Data-driven subgrid-scale parametrisation for a simple dynamical system Honours project update

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The complexity of atmospheric models makes parametrisation testing difficult.

I aim to:

- 1. Use a simpler dynamical system as a test-bed
- 2. Build data-driven parametrisation schemes from scratch, with/without
- i Stochasticity
- ii Memory
- 3. Evaluate the schemes' performance

Rayleigh-Bénard convection in 2D

$$egin{aligned} rac{\partial oldsymbol{u}}{\partial t} + oldsymbol{u} \cdot oldsymbol{
abla} oldsymbol{u} &= -oldsymbol{
abla} \pi + \left(rac{\mathsf{Pr}}{\mathsf{Ra}}
ight)^{1/2} \left(1 +
u^* |
abla^2 oldsymbol{u}|\right)
abla^2 oldsymbol{u} &+ oldsymbol{u} \cdot oldsymbol{
abla} oldsymbol{u} &= (\mathsf{Ra}\,\mathsf{Pr})^{-1/2} \left(1 +
u^* |
abla^2 oldsymbol{u}|\right)
abla^2 oldsymbol{u} &+ oldsymbol{u} \cdot oldsymbol{
abla} oldsymbol{u} &+ oldsymbol{u} \cdot oldsymbol{
abla} oldsymbol{u} &= 0 \end{aligned}$$

- ullet Rayleigh number ${\sf Ra}=10^9\Rightarrow {\sf turbulent}$
- Prandtl number Pr = 1
- ullet Nonlinear hyperdiffusion for model stability at low resolution $(
 u^*=10^{-3})$
- No-slip, isothermal top/bottom boundaries; periodic lateral boundaries

Numerical methods

- Equations solved using dedalus v3: spectral PDE solver for Python
 - Highly flexible
 - Designed to run in parallel
- Horizontal Fourier series expansion
- Vertical Chebyshev polynomial expansion

Animation

Integration for 100 time units with 1024 horizontal and 128 vertical modes:

https://youtu.be/SU8ColUYY1A

Key properties:

- Turbulent flow
- Large and small scale features
- Only two prognostic variables
- Comparatively inexpensive (13 cpu-hr for the above run)
- ⇒ A simple dynamical system analogue of the atmosphere

I need:

- 1. A reasonably accurate high-resolution model to treat as "truth"
- 2. A coarse-resolution model to correct using parametrisation
- 3. Evidence that under-resolution introduces biases that are:
 - i Systematic: consistent over- or under-estimation, not mere numerical error
 - ii Statistically significant

Quantities of interest

Nusselt number (dimensionless rate of vertical heat transport):

$$\mathsf{Nu}(z,t) \equiv (\mathsf{Ra}\,\mathsf{Pr})^{1/2} \langle w heta
angle_x - rac{\partial \langle heta
angle_x}{\partial z}$$

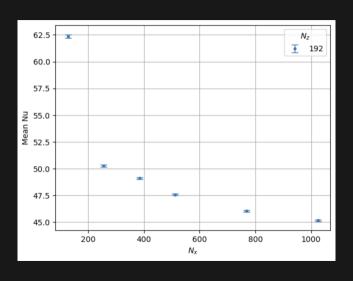
Root-mean-square velocity:

$$u_{\sf rms}(t) = \sqrt{\langle oldsymbol{u} \cdot oldsymbol{u}
angle_{ imes,z}}$$

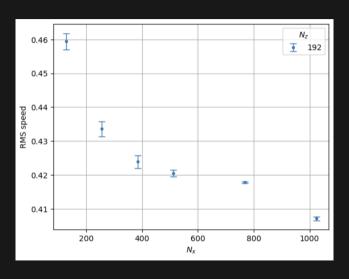
Mean values are obtained via:

- 1. 1000 time unit spin-up at coarse resolution
- 2. 300-time-unit integration at new resolution
- 3. Averaging over final 275 time units

Mean Nusselt number



Mean RMS velocity



Next steps: building the parametrisation

- 1. Given the state S_t of the fine model at time t,
 - i Advance it by Δt using the fine model $\mathcal{F}_{\Delta t}$, giving $\mathcal{F}_{\Delta t} S_t$
 - ii Coarse-grain it and advance it by Δt using the coarse model $\mathcal{C}_{\Delta t}$, giving $\mathcal{C}_{\Delta t}\langle S_t \rangle$
- 2. Determinine the *unresolved tendency*

$$\mathcal{T} = rac{\left\langle \mathcal{F}_{\Delta t} \mathcal{S}_t
ight
angle - \mathcal{C}_{\Delta t} \left\langle \mathcal{S}_t
ight
angle}{\Delta t}$$

- 3. Fit a model to the dataset of $(\langle S_t \rangle, T)$
- 4. Plug this into the coarse model as an external forcing and assess the results