Chapter 1

Governing Equations

The governing equations to be solved for the fluid system will now be delineated. A general discussion of conservation laws is presented first and then the specific thermohydraulic equations. The section will conclude with the constitutive relations used to close the system of thermohydraulic equations.

1.1 Conservation Laws

A conserved quantity q is taken to be any physical property whose time evolution within an arbitrary, closed volume exactly balances with its surface fluxes and volume sources. For the purpose of mathematical discussion and analysis, the conserved quantity q will simultaneously represent a function of space and time that conforms to the requirements of the physical property. These definitions define the following scalar conservation law over a volume Ω with

an enclosing surface Γ :

$$\frac{\mathrm{d}}{\mathrm{d}t} \int_{\Omega} q(x_i, t) \,\mathrm{d}\Omega = \int_{\Gamma} -F_i(q, x_i, t) n_i \,\mathrm{d}\Gamma + \int_{\Omega} S(q, x_i, t) \,\mathrm{d}\Omega, \tag{1.1}$$

where F_i is a vector of surface flux functions of q, n_i is the outward unit normal of the surface Γ , and S is the volume source of q. Such that the units in the equation agree, both q and S are taken on a per volume basis and F_i on a per area basis. The negative sign in the surface integral ensures that outward fluxes act as sinks and inward fluxes as sources to the time evolution of q's volume integral. Equation 1.1 is the most general scalar conservation law that will be presented and is always physically valid regardless of the functions' behaviors.

1.1.1 Differential Form

A general conservation law can also be presented in differential form (i.e., as a differential equation). First the Divergence Theorem is used to equate the surface integral in equation 1.1 to a volume integral:

$$\int_{\Gamma} -F_i(q, x_i, t) n_i \, d\Gamma = \int_{\Omega} -\partial_i F_i(q, x_i, t) \, d\Omega.$$
 (1.2)

For simplicity, the volume Ω is now taken to be fixed in space for all time, although a time-varying derivation with the identical result is possible via the Reynolds Transport Theorem. Substituting equation 1.2 into equation 1.1 and moving all terms to the left-hand side gives:

$$\int_{\Omega} \partial_t q(x_i, t) + \partial_i F_i(q, x_i, t) - S(q, x_i, t) d\Omega = 0,$$
(1.3)

where the time derivative operator could be moved into the integral since the volume is time-independent. Since the integration volume in equation 1.3 is arbitrary, taking the limit of the equality as the volume shrinks to a zero gives

$$\lim_{\Omega \to 0} \left[\int_{\Omega} \partial_t q(x_i, t) + \partial_i F_i(q, x_i, t) - S(q, x_i, t) \, d\Omega \right] = 0.$$
 (1.4)

In this limit, the enforcement of the equality changes from one over a finite volume into one that is enforced at a particular point. To ensure point-wise enforcement over the domain of x_i , the integrand itself is required to be equally zero at every point in the domain; this yields the differential conservation law

$$\partial_t q(x_i, t) + \partial_i F_i(q, x_i, t) - S(q, x_i, t) = 0. \tag{1.5}$$

When equipped with adequate boundary and initial data, this differential equation defines the requirement for all sufficiently smooth functions that describe how the conserved quantity evolves at every point in space-time.

The differential form may seem cleaner than the integral form, but it is not valid for all functional forms of q. In transitioning from equation 1.4 to equation 1.5, it is assumed that all functions in the integrand remain bounded. If any of the terms within the limit become infinite, the balance cannot be satisfied. In particular, if q has an area where it undergoes a discontinuous jump, called a shock, the gradient of the flux is infinite and the differential form is rendered invalid. While methods that handle shocks in a robust fashion (shock-capturing schemes) are not the focus of this work, it is mentioned for completeness and consideration.

1.1.2 Vector Form

The simulation of real world problems often involves the solution of a system of conservation laws. In general, the systems are nonlinear and tightly coupled. The actual solution of these problems will be discussed in future sections, but the notation used will be introduced now.

The integral conservation law for a vector of conserved quantities q_i is

$$\partial_t \int_{\Omega} q_i(x_i, t) d\Omega = \int_{\Gamma} -F_{ij}(q_i, x_i, t) n_j d\Gamma + \int_{\Omega} S_i(q_i, x_i, t) d\Omega, \qquad (1.6)$$

where F_{ij} is a matrix of surface fluxes of q_i in the j-th direction, n_j is the outward unit normal of the surface Γ , and S_i is the volume source of q_i . The integrals of the vector quantities represent element-wise integration. A process similar to the scalar case also yields the differential form of the system conservation law:

$$\partial_t q_i(x_i, t) + \partial_j F_{ij}(q_i, x_i, t) - S_i(q_i, x_i, t) = 0, \tag{1.7}$$

where $\partial_j F_{ij}$ represents a row-wise divergence operation.

1.2 Thermohydraulic Laws

The main equations of interest in thermohydraulics will be presented using the definitions above. Mass, momentum, and energy are the primary quantities of interest and will be derived for an assumed-flowing system. For all of the definitions, the following items are noted:

- As prescribed in section 1.1, all of the conserved quantities are taken to be continuous functions of space and time. Although fluids are actually composed of discrete, interacting molecules, taking the quantities to be everywhere-defined is a valid approximation as long as the length scales to be modeled are much greater than the mean free path of the local medium. NEED REFERENCE
- Also as prescribed in section 1.1, all of the conserved quantities will be considered on a per unit volume basis.

• For clarity, the dependence of all quantities and functions will be omitted; however, they are all assumed to be dependent of any variable in the system.

1.2.1 Conservation of Mass

The net mass in a control volume or, rather, the volume-integrated density is conserved as a simple scalar. It is controlled by the background flow field advecting the mass through the boundary. Therefore, the integral conservation law for density is

$$\frac{\mathrm{d}}{\mathrm{d}t} \int_{\Omega} \rho \,\mathrm{d}\Omega = \int_{\Gamma} -(u_j \rho) n_j \,\mathrm{d}\Gamma + \int_{\Omega} S^{\rho} \,\mathrm{d}\Omega,\tag{1.8}$$

1.2.2 Conservation of Momentum

A control volume's momentum balance is more complicated than the simple in-out mass balance. It is inherently a vector quantity, so the combination of mass and the background flow field feed the momentum in a non-linear fashion. Also, even though a continuum approximation has been made, physical molecular stresses on the boundary must still be modeled in some manner for physical solutions in all circumstance. Lastly, a body force is be added to account for buoyancy effects on the volume relative to the control volume's surroundings. With all of the above nuances considered, the integral conservation of momentum is taken to be

$$\frac{\mathrm{d}}{\mathrm{d}t} \int_{\Omega} \rho u_i \,\mathrm{d}\Omega = \int_{\Gamma} (-u_j \rho u_i) n_j \,\mathrm{d}\Gamma + \int_{\Gamma} (-\delta_{ij} P + \tau_{ij}) n_j \,\mathrm{d}\Gamma + \int_{\Omega} \rho g_i + S_i^u \,\mathrm{d}\Omega \tag{1.9}$$

where ρu_i is the momentum in direction i, u_j is the velocity in direction j, δ_{ij} is the Kronecker delta, P is the thermodynamic pressure, τ_{ij} is the viscous stress tensor, and g_i is the gravitational acceleration in direction i. A more detailed discussion of the pressure and viscous stress tensor is presented in section 1.3.

Each of the terms on the right-hand side of equation 1.9 stems from the considerations made. The first integral represents how the background flow field transports the momentum across the volume's surface. The second integral accounts for the ever-present thermodynamic pressure of the fluid and all other non-equilibrium molecular friction via the tensor τ_{ij} . The last integral has a source of momentum due to the density of the volume and a generic source S_i^u .

1.2.3 Conservation of Energy

The energy equation is similar to the momentum equation since there are non-linear and molecular components to the balance. The energy to be balanced is the *total energy* of the medium, where total means the sum of thermal, kinetic, and potential energies. The integral conservation for total energy is taken to be

$$\partial_t \int_{\Omega} \rho e \, d\Omega = \int_{\Gamma} \left[-(\rho e + P)u_j + u_i \tau_{ij} - q_j \right] n_j \, d\Gamma + \int_{\Omega} \rho g_j u_j + S^e \, d\Omega \tag{1.10}$$

where ρe is the total energy of the system and q_j is a conductive heat flux; all of the other terms have the same meaning as in the momentum equation. A more detailed discussion of the conductive heat flux is presented in section 1.3.

As in the momentum equation, all right-hand side terms have a purpose. The first term, $\rho e + P$, represents the advected energy in addition to the associated work due to pressure. The second term accounts for energy loss due to molecular friction. The third term arises due to molecular diffusion of heat across the boundary. The final term acts as a source or sink of energy due to the buoyancy force from momentum.

1.2.4 Differential Forms

All of the integral conservation laws shown above also have differential forms:

$$\partial_t \rho + \partial_i (u_i \rho) = S^{\rho} \tag{1.11}$$

$$\partial_t(\rho u_i) + \partial_j(u_j \rho u_i) = -\partial_i P + \partial_j \tau_{ij} + \rho g_i + S_i^u$$
(1.12)

$$\partial_t(\rho e) + \partial_i(u_i \rho e) = -\partial_i(u_i P) + \partial_i(u_i \tau_{ij}) + \rho g_i u_i + S^e$$
(1.13)

Given the same initial and boundary conditions, these differential equations and their integral counterparts are equivalent for sufficiently smooth solutions. As required by the equations listed, "sufficiently smooth" means that at least all first derivatives must be continuous over the entire solution domain; however, certain constitutive relations (e.g., the stress tensor) may place more stringent requirements on continuity.

1.2.5 Conservation of Bulk Momentum

All of the above equations are three-dimensional, but in thermohydraulic modeling, it is common practice to collapse the momentum equation to a single, dominant flow direction. The result of this dimensional collapse results in an equations that conserves bulk momentum. The bulk flow equation is created by dotting the integral momentum equation with a constant unit vector z_i . The unit vector represents the physically dominant flow direction for the given volume. Performing the dot operation and letting $a_z = a_i z_i$ for any vector a_i , the bulk flow equation is

$$\partial_t \int_{\Omega} \rho u_z \, d\Omega = \int_{\Gamma} (-u_j \rho u_z) n_j \, d\Gamma + \int_{\Gamma} -P n_z + \tau_{ij} z_i n_j \, d\Gamma + \int_{\Omega} \rho g_z + S_z^u \, d\Omega.$$
 (1.14)

A further implication called *channel flow conservation* is also employed by a number of

modeling tools. The channel flow equations are derived from bulk momentum conservation by placing the following restrictions on the integration volume:

- there are exactly two open surfaces that parallel to one another and perpendicular to the bulk flow direction;
- the outward normal the surface connecting the two openings and enclosing the volume is always perpendicular to the bulk flow direction.

By imposing these requirements, equation 1.14 takes on the simpler form

$$\partial_t \int_{\Omega} \rho u_z \, d\Omega = \int_{\Gamma_1 + \Gamma_2} \pm (P + u_z \rho u_z) \, d\Gamma + \int_{\Gamma_w} \tau_{ij} z_i n_j \, d\Gamma + \int_{\Omega} \rho g \, \operatorname{Cos}(\theta) + S_z^u \, d\Omega, \quad (1.15)$$

where Γ_1 and Γ_2 are open surfaces one and two, Γ_w is the third (wetted) surface, and θ is the angle between the bulk flow direction and gravity with counter-clockwise sense. The plus-minus accounts for the sign flip which occurs between surfaces one and two, which technically embody inflow and outflow once the sense of the bulk flow direction is chosen.

1.3 Constitutive Relations

In deriving these conservation laws, a number of unknowns were introduced to account for different thermodynamic, molecular, and non-equilibrium phenomena. These unknowns have to be related to the system variables in some manner to make the system solvable. Equations that perform this task as called *constitutive relations* or, less formally, models. The terms of interest are the thermodynamic pressure, viscous stress tensor, and conductive flux.

1.3.1 Pressure

Pressure is a macroscopic, thermodynamic property that arises from ensembles of particles exerting forces over some unit area. A constitutive relation that relates pressure to other system variables that define the system's state is called an *equation-of-state* (EOS). An EOS can be constructed from first principle (microscopic) approaches or thermodynamic potential (macroscopic) approaches and may be composed of one or more equations. Regardless of the EOS's construction, it should be able to calculate all relevant thermodynamic properties given two independent properties for a pure fluid NEED REFERENCE.

Since there are multiple avenues for creating an EOS, for any given fluid, several may exist; each one varying in complexity of formulation, range of validity, and fidelity to data. The choice of EOS is dependent on the importance of each of those three attributes. The EOS used for this work is the International Association for the Properties of Water and Steam (IAPWS) Formulation 1995 for the Thermodynamic Properties of Ordinary Water Substance for General and Scientific Use (IAPWS-95). The IAPWS-95 formulation is extremely complex, but the range of validity and fidelity to data is second-to-none for water. The details of the formulation is left to the appendix and its original publication.

1.3.2 Stress Tensor

The stress tensor requires a model that relates molecular friction, continuum deformation, and system variables. There are, like an EOS for a fluid, several different stress tensor models for any given system. These models vary in terms of the type of valid fluid, dimensionality of usage, and fidelity to data.

Newtonian fluid models are among the most popular stress tensor model choices and form the

basis for the Navier-Stokes equations. From experiment, Newton observed that a proportional relationship exists between the shear stress exerted on a fluid and the fluid's uni-directional velocity gradient for a wide array of fluids. Stokes then extended this proportional relationship to higher dimensions by making three assumptions: the tensor is a linear function of strain, the fluid is isotropic, and the tensor's divergence must be zero for a fluid at rest (such that pressure is the only extant surface force). With all of these observations and assumptions considered, the Newton stress tensor model (sometimes described as a compressible model) is defined by the equation:

$$\tau_{ij} = \mu \left(\frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) + \lambda \delta_{ij} \frac{\partial u_k}{\partial x_k}$$
(1.16)

where μ is the dynamic viscosity and λ is the bulk viscosity.

While the above relation is suitable to equation 1.9, the bulk and channel momentum conservation equations require a different relation since the multi-dimensional nature of the flow has been dotted away. One such relation is the Fanning friction factor which poses the traction from the walls as a loss of kinetic energy:

$$\tau_{ij} z_i n_j = \Delta P_{\text{fan}} = 2 f_{\text{F}} \frac{L_{\text{char}}}{D_{\text{eff}}} \rho v^2, \qquad (1.17)$$

where $f_{\rm F}$ is the Fanning friction factor, $L_{\rm char}$ is a characteristic (path) length, and $D_{\rm eff}$ is an effective flow diameter. The details of what densities and velocities to use is left to the numerical discretization portion that discusses the averaging process. The Fanning friction factor is well approximated by the Colebrook-White equation

$$\frac{1}{\sqrt{f_{\rm F}}} = -4 \operatorname{Log} \left(\frac{\varepsilon}{3.71 \, D_{\rm eff}} + \frac{1.26}{\operatorname{Re}\sqrt{f_{\rm F}}} \right),\tag{1.18}$$

where ε is the surface roughness and Re is the Reynolds number.

- 1.3.3 Conductive Flux
- 1.4 Numerical Methods
- 1.5 Stability Theory