Julia for adaptive high-order multi-physics simulations

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Acknowledgments

- Andrew Winters, Linköping University, Sweden (website)
 - ▶ especially coupled Euler-gravity simulations
- Hendrik Ranocha, KAUST, Saudia Arabia (website)
- Gregor Gassner, University of Cologne, Germany (website)

Back in January 2020...

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FLUXO: fast, parallel Fortran 3D-DGSEM code for curvilinear compressible
 Euler & MHD simulations (github.com/project-fluxo/fluxo)



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 Many one-trick ponies: codes with a singular purpose, often unusable for anything else, discarded after use







Images: ©MJ Boswell (top), ©Clive Williams (bottom)

What should we do? Write more code!

- Plan: use hackathon to write new simulation framework
 - should be useful to scientists and students
 - extensible and fast
- But we want even more:
 - easy to use for inexperienced users
 - hassle-free toolchain (= installation & postprocessing)
 - potential for HPC (maybe)
- Main question: Which language to use? Fortran? C++? Python? Julia?

What is Julia?

According to the official website Julia is ...

- fast
- dynamic
- reproducible
- composable
- general
- open source

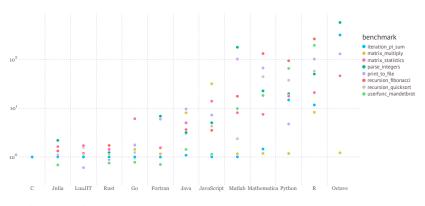


 $\verb|https://github.com/JuliaLang/julia-logo-graphics|$

Source: https://julialang.org

Julia is fast

- "Just-Ahead-Of-Time" (JAOT) compilation from Julia to machine code
- Facilitated by LLVM infrastructure
- Built-in support for parallelism (GPUs, shared memory, distributed)



https://julialang.org/benchmarks/

Julia is dynamic

- Dynamic type system
- Automatic memory management (garbage collection)
- Focus on interactive use

Julia's interactive read-eval-print-loop (REPL)

Julia is reproducible

- Create reproducible software environments across platforms
- Built-in package manager Pkg handles dependencies and versioning
- Automatic provisioning of pre-built binaries

Project.toml

```
name = "Trixi"
uuid = "a7f1ee26-1774-49b1-8366-f1abc58fbfcb"
authors = [...]
version = "0.3.9-pre"
[deps]
EllipsisNotation = "da5c29d0-fa7d-589e-88eb-ea29b0a81949"
LinearMaps = "7a12625a-238d-50fd-b39a-03d52299707e"
[compat]
EllipsisNotation = "1.0"
LinearMaps = "2.7, 3.0"
```

Which language should we use?

Battle of Languages

C++	Fortran	Julia
Pro 1. Auf allen emsthaften Maschinen installiert 2. Sehr schneil 3. Ausfrucksstark: komplexe Operationen lassen sich kompakt darstellen 4. Gute Standardbibliothek 5. Gute Kompatbillatt zu externen Libraries 6. Statlasch kompiliert -> Compiler hilft bei Fehlersuche	Pro 1. Auf allen ernsthaften Maschinen installiert 2. Sehr schneil 3. Elifache Syntax -> wenig Fehlerpotenzial 4. Alle in Gruppe sind "Experten" 5. Statisch kompiliert -> Compiler hilft bei Fehlersuche	Pro 1. "Sevy" für Studenten 2. Im Korn einfach zu seiremen (Malab-ähnlich) 3. Weniger Polierplate" Code (vs. Fortran) 4. Gut für schnelles Prototypring 5. Elinfach zu installieren (auch Laptop) 6. Große Paket-Bibliothek -> viel Funktionalität in Julia-only 7. Neuheitswert (ggr. auch wissenschafflich?) 8. Bei Erfog: ochtes Alleinstellungsmerkmal 9. Gute (nachgesagte) hybride Parallelisierung
Con 1. Kann kein Student 2. Mittelmäßig sexy für Studenten 3. Nur ein "Experfe" in der Gruppe 4. Vriele Wege, etwas falsch zu machen 5. Nicht memory-safe	Con 1. Kann kein Student 2. Super unsery für Studenten 3. Sehr viel Böllerplate 4. Mittelhallige Protterburkeit 5. Schlechtes Interfacing mit C Bibliotheken (Wrapper notwendig)	Con 1. Kann kein Student 2. Kein Experte in der Gruppe 3. Zwingt (I) bestimmte, ungewohnte Programmeleprantignen zu nutzen 4. Gehr) langsamer Startup von sogar teleinen Programmen, da immer enst kompiliert werden mit der

Which language should we use?

Battle of Languages C++ Fortran Julia Pro Pro Pro 1. Auf allen ernsthaften Maschinen installiert 1 Auf allen ernsthaften Maschinen installiert 1. "Sexy" für Studenten Im Kern einfach zu erlernen (Matlab-ähnlich) 3. Ausdrucksstark: komplexe Operationen lassen Einfache Syntax -> wenig Fehlerpotenzial Weniger "boilerplate" Code (vs. Fortran) sich kompakt darstellen Alle in Gruppe sind "Experten" Gut für schnelles Prototyping 4. Gute Standardbibliothek Statisch kompiliert -> Compiler hilft bei Einfach zu installieren (auch Laptop) 5. Gute Kompatibilität zu externen Libraries Fehlersuche Große Paket-Bibliothek -> viel Funktionalität in. 6. Statisch kompiliert -> Compiler hilft bei Julia-only Fehlersuche Neuheitswert (ggf. auch wissenschaftlich?) Bei Erfolg: echtes Alleinstellungsmerkmal Gute (nachgesagte) hybride Parallelisierupe too complicated maybe interesting too unsexy Con Mittelmäßig sexy für Studenten Super unsexy für Studenten Kein Experte in der Gruppe Nur ein "Experte" in der Gruppe Sehr viel Boilerplate Zwingt (!) bestimmte, ungewohnte Viele Wege, etwas falsch zu machen Mittelmäßige Portierbarkeit Programmierparadigmen zu nutzen Nicht memory-safe Schlechtes Interfacing mit C Bibliotheken (Sehr) langsamer Startup von sogar kleinen. (Wrapper notwendig) Programmen, da immer erst kompiliert werden 5. Programmierfehler erst zur Laufzeit sichtbar Unausgereifte Toolchain: z.B. Fehlermeldungen nicht so hilfreich Kleine Community, wenige Experten greifbar 8. Schlechtere Unterstützung auf ernsthaften Maschinen (wenig Erfahrung) 9. Viele Feinheiten, die nicht auf den ersten Blick offensichtlich sind 10. Mit Garbage Collection etc. sind a priori Performanceabschätzungen schwieriger

Outline of this talk

- 1. Should we use Julia?
- 2. Hyperbolic self-gravitating gas dynamics
- 3. Julia in practice: Trixi.jl
 - ▶ Live demonstration
- 4. Evaluating Julia for scientific computing
- 5. Conclusions and outlook

Hyperbolic self-gravitating gas dynamics

Goal: Approximate self-gravitating gas dynamics

Compressible Euler equations for hydrodynamics

$$\frac{\partial}{\partial t} \begin{bmatrix} \rho \\ \rho v_1 \\ \rho v_2 \\ E \end{bmatrix} + \frac{\partial}{\partial x} \begin{bmatrix} \rho v_1 \\ \rho v_1^2 + p \\ \rho v_1 v_2 \\ (E + p) v_1 \end{bmatrix} + \frac{\partial}{\partial y} \begin{bmatrix} \rho v_2 \\ \rho v_1 v_2 \\ \rho v_2^2 + p \\ (E + p) v_2 \end{bmatrix} = \begin{bmatrix} 0 \\ -\rho \phi_x \\ -\rho \phi_y \\ -(\vec{v} \cdot \vec{\nabla} \phi) \rho \end{bmatrix}$$

Newtonian potential equation for gravitation

$$-\vec{\nabla}^2\phi = -4\pi G\rho$$

- PDE for hydrodynamics is hyperbolic whereas gravity is elliptic
- Coupling of the two equations entirely through source terms

There is so much potential...

Let's manipulate the Poisson equation in 2D

$$-\nu \vec{\nabla}^2 u = f$$

Potential u is the steady state solution of a parabolic equation

$$\mathbf{u}_t - \nu \vec{\nabla}^2 \mathbf{u} = \mathbf{f}$$

• Introduce variables $\vec{\nabla} u = (q_1 \,,\, q_2)^T$ to have a parabolic system

$$u_t - \nu[q_1]_x - \nu[q_2]_y = f$$

$$u_x = q_1$$

$$u_y = q_2$$

There is so much potential...

- Diffusion equation has paradox of instant propogation
- Idea of Cattaneo, introduce (small) time scale T_r

$$u_t - \nu[q_1]_x - \nu[q_2]_y = f$$

$$T_r[q_1]_t - u_x = -q_1$$

$$T_r[q_2]_t - u_y = -q_2$$

to correct unphysical behavior

Yields the hyperbolic diffusion equations

$$\frac{\partial}{\partial t} \begin{bmatrix} u \\ q_1 \\ q_2 \end{bmatrix} + \frac{\partial}{\partial x} \begin{bmatrix} -\nu q_1 \\ -u/T_r \\ 0 \end{bmatrix} + \frac{\partial}{\partial y} \begin{bmatrix} -\nu q_2 \\ 0 \\ -u/T_r \end{bmatrix} = \begin{bmatrix} f(x,y) \\ -q_1/T_r \\ -q_2/T_r \end{bmatrix}$$

Numerical solver for hyperbolic conservation laws

 Select discontinuous Galerkin (DG) to approximate solution of general conservation law system

$$\mathbf{u}_t + \mathbf{f}_{\mathsf{x}}(\mathbf{u}) = \mathbf{s}(\mathbf{u})$$

- Divide spatial domain into elements, map to reference element
- Multiply by test function and integrate-by-parts (IBP) to obtain weak form

 → optionally use IBP again to get strong form
- Approximate solution, fluxes, sources with nodal polynomials
- Resolve discontinuities at element boundaries by numerical flux
- Approximate integrals with Gauss-type quadrature
- Interpolation and quadrature nodes are collocated

Numerical approximation: Discontinuous Galerkin

 Gives an ODE to integrate in time and update the approximate solution in each element

$$\boxed{\frac{\mathrm{d}\mathbf{U}_{j}}{\mathrm{d}t} = -\frac{2}{\Delta x_{i}} \left\{ \frac{\delta_{jN}}{\omega_{N}} \left[\mathbf{F}^{*} - \mathbf{F}_{N} \right] - \frac{\delta_{j0}}{\omega_{0}} \left[\mathbf{F}^{*} - \mathbf{F}_{0} \right] + \sum_{m=0}^{N} \mathcal{D}_{jm} \mathbf{F}_{m} \right\} + \mathbf{S}_{j}}$$

- Here $j=0,\ldots,N$ and $\mathcal{D}_{jm}=\ell'_m(\xi_j)$ is the polynomial derivative matrix
- Use explicit time integration via Runge-Kutta (RK) methods
- Stable DG time step has the form

$$\Delta t = rac{ extsf{CFL}}{ extsf{N}+1} rac{\Delta x}{|\lambda_{ extsf{max}}|}$$

with adjustable CFL constant

Verify high-order DG method for hyperbolic diffusion

- Integrate in time with "standard" five-stage, four-order low-storage RK method of Carpenter & Kennedy
- Take CFL = 0.5 such that spatial errors dominate
- ullet Threshold to define steady state taken as $tol = 10^{-10}$

$$u(x,y) = 2 + 2\cos(\pi x)\sin(2\pi y)$$
 $f(x,y) = 10\pi^2\cos(\pi x)\sin(2\pi y)$

Boundary conditions: Dirichlet in x-direction, periodic in y-direction

Convergence of hyperbolic diffusion

K	$L^2(u)$	$L^2(q_1)$	$L^2(q_2)$
4^{2}	3.15E-03	1.24E-02	2.19E-02
8^{2}	2.26E-04	8.83E-04	1.50E-03
16^{2}	1.50E-05	5.51E-05	9.68E-05
32^{2}	9.65E-07	3.32E-06	6.14E-06
avg. EOC	3.89	3.96	3.93

K	$L^2(u)$	$L^{2}(q_{1})$	$L^{2}(q_{2})$
4^{2}	2.51E-04	8.81E-04	1.63E-03
8^{2}	8.52E-06	2.88E-05	5.45E-05
16^{2}	2.77E-07	9.12E-07	1.76E-06
32^{2}	8.85E-09	2.85E-08	5.60E-08
avg. EOC	4.93	4.97	4.94

- Discrete L² errors computed on uniform Cartesian meshes of increasing resolution
- Demonstrate high-order accuracy of two polynomial orders for potential u and its gradient

Revisit equations of self-gravitating gas dynamics

Compressible Euler equations

$$\frac{\partial}{\partial t} \begin{bmatrix} \rho \\ \rho v_1 \\ \rho v_2 \\ E \end{bmatrix} + \frac{\partial}{\partial x} \begin{bmatrix} \rho v_1 \\ \rho v_1^2 + \rho \\ \rho v_1 v_2 \\ (E+\rho) v_1 \end{bmatrix} + \frac{\partial}{\partial y} \begin{bmatrix} \rho v_2 \\ \rho v_1 v_2 \\ \rho v_2^2 + \rho \\ (E+\rho) v_2 \end{bmatrix} = \begin{bmatrix} 0 \\ -\rho \mathbf{q}_1 \\ -\rho \mathbf{q}_2 \\ -(v_1 \mathbf{q}_1 + v_2 \mathbf{q}_2) \rho \end{bmatrix}$$

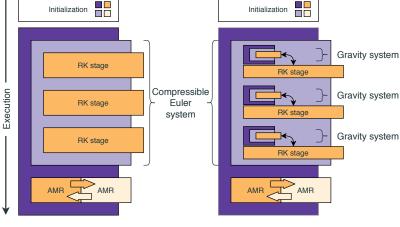
Recast gravitational potential equation

$$\frac{\partial}{\partial t} \begin{bmatrix} \phi \\ q_1 \\ q_2 \end{bmatrix} + \frac{\partial}{\partial x} \begin{bmatrix} -q_1 \\ -\phi/T_r \\ 0 \end{bmatrix} + \frac{\partial}{\partial y} \begin{bmatrix} -q_2 \\ 0 \\ -\phi/T_r \end{bmatrix} = \begin{bmatrix} -4\pi G \rho \\ -q_1/T_r \\ -q_2/T_r \end{bmatrix}$$

PDEs for dynamics are both hyperbolic and coupled via source terms

Schematic of volumetric coupling

Single-physics simulation
 Multi-physics simulation



Creating a multi-physics simulation

- Instantiate two DG solvers: One for hydrodynamics, one for gravity
- Independent and only "talk" through the source terms
- For simplicity assume the two solvers share a mesh
- Compressible Euler solver can have other features, shock capturing or AMR
- Use different time integrators:
 - Compressible Euler uses five-stage, fourth-order low-storage RK scheme of Carpenter & Kennedy
 - Gravity uses five-stage, second-order low-storage RK scheme optimized to allow large explicit time steps
 - ullet Two adjustable coefficients for time step selection: CFL $_{
 m Eu}$ \in (0, 1], CFL $_{
 m Gr}$

Verify high-order DG method for coupled problem

- ullet Take $\mathtt{CFL}_{\mathrm{Eu}} = \mathtt{CFL}_{\mathrm{Gr}} = 0.5$ such that spatial errors dominate
- Threshold to define steady state gravity is $tol = 10^{-10}$
- Domain $\Omega = [0,2]^2$ with manufactured solution

$$\rho = 2 + \frac{1}{10}\sin(\pi(x+y-t))$$
 $v_1 = v_2 = 1$
 $p = \frac{1}{\pi}\rho^2$
 $\phi = -\frac{2}{\pi}(\rho - 2)$

Gravitational constant G = 1

- Boundary conditions: Periodic in all directions
- Introduces additional residual terms added to the right-hand-side

Convergence of coupled self-gravity problem

■ *N* = 3

K	$L^2(\rho)$	$L^2(\rho v_1)$	$L^2(\rho v_2)$	$L^2(E)$	$L^2(\phi)$	$L^{2}(q_{1})$	$L^{2}(q_{2})$
4 ²	4.37E-04	4.69E-04	4.69E-04	9.72E-04	1.64E-04	8.33E-04	8.33E-04
8^2	2.43E-05	2.60E-05	2.60E-05	5.09E-05	9.90E-06	5.65E-05	5.65E-05
16^{2}	1.06E-06	1.37E-06	1.37E-06	2.65E-06	6.63E-07	3.77E-06	3.77E-06
32^{2}	4.73E-08	8.03E-08	8.03E-08	1.56E-07	4.33E-08	2.44E-07	2.44E-07
avg. EOC	4.39	4.17	4.17	4.20	3.96	3.91	3.91

■ N = 4

K	$L^2(\rho)$	$L^2(\rho v_1)$	$L^2(\rho v_2)$	$L^2(E)$	$L^2(\phi)$	$L^2(q_1)$	$L^2(q_2)$
4^{2}	3.50E-05	3.38E-05	3.38E-05	6.59E-05	1.15E-05	6.31E-05	6.31E-05
8^{2}	7.99E-07	9.00E-07	9.00E-07	1.71E-06	3.74E-07	2.11E-06	2.11E-06
16^{2}	1.95E-08	2.49E-08	2.49E-08	4.78E-08	1.23E-08	6.95E-08	6.95E-08
32^{2}	5.31E-10	7.73E-10	7.73E-10	1.44E-09	4.03E-10	2.25E-09	2.25E-09
avg. EOC	5.34	5.14	5.14	5.16	4.93	4.93	4.93

Retain high-order accuracy for all variables in the coupled system

A more physical self-gravitating setup

- Jeans instability models perturbations and interactions between a gas cloud and gravity
- Consider a background state in centimeter-gram-second (CGS) units

$$ho_0 = 1.5 \cdot 10^7 \; [\mathrm{g \; cm}^{-3}] \qquad p_0 = 1.5 \cdot 10^7 \; [\mathrm{dyn \; cm}^{-2}]$$

• Perturb a particular mode \vec{k} with small amplitude δ_0

$$ho =
ho_0 \left(1 + \delta_0 \cos(\vec{k} \cdot \vec{x}) \right) \qquad \vec{v} = \vec{0} \qquad p = p_0 \left(1 + \delta_0 \gamma \cos(\vec{k} \cdot \vec{x}) \right)$$

Gravitational field responds accordingly

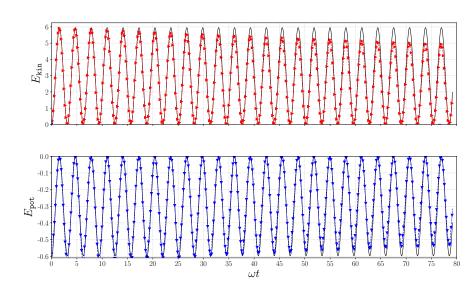
$$-\vec{\nabla}^2 \phi = -4\pi G(\rho - \rho_0)$$
 $G = 6.674 \cdot 10^{-8} \text{ [cm}^3 \text{ g}^{-1} \text{ s}^{-2}]$

A more physical self-gravitating setup (continued)

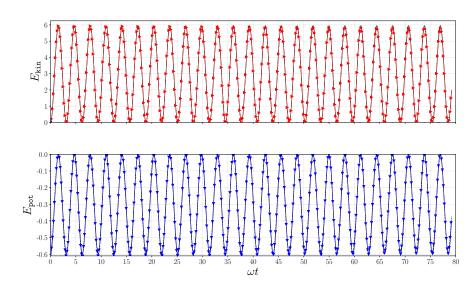
- Domain $\Omega = [0,1]^2$ with periodic boundary conditions
- \bullet Take $\mathtt{CFL}_{\mathrm{Eu}} = 0.5$ and $\mathtt{CFL}_{\mathrm{Gr}} = 1.2$
- Uniform 16×16 Cartesian mesh with polynomial order N=3
- Threshold to define steady state taken as $tol = 10^{-4}$
- Offers a convenient bridge to examine numerics: More physically relevant but still has analytical expressions for energy evolution

$$E_{ ext{kin}} = \int\limits_{\Omega} rac{
ho}{2} (v_1^2 + v_2^2) \, \mathrm{d}\Omega \qquad E_{ ext{int}} = \int\limits_{\Omega} rac{p}{\gamma - 1} \, \mathrm{d}\Omega \qquad E_{ ext{pot}} = \int\limits_{\Omega}
ho \phi \, \mathrm{d}\Omega$$

Energy evolution of Jeans instability (couple every RK step)

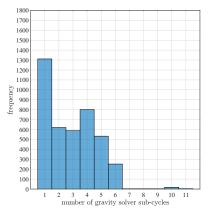


Energy evolution of Jeans instability (couple every RK stage)



Visualize cost of gravity solver for Jeans instability

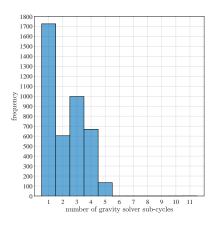
 Gravity sub-cycle: Assembly and evolution of hyperbolic gravity system by one complete time step



- Explicit time integration with low storage RK
- Compressible Euler solver uses five-stage, fourth order scheme of Carpenter & Kennedy ${\tt CFL}_{\rm Eu} = 0.5$
- Hyperbolic gravity uses the same time integration scheme as compressible Euler ${\tt CFL_{Gr}}=0.8$

Visualize cost of gravity solver for Jeans instability

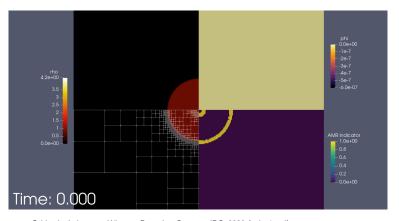
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- Explicit time integration with low storage RK
- $\begin{tabular}{ll} \hline & \textbf{Compressible Euler} & \text{solver uses} \\ \hline & \text{five-stage, fourth order scheme of} \\ \hline & \text{Carpenter \& Kennedy} \\ \hline & \text{CFL}_{\rm Eu} = 0.5 \\ \hline \end{tabular}$
- Hyperbolic gravity uses five-stage, second order RK scheme optimized to take larger times steps
 CFL_{Gr} = 1.2

Bring everything together

- Sedov explosion problem with self-gravity
- Localize explosion to occur within a dense disc of radius one
- Contains strong shocks and necessitates AMR

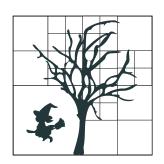


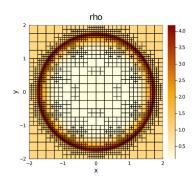
 $Schlottke-Lakemper,\ Winters,\ Ranocha,\ Gassner,\ JPC,\ 2020\ (submitted).\ ar \texttt{Xiv}: 2008.10593$

Julia in practice: Trixi.jl

Trixi.jl: A tree-based numerical simulation framework for (hyperbolic) PDEs

- Adaptive hierarchical quadtree/octree grids
- Nodal discontinuous Galerkin spectral element methods
- Explicit time integration with SciML's OrdinaryDiffEq. jl
- Multiple governing equations
- Available at github.com/trixi-framework/Trixi.jl (open source)

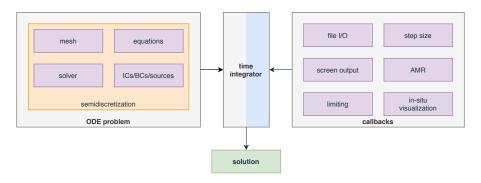




Launch your Binders!

| launch binder

Overall structure of Trixi

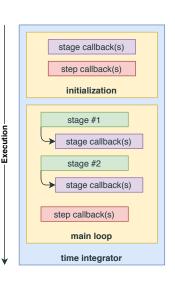


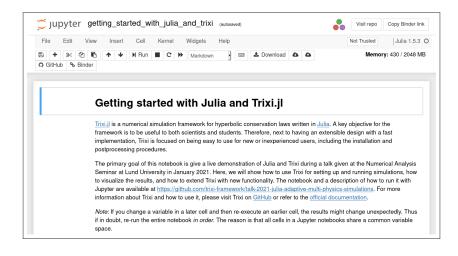
Guiding principles for the framework:

- Modular architecture ("Trixi as a library")
- Only read code that you actually use
- Implement first, ask questions later

Extended functionality in callbacks

- Callbacks provide additional functionality
 - file I/O
 - step size
 - adaptive mesh refinement
 - limiting
 - in-situ visualization
 -
- Flexible: combine callbacks for applications
- Extensible: easily add new features





Evaluating Julia for scientific computing

Evaluating Julia for scientific computing

- What about performance?
- What about ease of use?
- What about reproducibility?
- What about composability?
- What else is there?

What about *performance*?

- Serial performance is good
 - FLUXO vs. FLUXO-in-Julia-v0.6: within factor of $2-3\times$
 - FLUXO vs. Trixi: Trixi can be faster (caveat: no metric terms)
- Requires new performance intuition (when coming from C/C++/Fortran)
 - use many small functions (function barriers)
 - ullet type instabilities o Python-like performance
 - benchmarking new code is a must
- Parallel performance for simulation science? The verdict is still out...
 - no true MPI HPC codes out there (yet)
 - possibly first large-scale project: https://github.com/CliMA/ClimateMachine.jl
 - petascale showcase (Celeste.jl) is not a simulation

Just-ahead-of-time compilation has its quirks – and perks!

- Very long startup times: only usable with REPL (caching)
 - Time to first result in Trixi: ~20 seconds
 - Time to second result: 60 *milli*seconds
- Compilation is serial

Just-ahead-of-time compilation has its quirks — and perks!

- Very long startup times: only usable with REPL (caching)
 - Time to first result in Trixi: ~20 seconds
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- Compilation is serial
- Kernels can be written as in C/C++/Fortran (loops are fast!)
- Extremely high degree of function specialization (compare to fully templated C++)

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ClangJIT: Enhancing C++ with Just-in-Time Compilation

Hal Finkel
Lead, Compiler Technology and
Programming Languages
Leadership Computing Facility
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David Poliakoff Lawrence Livermore National Laboratory Livermore, CA, USA poliakoff1@llnl.gov David F. Richards Lawrence Livermore National Laboratory Livermore, CA, USA richards12@llnl.gov

ABSTRACT

The C++ programming language is not only a keystone of the

body of C++ code, but critically, defer the generation and optimization of template specializations until runtime using a relativelynatural extension to the core C++ programming language.

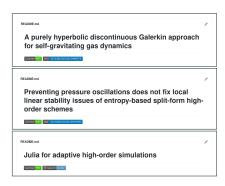
What about ease of use?

- Very easy to get up and running (see reproducibility)
- Code can be very simple (MATLAB-like) if desired
 - but be careful about unwanted allocations (type instabilities)
- Lack of a mature toolchain
 - Plotting is only OK
 - No widespread IDE support (Julia plugin for VS Code)
 - Only one debugger
- No surprise: truly fast code looks virtually the same everywhere

What about reproducibility?

- Provisioning reproducible compute environments is straightforward
- Two files provide all relevant information (Project.toml, Manifest.toml)
- Recreation with only a few lines of code
- Excellent for reproducible science:
 paper 1, paper 2, this talk's repo

```
import Pkg
Pkg.activate()
Pkg.instantiate()
using Trixi # ... and enjoy!
```



What about composability?

3

- Multiple dispatch: "function overloading on runtime types"
- No difference between standard library code, package code, own code
- lacktriangledown Only implement what you need ightarrow rapid prototyping
- Increased code reuse invites collaboration

What else is there?

- Julia community is focused on data science, not simulation science
- Different notion of "HPC" than computional science community
 - $\, \bullet \, \to \, {\sf think \ big \ data}, \ {\sf not \ necessarily \ exascale \ computing} \,$



Conclusions and outlook: hyperbolic gravity

- Compute solution to elliptic problem via hyperbolic framework
 - Verified experimental order of convergence
 - High-order gradient computations
- Multi-physics coupling for flow-gravity simulations works
 - Reuse hyperbolic schemes without modifications
 - Supports adaptive mesh refinement
- Next up: speed up gravity solver, add more physics

Conclusions and outlook: Julia for scientific computing

- Performance is good, but predictability not so much
 - "no free lunch" for C/C++/Fortran-like performance
 - Requires new performance intuition
 - No verdict on large-scale MPI parallelization yet
- Ad-hoc compilation and multiple dispatch can be great assets
 - Facilitates rapid prototyping
 - More code sharing and code reuse
- Ease of use
 - Minimal setup time for new users
 - Invites collaboration
 - Great for reproducibility of scientific findings
- Next up: fully parallelize Trixi and scale to 10,000+ cores

Thank you for your interest!

Are there any questions?