that's very close to 0) as $N \rightarrow \infty$. We conjecture that it does approach 0. Here's a non-rigorous, hand-wavy reasoning. As seen in Figure 6, the eigenvalues are symmetric about the origin. So when we compute the successive differences, consider approaching it from left and right i.e. moving from the most positive eigenvalue in the direction of decreasing eigenvalues and at the same time, moving from the most negative eigenvalue towards decreasingly negative eigenvalues. Because of the inherent symmetry in the eigenvalues, we must approach the same limiting value. However, if this limit positive, that would mean that we'd have to hit 0 and move towards positive differences. But all the eigenvalues to the left of 0 are negative so the differences can never be positive. Same with while moving from right to left, the differences can never be negative. But, since we have to approach the same limiting value for either case, this limit must be 0.

But if the eigenvalue difference does hit 0 when H is infinite-dimensional, that would mean there is a degeneracy. However, because the eigenvalues of H would

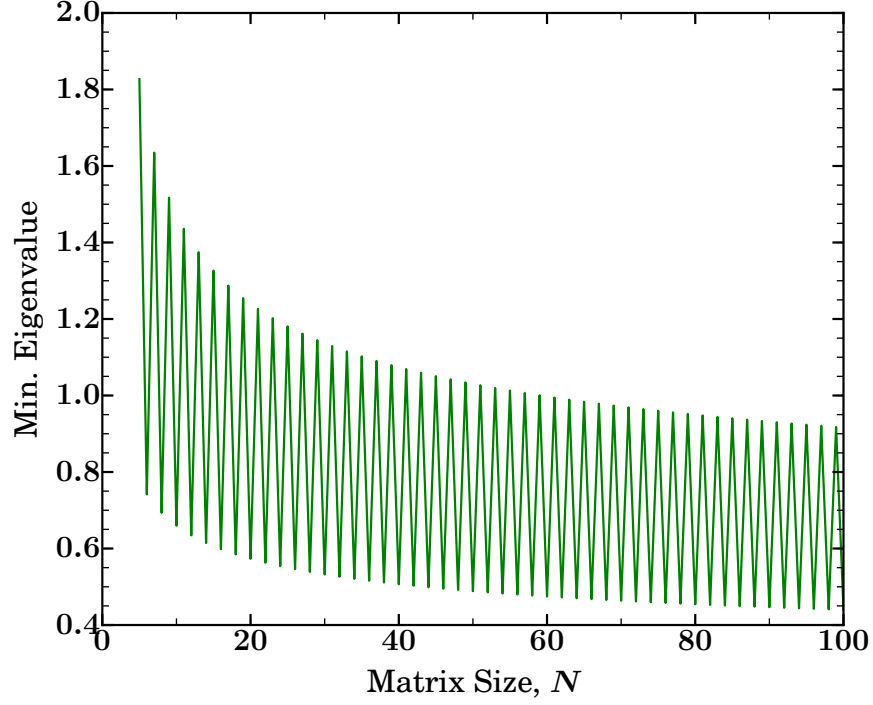


FIGURE 7. A plot of how the minimum eigenvalues λ_{+1} varies with the size of the matrix we choose to represent our Hamiltonian H . Note that for odd N 's, the minimum would be 0, so we choose the next largest minimum. In other words, we're plotting the minimum non-zero eigenvalues of H against the size of its matrix representation N .

correspond to the imaginary parts of the non-trivial roots of $\zeta(t)$ and so a degeneracy in the eigenvalues of H implies a degeneracy in the non-trivial roots of $\zeta(t)$. Since no such degeneracy is known, any quantization of H that would be a prospective Hilbert-Polya Hamiltonian would have to have something to break this degeneracy. Note that even though it seems like the difference in roots of $\zeta(t)$ in Figure 4 seems to be going all the way down to 0, it does not. For the plot produced, which contains the first 2 million zeros, it goes all the way down to ~ 0.0052 (in normalized units).

9. Conclusion

So this is the conclusion. Haha

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