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## 1 Introduction

## 2 High level goals

Juno is a NASA spacecraft orbiting Jupiter. Built by Lockheed Martin and operated by NASA, it was launched by Atlas V551 on 5<sup>th</sup> August 2011. After 5 years, during which many manoeuvres occurred, including an Earth flyby, Juno entered a polar orbit around Jupiter and started its observation, which lasts to this day. Its aim is to study the planet to understand its composition and evolution, analyzing its gravitational and magnetic fields and its atmosphere dynamics. The mission should have ended in 2017, but is still ongoing and it will end in a deorbit destroying the spacecraft into the planet's atmosphere to avoid contaminating the environment.

Through an analysis of the mission and payload, the main goals of the mission can be highlighted.

LA TABELLA CON I GOALS O I SCIENTIFIC OBJECTIVES? CHE PERO SONO PRATICAMENTE UGUALI PER TUTTO

### 1. How did Jupiter form and influence the solar system?

Since Jupiter is the biggest planet of the solar system, it has influenced the formation of all other planets. Its composition has remained unchanged ever since, making it like a time capsule: understanding how and where it has formed can give knowledge on Earth and the whole solar system's origin, evolution and characteristics.

### 2. What's Jupiter's deep structure?

One important aspect of the mission is the analysis of Jupiter's deep structure through the measurement of radiations, magnetic and gravitational fields. This allows to comprehend whether or not the planet has a solid nucleus, if so how large it is, and to analyze the hypothesized layer of metallic hydrogen, compressed so much that it loses its electrons creating a conducting layer. Moreover, Juno will possibly reveal if Jupiter is rotating as a solid body or if the rotating interior is made up of concentric cylinders.

### 3. What's the structure of Jupiter's atmosphere?

One of Juno's goals is to study the composition and dynamics of Jupiter's atmosphere, composed by stripes and dots made of different gasses and vapors, including water, whose percentage has to be defined. A significant aspect of the analysis is the great red spot, a swirling mass of gas bigger than Earth, which resembles a hurricane but is very different in the way it works. The movement of stripes and dots is dictated by the weather, characterized by lightning and thunderstorms, which are observed by Juno.

### 4. What do the poles look like and what are the physical processes generating auroras?

The observation of Jupiter's poles had never been possible before because of the absence of a polar orbiting spacecraft. Thanks to Juno we have been able to study Jupiter's auroras, representative of the interaction between charged particles and the atmosphere. This can allow a better understanding of the atmospheric composition and the magnetic field's structure and extension.

Guiding questions	Science objectives
How did Jupiter form and influence the solar system?	Map Jupiter's composition and image the whole planet
What's Jupiter's deep structure?	Analyze gravitational field, magnetic field and radiations
What's the structure of Jupiter's atmosphere?	Screen the planet, analyze atmospheric processes and emissions
What do the poles look like and what are the physical processes generating auroras?	Image Jupiter, study interactions between atmosphere and magnetic field

Table 1: Guiding questions for high level goals

## 3 Mission drivers

Being JUNO an interplanetary mission starting from a distance of around 1 AU, with a final nominal distance from the Sun of 5.2 AU, and operating in an highly radiation intense environment, the following drivers have been identified:

### 1. Using proven technologies THE JUNO Mission

The total program is financed with 1.1 Billion \$ for 74 months from the launch date and includes development of the S/C, science instruments, launch services, mission operations, science processing and relay support. The simplicity and the need of proven technologies was thus fundamental. The S/C is mainly maintained stable during the maneuvers thanks to its spin, raised to 5 RPM from 2, nominal condition during science operations, reducing the need of active stabilization methods.

### 2. Providing enough electricity during the duration of the mission

The journey of JUNO is long and passes through different regions of the solar system. Solar panels were chosen to provide electric energy across the mission over a nuclear source, since it has been decided that it was better to advance technology of solar cells rather than developing a new reactor. It is the first S/C to operate with solar panel at such distance from the Sun. The system needed is thus oversized at 1 AU: the solar radiation on Jupiter is in fact up to 96% lower than on Earth. Furthermore the operations are scheduled to begin around 5 years into the mission, so degradation of the solar cells must be taken into account. The final design consists in 3 solar arrays, 2 are 9 by 2.65 *m* each, meanwhile the third one is shorter at about and only 2 *m* wide, resulting in about 60 *m*<sup>2</sup> and granting a maximum power of 14 *KW* around Earth and up to 500 *W* around Jupiter. This configurations is needed to mount the MAG faraway from the electronics and store everything correctly inside the fairing. Its solar pannels are mounted on the side of exagonal body of the S/C at 120° one by the other and are deployed after second stage separation. The spasecraft is spinning at around 1.4 *RPM* during this phase and the deployment of the solar arrays slows it down. A crucial element must be considered during the cruise phases: nominally the high gain antenna, that is mounted with the same positive direction of the solar arrays, is doing Earth pointing to communicate data but, at some point during the cruise phase, due to the proximity of the Sun, thermal requirments dictate a change in attitude and thus the spacecraft won't be pointing directly the Sun. Moreover, since the fly-by around Earth is done to gain  $\Delta V$ , the S/C will be in an eclipse for around 10 minutes: attention must be paid in battery sizing. Two litium-ions battery of 55 *Ah* each are present to make sure power is always provided. The nominal polar orbit around Jupiter allows Sun pointing during the majority of nominal Science Operation phase. INSERIRE RIFERIMENTI: VIDEO, SLIDES LAVAGNA, JUNO Mission, batteries

### 3. Shielding the instruments from the harsh enviornment of Jupiter

To accomplish its goals, JUNO will need to cross the Jupiter radiation belts: a heavy shieldind structure is needed. The magnetosphere represents a great challenge for JUNO: the value of the magnetic field measured at its perijove is 776  $\mu T$ , 50 % higher than expected. The main issue with JUNO orbit is represented by the ionizin particles present in the belt around Jupiter: with measured value up to of tens and hundreds of MeV ions located between 2 and 4  $R_J$ , order of GeV were expected under 2  $R_J$ , where JUNO should pass through to reach a lower altitude, thanks to its highly elliptical orbit, where radiations are lower, to perform science. The voutl in which all the electronics is preserved is cubed shaped and it is made of 1 *cm* thick titanium alloy, 144 *Kg* in total. The top deck of Juno is planned to receive a radiation dose of 22 *Mrad*. Moreover, startrackers are also heavily shielded.

### 4. Mantaining communication during the journey and the science operations

The attitude of the S/C is defined in a way to point Earth during most of the cruise and science operations. This configurations, given the distance from the Sun and the Earth, grants also a sufficient inclinations of the solar pannels with respect to the Sun to provide enought electric power. The ground equipment used by JUNO is NASA's DSN.

## 4 Functional analysis

Functional analysis is performed in order to identify the functionalities that the spacecraft must perform during the mission. The identified functionalities are schematized in [Figure 1](#). [1]



Figure 1: Functional analysis for Juno mission

## 5 Main mission phases

The JUNO mission was divided into five phases: Launch, Cruise, Insertion, Science Operations and Deorbit.

### 5.1 Launch

The launch took place on August 5 2011, 16:25 UTC from Cape Carnival, in Florida. The selected launcher was Atlas V551, which injected the spacecraft in the desired trajectory. This phase began with the ignition of the booster engine system, and ended with spacecraft separation from the upper stage, about 54 minutes after. The solar panels deployment was performed about five minutes after the spacecraft separation, and it took approximately five minutes.

### 5.2 Cruise

At the beginning of this phase, the spacecraft was injected in an interplanetary trajectory with the aim of reaching Jupiter. The cruise had a duration of about five years, during which two deep space manoeuvres and an Earth fly-by were performed. All manoeuvres will be better described in [section 8](#). This phase also included instruments testing and verification, to ensure they were functioning properly and ready for the usage during the mission. This phase was also characterized by initial science observations of Jupiter.

### 5.3 Insertion

This phase began four days before the start of orbit insertion manoeuvre and ended one hour after the start of the orbit insertion manoeuvre. The latter occurred at closest approach to Jupiter and slowed the spacecraft enough to let it be captured by Jupiter in a 53 days period orbit. The Jupiter orbit insertion burn was performed by the main engine, and it lasted 30 minutes. After the burn, the spacecraft was in a polar orbit around Jupiter. The 53-days orbit provided substantial propellant savings with respect to the direct insertion in the operational orbit.

### 5.4 Science Operations

Before starting with the science operations, it was planned to reduce the period of the orbit from 53 to 14 days. However, due to problems with helium valves in the propulsion system, it was decided to remain in the 53-day orbit. The Juno polar, highly eccentric orbit, was designed to facilitate the close-in measurements and to minimize the time spent in the Jupiter radiation belts.

### 5.5 Deorbit

The Deorbit phase occurs during the final orbit of the mission. The latter was designed to satisfy NASA's planetary protection requirements and ensure that JUNO doesn't impact any of Jupiter's moons. A deorbit burn will be performed, placing the spacecraft on a trajectory towards Jupiter surface. JUNO was not designed to operate in the atmosphere and will burn up.

## 6 ConOps

The mission Conceptual Operations are summarized in Figure 2.

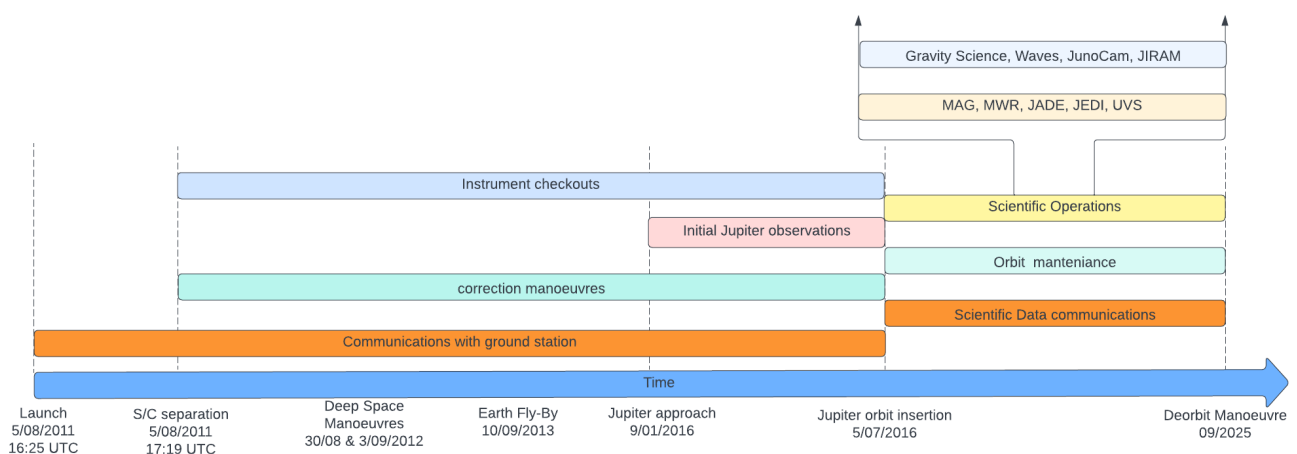


Figure 2: Conceptual Operations, time not in scale

## 7 Payload analysis

### 7.1 Instruments overview

As previously described in section 2 the mission scientific goals are quite numerous and diverse. Thus, to achieve all of them the payload consists of several instruments, nine to be precise, covering a wide spectrum of experimentations. In its entirety the payload has a mass of around 174 kg and consumes approximately 125 W excluding the Gravity science experiment. Here we have a brief overview of all the singular instruments where, unless otherwise specified, only the sensors are mounted on the exterior of the spacecraft while all the relevant electronics are located inside the radiation vault.

- **Magnetometer (MAG):** As the name implies its objective is to accurately measure Jupiter's magnetic field, achieved by employing a fluxgate magnetometer, a scalar helium magnetometer and star cameras. All the sensors are mounted on the magnetometer boom located at the end of one of the solar array wings to reduce the interference from the spacecraft itself. Even then the presence of two magnetometers allows to subtract this contribution from the measurement.
- **Microwave Radiometer (MWR):** It consists of six antennas which measure six different frequencies (600 MHz, 1.2 GHz, 2.4 GHz, 4.8 GHz, 9.6 GHz and 22 GHz) in order to investigate the Jovian atmosphere below the visible external layer. A key objective of this analysis is also the determination of the abundance of water inside the planet.

The antennas are mounted on two sides of the exagonal prism that constitutes the main body of the spacecraft relying on its spin to survey Jupiter.

- **Gravity science:** It's quite a unique instrument as it's composed both by a space and a ground elements, which mainly consists with the telecomm systems of both the spacecraft and ground stations. This is because this experiment is based on measuring the doppler shift in the returning signal from Juno which allows to characterize Jupiter's gravitational field. Thus the instrument can't really be separated from the telecomm hardware which is the reason its weight and power requirment were omitted in the previuosly shown totals.
- **Jupiter Energetic-particle Detector Instrument (JEDI):** It detects high energy electrons and ions present in the Jovian magnetosphere which are discriminated by composition. Each sensor is characterized by six electron and six ion viewing directions that together cover a  $160^\circ \times 12^\circ$  field of view. In total three sensors are present on Juno, two arranged to obtain an almost compete  $360^\circ$  view perpendicular to the spacecraft spin axis while the third one is instead aligned with it to achieve a full scan of the sky over one spin period. As the JEDI sensors are self-contained units no electronic hardware is present within the radiation vault.
- **Jovian Auroral Distribution Experiment (JADE):** It detects low energy electrons and ions with the same goal of characterizing the magnetosphere as JEDI. The instrument comprises of three identical electron energy per chrage analyzers (JADE-E) and a single ion mass spectrometer (JADE-I). The electron sensors are located on the three sides of the spacecraft that do not house the solar arrays pointing outwards, to again obtain a complete view normal to the spin axis. The spectrometer field of view, instead, contains the spin axis and like the third JEDI sensor it scans all the sky over a full rotation.
- **Ultraviolet Spectrograph (UVS):** This instrument images and measures the spectrum of the Jovian aurora in order to understand its morphology and source. The chosen ultraviolet range of  $68 \div 210 \text{ nm}$  covers all of the most important UV emissions form the aurora, mainly the H Lyman series and longer wavelenghts from hydrocarbons. The sensor is mounted on the side of Juno, relying once more on the spinning of the spacecraft to achive a full sweep of the planet.
- **Radio and Plasma Waves (Waves):** Its objective is to study both components of the electromagnetic field generated by plasma and radio waves inside the polar regions of Jupiter's magnetosphere to understand its interaction with the atmosphere and magnetic field. To detect the electric component a V-shaped dipole antenna is used, while for the magnetic component amuch smaller magnetic search coil is employed. Both sensors cover a vast range of frequencies, namely from  $50 \text{ Hz}$  up to  $40 \text{ MHz}$ .
- **Visible-spectrum Camera (JunoCam):** It's designed to provide highly detailed color images of Jupiter to help and support public engagement of the mission. The instrument is thus only comprised of the camera itself, mounted on the side of the spacecraft, and all the necessary electronics which, given the less crital objective and relaxed radiation tolerance requirments, aren't housed in the radiation vault.
- **Juno Infra-Red Auroral Mapper (JIRAM):** It's an infra-red imager and spectrometer that studies the Jovian atmosphere in the  $2 \div 5 \mu\text{m}$  range complementing both the atmospheric and magnetospheric experiments. This instrument is also completely housed outside of the radiation vault since it is a late addition after mission selection, reason for both the relaxed radiation requirments and less than ideal positioning of the sensor on the aft deck of the spacecraft.

## 7.2 Payload and Goals correlation

There is a notable overlap in the main objectives of the payload instruments, both in the sense that multiple ones collaborate towards a single scientific goal, but also in the sense that a single instrument can addres multiple goals. All of these relations are exemplified in the following table.

Mission goals	Instrument objectives
a	b
c	d

Table 2: Mission goals and instrument objectives correlation

## 7.3 Payload and Phases/ConOps correlation

Another high-level correlation can be highlighted between the mission phases/ConOps and the activities of the payload as shown below.

Mission phases/ConOps	Payload activities
a	b
c	d

Table 3: Mission phases/ConOps and Payload activities correlation

## 8 Mission analysis

### Bibliography

- [1] Howard D. Curtis. *Orbital Mechanics for Engineering Students*. Elsevier, 2014.