### Chapter 1

# Theory

The theory encompassed by this work can be divided into four basic parts: conservation laws, thermohydraulics, numerical methods, and stability. Conservation laws, such as those for momentum and energy, form a fundamental basis of analysis in all branches of science. Therefore, a precise definition for a generic conserved quantity and an associated conservation law will be given. Following this generic treatment, the specific conservation laws for thermohydraulics will be presented. Numerical methods for solving the thermohydraulic conservation laws will then be given. Finally, the stability theory for the thermohydraulic system will be discussed.

### 1.1 Conservation Laws

#### 1.1.1 Scalar Form

A conserved quantity q refers to any physical property whose time evolution within an arbitrary, closed volume exactly balances with its surface fluxes and volume sources. For the purpose of mathematical discussion and analysis, the conserved quantity q will simultaneously represent a function of space and time that conforms to the requirements of the physical property. These definitions define the following scalar conservation law over a volume  $\Omega$  with a closed surface  $\Gamma$ :

$$\partial_t \int_{\Omega} q(x_i, t) \, \partial\Omega = \int_{\Gamma} -F_i(q, x_i, t) n_i \, \partial\Gamma + \int_{\Omega} S(q, x_i, t) \, \partial\Omega, \tag{1.1}$$

where  $F_i$  is the surface flux of q,  $n_i$  is the outward unit normal of the surface  $\Gamma$ , and S is the volume source of q. Such that the units in the equation agree, both q and S are taken on a per volume basis and  $F_i$  on a per area basis. The negative sign in the surface integral ensures that outward fluxes act as sinks and inward fluxes as sources to the time evolution of q's volume integral. Equation 1.1 is the most general scalar conservation law that will be presented and is always physically valid regardless of the functions' behaviors.

A general conservation law can also be presented in differential form (i.e., as a differential equation). First the Divergence Theorem is used to equate the surface integral in equation 1.1 to a volume integral:

$$\int_{\Gamma} -F_i(q, x_i, t) n_i \, \partial \Gamma = \int_{\Omega} -\partial_i F_i(q, x_i, t) \, \partial \Omega. \tag{1.2}$$

Substituting equation 1.2 into equation 1.1 and moving all terms to the left-hand side gives:

$$\int_{\Omega} \partial_t q(x_i, t) + \partial_i F_i(q, x_i, t) - S(q, x_i, t) \, \partial\Omega = 0, \tag{1.3}$$

where the time derivative operator could be moved into the integral since the volume is taken to be time-independent. Since the integration volume in equation 1.3 is arbitrary, we take limit of the equality as the volume shrinks to a zero.

$$\lim_{\Omega \to 0} \left[ \int_{\Omega} \partial_t q(x_i, t) + \partial_i F_i(q, x_i, t) - S(q, x_i, t) \, \partial\Omega \right] = 0.$$
 (1.4)

In this limit, the enforcement of the equality changes from one over a finite volume into one that is enforced at a particular point. To ensure point-wise enforcement over a given domain, the integrand itself is required to be equally zero at every point in the domain; this yields the differential conservation law

$$\partial_t q(x_i, t) + \partial_i F_i(q, x_i, t) - S(q, x_i, t) = 0. \tag{1.5}$$

When equipped with adequate boundary and initial data, this differential equation defines the requirement for all sufficiently smooth functions that describe how the conserved quantity evolves at every point in space-time.

The differential form may seem cleaner than the integral form, but it is not valid for all functional forms of q. In transitioning from equation 1.4 to equation 1.5, it is assumed that all functions in the integrand remain bounded. If any of the terms within the limit become infinite, the balance cannot be satisfied. In particular, if q has an area where it undergoes a discontinuous jump, called a shock, the gradient of the flux is infinite and the differential form is rendered invalid. While methods that handle shocks in a robust fashion (shock-capturing schemes) are not the focus of this work, it is mentioned for completeness and consideration.

#### 1.1.2 Vector Form

The simulation of real world problems often involves the solution of a system of conservation laws. In general, the systems are nonlinear and tightly coupled. The actual solution of these problems will be discussed in future sections, but the notation used will be introduced now.

The integral conservation law for a vector of conserved quantities  $q_i$  is

$$\partial_t \int_{\Omega} q_i(x_i, t) \, \partial\Omega = \int_{\Gamma} -F_{ij}(q_i, x_i, t) n_j \, \partial\Gamma + \int_{\Omega} S_i(q_i, x_i, t) \, \partial\Omega, \tag{1.6}$$

where  $F_{ij}$  is a matrix of surface fluxes of  $q_i$  in the j-th direction,  $n_j$  is the outward unit normal of the surface  $\Gamma$ , and  $S_i$  is the volume source of  $q_i$ . The integrals of the vector

quantities represent element-wise integration. A process similar to the scalar case also yields the differential form of the system conservation law:

$$\partial_t q_i(x_i, t) + \partial_i F_{ii}(q_i, x_i, t) - S_i(q_i, x_i, t) = 0,$$
 (1.7)

where  $\partial_j F_{ij}$  represents a row-wise divergence operation.

## 1.2 Thermohydraulic Laws

The main equations of interest in thermohydraulics will be presented using the definitions above. Mass, momentum, and energy are the primary quantities of interest and will be derived for an assumed-flowing system. For all of the definitions, the following items are noted:

- As prescribed in section 1.1, all of the conserved quantities are taken to be continuous functions of space and time. Although fluids are actually composed of discrete, interacting molecules, taking the quantities to be everywhere-defined is a valid approximation as long as the length scales to be modeled are much greater than the mean free path of the local medium. NEED REFERENCE
- Also as prescribed in section 1.1, all of the conserved quantities will be considered on a per unit volume basis.
- For clarity, the dependence of all quantities and functions will be omitted; however, they are all assumed to be dependent of any variable in the system.

#### 1.2.1 Conservation of Mass

The net mass in a control volume or, rather, the volume-integrated density is conserved as a simple scalar. It is controlled by the background flow field advecting the mass through the boundary. Therefore, the integral conservation law for density is

$$\partial_t \int_{\Omega} \rho \, \partial\Omega = \int_{\Gamma} -(u_j \rho) n_j \, \partial\Gamma + \int_{\Omega} S^{\rho} \, \partial\Omega, \tag{1.8}$$

#### 1.2.2 Conservation of Momentum

A control volume's momentum balance is more complicated than the simple in-out mass balance. It is inherently a vector quantity, so the combination of mass and the background flow field feed the momentum in a non-linear fashion. Also, even though a continuum approximation has been made, there still physical molecular stresses on the boundary must be modeled in some way. Lastly, a body force must be added to account for buoyancy effects on the volume relative to the outside universe. With all of the above nuances considered, the integral conservation for momentum is taken to be

$$\partial_t \int_{\Omega} \rho u_i \, \partial\Omega = \int_{\Gamma} (-u_j \rho u_i) n_j \, \partial\Gamma + \int_{\Gamma} (-\delta_{ij} P + \tau_{ij}) n_j \, \partial\Gamma + \int_{\Omega} \rho g_i + S_i^u \, \partial\Omega \tag{1.9}$$

where  $\rho u_i$  is the momentum in direction i,  $u_j$  is the velocity in direction j,  $\delta_{ij}$  is the Kronecker delta, P is the thermodynamic pressure,  $\tau_{ij}$  is the viscous stress tensor, and  $g_i$  is the gravitational acceleration in direction i. A more detailed discussion of the pressure and viscous stress tensor is presented in section 1.3.

Each of the terms on the right-hand side of equation 1.9 stems from the considerations made. The first integral represents how the background flow field transports the momentum across the volume's surface. The second integral accounts for all of the molecular stresses present in the volume, both normal and tangential. The pressure is an ever-present, thermodynamic property of the medium, and the tensor  $\tau_{ij}$  accounts for all molecular friction and other deviations from thermodynamic equilibrium. The last integral has a source of momentum due to the mass of the volume in addition to the general source  $S^u$ .

## 1.2.3 Conservation of Energy

The energy equation is similar to the momentum equation since there are non-linear and molecular components to the balance. The energy to be balanced is the *total energy* of the medium, where total means the sum of thermal, kinetic, and potential energies. The integral conservation for total energy is taken to be

$$\partial_t \int_{\Omega} \rho e \, \partial\Omega = \int_{\Gamma} \left[ -(\rho e + P)u_j + u_i \tau_{ij} - q_j \right] n_j \, \partial\Gamma + \int_{\Omega} \rho g_j u_j + S^e \, \partial\Omega \tag{1.10}$$

where  $\rho e$  is the total energy of the system and  $q_j$  is a conductive heat flux; all of the other terms have the same meaning as in the momentum equation. A more detailed discussion of the conductive heat flux is presented in section 1.3.

As in the momentum equation, all right-hand side terms have a purpose. The first term,  $\rho e + P$ , represents the advected energy in addition to the associated work due to pressure. The second term accounts for energy loss due to molecular friction. The third term arises due to molecular diffusion of heat across the boundary. The final term acts as a source or sink of energy due to the buoyancy force from momentum.

#### 1.2.4 Differential Forms

All of the integral conservation laws shown above also have differential forms:

$$\partial_t \rho + \partial_i (u_i \rho) = S^{\rho} \tag{1.11}$$

$$\partial_t(\rho u_i) + \partial_j(u_j \rho u_i) = -\partial_i P + \partial_j \tau_{ij} + \rho g_i + S_i^u$$
(1.12)

$$\partial_t(\rho e) + \partial_j(u_j \rho e) = -\partial_j(u_j P) + \partial_j(u_i \tau_{ij}) + \rho g_i u_i + S^e$$
(1.13)

Given the same initial and boundary conditions, these differential equations and their integral counterparts are equivalent for sufficiently smooth solutions. As shown here, "sufficiently smooth" means that at least all first derivatives must be continuous over the entire solution domain; however, certain constitutive relations (e.g., for the stress tensor) may place more stringent requirements on continuity.

#### 1.2.5 Channel Flow

All of the above equations are three-dimensional, but in thermohydraulic modeling, it is common practice to collapse the momentum equation to a single, dominant flow direction. The result of this dimensional collapse will be called *channel flow* momentum equation. The channel flow equation can be created by simply dotting the integral momentum equation with a constant unit vector  $z_i$ . The unit vector represents the physically dominant flow direction for the given volume. Performing the dot operation and letting  $a_z = a_i z_i$  for any vector  $a_i$ , the channel flow equation is of the form

$$\partial_t \int_{\Omega} \rho u_z \, \partial\Omega = \int_{\Gamma} (-u_j \rho u_z) n_j \, \partial\Gamma + \int_{\Gamma} -P n_z + \tau_{ij} z_i n_j \, \partial\Gamma + \int_{\Omega} \rho g_z + S_z^u \, \partial\Omega \tag{1.14}$$

#### 1.3 Constitutive Relations

In deriving these conservation laws, a number of unknowns were introduced to account for different thermodynamic, molecular, and non-equilibrium phenomena. These unknowns have to be related to the system variables in some manner to make the system solvable. Equations that perform this task as called *constitutive relations* or, less formally, models. The terms of interest are the thermodynamic pressure, viscous stress tensor, and conductive flux.

#### 1.3.1 Pressure

Pressure is a macroscopic, thermodynamic property that arises from ensembles of particles exerting forces over some unit area. A constitutive relation that relates pressure to other system variables that define the system's state is called an *equation-of-state* (EOS). An EOS can be constructed from first principle (microscopic) approaches or thermodynamic potential (macroscopic) approaches and may be composed of one or more equations. Regardless of the EOS's construction, it should be able to calculate all relevant thermodynamic properites given two independent properites for a pure fluid NEED REFERENCE.

Since there are multiple avenues for creating an EOS, for any given fluid, several may exist; each one varying in complexity of formulation, range of validity, and fidelity to data. The choice of EOS is dependent on the importance of each of those three attributes. The EOS used for this work is the International Association for the Properties of Water and Steam (IAPWS) Formulation 1995 for the Thermodynamic Properties of Ordinary Water Substance for General and Scientific Use (IAPWS-95). The IAPWS-95 formulation is extremely complex, but the range of validity and fidelity to data is is second-to-none for water. The details of the formulation is left to the appendix and its original publication.

#### 1.3.2 Stress Tensor

The stress tensor requires a model that relates molecular friction, continuum deformation, and system variables. There are, like an EOS for a fluid, several different stress tensor models for any given system. These models vary in terms of the type of valid fluid, dimensionality of usage, and fidelity to data.

Newtonian fluid models are among the most popular stress tensor model choices and forms the basis for the Navier-Stokes equations. From experiment, Newton observed that a proportional relationship exists between the shear stress exerted on a fluid and the fluid's uni-directional velocity gradient for a wide array of fluids. Stokes then extended this proportional relationship to higher dimensions by making three assumptions: the tensor is a linear function of strain, the fluid is isotropic, and the tensor's divergence must be 0 for a fluid at rest (such that only pressure is the only extant surface force). With all of these observations and assumptions considered, the Newtons stress tensor model (sometimes described as a compressible model) is defined by the equation:

$$\tau_{ij} = \mu \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) + \lambda \delta_{ij} \frac{\partial u_k}{\partial x_K}$$
 (1.15)

#### 1.3.3 Conductive Flux

## 1.4 Numerical Methods

## 1.5 Stability Theory