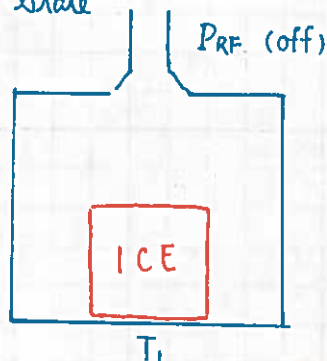


Plasma Physics.

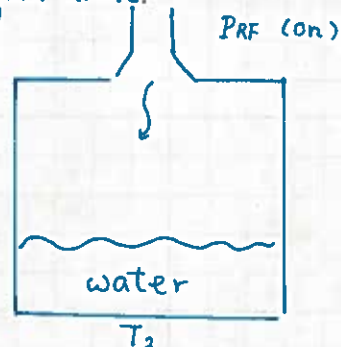
1. What is a plasma?

- 4th state of matter.

(i) Solid State



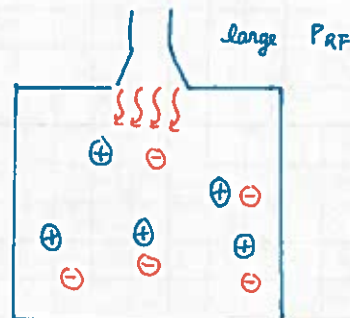
(ii) Liquid State



(iii) Gas



(iv)

* Temperature alone does not define a plasma. ($1 \text{ eV} \approx 11.6 \cdot 10^3 \text{ } ^\circ\text{K}$)

- white dwarf \longrightarrow not a plasma
 $T \sim 10 \text{ eV}$
 $n \sim 10^{30} \text{ cm}^{-3}$

- ionosphere \longrightarrow a plasma.
 $T \sim 0.1 \text{ eV}$
 $n \sim 10^5 \text{ cm}^{-3}$

2. What parameter determines plasma behavior.

- Temperature
- Density.

• DEGREE OF CORRELATION $\approx \frac{\text{Interaction Energy}}{\langle KE \rangle}$

DEF: $\Gamma = \frac{\phi}{kT}$

- Estimation of plasma parameter.

$$\langle \phi \rangle \approx \frac{e^2}{\langle r \rangle} \approx e^2 n^{+1/3}$$

$$\langle r \rangle \approx n^{-1/3}$$

$$\therefore \Gamma \equiv \frac{\langle \phi \rangle}{\langle KE \rangle} = \left(\right) \frac{e^2 n^{1/3}}{T}$$

↑
numerical constant.

\therefore 2 plasma regimes

- low density.
- high temperature.

$$\Gamma \ll 1 \longrightarrow \text{plasma.}$$

$$\Gamma \gg 1 \longrightarrow \text{degenerate QM Fluid}$$

* $\Gamma \approx 1 \longrightarrow$ strongly coupled plasma.

** in 2-D, crystallization occurs @ $\Gamma \approx 170$ (Exp. and simulation)

$$\Gamma = \frac{e^2 n^{1/3}}{T} = \frac{4\pi e^2 (n^{2/3} n^{1/3})}{4\pi n^{2/3} T \cdot \frac{m}{m}} = n$$

Recall.

$$\omega_p^2 = \frac{4\pi n e^2}{m_{\text{electron mass}}} = \text{"electron" plasma frequency.}$$

$$\omega_p \approx 5.6 \cdot 10^4 \cdot \sqrt{n} \frac{\text{rad}}{\text{sec.}} \quad (n \text{ in } \text{cm}^{-3})$$

$$\bar{v}_T^2 = \frac{T}{m.}$$

$$\Gamma = \frac{1}{4\pi} \frac{\omega_{pe}^2}{\bar{v}_e^2} \cdot \frac{1}{n^{2/3}}$$

* $\frac{\bar{v}_e}{\omega_p} \approx$ distance traveled in one plasma cycle = λ_D .

$$\lambda_D \equiv \frac{\bar{v}_e}{\omega_p} = \text{DEBYE LENGTH.}$$

$$\lambda_D = 7.4 \cdot 10^2 \sqrt{T} n^{-1/2} \text{ cm}$$

$$\therefore \Gamma = \left(\frac{1}{4\pi} \right) \frac{1}{\lambda_D^2 n^{2/3}}$$

recall

$$N = \left(\frac{4\pi}{3} \right) \lambda_D^3 n. \longrightarrow \# \text{ of particles in a debye sphere.}$$

$$\Gamma = (\#) \frac{1}{N^{2/3}} \longrightarrow$$

∴ plasma is determined by # of particles in a Debye sphere.

plasma requires $N \gg 1$

• in an experiment, $N \approx 10^6$.

3. How large a container do you need to see plasma behavior.



l

$$* T_{\text{TRANSIT}} = \frac{l}{\bar{v}_e} = \text{transit.}$$

$$* \text{Typical Time in a plasma is } \approx \frac{2\pi}{\omega_p}$$

$$l \gg \lambda_D = \frac{\bar{v}_e}{\omega_p}$$

Dual Constrains.

(i) $N \gg 1$

(ii) $l \gg \lambda_D \rightarrow$ in practice: $l \approx 10 \lambda_D$ is good enough.

Recall: any external influence is wiped out by DEBYE SHIELDING.

Subtleties of plasma physics \longrightarrow Why don't we use Thermodynamics for plasmas?

(i)



classical.



Plasma

(ii) in an ideal gas



However, for away interactions $(\frac{e^2}{r})$ decays slowly in a plasma.
 \Rightarrow plasma is interaction w/ collective oscillation, i.e., "waves"

\Rightarrow a plasma is really a "classical duality"

• particle behavior exists.

• wave behavior exists.

i.e., particle interacts w/ wave, and it is the wave itself.

222 A.

10/2.

1. historical perspective.

Date (~)	Event	Name.
1879	Exp. on gas discharges	Crookes
1906	Plasma Disc. \rightarrow in jellium model of atom.	Lord Rayleigh.
1926	Named "Plasma" invented + basic properties explained	Langmuir Tonks.
1936	MHD waves. \rightarrow couples matter w/ EM waves	Alfven
1938	Collision model \rightarrow strange for an strongly interacting system.	Vlasov.
1945	Collisionless damping (!!!) / Linear response of plasmas	Landau.
1944-50	Isotope separation + H bomb studies. (i.e., <u>large release of energy</u>)	Classified (Bohm)
1957	Fusion research de-classified. - MHD Theory - Fokker-planck description.	USSR, UK, USA.

 \rightarrow Papers available.

Modern Age of Plasma Physics

1960	Q-machine.	Rynn + D'angelo.
1962	Exp. observation of Landau damping.	Malmberg. Wharton.
1965-70	Linear waves, mirror machines, stellarators	
1970-80	NL-waves. (pondermotive force, para. instability). Tokamak.	
1980-90	Tokamak confinement. Application to space science. Non-neutral plasmas. Non-uniform properties	
Today.	Transport... \rightarrow Scott's work. Non-equilibrium.	

 \rightarrow Theory done @ 1945.

2. System of units. \rightarrow Gaussian.

$$\nabla \times \underline{E} = -\frac{1}{c} \frac{\partial}{\partial t} \underline{B}$$

$$\nabla \times \underline{B} = \frac{4\pi}{c} \underline{j} + \frac{1}{c} \frac{\partial}{\partial t} \underline{E}$$

$$\nabla \cdot \underline{E} = 4\pi \rho.$$

$$* c = '3' \cdot 10^{10} \text{ cm/sec.}$$

$$[E] = \frac{\text{stat-volt}}{\text{cm}}$$

Exp. unit.

$$[E] = \frac{\text{volt}}{\text{meter}} = \frac{1}{3} \cdot 10^{-9} \frac{\text{statvolt}}{\text{cm}}$$

$$[I] : \text{stat-amps.} \quad * 1 \text{ amp} = 3 \cdot 10^9 \text{ stat-amp.}$$

$$[B] : \text{in Gaussian in lab.}$$

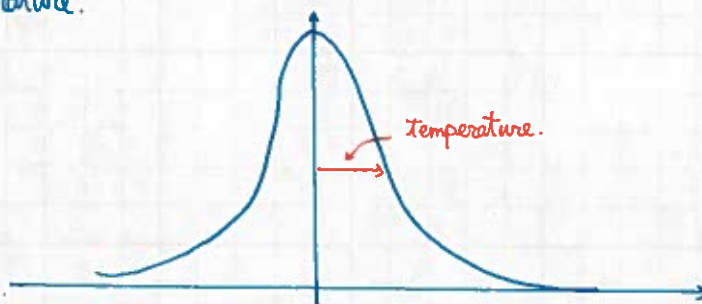
$$[e] : \text{in stat-coulomb} \quad * 1 \text{ coulomb} = 3 \cdot 10^9 \text{ stat-coulomb.}$$

• temperature in eV

$$1 \text{ eV} \approx 11.6 \cdot 10^3 \text{ K}$$

3. Macroscopic parameters.

• temperature.



• density. = n

• L_T = temperature scale length.

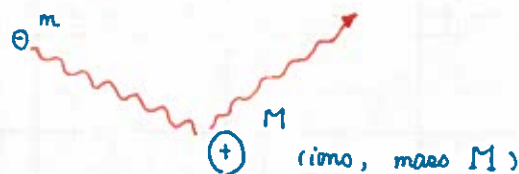
• L_n = density scale length.

$$L_T^{-1} = \frac{\partial}{\partial x} \ln(T)$$

$$L_n^{-1} = \frac{\partial}{\partial x} \ln(n)$$

\rightarrow where is L_T , L_n , measured in a plasma.

Equilibrium Problem



$$\frac{M}{m} > 1836.$$

- fastest. \rightarrow electron - electron equilibration: τ_{ee} .
- ion - ion equilibration. $\tau_{ii} = \sqrt{\frac{M}{m}} \tau_{ee}$.
- slowest \rightarrow electron - ion equilibration: $\tau_{ei} = \frac{M}{m} \tau_{ee}$

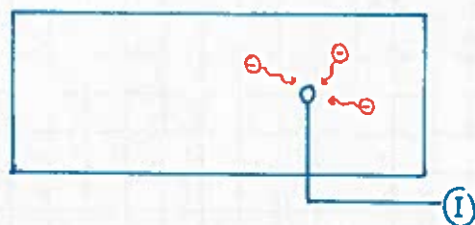
\therefore in general, $T_e \neq T_i$ (except in well-confined plasmas)

How to measure macroscopic parameters

I. for low density, low temperature plasma \rightarrow Langmuir Probe.

$$n < 10^{12} \text{ cm}^{-3}$$

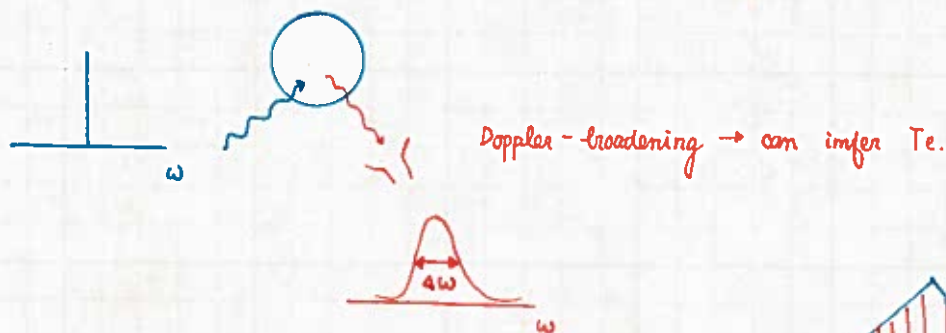
$$T < 20 \text{ eV}$$



- measure current for a given voltage.
- can extract n_e , T_e .

II High density, high temperature.

- Te
 - To get T_e : in magnetized plasma \rightarrow get T_\perp from cyclotron emission.
 - Non-magnetized plasma \rightarrow Use Thomson scattering.



2. T_i

- Charge X-change.



cold neutral



Hot ion.



fast neutral

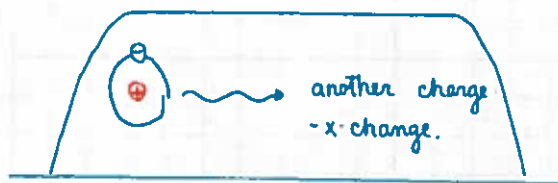


cold ~~neutral~~ ion



Neutral Particle Detector \rightarrow can reconstruct T_i

* limitations of the charge x-change method.



multiple charge x-change
→ optically thick

in high density : need Doppler broadening of line-radiation.

222 Reading Notes.

(i) Plasma parameter (g) (p. 57)

$$g = \frac{1}{n_0 \lambda_D^3}$$

when g is small, \exists many particles in a Debye sphere.

$$\langle PE \rangle \ll \langle KE \rangle \quad (g \ll 1)$$

and the interaction can be neglected, i.e., plasma acts like ideal gas.

2. Gibbs distribution. (N = total # of particles, $N/2$ ions, $N/2$ electrons)

$$D(\vec{x}_1, \vec{x}_2, \dots, \vec{x}_N) = \frac{1}{Z} \exp\left(-\frac{\sum_k \sum_{i>k} W_{ik}}{kT}\right)$$

$$W_{ik} = \frac{q_i q_k}{|\vec{x}_i - \vec{x}_k|} + \phi_{ext}$$

Z = partition fcn.

$$* F_1(\vec{x}_1) = \int D d^3\vec{x}_2 d^3\vec{x}_3 \dots d^3\vec{x}_N$$

where $F(\vec{x}_1)$ = probability density of finding particle 1 @ $\vec{x} = \vec{x}_1$

* special case $W_{ik} \ll kT$

$$F_1(\vec{x}_1) = \frac{1}{V} \rightarrow \text{i.e., } \exists \text{ no preference}$$

$$F_2(\vec{x}_1, \vec{x}_2) = \left\{ 1 + \underline{P_{12}(\vec{x}_1, \vec{x}_2)} \right\} F_1(\vec{x}_1) F_1(\vec{x}_2)$$

$$F_3(\vec{x}_1, \vec{x}_2, \vec{x}_3) = \left\{ 1 + \overset{\text{2-particle correlation function.}}{P_{12}(\vec{x}_1, \vec{x}_2) + P_{12}(\vec{x}_2, \vec{x}_3) + P_{12}(\vec{x}_1, \vec{x}_3)} + T_{123} \right\} F_1(\vec{x}_1) F_1(\vec{x}_2) F_1(\vec{x}_3)$$

$$* T_{123} \ll P_{12} \ll 1$$

3. Correlation fcn. (P_{12})

(i) we assume $P_{12} \gg T_{123}$, i.e., only consider 2 particle correlation.

4 Plasma density.

$$n_{\alpha}(\vec{x}) = \bar{n}_{\alpha} \exp\left(-\frac{q_{\alpha} \phi_P + q_{\alpha} \phi_{\text{ext}} + W_{\text{others}}}{kT}\right) \quad \text{--- } \textcircled{*}$$

* Ideal plasma: if plasma is divided in P pieces (p. 68)

$$n^* = P n_0, \quad q^* = \frac{1}{P} q_0$$

$$m^* = \frac{m_0}{P}, \quad T^* = \frac{T_0}{P}$$

$$\lambda_D^* = \lambda_{D0}, \quad \omega_p^* = \omega_{p0}$$

$$g^* = g_0 \cdot \frac{1}{P}$$

$$P_{12}^* = (P_{12})_0 \cdot \frac{1}{P}$$

$$P^* = P_0$$

$$\text{as } P \rightarrow \infty$$

$$\begin{array}{l} P_{12}^* \rightarrow 0 \\ g^* \rightarrow 0 \\ F \rightarrow F_0 \end{array}$$

→ single-particle properties of plasma

from (*), we have.

$$\nabla^2 \phi = -4\pi \rho_{\text{ext}} - 4\pi \sum_{\alpha} q_{\alpha} n_{\alpha}(\vec{x}) \quad \text{--- } (2)$$

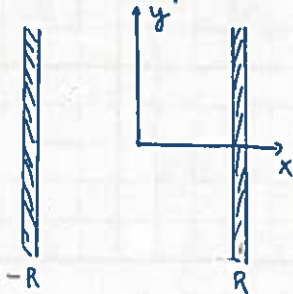
$$\exp\left(-\frac{q\phi}{kT}\right) \approx 1 - \frac{q\phi}{kT}$$

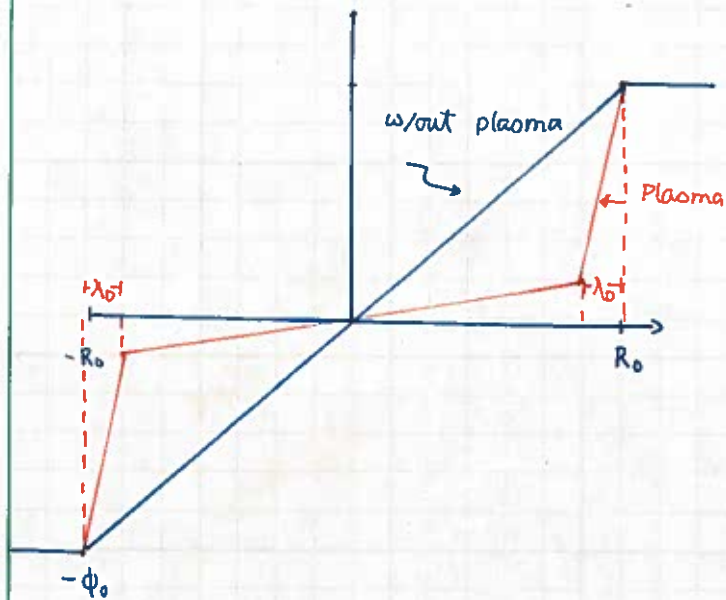
eq. (2) becomes

$$\nabla^2 \phi \approx -4\pi \rho_{\text{ext}} - 4\pi \sum_{\alpha} q_{\alpha} n_{\alpha 0} \left(1 - \frac{q_{\alpha} \phi}{kT_{\alpha}} - \frac{W_{\text{others}}}{kT_{\alpha}}\right)$$

$$\text{neutral plasma} \rightarrow \sum_{\alpha} q_{\alpha} n_{\alpha 0} = 0$$

Plasma Capacitor.





(p. 76) The ratio of electrostatic energy to thermal energy is.

$$\frac{W(\text{Coulomb})}{n_0 k T} = \frac{1}{n_0 \lambda_D^3} = g$$

$g = 0 \rightarrow$ ideal gas...

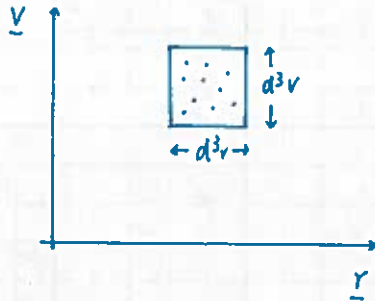
Ch. 3.

Liouville's th.

$$\frac{\partial F}{\partial t} + \sum_i \left(\frac{\partial F}{\partial \vec{x}_i} \cdot \vec{v}_i + \frac{\partial F}{\partial \vec{v}_i} \cdot \vec{a}_i^T \right) = 0$$

$$i = 1 \dots N \quad (\# \text{ of particles})$$

Today.

μ-scopic quantities.1. Most basic quantity. : $f_j(\underline{r}, \underline{v}; t)$
~ species.

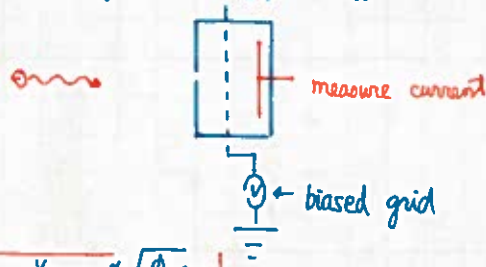
$$dn_j = d^3v d^3r f_j(\vec{v}, \vec{r}, t)$$

• density: $n_j(\vec{r}, t) = \int d^3v f_j(\vec{v}, \vec{r}, t) \rightarrow$ in $\frac{\# \text{ of particles}}{\text{cm}^3}$
(velocity + temperature can be obtained in a similar manner)

$$f \rightarrow f_{\text{maxwell}} = \frac{n_0}{(2\pi \bar{v}^2)^{3/2}} \exp\left(-\frac{v^2}{2\bar{v}^2}\right) \rightarrow \text{Maxwell's distribution.}$$

Dr. M's convention.

$$\bar{v}^2 = \frac{T}{m} \rightarrow \text{temp. defined.}$$

2. Exp. measurements.• low temp, low density. \rightarrow energy analyzer. \rightarrow a 1-D device.

$$I \sim \int_0^{v_{\text{max}} \propto \sqrt{\phi_{\text{bias}}}} v f(v) dv$$

$$\frac{dI}{dv_{\text{max}}} \approx v f(v) \rightarrow \text{Energy Analyzer. Formula.}$$

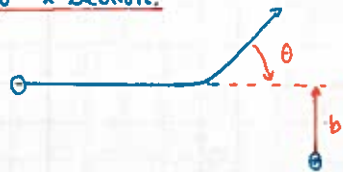
• for high n , high T . \rightarrow from spectrum of δ 's, try to find (infer) $f(v)$

- for cyclotrons (magnetized plasma)



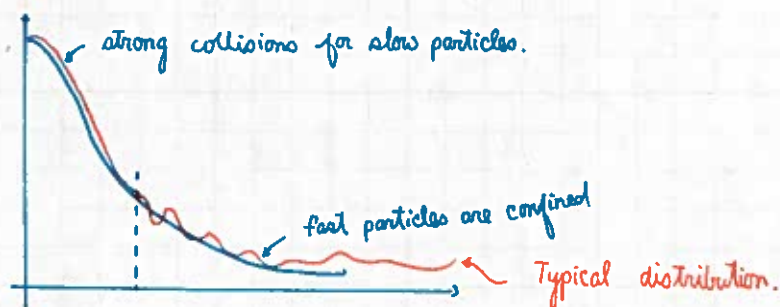
- charge x-change (line of sight measurement).

3. Coulomb x-section.

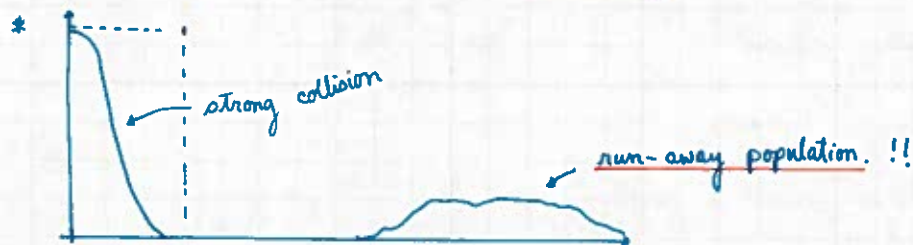


$$\sigma_{\text{coll}} (\theta = 90^\circ) = \pi \left(\frac{e^2}{m v^2} \right)^2 \sim \frac{1}{v^4} \quad (\text{Ask about Debye Shielding...})$$

$$\begin{aligned} \therefore \nu_{\text{coll}} &\sim v \sigma_{\text{coll}} & (* \nu_{\text{coll}} = n v \sigma_{\text{coll}}) \\ &\sim \frac{1}{v^3} & \leftarrow \text{for Coulomb collision.} \end{aligned}$$



\therefore if one applies a force \rightarrow can develop run-aways.



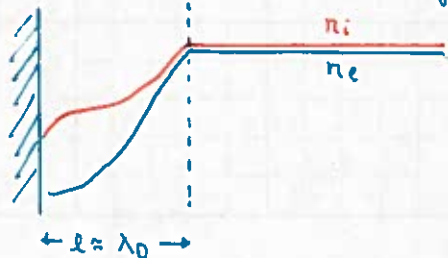
- radiation.
- carries most of the current.
- can be easily dumped by the system \rightarrow it suffers very little collision.

4. Charge neutralization

\therefore if a plasma is made by ionizing neutrals



$n_e \approx n_i$ \rightarrow except @ interface.

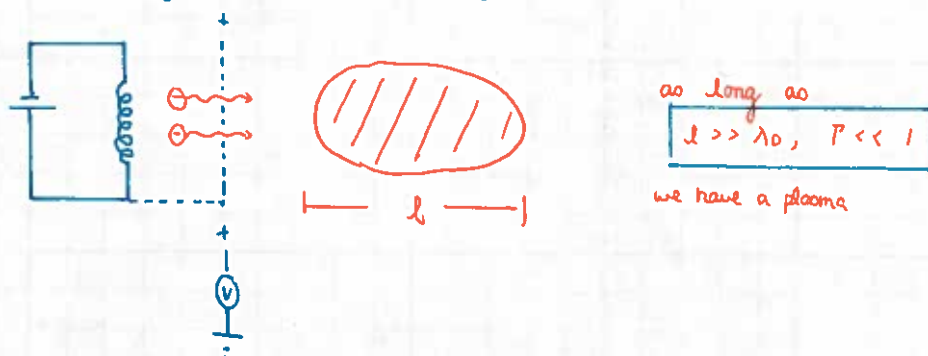


- potential develops such that

$$T_i = T_e$$

hence, sheath occurs (hi Ned! @)

- ii. Plasma can be made of a single species.
e.g. emitting electron from a filament.



5. Collective parameters.

i. $\omega_{pj}^2 = \frac{4\pi e^2 n_j}{m_j} \rightarrow$ represents collective oscillation.

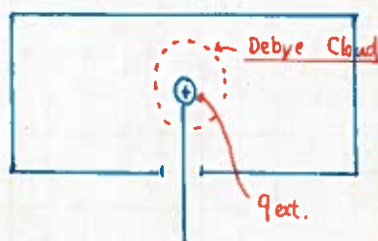


- ω_{pe} : electron plasma frequency. \rightarrow highest frequency in unmagnetized plasma
- ω_{pi} : ion plasma frequency \rightarrow] no oscillation @ ω_{pi} . because when we try to excite an ion plasma wave, the electrons will short out the oscillation.

ii. Debye shielding.

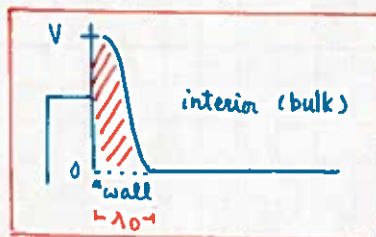
$$\lambda_{D,j} = \frac{\overline{V_j^2} \frac{T}{m}}{\omega_{p,j}^2 \frac{e^2 n \cdot 4\pi}{m}} \rightarrow \text{mass independent.}$$

6. Debye shielding.



$$\phi(r) = \frac{q_{ext}}{r} e^{-r/\lambda_D}$$

Yukawa Type



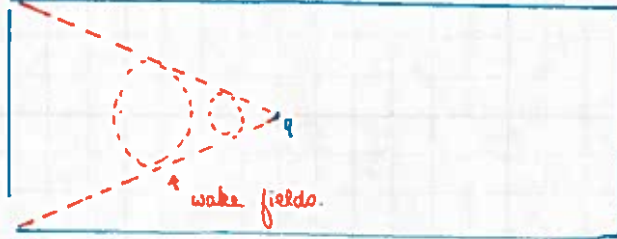
7. Dressed Test Particles (New stuff)

- modify plasma properties
- Test Particle. - its property is not affected by the plasma

(i) $V = 0$ 

→ fully dressed test charge.

fully dressed test charge.

(ii) $V > \bar{v}$ if $v_{\text{test}} > \bar{v} \rightarrow$ it leaves a wake field. (Cherenkov radiation)

$$\vec{k} \cdot \vec{v} = \omega_p$$

• In a plasma, each particle can be considered as a test charge.6. Landau damping

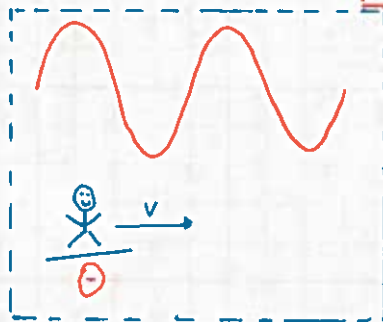
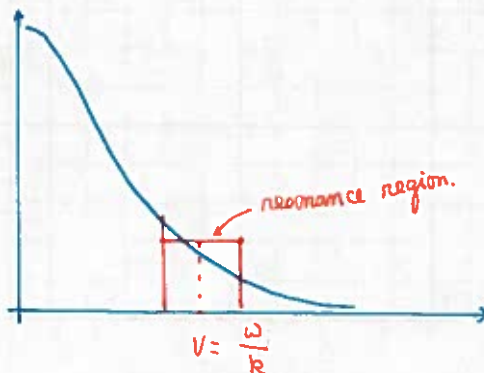
(i) Given: a travelling wave.

$$E(x, t) \sim \cos(kx - \omega t)$$

(ii) in a particle frame.

$$x \rightarrow v(t - t_0)$$

$$\begin{aligned} E'(x, t) &\rightarrow \cos(kvt - kv t_0 - \omega t) \\ &\rightarrow \cos((kv - \omega)t - \underline{kv t_0}) \end{aligned}$$

if $kv = \omega$ (resonance condition), we can have strong acceleration and/or deceleration (depending on phase) @ $v = \omega/k$ 

$$E(\vec{x}, t) \sim E_0 e^{-i\gamma t} \cos(\vec{k} \cdot \vec{x} - \omega t)$$

hence. Equilibrium in plasma is a 2-step process.

1. Cherenkov radiation by a dressed charge.
2. Landau Damping by another charge.

* Detuning occurs when $\Delta V \sim \frac{v_L}{k}$

Today.

Waves in unmagnetized plasma.

1. Given.

$$E \sim \exp(i(\vec{k} \cdot \vec{r} - \omega t))$$

• Linear theory gives....

$$k = k(\omega) \rightarrow \text{dispersion relation.}$$

$$\bullet \text{ EM waves: } \omega^2 = \omega_p^2 + k^2 c^2$$

• ES waves.

- Langmuir wave.

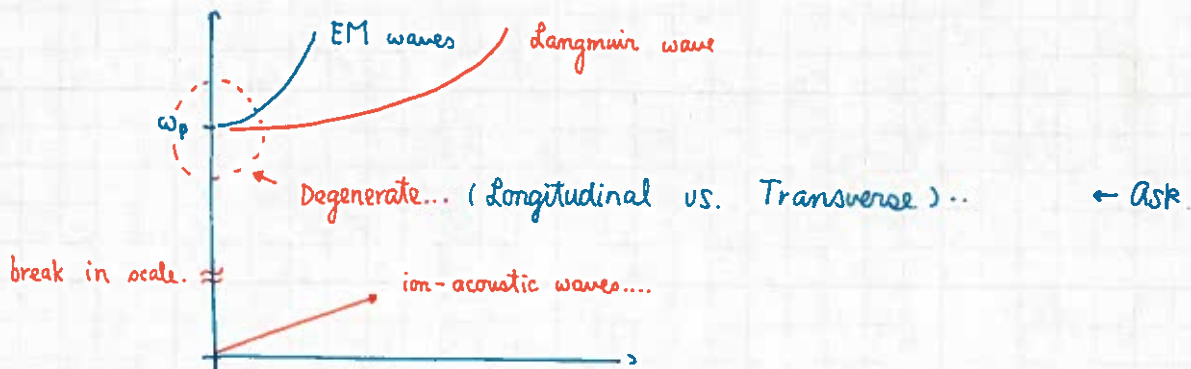
$$\omega^2 \approx \omega_p^2 + 3 k^2 \bar{v}_e^2$$

- Ion acoustic waves.

$$\omega^2 = k^2 c_s^2$$

sound velocity.

$$c_s = \sqrt{\frac{T_e}{M}}$$



$$\bullet \text{ Langmuir wave regime: } \omega < \omega_{pe} \cdot 1.5$$

$$\bullet \text{ ion acoustic wave: } \omega < 0.2 \omega_{pi}$$

2 Roles of the waves.

- Ion Acoustic: Equalize pressure difference....

- Langmuir Wave: Equalize distribution, $f_e(v)$

- EM wave: Communicate w/outside.

• @ $\omega = \omega_p \rightarrow$ degeneracy \rightarrow energy transfer between modes.

3 for an EM wave.

$$k = \sqrt{\frac{\omega^2 - \omega_p^2}{c^2}}$$

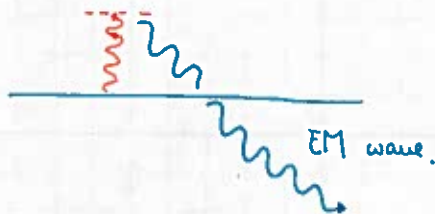
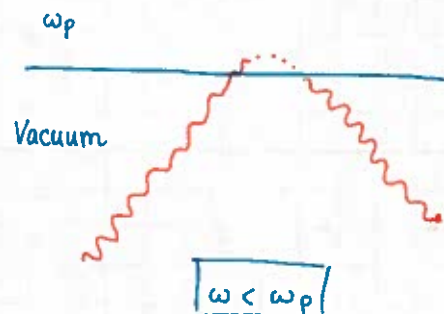
$$\bullet \omega > \omega_p \rightarrow \text{Propagation.}$$

$$\bullet \omega = \omega_p \rightarrow \text{cut-off}$$

$$\bullet \omega < \omega_p \rightarrow \text{Evanescent.}$$

Ex: Ionosphere, again...

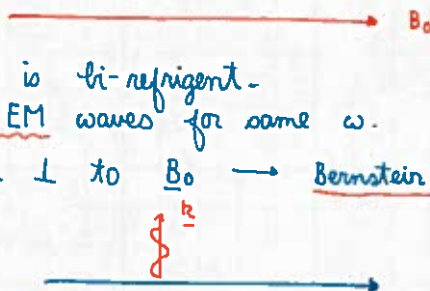
- if one works out this problem carefully, we can see the mode conversion process happening @ $\omega = \omega_p$
- Conversely, one can generate an ES (Langmuir) wave in a plasma and watch EM emission on the ground.



- Cerenkov condition: $\underline{k} \cdot \underline{v} = \omega$ but $v < c$.
→ Can excite ES waves but not EM waves.

4. For magnetized plasmas → a lot more structures.

- Medium is bi-refrigent.
→ 2 EM waves for same ω .
- Propagation \perp to \underline{B}_0 → Bernstein modes.



5. Plasma Production.

- Develop conditions which rate of ionization exceeds rate of recombination.

- Need to confine plasma for a time. $\tau > t_{\text{physics}}$

- Particle interaction w/ Langmuir wave: $t_{\text{physics}} \approx N \cdot \frac{1}{\omega_p}$
- Plasma transport: $t_{\text{physics}} \approx N \cdot \frac{1}{\omega_{i.a.}}$

i. Ionization. → Need energy larger than ionization potential

- Hydrogen: 13.6 eV
- Argon: 15.7 eV
- Cesium: 3.9 eV → low for Alkali metals

- Methods of ionization
 - electron impact,



- Photo-ionization.



- for electron impact. the creation rate.

$$\frac{d}{dt} n_- = n_0 n_- P_{oe}$$

↑
free electron population

P_{oe} = probability of ionization.

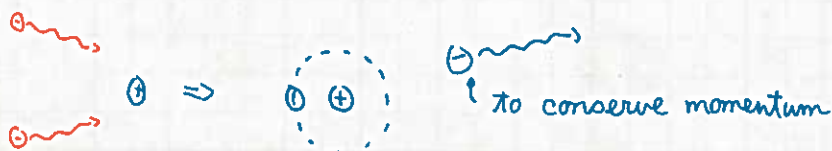
- for photo-ionization, the rate can be found as followed.

$$\frac{d}{dt} n_- = n_0 \phi P_{oi} \rightarrow \text{e.g.: ionization in ionosphere.}$$

P_{oi} = photo-ionization cross section.

- Recombination \rightarrow Several channels, but the 2 most prominent ones are the ~~most prominent~~ a time-reversed creation process.

- 3-body recombination.



$$\frac{d}{dt} n_- \approx -n_-^2 n_+ P_{eo}$$

$$n_+ \approx n_-$$

$$\rightarrow \frac{d}{dt} n_- = -n_-^3 P_{eo} \propto n_-^3$$

hence, we'd expect the equilibrium density to be low....

• Radioactive Recombination.

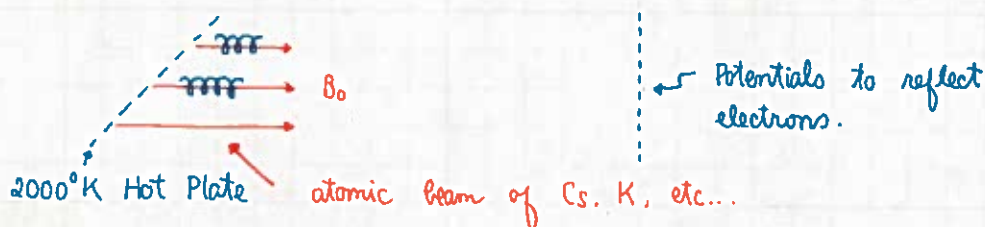


$$\frac{d}{dt} n_- = n_+ n_- P_{\gamma 0} = \mathcal{O}(n_-^2)$$

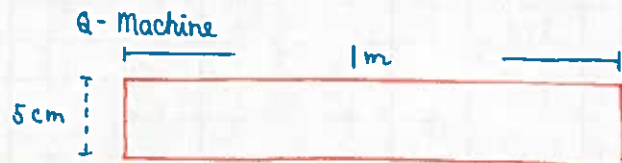
Recombination Rate...

• Confinement time → BIG PROBLEM.

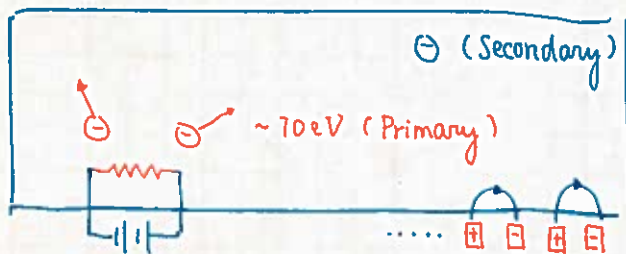
- Use of Alkali Metal Vapor → Low Ionization Pot (4 eV)
(Rynn + D'Angelo, → Q-machine")



- Fully ionized. (Good)
- Low temperatures (determined by melting point of hot plate)
- $T_i \approx T_e$ (Good)
- small plasma → determined by the size of the beam.



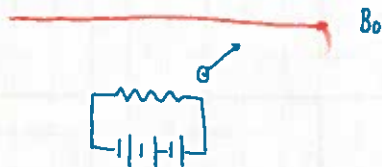
- Mackenzie Bucket (UCLA) → 180E Lab.



Cheap Permanent Magnets Around the machine.

- Surface $B \sim 2 \text{ KGauss}$
- $n \sim 10^{10} \text{ cm}^{-3}$
- $T_e \sim 2 \sim 3 \text{ eV}$
- $T_i \sim 0.2 \text{ eV}$

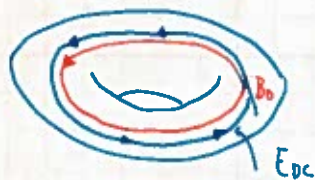
Side note:



In a highly magnetized plasma, e^- cannot go across B_0 , then, the filament will acquire a negative potential and stop injecting.

$$\tau_{\text{confinement}} = \frac{L}{c_s} = \frac{10^2}{10^5} \approx 10^{-3} \text{ sec.} \quad (* c_s = 10^5 \text{ cm/sec})$$

• Tokamak



- Toroidal Device
- Strong mag field. ($\sim 10 \sim 50 \text{ K Gauss}$)
- $E_{DC} \rightarrow$ DC electric field.
 - to create secondary electrons

- $\tau_{\text{confinement}} = 1 \text{ sec.}$

- Temperature $\sim 5 \text{ KeV}$

- Density $\sim 10^{14} \text{ cm}^{-3}$

- Volume $\sim 20 \text{ m}^3$

• fusion parameter $n\tau \approx 10^{14}$ (thermo-nuclear range)

Today

Basic Equations of Plasma.

1. We can have an universal equation

$$\frac{D}{Dt} F = 0 \rightarrow \text{is all and is useless.}$$

too general.

- ask a question \rightarrow identify τ, l
 τ = time scale
 l = length scale.

Options

- Collisionless \rightarrow Vlasov Equation
- Collisional + Kinetic \rightarrow Fokker Planck
- Fluid description.
- MHD.

2 Root of all evil \rightarrow Coulomb force

$$|F_{1,2}| = \frac{q_1 q_2}{r^2}$$

(Range is too long)

- What is a collision in a plasma?

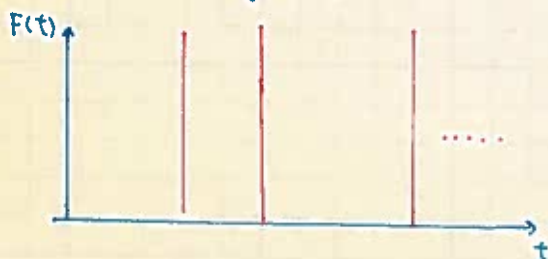
- hard sphere

$$F(r) = \begin{cases} 0 & \text{if } \Delta r \geq 2a \\ \infty & \text{if } \Delta r < 2a \end{cases}$$

Test Particle



Force experience by the particle



In this case, we have.

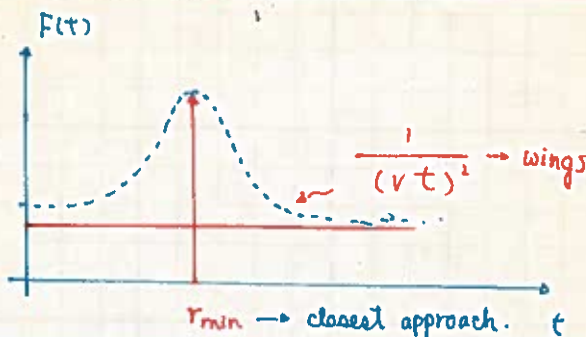
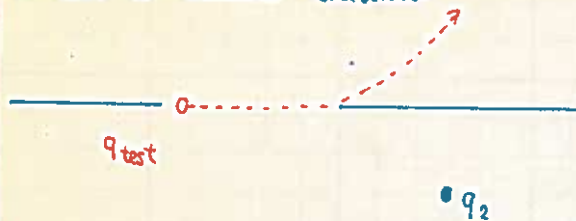
(i) $\sigma = \pi a^2$

(ii) $\nu_c = v n \sigma = [n \pi a^2] v$

$\nu_c \propto v$

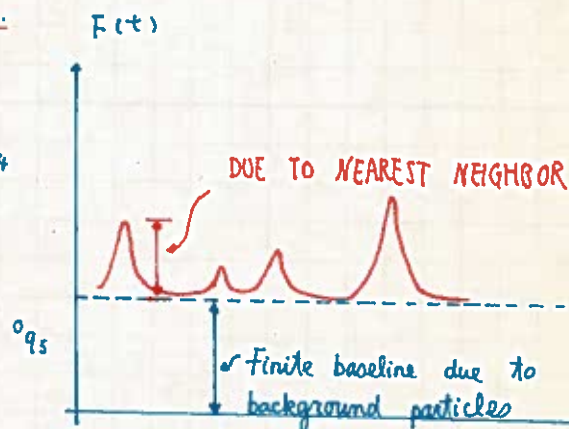
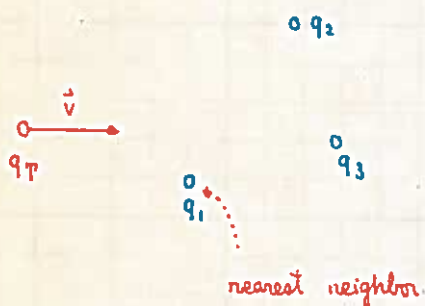
DIFFERENT THAN COULOMB COLLISION.

In a Coulomb Collision.



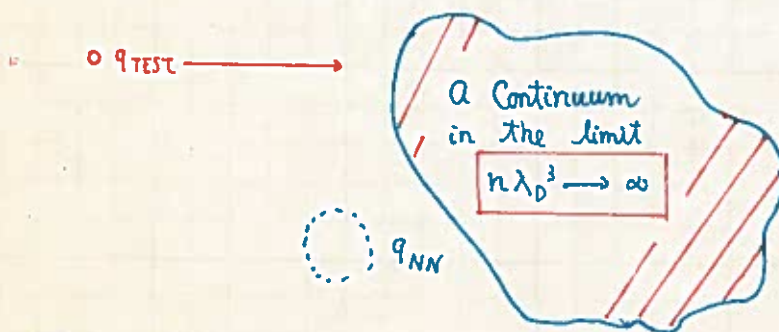
In a plasma

- Never any binary interaction.



- As # of particles increases, the baseline force dominates !!!

- Ideal limit



Electric Field.

$$\vec{E} = \vec{E}_{NN} + \langle \vec{E} \rangle$$

Continuum \rightarrow smooth charge distribution

$$\nabla \cdot \langle \vec{E} \rangle = 4\pi \rho$$