

UNIVERSITY OF OSLO  
Faculty of Mathematics and Natural Sciences

**Exam in Cosmology I**

**Day of exam: Thursday 13th of June 2013**

**Exam hours: 09.00 – 13.00**

**This examination paper consists of 4 pages.**

**Appendices: None**

**Permitted materials: All non-communicative aids.**

*Make sure that your copy of this examination paper is complete before answering.*

Throughout the equation set you can assume that today's value of the Hubble constant is:

$$\begin{aligned} H_0 &= 70 \text{kms}^{-1} \text{Mpc}^{-1} \\ &= (14 \times 10^9 \text{yr})^{-1} \end{aligned}$$

and that the scale factor today is set to unity

$$a_0 = 1$$

## Problem 1

### Cosmological Structure Formation

- a) What is the meaning of the Jeans-wavelength and of the Jeans-mass?
- b) Discuss the Jeans-length before and after recombination.

Consider the growth of small (linear) perturbations of scales large enough so that pressure is negligible.

- c) Show that in the radiation dominated regime,  $\Omega_{rad} = 1$  ( $\Omega_m, \Omega_\lambda \ll \Omega_{rad}$ ), the growth of perturbations to the radiation density can be described as  $\delta \propto \tau^2$ , where  $\tau$  is the so-called conformal time,  $d\tau = dt/a(t)$ .
- d) Show that the same conformal time relation as for radiation in c) can describe the growth of matter perturbations in the matter dominated regime,  $\Omega_m = 1$  ( $\Omega_{rad}, \Omega_\lambda \ll \Omega_m$ ).

## Problem 2

### **Primordial Nucleosynthesis**

- a) Why is there no C, N, O etc. fusion in the early Universe?
- b) What is the binding energy of the hydrogen atom? Explain qualitatively why recombination starts only when the temperature drops below 0.5 eV.

## Problem 3

### **The Inflationary Universe**

- a) Describe 3 severe problems of the Standard Big-Bang Model of the Universe without Inflation.
- b) Describe the basic elements of the Inflationary Universe (Friedmann equation, inflaton field, dynamics etc.).
- c) What is the meaning of the slow-roll approximation?
- d) Describe, how the problems you mentioned in a) are solved by the Inflationary Universe.

## Problem 4

### Modified Chaplygin gas

The modified Chaplygin gas obeys the equation of state

$$p = A\rho - \frac{B}{\rho^\alpha}$$

- a) Write down the adiabatic expansion equation for the modified Chaplygin gas.
- b) Show that

$$\rho = \left[ C \left( \frac{a_0}{a} \right)^{3(1+\alpha)(1+A)} + \frac{B}{1+A} \right]^{\frac{1}{1+\alpha}} \quad (1)$$

where  $C$  is a constant, solves the adiabatic equation for the modified Chaplygin gas if  $\rho_0$  is the density of the gas at time  $a_0$ .

- c) Discuss the modified Chaplygin gas in the limit where  $B = 0$ . What does  $A$  then signify in terms of a normal ideal gas?
- d) Discuss the modified Chaplygin gas in the limit where  $A = 0$  and  $\alpha = -1$ . What does  $B$  then signify in terms of a normal ideal gas?