Bayesian Inference on Single-Parameter Models

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1 Prior to Posterior Update for Bernoulli Trials

The model for the trial data is that $Y_i|\theta \stackrel{iid}{\sim} Bern(\theta)$. If the chosen distribution for the prior is $\theta \sim Beta(\alpha_0, \beta_0)$, then the posterior distribution will be $\theta|y_1...y_n \sim Beta(\alpha_0 + s, \beta_0 + f)$ where s and f denote the number of successes and failures (ones and zeros) in our sample data. This is because the Beta distribution is a conjugate prior to the Bernoulli distribution.

With $\alpha_0 = \beta_0 = 3$, s = 8 and n = 24, i.e. f = 16, the analytical mean and standard deviation of the posterior distribution is $\mu = 0.3666...$ and $\sigma = 0.0865...$ The mean and standard deviation can also be evaluated by drawing random samples from the posterior distribution, given that the sample size is large enough. This is illustrated in figure 1, where one can see that the two values converge towards their analytically derived values.

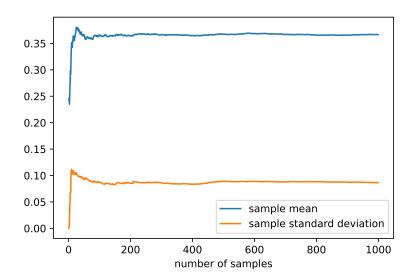


Figure 1: Convergence of the sample mean and standard deviation

From the samples, one can also evaluate probabilities for the parameter θ . For instance $P[\theta > 0.4|y_1...y_n] \approx \frac{n_{0.4}}{n} = 0.3445$ when $n_{0.4}$ denotes the number of samples greater than or equal to 0.4 for n = 10000 samples. This can be compared to the true probability, given the posterior distribution 1 - F(0.4) = 0.3426 where F denotes the cdf of the posterior distribution.

Another way to illustrate the distribution of the posterior is to transform it to the log-odds: $\phi = log(\frac{\theta}{1-\theta})$. The histogram for ϕ is illustrated in figure 2, reusing the simulated draws from before.

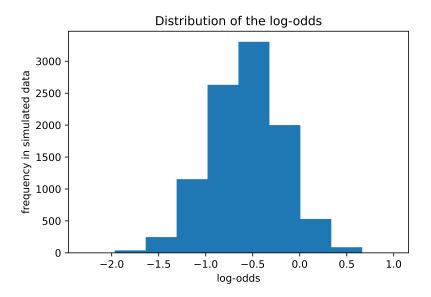


Figure 2: Histogram of the log-odds

```
from scipy import stats as st
  import numpy as np
2
  import matplotlib.pyplot as plt
3
  np.random.seed(123) # set seed for reproducability
  # Parameters for beta prior distribution
   alpha0 = 3
   beta0 = 3
9
10
  n = 24 # number of samples from Bernoulli trial data
11
  s = 8 \# number of sucesses in sample
12
   f = n - s \# number of failiures
13
14
  # Analytical mean and standard deviation for posterior
15
  mean, var = st.beta.stats(alpha0+s, beta0+f, moments="mv")
16
   std = np.sqrt(var)
17
18
  # Draw 10 000 random samples from posterior
19
   post\_sim = st.beta.rvs(alpha0+s, beta0+f, size=10000)
20
21
  # Check convergence of mean, standard deviation
22
  mean\_sim = np.zeros(1000)
23
   std_sim = np.zeros(1000)
24
25
   for i in range (1000):
26
       mean\_sim[i] = np.mean(post\_sim[0:i])
27
       std_sim[i] = np.std(post_sim[0:i])
28
29
   plt.plot(mean_sim)
30
   plt.plot(std_sim)
  plt.legend(["sample mean", "sample standard deviation"])
```

```
plt.xlabel("number of samples")
   plt.savefig("mean_sd_convergence.png", dpi=1500)
34
   plt.show
35
36
  #Compare simulated probability to exact posterior
37
  geq04 = post\_sim >= 0.4
38
  # ratio of number of samples >= 0.4 and total samples
39
  p_sim = float(sum(geq04))/float(len(post_sim))
40
  p_{true} = 1 - st.beta.cdf(0.4, alpha0 + s, beta0 + f)
41
42
  # Explore the distribution of the log-odds
43
  logodds_sim = np.log(post_sim/(1-post_sim))
44
  plt.hist(logodds_sim)
45
  plt.title("Distribution of the log-odds")
  plt.xlabel("log-odds")
47
  plt.ylabel("frequency in simulated data")
  plt.savefig("logodds_hist", dpi = 1500)
49
  plt.show
```

2 The Gini Coefficient and the log-normal distribution

The model used for the data of income levels is $Y_i|\mu, \sigma \stackrel{iid}{\sim} \log N(\mu, \sigma^2)$, or equivalently $\log(Y)|\mu, \sigma \stackrel{iid}{\sim} N(\mu, \sigma^2)$. Assuming that $\mu = 3.8$ is known, and using a non informative prior for σ , i.e. $p(\sigma^2) \propto 1/\sigma^2$, the posterior distribution will be $\sigma^2|y_1,...,y_n,\mu \sim \text{Scale-Inv-}\chi^2(n,\tau^2)$, where τ^2 is defined as:

$$\tau^{2} = \frac{\sum_{1}^{n} (\ln(y_{i}) - \mu)^{2}}{n}$$

To simulate from this distribution, the algorithm is the following:

- 1. Draw n samples from $X \sim \chi^2(n-1)$
- 2. Calculate τ^2 using the data and the formula above
- 3. Calculate $\sigma^2 = \frac{(n-1)\tau^2}{X}$
- 4. The result is $\sigma^2 \sim \text{Scale-Inv-}\chi^2(n,\tau^2)$

If the population income follows a lognormal distribution, then the Gini coefficient takes on the value $G = 2\Phi(\sigma/\sqrt{2})$ where Φ is the cdf of the standard normal. By drawing samples of σ following the algorithm above, we can explore the distribution of the Gini coefficient. Figure 3 illustrates a histogram of the simulated Gini coefficients.

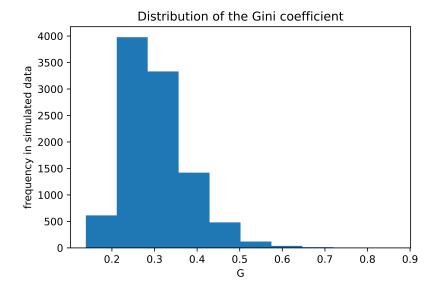


Figure 3: Histogram of the Gini coefficient

From the simulated data, one can also fit a density kernel to explore further properties of the posterior distribution, such as the equal tail credible interval and the highest posterior density interval (HPDI). The resulting intervals, on a 90% level, were:

$$HPDI_{0.9} = [0.1892; 0.4184]$$

 $CI_{0.9} = [0.2052; 0.4464]$

One can see that allthough both intervals contain 90% of the probability mass, they are not equal, which follows from the fact that the distribution is skewed. The fitted density, together with the corresponding intervals, is shown in figure 4.

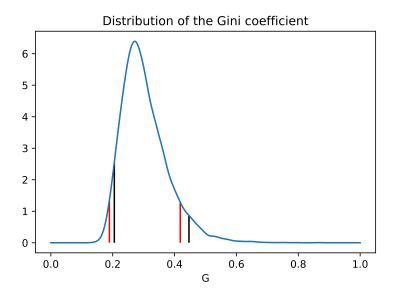


Figure 4: Density, equal tail credible interval (black) and HPDI (red) for the Gini distribution

```
from scipy import stats as st
   import numpy as np
   import matplotlib.pyplot as plt
3
   from scipy.integrate import quad
  np.random.seed(123) # set seed for reproducability
  # Dataset of monthly income in KSEK/mo.
  y = np.array([38, 20, 49, 58, 31, 70, 18, 56, 25, 78])
  n = len(y)
10
11
  # Known parameter of log-normal distribution
12
  mean = 3.8
13
  # Draw 10 000 random samples from the scaled inverse chi2
15
  # by first drawing from a chi2 distribution with n-1 df's
   tau2 = np.sum((np.log(y)-mean)**2)/n
17
  X = st.chi2.rvs(n-1, size=10000)
18
   var_sim = (n-1)*tau2/X
19
20
  # Explore the distribution of the Gini-coefficient
^{21}
  # By fitting a gaussian density kernel to the data
22
  G = 2*st.norm.cdf(np.sqrt(var_sim/2))-1
23
   G_{-pdf} = st.gaussian_kde(G)
24
25
   \operatorname{def} G_{-}\operatorname{cdf}(x):
26
       return quad(G_pdf, 0, x)[0]
27
28
  # Plot histogram from the simulated variances
29
   plt.hist(G)
30
   plt.title ("Distribution of the Gini coefficient")
```

```
plt.xlabel("G")
   plt.ylabel ("frequency in simulated data")
33
   plt.savefig("gini_hist.png", dpi=1500)
   plt.show
35
36
  # Evaluate the pdf and cdf for a dense grid of values.
37
  x = np. linspace(0,1,1000)
38
   G_pdf_vals = G_pdf(x)
39
   G_{cdf_vals} = np.zeros_{like}(x)
40
41
   for i, vals in enumerate(x):
42
       G_{cdf_vals}[i] = G_{cdf}(vals)
43
44
  # Generate 90% highest posterior density interval
45
   sorted_index = np. argsort(G_pdf_vals)[::-1]
46
  HPDI = [x[sorted\_index[1]]]
47
48
  # By adding points until they integrate to 0.9
   for i in range(1,len(sorted_index)):
50
       if quad(G_pdf, min(HPDI), max(HPDI))[0] < 0.9:
51
           HPDI. append (x [sorted_index [i]])
52
53
  HPDI = [\min(HPDI), \max(HPDI)]
54
55
  # Generate 90% equal tail credible interval from the cdf values
56
  # By having 5% of the mass to the left and to the right
57
  lb\_index = np.argmin(np.abs(G\_cdf\_vals-0.05))
58
   ub\_index = np.argmin(np.abs(G\_cdf\_vals - 0.95))
59
60
   eq_tail_CI = [x[lb_index], x[ub_index]]
61
62
  # Plot density
63
  # With the equal tail CI and HPDI
  plt.plot(x, G_pdf_vals)
65
  plt.title("Distribution of the Gini coefficient")
   plt.xlabel("G")
67
   plt.vlines(eq_tail_CI[0], 0, G_pdf(eq_tail_CI[0]))
  plt.vlines(eq_tail_CI[1], 0, G_pdf(eq_tail_CI[1]))
69
  plt.vlines(HPDI[0], 0, G_pdf(HPDI[0]), color = "red")
  plt.vlines(HPDI[1], 0, G_pdf(HPDI[1]), color = "red")
   plt.savefig("gini_density.png", dpi = 1500)
  plt.show
```

3 Bayesian Inference on the von-Mises distribution

The model used for the data of wind directions (in radians) is $Y_i|\mu,\kappa \stackrel{iid}{\sim} \text{von-Mises}(\mu,\kappa)$. The pdf of this distribution is:

$$\frac{e^{(\kappa\cos(y-\mu))}}{2\pi I_0(\kappa)} \quad \forall -\pi \le y \le \pi$$

Where I_{ν} is the modified Bessel function of the first kind of order ν . By assuming $\mu = 2.39$ to be known, and using a prior distribution for $\kappa \sim Exp(\lambda)$, we can derive the posterior through:

$$p(\kappa|y_1,..,y_n,\mu) \propto \lambda e^{-\lambda\kappa} \prod_{1}^{n} \frac{e^{\kappa\cos(y-\mu)}}{2\pi I_0(\kappa)} \propto \left(\frac{\lambda}{I_0(\kappa)}\right)^n \exp\left[\kappa\left(\sum_{1}^{n} \cos(y_i - \mu) - \lambda\right)\right]$$

To make this a propper pdf (i.e. that it integrates to one), it is normalized by dividing the expression above with its integral. The resulting distribution when $\lambda = 1$ is plotted in figure 5.

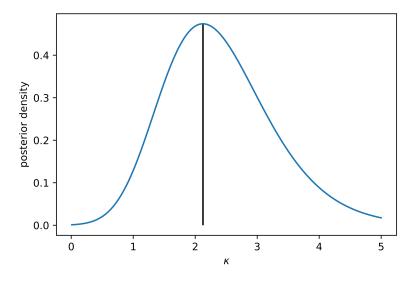


Figure 5: Posterior distribution of κ

Finally, the approximate posterior mode is found, as the value of κ that yields the highest pdf-value, as $\kappa = 2.1249$, which is also highlighted by the vertical line in the plot above.

```
import numpy as np
  import matplotlib.pyplot as plt
  from scipy.special import iv as bessel
3
  from scipy.integrate import quad
  np.random.seed(123) # set seed for reproducability
  # Dataset of wind directions in radians
  y = \text{np.array}([-2.44, 2.14, 2.54, 1.83, 2.02, 2.33, -2.79, 2.23, 2.07, 2.02])
        = 2.39
11
  lamda = 1
13
  # Function proportional to the posterior distribution
  # Given von-Mises(kappa, mu) likelihood and Exp(lamda) prior
15
  def posterior (kappa, y, lamda, mu):
      n = len(y)
17
       return (lamda/bessel (0, kappa)) **n*np.exp(kappa*(np.sum(np.cos(y-mu))-lamda
18
          ))
19
  # Find posterior for different values of kappa
20
  kappa = np. linspace (0.01, 5, 10000)
21
  posterior_kappa = np.zeros_like(kappa)
22
23
  posterior_kappa = posterior(kappa, y, lamda, mu)
24
25
  # Normalize the posterior so that it integrates to 1
26
  posterior_kappa = posterior_kappa/(quad(posterior, 0.01, 5, args = (y, lamda,
27
      mu))[0])
28
  # Find the posterior mode from the density
  mode = kappa [np.argmax(posterior_kappa)]
30
31
  # Plot posterior
32
  plt.plot(kappa, posterior_kappa)
  plt.xlabel(r"$\kappa$")
34
  plt.ylabel("posterior density")
35
  plt.vlines(mode,0,np.max(posterior_kappa))
  plt.savefig("kappa_posterior_pdf.png", dpi=1500)
37
  plt.show
```