

# Long Title

T. Taro Shimizu<sup>1</sup><sup>★</sup>, Author 2<sup>2</sup>, Author 3<sup>3</sup>

<sup>1</sup>*Max-Planck-Institut für extraterrestrische Physik, Postfach 1312, 85741, Garching, Germany*

<sup>2</sup>*Institution 2*

<sup>3</sup>*Institution 3*

Accepted XXX. Received YYY; in original form ZZZ

## ABSTRACT

**Key words:** galaxies: active – galaxies: Seyfert – infrared: galaxies – galaxies: star formation – galaxies: evolution

## 1 INTRODUCTION

## 2 SECTION 2

## 3 SECTION 3

## 4 DISCUSSION

## 5 SUMMARY AND CONCLUSIONS

## ACKNOWLEDGEMENTS

This research has made use of the NASA/IPAC Extragalactic Database (NED), which is operated by the Jet Propulsion Laboratory, California Institute of Technology, under contract with the National Aeronautics and Space Administration. This research made use of Astropy, a community-developed core Python package for Astronomy ([Astropy Collaboration et al. 2013](#)). Figures in this publication were created with the Python package MATPLOTLIB ([Hunter 2007](#)).

## REFERENCES

Astropy Collaboration et al., 2013, [A&A](#), **558**, A33  
Hunter J. D., 2007, Computing In Science & Engineering, 9, 90

## APPENDIX A: APPENDIX 1

## APPENDIX B: APPENDIX 2

This paper has been typeset from a T<sub>E</sub>X/L<sup>A</sup>T<sub>E</sub>X file prepared by the author.

★ Email: shimizu@mpe.mpg.de

**Table 1.** Example Table Title

$F_1/F_2$	$m$	$b$	$\sigma_{\text{int}}$	$\rho$	75th %tile (%tile for HRS)	50th %tile (%tile for HRS)	25th %tile (%tile for HRS)
12/70	$0.67 \pm 0.04$	$1.02 \pm 0.03$	$0.016 \pm 0.004$	$0.89 \pm 0.02$	-1.23 (90)	-0.97 (50)	-0.66 (21)
12/160	$0.58 \pm 0.02$	$1.05 \pm 0.03$	$0.002 \pm 0.002$	$0.99 \pm 0.01$	-1.51 (31)	-1.21 (5.2)	-0.85 (2.2)
12/250	$0.51 \pm 0.03$	$0.84 \pm 0.03$	$0.019 \pm 0.004$	$0.87 \pm 0.03$	-1.22 (28)	-0.91 (5.2)	-0.57 (1.5)
22/70	$0.79 \pm 0.03$	$0.86 \pm 0.02$	$0.001 \pm 0.001$	$0.99 \pm 0.01$	-0.89 (32)	-0.60 (16)	-0.32 (12)
22/160	$0.58 \pm 0.02$	$0.85 \pm 0.02$	$0.0004 \pm 0.0005$	$0.997 \pm 0.003$	-1.14 (5.6)	0.83 (1.2)	-0.45 (0)
22/250	$0.49 \pm 0.03$	$0.65 \pm 0.02$	$0.016 \pm 0.004$	$0.90 \pm 0.02$	0.87 (6.8)	-0.51 (0.02)	-0.18 (0)