[Team Whitespace 2024]

Seismic Detection Across the Solar System

The Challenge:

- Limited transmission bandwidth demands efficient data selection.
- Real-world seismic data is inherently noisy.
- Classical STA/LTA methods require extensive tuning for different planetary environments.

Our Solution: A Simple Vision-Inspired Seismic Detection and Package

Tuning STA/LTA:

- Optimizes the STA/LTA method for lunar and mars missions
- High recall version: 84% (but #FP ~= 40x#GT)
- FP-fewer version: 76% (#FP ~= 13x#GT)
- If we would like to suppress FP really
 - Method STA/LTA+StaticGaussian
 - Achieves acceptable recall of significant seismic events while minimizing false detections.
 - Recall 55% (#FP ~= 1.5#GT)
- Initial study using NASA data demonstrates the effectiveness of:
 - Tuned STA/LTA, Dynamic Gaussian Modeling, Static Gaussian Modeling
 - o Finally suppress 93% FP to obtain acceptable method: STA/LTA+StaticGaussian
- Open-Source Collaboration: We provide a complete script to empower amateurs and researchers to
 explore and contribute to this cutting-edge technology.

Vision-Inspired? Background Subtraction & Object Detection Are Like Seismic Detection

 Similar Approach: In vision, we separate moving objects from the background, just like detecting quakes from background noise in seismic data.

Why Use a Background Model?

- Object Detection Requires Specific Knowledge: We'd need to define guake characteristics, which we don't have.
- **General Approach**: Background models allow us to detect anomalies without specific quake properties.
- Amateurs in Seismology: We're still learning about quakes, so we prefer a simpler, flexible method.

Why Not Use Deep Learning?

- Limited Seismology Expertise: We use DL in vision, but applying it here without enough knowledge could lead to bad results.
- **Noise Modeling**: DL might struggle without proper noise understanding.

Evolution of Background Subtraction

- Static Gaussian: Fixed background.
- **Dynamic Gaussian**: Adapts to changes.
- Mixture of Gaussians: Handles sudden shifts.

Experimental results show Static Gaussian could help STA/LTA to suppress large amount false positive

Enhancing Seismic Event Detection with STA/LTA

- STA/LTA: A Classical Approach
 - Widely used in seismology for detecting events.
 - Relies on comparing short-term averages to long-term averages of signal amplitude.
 - Requires careful tuning of parameters:
 - `sta_len` (Short-Term Average length)
 - `lta_len` (Long-Term Average length)
 - `thr_on` (Threshold for triggering detection)
 - `thr_off` (Threshold for ending detection)
- Reference Settings: In NASA packages, a common starting point is:

```
o `sta_len = 120`, `lta_len = 600`, `thr_on = 4.0`, `thr_off = 1.5`.
```

Our Package for Optimized STA/LTA Performance

- Flexibility: Separate training and validation datasets.
- Evaluation: Function to calculate recall and false positive rate.
- Parameter Tuning: Use the training data to explore different parameter settings through grid search (`sta_len`, `lta_len`, `thr_on`, `thr_off`).

- Training data is split into non-overlapping train and val
- We use train part to pick three candidate settings
 Candidate 1: has too many false positive
 Candidate 2: is another setting with fewer false positive
- And try to use the same conf for lunar and mars

 | Description | De

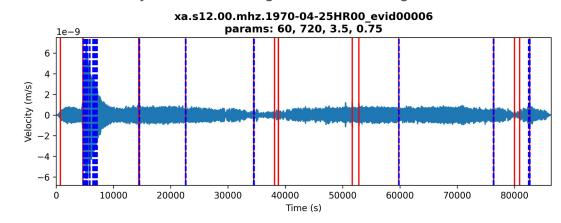
| | the same con | i ioi iuiia | i and ma | 10 | | |
|------------------------|---------------------|-------------|-------------|---------|--------------------------|---|
| lunar | Recall(train) FP(tr | rain) Rec | all(val) FP | (val) | | |
| [120, 600, 4.0, 1.5] | 60.53 | 259 | 47.37 | 422 ref | ference setting in ipynb | 1 |
| | | | | hi | ighest recall in train | |
| [60, 720, 3.5, 0.75] | 97.37 | 987 | 84.21 | 1352 ca | andidate 1 | |
| [160, 680, 3.25, 0.75] | 68.42 | 301 | 63.16 | 752 | | |
| [100, 760, 3.75, 1.0] | 78.95 | 488 | 76.32 | 488 ca | ndidate 2 | |
| | | | | | | |

| mars | Recall(train) FP(train) | Recall | (val) FP(val) | |
|----------------------------------|-------------------------|--------------|----------------|---------------------------------|
| [120, 600, 4.0, 1.5] | 0 | 0 | 0 | 0 reference setting in ipynb |
| [60, 400, 3.25, 0.75] | 100 | 2 | 100 | 1 highest recall in mars train |
| [60, 400, 4.75, 0.75] | 100 | θ | θ | θ |
| [60, 720, 3.5, 0.75] | 100 | 1 | 100 | 0 highest recall in lunar train |
| [160, 680, 3.25, 0.75] | 100 | 0 | 100 | 0 |
| [100, 760, 3.75, 1.0] | 100 | 0 | 100 | 1 candidate 2 from lunar train |
| | | | | |

Applying background modeling

- Based on train part we generate two candidate settings
 - o [60, 720, 3.5, 0.75]
 - o [100, 760, 3.75, 1.0]
- We would like to further improve candidate 1 by using background moding to suppress FP
 - We found noise could not be stable, hence need dynamic or static gaussian modeling

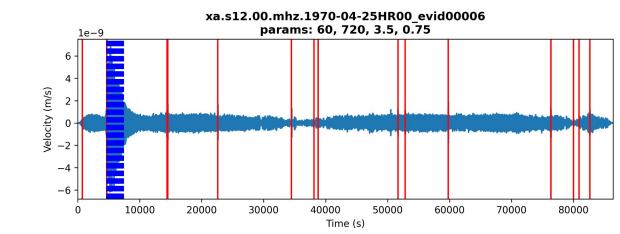
STA/LTA (red solid line)
DynamicGaussian
(blue dot line)

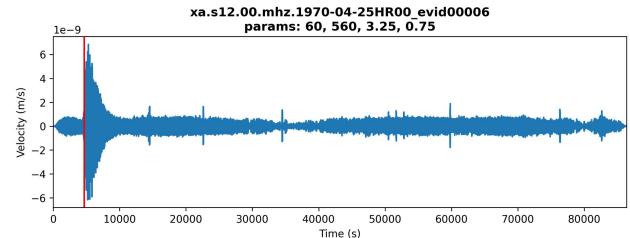


STA/LTA (red solid line) STGaussian (blue dot line)

STGaussian: only one static gaussian for background noise

Final: STA/LTA+STGaussian





Final val results

| lunar | Recall(train) | FP(train) | Recall(val) | FP(val) | | |
|------------------------|---------------|-----------|-------------|---------|----------------------------|--|
| [120, 600, 4.0, 1.5] | 60.53 | 259 | 47.37 | 422 | reference setting in ipynb | |
| | | | | | highest recall in train | |
| [60, 720, 3.5, 0.75] | 97.37 | 987 | 84.21 | 1352 | candidate 1 | |
| [160, 680, 3.25, 0.75] | 68.42 | 301 | 63.16 | 752 | | |
| [100, 760, 3.75, 1.0] | 78.95 | 488 | 76.32 | 488 | candidate 2 | |
| | | | | | | |
| [60, 720, 3.5, 0.75] | | | 55.26 | 58 | FP suppressed | |

Thanks

https://github.com/tungtylee/seismic-viewer