1 High Frequency Linear Waves in a Hot Uniform Plasma

This follows Section 2 from the thesis of Decker, referencing heavily Fundamentals of Plasma Physics by Paul Bellan [1].

1.1 Maxwell Equations

The Maxwell equations are (in Heaviside Form)

$$\nabla \cdot \mathbf{E} = \frac{\rho}{\epsilon_0} \tag{1.1.1}$$

$$\nabla \cdot \mathbf{B} = 0 \tag{1.1.2}$$

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t} \tag{1.1.3}$$

$$\nabla \times \mathbf{B} = \mu_0 \mathbf{j} - \epsilon_0 \mu_0 \frac{\partial \mathbf{E}}{\partial t}$$
 (1.1.4)

When considering waves its more convenient to Fourier transform the equations. The Fourier transformed Maxwell equations are

$$i\mathbf{k} \cdot \mathbf{E}_k = \frac{\rho_k}{\epsilon_0} \tag{1.1.5}$$

$$i\mathbf{k} \cdot \mathbf{B}_k = 0 \tag{1.1.6}$$

$$i\mathbf{k} \times \mathbf{E}_k = i\omega \mathbf{B}_k$$
 (1.1.7)

$$i\mathbf{k} \times \mathbf{B}_k = \mu_0 \mathbf{j} - \epsilon_0 \mu_0 i\omega \mathbf{E}_k$$
 (1.1.8)

1.2 Wave Equation

The wave equation for a linear wave in a weakly inhomogeneous medium, using the WKB approximation, is

$$\underline{\mathcal{D}} \cdot \mathbf{E} = 0 \tag{1.2.1}$$

$$\underline{\underline{\mathcal{D}}} := \mathbf{N} \otimes \mathbf{N} - N^2 \underline{\underline{I}} + \underline{\underline{\epsilon}}$$
 (1.2.2)

where **N** is the refractive index, \otimes is the Tensor/Outer Product, $\underline{\underline{I}}$ is the Identity Tensor, $\underline{\underline{\epsilon}}$ is the Dielectric Tensor, **E** is the wave electric field and $\underline{\underline{\mathcal{D}}}$ is

the Dispersion Tensor. The Dielectric Tensor is related to the Susceptibility Tensor $\underline{\chi}$ and the Conductivity Tensor $\underline{\sigma}$ by

$$\underline{\underline{\epsilon}} = \underline{\underline{I}} + \underline{\underline{\chi}} = \underline{\underline{I}} + \frac{i}{\omega \epsilon_0} \underline{\underline{\sigma}} \tag{1.2.3}$$

The conductivity tensor also appears in linear Ohm's law to relate ${\bf E}$ and ${\bf j}$

$$\mathbf{j} = \underline{\sigma} \cdot \mathbf{E} \tag{1.2.4}$$

The condition (1.2.1) can be left multiplied by **E** to get a quadratic form

$$\mathcal{D} = \mathbf{E}^* \cdot \underline{\mathcal{D}} \cdot \mathbf{E} = 0 \tag{1.2.5}$$

When there is weak absorption, $\underline{\underline{\epsilon}}$ and hence $\underline{\underline{\mathcal{D}}}$ are Hermitian. We define the Hermitian part $\underline{\underline{\mathcal{D}}}^H := \frac{1}{2} \left(\underline{\underline{\mathcal{D}}} + \underline{\underline{\mathcal{D}}}^\dagger\right)$ and anti-Hermitian part $\underline{\underline{\mathcal{D}}}^A := \frac{1}{2} \left(\underline{\underline{\mathcal{D}}} - \underline{\underline{\mathcal{D}}}^\dagger\right)$ such that $\underline{\underline{\mathcal{D}}} = \underline{\underline{\mathcal{D}}}^H + \underline{\underline{\mathcal{D}}}^A$. We can then define the condition (1.2.1) as

$$\mathcal{D} = \mathbf{E}^* \cdot \underline{\underline{\mathcal{D}}}^H \cdot \mathbf{E} + \mathbf{E}^* \cdot i\underline{\underline{\mathcal{D}}}^A \cdot \mathbf{E} = 0$$
 (1.2.6)

1.3 Energy Equation for Linear Waves

From the continuity equation we can define the conservation of energy for a weakly inhomogeneous, weakly dissipative plasma for the linear mode \mathbf{E}_k

$$\frac{\partial \omega_k}{\partial t} + \nabla \cdot \mathbf{s}_k = -P_k^{lin} \tag{1.3.1}$$

where ω_k is the time averaged energy density, \mathbf{s}_k is the time averaged energy flow and P_k^{lin} is the density of power dissipated. This equation is a function of the real parts of ω and \mathbf{k} .

1.3.1 Poynting's Theorem

Poynting's Theorem describes the energy flow of electromagnetic waves. It is derived by first considering the work done by an electromagnetic field on a charged particle dw_i travelling at velocity \mathbf{v} in a time dt

$$dw_i = \mathbf{F} \cdot d\mathbf{l} = q \left(\mathbf{E} + \mathbf{v} \times \mathbf{B} \right) \cdot \mathbf{v} dt = q \mathbf{E} \cdot \mathbf{v} dt \tag{1.3.2}$$

We get the total work done dw by summing over all particles and recognising $\mathbf{j} = \sum_{i} q\mathbf{v}$, where \mathbf{j} is the current density.

$$dw = \sum_{i} \mathbf{E} \cdot q\mathbf{v}dt = \mathbf{E} \cdot \mathbf{j}dt \tag{1.3.3}$$

Re-arranging we get an expression for the rate of transfer of energy density from the electromagnetic field to the particles w_P , i.e. the kinetic energy associated with particles coherent motion in the electromagnetic field.

$$\frac{dw_P}{dt} = \mathbf{E} \cdot \mathbf{j} \tag{1.3.4}$$

We see this term is related to kinetic energy as

$$\mathbf{F} \cdot \mathbf{v} = m \frac{d\mathbf{v}}{dt} \cdot \mathbf{v} = \frac{m}{2} \frac{dv^2}{dt} = \frac{d}{dt} \left(\frac{1}{2} m v^2 \right)$$
 (1.3.5)

 $\mathbf{E} \cdot \mathbf{j}$ can be expressed using (1.1.4), re-arranged for \mathbf{j} , and dotting with \mathbf{E}

$$\mathbf{j} = \frac{1}{\mu_0} \nabla \times \mathbf{B} - \epsilon_0 \frac{\partial \mathbf{E}}{\partial t}$$

$$\mathbf{E} \cdot \mathbf{j} = \frac{1}{\mu_0} \mathbf{E} \cdot (\nabla \times \mathbf{B}) - \epsilon_0 \mathbf{E} \cdot \frac{\partial \mathbf{E}}{\partial t}$$
(1.3.6)

Now use the vector expression $\nabla \cdot (\mathbf{A} \times \mathbf{B}) = (\nabla \times \mathbf{A}) \cdot \mathbf{B} - (\nabla \times \mathbf{B}) \cdot \mathbf{A}$ allowing us to replace $\mathbf{E} \cdot (\nabla \times \mathbf{B})$ with $-\nabla \cdot (\mathbf{E} \times \mathbf{B}) + (\nabla \times \mathbf{E}) \cdot \mathbf{B}$

$$\mathbf{E} \cdot \mathbf{j} = -\frac{1}{\mu_0} \nabla \cdot (\mathbf{E} \times \mathbf{B}) + (\nabla \times \mathbf{E}) \cdot \mathbf{B} - \epsilon_0 \mathbf{E} \cdot \frac{\partial \mathbf{E}}{\partial t}$$
(1.3.7)

Use (1.1.3) to replace $\nabla \times \mathbf{E}$

$$\mathbf{E} \cdot \mathbf{j} = -\frac{1}{\mu_0} \nabla \cdot (\mathbf{E} \times \mathbf{B}) - \frac{1}{\mu_0} \frac{\partial \mathbf{B}}{\partial t} \cdot \mathbf{B} - \epsilon_0 \mathbf{E} \cdot \frac{\partial \mathbf{E}}{\partial t}$$
(1.3.8)

Rearranging we get

$$\mathbf{E} \cdot \mathbf{j} + \epsilon_0 \mathbf{E} \cdot \frac{\partial \mathbf{E}}{\partial t} + \frac{1}{\mu_0} \frac{\partial \mathbf{B}}{\partial t} \cdot \mathbf{B} + \frac{1}{\mu_0} \nabla \cdot (\mathbf{E} \times \mathbf{B}) = 0$$
 (1.3.9)

This is Poynting's Theorem, which is really just a statement of the continuity equation for electromagnetic waves. We define the Poynting Vector ${\bf S}$

$$\mathbf{S} := \frac{1}{\mu_0} \mathbf{E} \times \mathbf{B} \tag{1.3.10}$$

which is the energy flux carried by the electromagnetic wave. Defining the total energy of the wave \mathcal{U} , we can rewrite Poynting's Theorem as

$$\frac{\partial \mathcal{U}}{\partial t} + \nabla \cdot \mathbf{S} = 0 \tag{1.3.11}$$

where

$$\frac{\partial \mathcal{U}}{\partial t} = \mathbf{j} \cdot \mathbf{E} + \epsilon_0 \mathbf{E} \cdot \frac{\partial \mathbf{E}}{\partial t} + \frac{1}{\mu_0} \frac{\partial \mathbf{B}}{\partial t} \cdot \mathbf{B}$$
 (1.3.12)

The total energy $\mathcal{U} = \mathcal{U}_P + \mathcal{U}_{EM}$, where \mathcal{U}_P is the energy associated with coherent motion of charged particles within the field and \mathcal{U}_{EM} is the energy associated with the electromagnetic field. Earlier we associated $\mathbf{E} \cdot \mathbf{j}$ with \mathcal{U}_P so we can write

$$\frac{\partial \mathcal{U}_P}{\partial t} = \mathbf{E} \cdot \mathbf{j} \tag{1.3.13}$$

$$\frac{\partial \mathcal{U}_{EM}}{\partial t} = \epsilon_0 \mathbf{E} \cdot \frac{\partial \mathbf{E}}{\partial t} + \frac{1}{\mu_0} \frac{\partial \mathbf{B}}{\partial t} \cdot \mathbf{B}$$
 (1.3.14)

We can also write the energy of the wave $\mathcal{U}(t)$ using (1.3.12)

$$\mathcal{U}(t) = \mathcal{U}(-\infty) + \int_{-\infty}^{t} \left\langle \mathbf{j} \cdot \mathbf{E} + \epsilon_0 \mathbf{E} \cdot \frac{\partial \mathbf{E}}{\partial t} + \frac{1}{\mu_0} \frac{\partial \mathbf{B}}{\partial t} \cdot \mathbf{B} \right\rangle dt' \qquad (1.3.15)$$

where $\langle ... \rangle$ represents time averaging over a wave period. Time averaging will remove the rapidly varying component from the oscillation of the fields and keep the slowly varying component from the evolution of the wave, which is what we are interested in. Likewise we would need to take a time average of the Poynting Flux S to see evolution of the wave energy flow.

Also note the background energy term $\mathcal{U}(-\infty)$, indicating the integral term is the difference in energy between when the wave is present and when it is absent. This implies the existence of 'negative energy waves', i.e. the plasma is unstable to the creation of electromagnetic waves.

1.3.2 Time Averaged Poynting's Theorem

In order to make progress we need to take the time average of the terms in Poynting's Theorem over a wave period. This is complicated as some of the terms are the product of two complex oscillating fields, e.g. $\mathbf{E} \cdot \mathbf{j}$. As only the real parts correspond to physical quantities we need to be careful, this is covered in 4.1, to summarise we need to use (4.1.5) to correctly calculate time averages for products of complex oscillating fields.

We will let $\omega = \omega_R + i\omega_I$ and $\mathbf{k} = \mathbf{k}_R + i\mathbf{k}_I$, essentially allowing for wave dissipation in space and time. We only consider weak dissipation i.e. small ω_I and ω_I . We will define our time averaged energy density as w and the time averaged energy flow as \mathbf{s} . We also define their Fourier Transforms as w_k and \mathbf{s}_k , giving the respective energy density and energy flow for each mode \mathbf{k} .

The first thing to note is (4.1.5) implies all our time averaged quantities will be proportional to the factor $e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})}$. Therefore our time averaged Poynting's Theorem will look like

$$\frac{\partial}{\partial t} \left(w e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})} \right) + \nabla \cdot \left(\mathbf{s} e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})} \right) = 0$$

$$2\omega_I w e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})} - 2\mathbf{k}_I \cdot \mathbf{s} e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})} = 0$$
(1.3.16)

We will get rid of $e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})}$ factors in final results by assuming ω_I and \mathbf{k}_I are so small $e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})} \approx 1$. However they are important as through the time derivative and divergence they introduce factors of $2\omega_I$ and $2\mathbf{k}_I$ we need to divide by to get the time averaged energy density and energy flow.

Fourier Transforming and letting $e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})} \approx 1$ we get a generalised form of Poynting's Theorem

$$2\omega_I w_k - 2\mathbf{k}_I \cdot \mathbf{s}_k = 0 \tag{1.3.17}$$

Now we will calculate the time averaged quantities. To start with Fourier Transform (1.3.9)

$$\mathbf{E}_k \cdot \mathbf{j}_k + \epsilon_0 \mathbf{E}_k \cdot (-i\omega \mathbf{E}_k) + \frac{1}{\mu_0} \mathbf{B}_k \cdot (-i\omega \mathbf{B}_k) + \frac{1}{\mu_0} i\mathbf{k} \cdot (\mathbf{E}_k \times \mathbf{B}_k) = 0 \quad (1.3.18)$$

It's convenient to split this calculation into 3 parts. To start with we will calculate $\langle \mathbf{E}_k \cdot \mathbf{j}_k + \epsilon_0 \mathbf{E}_k \cdot (-i\omega \mathbf{E}_k) \rangle$. Firstly we note using the definition of the dielectric tensor (1.2.3) and Ohm's Law (1.2.4)

$$\mathbf{E}_{k} \cdot \mathbf{j}_{k} + \epsilon_{0} \mathbf{E}_{k} \cdot (-i\omega \mathbf{E}_{k}) = \mathbf{E}_{k} \cdot (\underline{\underline{\sigma}} - i\omega \epsilon_{0} \underline{\underline{I}}) \cdot \mathbf{E}_{k}$$

$$= \mathbf{E}_{k} \cdot (-i\omega \epsilon_{0}) \left(\frac{i}{\omega \epsilon_{0}} \underline{\underline{\sigma}} + \underline{\underline{I}} \right) \cdot \mathbf{E}_{k} = \mathbf{E}_{k} \cdot (-i\omega \epsilon_{0} \underline{\underline{\epsilon}}) \cdot \mathbf{E}_{k}$$

$$\implies \langle \mathbf{E}_{k} \cdot \mathbf{j}_{k} + \epsilon_{0} \mathbf{E}_{k} \cdot (-i\omega \mathbf{E}_{k}) \rangle = \langle \mathbf{E}_{k} \cdot (-i\omega \epsilon_{0} \underline{\underline{\epsilon}}) \cdot \mathbf{E}_{k} \rangle$$

$$(1.3.19)$$

Using (4.1.5) and substituting in $i\omega^* = i\omega_R + \omega_I$ and $-i\omega = -i\omega_R + \omega_I$

$$\langle \mathbf{E}_{k} \cdot \left(-i\omega\epsilon_{0}\underline{\underline{\epsilon}} \right) \cdot \mathbf{E}_{k} \rangle = \frac{1}{4} \left[\mathbf{E}_{k}^{*} \cdot \left(-i\omega\epsilon_{0}\underline{\underline{\epsilon}} \right) \cdot \mathbf{E}_{k} + \mathbf{E}_{k} \cdot \left(i\omega^{*}\epsilon_{0}\underline{\underline{\epsilon}^{*}} \right) \cdot \mathbf{E}_{k}^{*} \right] e^{2(\omega_{I}t - \mathbf{k}_{I} \cdot \mathbf{x})}$$

$$= \frac{\epsilon_{0}}{4} \left[\mathbf{E}_{k}^{*} \cdot \left(-i\omega_{R} + \omega_{I} \right) \underline{\underline{\epsilon}} \cdot \mathbf{E}_{k} + \mathbf{E}_{k} \cdot \left(i\omega_{R} + \omega_{I} \right) \underline{\underline{\epsilon}^{*}} \cdot \mathbf{E}_{k}^{*} \right] e^{2(\omega_{I}t - \mathbf{k}_{I} \cdot \mathbf{x})}$$

$$= \frac{\epsilon_{0}}{4} \left[i\omega_{R} \left(\mathbf{E}_{k} \cdot \underline{\underline{\epsilon}^{*}} \cdot \mathbf{E}_{k}^{*} - \mathbf{E}_{k}^{*} \cdot \underline{\underline{\epsilon}} \cdot \mathbf{E}_{k} \right) + \omega_{I} \left(\mathbf{E}_{k} \cdot \underline{\underline{\epsilon}^{*}} \cdot \mathbf{E}_{k}^{*} + \mathbf{E}_{k}^{*} \cdot \underline{\underline{\epsilon}} \cdot \mathbf{E}_{k} \right) \right] e^{2(\omega_{I}t - \mathbf{k}_{I} \cdot \mathbf{x})}$$

$$(1.3.20)$$

We can now use a quadratic form identity to replace $\mathbf{E}_k \cdot \underline{\underline{\epsilon}}^* \cdot \mathbf{E}_k^*$ with $\mathbf{E}_k^* \cdot \underline{\epsilon}^\dagger \cdot \mathbf{E}_k$, where \dagger is the conjugate transpose.

$$\mathbf{E} \cdot \underline{\underline{\epsilon}}^* \cdot \mathbf{E}^* = \sum_{p,q} E_p \epsilon_{pq}^* E_q^* = \sum_{p,q} E_p \epsilon_{qp}^{\dagger} E_q^* = \mathbf{E}^* \cdot \underline{\underline{\epsilon}}^{\dagger} \cdot \mathbf{E}$$
 (1.3.21)

Applying this identity above we get, also taking the opportunity to assume $e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})} \approx 1$

$$\langle \dots \rangle = \frac{\epsilon_0}{4} \left[i \omega_R \left(\mathbf{E}_k^* \cdot \left(\underline{\underline{\epsilon}}^{\dagger} - \underline{\underline{\epsilon}} \right) \cdot \mathbf{E}_k \right) + \omega_I \left(\mathbf{E}_k^* \cdot \left(\underline{\underline{\epsilon}}^{\dagger} + \underline{\underline{\epsilon}} \right) \cdot \mathbf{E}_k \right) \right]$$
 (1.3.22)

Now we use the fact ω_I and \mathbf{k}_I are small to Taylor expand $\underline{\epsilon}$ about ω_R and \mathbf{k}_R

$$\underline{\underline{\epsilon}}(\omega_R + i\omega_I, \mathbf{k}_R + \mathbf{k}_I) \approx \underline{\underline{\epsilon}}(\omega_R, \mathbf{k}_R) + i\omega_I \left. \frac{\partial \underline{\underline{\epsilon}}}{\partial \omega} \right|_{\omega_R, \mathbf{k}_R} + i\mathbf{k}_I \cdot \left. \frac{\partial \underline{\underline{\epsilon}}}{\partial \mathbf{k}} \right|_{\omega_R, \mathbf{k}_R} (1.3.23)$$

$$\left[\underline{\underline{\epsilon}}\left(\omega_R + i\omega_I, \mathbf{k}_R + \mathbf{k}_I\right)\right]^{\dagger} \approx \underline{\underline{\epsilon}}\left(\omega_R, \mathbf{k}_R\right) - i\omega_I \left.\frac{\partial\underline{\underline{\epsilon}}}{\partial\omega}\right|_{\omega_R, \mathbf{k}_R} - i\mathbf{k}_I \cdot \left.\frac{\partial\underline{\underline{\epsilon}}}{\partial\mathbf{k}}\right|_{\omega_R, \mathbf{k}_R} (1.3.24)$$

Recognising that $\underline{\underline{\epsilon}}(\omega_R, \omega_I) = \underline{\underline{\epsilon}}^H$ as it is the small imaginary components of ω and \mathbf{k} which make $\underline{\epsilon}$ non-hermitian, we have

$$\underline{\underline{\epsilon}}^{H} = \frac{1}{2} \left(\underline{\underline{\epsilon}} + \underline{\underline{\epsilon}}^{\dagger} \right) = \underline{\underline{\epsilon}} \left(\omega_{R}, \mathbf{k}_{R} \right)$$
 (1.3.25)

$$\underline{\underline{\epsilon}}^{A} = \frac{1}{2} \left(\underline{\underline{\epsilon}} - \underline{\underline{\epsilon}}^{\dagger} \right) = i \left[\omega_{I} \frac{\partial \underline{\underline{\epsilon}}}{\partial \omega} + \mathbf{k}_{I} \cdot \frac{\partial \underline{\underline{\epsilon}}}{\partial \mathbf{k}} \right]_{\omega_{B}, k_{B}}$$
(1.3.26)

Substituting in these expressions and cancelling $-i^2$ we get

$$\langle \dots \rangle = \frac{\epsilon_0}{2} \left[\omega_R \left(\mathbf{E}_k^* \cdot \left(\omega_I \frac{\partial \underline{\epsilon}}{\partial \omega} + \mathbf{k}_I \cdot \frac{\partial \underline{\epsilon}}{\partial \mathbf{k}} \right) \cdot \mathbf{E}_k \right) + \omega_I \left(\mathbf{E}_k^* \cdot \underline{\epsilon} \cdot \mathbf{E}_k \right) \right] \Big|_{\omega_R, \mathbf{k}_R}$$
(1.3.27)

Gathering terms containing ω_I and terms containing \mathbf{k}_I

$$\langle \dots \rangle = \frac{\epsilon_0}{2} \left[\omega_I \mathbf{E}^* \cdot \left(\omega_R \frac{\partial \underline{\epsilon}}{\partial \omega} + \underline{\epsilon} \right) \cdot \mathbf{E} + \omega_R \mathbf{E}^* \cdot \left(\mathbf{k}_I \cdot \frac{\partial \underline{\epsilon}}{\partial \mathbf{k}} \right) \cdot \mathbf{E} \right] \Big|_{\omega_R, \mathbf{k}_R}$$
(1.3.28)

Using the product rule and the fact we are evaluating all terms at ω_R , \mathbf{k}_R we can replace ω_R with ω and write

$$\omega_R \frac{\partial \underline{\epsilon}}{\partial \omega} + \underline{\epsilon} = \frac{\partial \left(\omega_{\underline{\epsilon}}\right)}{\partial \omega} \tag{1.3.29}$$

Applying this substitution, we also rewrite to make comparison with (1.3.17)

$$\langle \dots \rangle = 2\omega_I \left(\frac{\epsilon_0}{4} \mathbf{E}^* \cdot \frac{\partial \left(\omega \underline{\underline{\epsilon}} \right)}{\partial \omega} \cdot \mathbf{E} \right) - 2\mathbf{k}_I \cdot \left(-\frac{\omega \epsilon_0}{4} \mathbf{E}^* \cdot \frac{\partial \underline{\underline{\epsilon}}}{\partial \mathbf{k}} \cdot \mathbf{E} \right)$$
(1.3.30)

Notice we've picked up an energy flow term despite time averaging terms we've associated with energy density. This flow is associated with particle motion due to the wave.

Now time average the third term in (1.3.18)

$$\left\langle \frac{1}{\mu_0} \mathbf{B}_k \cdot (-i\omega \mathbf{B}_k) \right\rangle = \frac{1}{4\mu_0} \left[\mathbf{B}_k \cdot (i\omega^*) \cdot \mathbf{B}^* - \mathbf{B}_k^* \cdot (i\omega) \cdot \mathbf{B} \right] e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})}$$
(1.3.31)

Again assuming $e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})} \approx 1$ and noting $\mathbf{B}^* \cdot \mathbf{B} = ||\mathbf{B}||^2$

$$\left\langle \frac{1}{\mu_0} \mathbf{B}_k \cdot (-i\omega \mathbf{B}_k) \right\rangle = \frac{i\|\mathbf{B}\|^2}{4\mu_0} \left(\omega^* - \omega \right) = 2\omega_I \frac{\|\mathbf{B}_k\|^2}{4\mu_0}$$
 (1.3.32)

Now we will time average the Poynting Flux term $\frac{1}{\mu_0}\nabla \cdot \langle (\mathbf{E} \times \mathbf{B}) \rangle$. We can use the standard method to write

$$\frac{1}{\mu_0} \nabla \cdot \langle \mathbf{E} \times \mathbf{B} \rangle = \frac{1}{\mu_0} \nabla \cdot \frac{1}{2} \mathcal{R} \left[\mathbf{E}^* \times \mathbf{B} \right] e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})}$$

$$= -2\mathbf{k}_I \cdot \frac{1}{2\mu_0} \mathcal{R} \left[\mathbf{E}^* \times \mathbf{B} \right] e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})} \tag{1.3.33}$$

Using the assumption $e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})} \approx 1$ and rewriting we get

$$\frac{1}{\mu_0} \nabla \cdot \langle \mathbf{E} \times \mathbf{B} \rangle = -2\mathbf{k}_I \cdot \frac{1}{2\mu_0} \mathcal{R} \left[\mathbf{E}^* \times \mathbf{B} \right]$$
 (1.3.34)

Combining (1.3.30), (1.3.32) and (1.3.34) we get

$$2\omega_{I} \left(\frac{\epsilon_{0}}{4} \mathbf{E}^{*} \cdot \frac{\partial \left(\omega \underline{\underline{\epsilon}} \right)}{\partial \omega} \cdot \mathbf{E} + \frac{\|\mathbf{B}_{k}\|^{2}}{4\mu_{0}} \right) - 2\mathbf{k}_{I} \cdot \left(-\frac{\omega \epsilon_{0}}{4} \mathbf{E}^{*} \cdot \frac{\partial \underline{\underline{\epsilon}}}{\partial \mathbf{k}} \cdot \mathbf{E} + \frac{1}{2\mu_{0}} \mathcal{R} \left[\mathbf{E}^{*} \times \mathbf{B} \right] \right) = 0$$

$$(1.3.35)$$

Comparing with (1.3.17) we see immediately

$$w_k = \frac{\epsilon_0}{4} \mathbf{E}^* \cdot \frac{\partial \left(\omega_{\underline{\underline{\epsilon}}}\right)}{\partial \omega} \cdot \mathbf{E} + \frac{\|\mathbf{B}_k\|^2}{4\mu_0}$$
 (1.3.36)

$$\mathbf{s}_{k} = -\frac{\omega \epsilon_{0}}{4} \mathbf{E}^{*} \cdot \frac{\partial \underline{\epsilon}}{\partial \mathbf{k}} \cdot \mathbf{E} + \frac{1}{2\mu_{0}} \mathcal{R} \left[\mathbf{E}^{*} \times \mathbf{B} \right]$$
 (1.3.37)

1.3.3 Normalised Energy Density and Energy Flow

We define the normalised time averaged energy density Σ_k and the normalised time averaged energy flow Φ_k for mode \mathbf{k} as

$$w_k = \frac{\epsilon_0}{2} \|\mathbf{E}_k\|^2 \Sigma_k \tag{1.3.38}$$

$$\mathbf{s}_k = \frac{\epsilon_0 c}{2} \|\mathbf{E}_k\|^2 \Phi_k \tag{1.3.39}$$

Define the Polarisation vector \mathbf{e}_k

$$\mathbf{e}_k = \frac{\mathbf{E}_k}{\|\mathbf{E}_k\|} \tag{1.3.40}$$

Starting with Σ_k , using (1.3.36) and (1.3.38) we can write

$$\Sigma_k = \frac{1}{2} \mathbf{e}_k \cdot \frac{\partial \left(\omega_{\underline{\epsilon}}\right)}{\partial \omega} \cdot \mathbf{e}_k + \frac{c^2}{2} \frac{\|B_k\|^2}{\|\mathbf{E}_k\|^2}$$
 (1.3.41)

Taking the norm of (1.1.7) we get

$$\|i\mathbf{k} \times \mathbf{E}\| = i\omega \|\mathbf{B}_k\|$$

$$\implies \frac{\|\mathbf{B}_k\|}{\|\mathbf{E}_k\|} = \frac{\|\frac{\mathbf{k}}{\omega} \times \mathbf{E}_k\|}{\|E_k\|} = \frac{1}{c} \|\mathbf{N} \times \mathbf{e}_k\|$$
(1.3.42)

Substituting in we get

$$\Sigma_k = \frac{1}{2} \mathbf{e}_k \cdot \frac{\partial \left(\omega_{\underline{\epsilon}}\right)}{\partial \omega} \cdot \mathbf{e}_k + \frac{1}{2} \|\mathbf{N} \times \mathbf{e}_k\|^2$$
 (1.3.43)

Using the vector identity $\|\mathbf{N} \times \mathbf{e}_k\|^2 = \mathbf{e}_k^* \cdot (N^2 \underline{\underline{I}} - \mathbf{N} \otimes \mathbf{N}) \cdot \mathbf{e}_k$ (see 4.2.1) we get

$$\Sigma_{k} = \frac{1}{2} \mathbf{e}_{k}^{*} \cdot \left(\frac{\partial \left(\omega_{\underline{\underline{\epsilon}}} \right)}{\partial \omega} + N^{2} \underline{\underline{I}} - \mathbf{N} \otimes \mathbf{N} \right) \cdot \mathbf{e}_{k}$$
 (1.3.44)

We want to move the last 2 terms inside the ω derivative. We can do this by seeing

$$\mathbf{N} = \frac{c\mathbf{k}}{\omega} \implies \frac{\partial}{\partial \omega} (\omega^2 \mathbf{N}^2) = 0$$

$$\implies \frac{\partial}{\partial \omega} \left[\omega^2 \left(\mathbf{N} \otimes \mathbf{N} - N^2 \underline{\underline{I}} \right) \right] = 0$$
(1.3.45)

Splitting this term into $\omega \times \omega$ ($\mathbf{N} \otimes \mathbf{N} - N^2 \underline{\underline{I}}$) and using the Product rule we get

$$\omega \left(\mathbf{N} \otimes \mathbf{N} - N^{2} \underline{\underline{I}} \right) + \omega \frac{\partial}{\partial \omega} \left[\omega \left(\mathbf{N} \otimes \mathbf{N} - N^{2} \underline{\underline{I}} \right) \right] = 0$$

$$\implies \left(N^{2} \underline{\underline{I}} - \mathbf{N} \otimes \mathbf{N} \right) = \frac{\partial}{\partial \omega} \left[\omega \left(\mathbf{N} \otimes \mathbf{N} - N^{2} \underline{\underline{I}} \right) \right]$$
(1.3.46)

Substituting in we get, using the definition of \mathcal{D} (1.2.2)

$$\Sigma_k = \frac{1}{2} \mathbf{e}_k^* \cdot \frac{\partial}{\partial \omega} \left[\omega \left(\underline{\underline{\epsilon}} + \mathbf{N} \otimes \mathbf{N} - N^2 \underline{\underline{I}} \right) \right] \cdot \mathbf{e}_k$$

$$= \frac{1}{2} \mathbf{e}_{k}^{*} \cdot \frac{\partial}{\partial \omega} (\omega \mathcal{D}) \cdot \mathbf{e}_{k}$$
 (1.3.47)

Expanding the derivative we get

$$\Sigma_k = \frac{1}{2} \mathbf{e}_k^* \cdot \mathcal{D} \cdot \mathbf{e}_k + \frac{\omega}{2} \mathbf{e}_k^* \cdot \frac{\partial \mathcal{D}}{\partial \omega} \cdot \mathbf{e}_k$$
 (1.3.48)

Using $\mathbf{e}_k^* \cdot \mathcal{D} \cdot \mathbf{e}_k = 0$ in the limit of weak dissipation we get the final expression

$$\Sigma_k = \frac{\omega}{2} \mathbf{e}_k^* \cdot \frac{\partial \mathcal{D}}{\partial \omega} \cdot \mathbf{e}_k \tag{1.3.49}$$

Now we move on to Φ_k . Using (1.3.37) and (??) we can write

$$\Phi_{k} = -\frac{\omega}{2c} \mathbf{e}_{k}^{*} \cdot \frac{\partial \underline{\underline{\epsilon}}}{\partial \mathbf{k}} \cdot \mathbf{e}_{k} + \frac{1}{\|\mathbf{E}_{k}\|^{2}} \mathcal{R} \left[\mathbf{E}_{k}^{*} \times \mathbf{B}_{k} \right]$$

$$= -\frac{1}{2} \mathbf{e}_{k}^{*} \cdot \frac{\partial \underline{\underline{\epsilon}}}{\partial \mathbf{N}} \cdot \mathbf{e}_{k} + \frac{1}{\|\mathbf{E}_{k}\|^{2}} \mathcal{R} \left[\mathbf{E}_{k}^{*} \times \mathbf{B}_{k} \right] \tag{1.3.50}$$

Taking $\mathbf{E}_k^* \times (1.1.7)$ we get

$$\mathbf{E}_k^* \times (i\mathbf{k} \times \mathbf{E}_k) = i\omega \mathbf{E}_k^* \times \mathbf{B}_k$$

$$\implies \|\mathbf{E}_k^* \times \left(\frac{\mathbf{k}}{\omega} \times \mathbf{E}_k\right)\| = \|\mathbf{E}_k^* \times \mathbf{B}_k\| \tag{1.3.51}$$

Dividing through by $\|\mathbf{E}_k\|^2$ and using the definition of \mathbf{N} we get

$$\frac{1}{\|\mathbf{E}_k\|^2} \mathcal{R} \left[\mathbf{E}_k^* \times \mathbf{B}_k \right] = \frac{1}{c} \mathcal{R} \left[\mathbf{e}_k^* \times (\mathbf{N} \times \mathbf{e}_k) \right]$$
 (1.3.52)

We can expand the double cross product as

$$\mathbf{e}_{k}^{*} \times (\mathbf{N} \times \mathbf{e}_{k}) = (\mathbf{e}_{k}^{*} \cdot \mathbf{e}_{k}) \,\mathbf{N} - (\mathbf{e}_{k}^{*} \cdot \mathbf{N}) \,\mathbf{e}_{k} \tag{1.3.53}$$

Using the vector identity $\mathbf{N} - \mathcal{R}\left[\left(\mathbf{e}_{k}^{*} \cdot \mathbf{N}\right) \mathbf{e}_{k}\right] = -\frac{1}{2}\mathbf{e}_{k}^{*} \cdot \frac{\partial\left(\mathbf{N} \otimes \mathbf{N} - N^{2}\underline{\underline{I}}\right)}{\partial \mathbf{N}} \cdot \mathbf{e}_{k}$ (see 4.2.2) we can rewrite the expression for Φ_{k}

$$\Phi_{k} = -\frac{1}{2} \mathbf{e}_{k}^{*} \cdot \frac{\partial \underline{\underline{\epsilon}}}{\partial \mathbf{N}} \cdot \mathbf{e}_{k} - \frac{1}{2} \mathbf{e}_{k}^{*} \cdot \frac{\partial \left(\mathbf{N} \otimes \mathbf{N} - N^{2} \underline{\underline{I}} \right)}{\partial \mathbf{N}} \cdot \mathbf{e}_{k}$$

$$= -\frac{1}{2} \mathbf{e}_{k}^{*} \cdot \left(\underline{\underline{\epsilon}} + \mathbf{N} \otimes \mathbf{N} - N^{2} \underline{\underline{I}} \right) \cdot \mathbf{e}_{k} \tag{1.3.54}$$

Again using the definition for \mathcal{D} (1.2.2) we get

$$\Phi_k = -\frac{1}{2} \mathbf{e}_k^* \cdot \frac{\partial \mathcal{D}}{\partial \mathbf{N}} \cdot \mathbf{e}_k \tag{1.3.55}$$

1.3.4 Group Velocity

We can get a short definition of the group velocity in terms of our normalised energy flows. Using the standard definition of group velocity

$$\mathbf{v}_g = \frac{\partial \mathbf{k}}{\partial \omega} = \frac{\frac{\partial \mathcal{D}}{\partial \omega}}{\frac{\partial \mathcal{D}}{\partial \mathbf{k}}} = \frac{c\Phi_k}{\Sigma_k}$$
 (1.3.56)

1.4 Absorption Coefficient

Poynting's Theorem covers the energy density and flow in both the wave and the charged particles. However, energy flow from the wave fields to the particles is the definition of wave absorption. This term appeared from the $\mathbf{E} \cdot \mathbf{j}$ term. If we were to rearrange Poynting's Theorem to isolate this term we get

$$\epsilon_0 \mathbf{E} \cdot \frac{\partial \mathbf{E}}{\partial t} + \frac{1}{\mu_0} \mathbf{B} \cdot \frac{\partial \mathbf{B}}{\partial t} + \frac{1}{\mu_0} \nabla \cdot \mathbf{S} = -\mathbf{E} \cdot \mathbf{j}$$
 (1.4.1)

Writing the time average $\langle \mathbf{E} \cdot \mathbf{j} \rangle$, we can use Ohm's Law (1.2.4) and the standard method for time averaging

$$\langle \mathbf{E} \cdot \underline{\underline{\sigma}} \cdot \mathbf{E} \rangle = \frac{1}{4} \left[\mathbf{E}^* \cdot \underline{\underline{\sigma}} \cdot \mathbf{E} + \mathbf{E} \cdot \underline{\underline{\sigma}}^* \cdot \mathbf{E}^* \right] e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})}$$

$$= \frac{1}{4} \left[\mathbf{E}^* \cdot \left(\underline{\underline{\sigma}}^\dagger + \underline{\underline{\sigma}} \right) \cdot \mathbf{E} \right] e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})}$$
(1.4.2)

Assuming $e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})} \approx 1$ and recognising $\underline{\underline{\sigma}}^{\dagger} + \underline{\underline{\sigma}} = 2\underline{\underline{\sigma}}^H$ we get

$$\langle \mathbf{E} \cdot \underline{\underline{\sigma}} \cdot \mathbf{E} \rangle = \frac{1}{2} \mathbf{E}^* \cdot \underline{\underline{\sigma}}^H \cdot \mathbf{E}$$
 (1.4.3)

We can associate this with the power dissipation. Taking the Fourier transform we get the power dissipation P_k^{lin} associated with mode k

$$P_k^{lin} = \frac{1}{2} \mathbf{E}_k^* \cdot \underline{\underline{\sigma}}^H \cdot \mathbf{E}_k \tag{1.4.4}$$

Consider the anti-hermitian part of the susceptibility tensor $\underline{\chi}$

$$\underline{\underline{\chi}}^{A} = \frac{1}{2} \left(\underline{\underline{\chi}} - \underline{\underline{\chi}}^{\dagger} \right) = \frac{i}{\omega \epsilon_{0}} \frac{1}{2} \left(\underline{\underline{\sigma}} + \underline{\underline{\sigma}}^{\dagger} \right) = \frac{i}{\omega \epsilon_{0}} \underline{\underline{\sigma}}^{H}$$

$$\Longrightarrow \underline{\underline{\sigma}}^{H} = -i\omega \epsilon_{0} \underline{\underline{\chi}}^{A}$$
(1.4.5)

Substituting in, we can rewrite P_k^{lin}

$$P_k^{lin} = -\frac{i\omega\epsilon_0}{2} \|\mathbf{E}_k\|^2 \mathbf{e}_k^* \cdot \underline{\underline{\chi}}^A \cdot \mathbf{e}_k$$
 (1.4.6)

We almost have the normalisation constant between \mathbf{s}_k and Φ_k from (1.3.39). Substituting in we get

$$P_k^{lin} = -\frac{i\omega}{c} \frac{\|\mathbf{s}_k\|}{\|\Phi_k\|} \mathbf{e}_k^* \cdot \underline{\chi}^A \cdot \mathbf{e}_k$$
 (1.4.7)

Defining the absorption coefficient α_k^{lin}

$$\alpha_k^{lin} = \frac{P_k^{lin}}{\|\mathbf{s}_k\|} = -\frac{i\omega}{c} \frac{1}{\|\Phi_k\|} \mathbf{e}_k^* \cdot \underline{\chi}^A \cdot \mathbf{e}_k$$
 (1.4.8)

2 Hot Plasma Dielectric Tensor

This follows Richard Fitzpatrick's explanation [2] except I've filled in some of the gaps. The aim is to derive a non-relativistic form of the dielectric tensor accounting for finite temperature and hence finite Larmor radius.

Start by considering small amplitude waves propagating through a plasma in a uniform magnetic field $\mathbf{B} = B\hat{z}$. Consider the collisionless Vlasov equation

$$\frac{\partial f}{\partial t} + \mathbf{v} \cdot \nabla f + \frac{e}{m} \left(\mathbf{E} + \mathbf{v} \times \mathbf{B} \right) \cdot \frac{\partial f}{\partial \mathbf{v}} = 0 \tag{2.0.1}$$

We linearise by defining $\mathbf{E} = \mathbf{E}_1$, $\mathbf{B} = \mathbf{B}_0 + \mathbf{B}_1$ and $f(\mathbf{x}, \mathbf{v}, t) = f_0(\mathbf{v}) + f_1(\mathbf{x}, \mathbf{v}, t)$. As f_0 is a solution of (2.0.1) but the time and space derivatives vanish we get

$$(\mathbf{v} \times \mathbf{B}_0) \cdot \frac{\partial f_0}{\partial \mathbf{v}} = 0 \tag{2.0.2}$$

The cross product expands as $v_{\perp}B_0 \sin\theta \hat{\mathbf{x}} - v_{\perp}B_0 \sin\theta \hat{\mathbf{y}}$. We want to transform to polar co-ordinates (r, θ, z) , so use $\hat{\mathbf{r}} = \cos\theta \hat{\mathbf{x}} + \sin\theta \hat{\mathbf{y}}$ and $\hat{\theta} = -\sin\theta \hat{\mathbf{x}} + \cos\theta \hat{\mathbf{y}}$. Under this transformation the cross product becomes $v_{\perp}B_0\hat{\theta}$ so (2.0.2) implies

$$\frac{\partial f_0}{\partial \theta} = 0 \tag{2.0.3}$$

i.e. $f_0 = f_0 \left(v_{\perp}, v_{\parallel} \right)$ only.

Linearising (2.0.1) and re-arranging we get

$$\frac{\partial f_1}{\partial t} + \mathbf{v} \cdot \nabla f_1 + \frac{e}{m} \left(\mathbf{v} \times \mathbf{B}_0 \right) \cdot \frac{\partial f_1}{\partial \mathbf{v}} = -\frac{e}{m} \left(\mathbf{E}_1 + \mathbf{v} \times \mathbf{B}_1 \right) \cdot \frac{\partial f_0}{\partial \mathbf{v}}$$
(2.0.4)

We can recognise the LHS as the total rate of change of f_1 following the unperturbed particle trajectories. Therefore

$$\frac{Df_1}{Dt} = -\frac{e}{m} \left(\mathbf{E}_1 + \mathbf{v} \times \mathbf{B}_1 \right) \cdot \frac{\partial f_0}{\partial \mathbf{v}}$$
 (2.0.5)

Assuming f_1 vanishes at $t = -\infty$ we can write (dropping the 1 subscript as all vacuum fields are now on the LHS)

$$f_{1}(\mathbf{r}, \mathbf{v}, t) = -\frac{e}{m} \int_{-\infty}^{t} \left[\mathbf{E}(\mathbf{r}', t') + \mathbf{v} \times \mathbf{B}(\mathbf{r}', t') \right] \cdot \frac{\partial f_{0}(\mathbf{v}')}{\partial \mathbf{v}} dt'$$
(2.0.6)

where $(\mathbf{r}', \mathbf{v}')$ is the unperturbed trajectory that passes through (\mathbf{r}, \mathbf{v}) when t' = t. We know the dielectric tensor $\underline{\epsilon}$ is defined as

$$\underline{\underline{\epsilon}} \cdot \mathbf{E} = \mathbf{E} + \frac{i}{\omega \epsilon_0} \mathbf{j} = \mathbf{E} + \frac{i}{\omega \epsilon_0} \sum_{s} e_s \int \mathbf{v} f_{1s} d^3 \mathbf{v}$$
 (2.0.7)

relating the current **j** to the moments of f_1 . Here s refers to a sum over species we will drop for now to keep things simpler. So we need to do some work to evaluate f_1 , and then calculate the time and velocity integrals to get the dielectric tensor elements.

The Cartesian components of the velocity are

$$\mathbf{v} = (v_{\perp} \cos \theta, v_{\perp} \sin \theta, v_{\parallel}) \tag{2.0.8}$$

which implies (Ω is the gyrofrequency)

$$\mathbf{v} = \left(v_{\perp} \cos\left[\Omega\left(t - t'\right) + \theta\right], v_{\perp} \sin\left[\Omega\left(t - t'\right) + \theta\right], v_{\parallel}\right) \tag{2.0.9}$$

This expression is a bit long so define $\chi = \Omega(t - t') + \theta$ so

$$\mathbf{v} = (v_{\perp} \cos \chi, v_{\perp} \sin \chi, v_{\parallel}) \tag{2.0.10}$$

We can integrate in time to get the electron position

$$x' - x = -\frac{v_{\perp}}{\Omega} \left(\sin \chi - \sin \theta \right) \tag{2.0.11}$$

$$y' - y = \frac{v_{\perp}}{\Omega} \left(\cos \chi - \cos \theta \right) \tag{2.0.12}$$

$$z' - z = v_{\parallel} (t' - t) \tag{2.0.13}$$

Both v_{\perp} and v_{\parallel} are constants of the motion, implying $f_0(\mathbf{v}') = f_0(\mathbf{v})$. Using $v_{\perp} = \sqrt{v_x^2 + v_y^2}$ we get

$$\frac{\partial f_0}{\partial v_x'} = \frac{\partial v_\perp}{\partial v_x'} \frac{\partial f_0}{\partial v_\perp} = \frac{v_x'}{\partial v_\perp} \frac{\partial f_0}{\partial v_\perp} = \cos \chi \frac{\partial f_0}{\partial v_\perp}$$
(2.0.14)

$$\frac{\partial f_0}{\partial v_y'} = \frac{\partial v_\perp}{\partial v_y'} \frac{\partial f_0}{\partial v_\perp} = \frac{v_y'}{\partial v_\perp} \frac{\partial f_0}{\partial v_\perp} = \sin \chi \frac{\partial f_0}{\partial v_\perp}$$
(2.0.15)

$$\frac{\partial f_0}{\partial v_z'} = \frac{\partial f_0}{\partial v_{\parallel}} \tag{2.0.16}$$

If we expand the integrand in (2.0.4) we get

$$\left[\mathbf{E}\left(\mathbf{r}',t'\right) + \mathbf{v} \times \mathbf{B}\left(\mathbf{r}',t'\right)\right] \cdot \frac{\partial f_{0}\left(\mathbf{v}'\right)}{\partial \mathbf{v}} = \left(E_{x} + v'_{y}B_{z} - v'_{z}B_{y}\right) \frac{\partial f_{0}}{\partial v'_{x}} + \left(E_{x}y + v'_{z}B_{x} - v'_{x}B_{z}\right) \frac{\partial f_{0}}{\partial v'_{y}} + \left(E_{z} + v'_{x}B_{y} - v'_{y}B_{x}\right) \frac{\partial f_{0}}{\partial v'_{y}}$$

$$(2.0.17)$$

Using (2.0.14), (2.0.15) and (2.0.16)

$$= (E_x - v_{\parallel} B_y) \cos \chi \frac{\partial f_0}{\partial v_{\perp}} + (E_y + v_{\parallel} B_x) \sin \chi \frac{\partial f_0}{\partial v_{\perp}} +$$

$$+ \left(v_y' \cos \chi - v_x' \sin \chi\right) B_z \frac{\partial f_0}{\partial v_{\perp}} + \left(E_z + v_{\perp} \cos \chi B_y - v_{\perp} \sin \chi B_x\right) \frac{\partial f_0}{\partial v_{\parallel}} (2.0.18)$$

Using $v_y' \cos \chi - v_x' \sin \chi = v_{\perp} \cos \chi \sin \chi - v_{\perp} \cos \chi \sin \chi = 0$ the third term vanishes. Now assuming all linearised terms have a $e^{i(\mathbf{k}\cdot\mathbf{x}-\omega t)}$ dependence the integrand becomes

$$= \left[\left(E_x - v_{\parallel} B_y \right) \cos \chi \frac{\partial f_0}{\partial v_{\perp}} + \left(E_y + v_{\parallel} B_x \right) \sin \chi \frac{\partial f_0}{\partial v_{\perp}} + \right.$$

$$\left. + \left(E_z + v_{\perp} \cos \chi B_y - v_{\perp} \sin \chi B_x \right) \frac{\partial f_0}{\partial v_{\parallel}} \right] e^{i(\mathbf{k} \cdot \mathbf{x}' - \omega t')} e^{i(\mathbf{k} \cdot \mathbf{x} - \omega t)}$$

$$(2.0.19)$$

where the first complex exponential is from the perturbed electric and magnetic field and the second is from the perturbed f_1 . We can combine these to get a term like

$$\exp\left[i\mathbf{k}\cdot(\mathbf{x}'-\mathbf{x})\right]\exp\left[-i\omega\left(t'-t\right)\right] \tag{2.0.20}$$

Taking **k** to lie in the x-z plane we can use (2.0.11) and (2.0.13) to get

$$\exp\left[i\frac{k_{\perp}v_{\perp}}{\Omega}\left(\sin\chi - \sin\theta\right) + ik_{\parallel}v_{\parallel}\left(t'-t\right)\right] \exp\left[-i\omega\left(t'-t\right)\right]$$

$$= \exp\left[i\mu\sin\chi\right] \exp\left[-i\mu\sin\theta\right] \exp\left[i\left(k_{\parallel}v_{\parallel} - \omega\right)\left(t'-t\right)\right] \tag{2.0.21}$$

where we define $\mu = \frac{k_{\perp}v_{\perp}}{\Omega}$. We can use the complex exponential dependence of **E** and **B** to eliminate **B** using the (1.1.7)

$$\mathbf{k} \times \mathbf{E} = \omega \mathbf{B} \implies \mathbf{B} = \frac{\mathbf{k} \times \mathbf{E}}{\omega}$$
 (2.0.22)

Again using the fact **k** lies in the x-z plane to get

$$B_x = -\frac{k_{\parallel} E_y}{\omega} \tag{2.0.23}$$

$$B_y = \frac{k_{\parallel} E_x - k_{\perp} E_z}{\omega} \tag{2.0.24}$$

Using these expressions we can write

$$E_{x} - v_{\parallel} B_{y} = \frac{\omega - v_{\parallel} k_{\parallel}}{\omega} E_{x} + \frac{k_{\perp} v_{\parallel}}{\omega} E_{z}$$
 (2.0.25)

$$E_y + v_{\parallel} B_x = \frac{\omega - k_{\parallel} v_{\parallel}}{\omega} E_y \tag{2.0.26}$$

Substituting into (2.0.19) we get

$$\frac{1}{\omega} \left(\left[\left(\omega - k_{\parallel} v_{\parallel} \right) E_{x} + k_{\perp} v_{\parallel} E_{z} \right] \cos \chi \frac{\partial f_{0}}{\partial v_{\perp}} + \left(\omega - k_{\parallel} v_{\parallel} \right) \sin \chi \frac{\partial f_{0}}{\partial v_{\perp}} E_{y} \right. \\
+ \left[\omega E_{z} + k_{\parallel} v_{\perp} \cos \chi E_{c} - k_{\perp} v_{\perp} \cos \chi E_{z} + k_{\parallel} v_{\perp} \sin \chi E_{y} \right] \frac{\partial f_{0}}{\partial v_{\parallel}} \right) \times \\
\exp \left[i \mu \sin \chi \right] \exp \left[-i \mu \sin \theta \right] \exp \left[i \left(k_{\parallel} v_{\parallel} - \omega \right) (t' - t) \right] \tag{2.0.27}$$

In a little bit we will use moments of this to find the elements of the dielectric tensor, i.e. the coefficients of E_x , E_y and E_z so it is convenient to group these coefficients now. Doing this we obtain

$$\frac{1}{\omega} \left(\left[\left(\omega - k_{\parallel} v_{\parallel} \right) \frac{\partial f_{0}}{\partial v_{\perp}} + k_{\parallel} v_{\perp} \frac{\partial f_{0}}{\partial v_{\parallel}} \right] \cos \chi E_{x} + \left[\left(\omega - k_{\parallel} v_{\parallel} \right) \frac{\partial f_{0}}{\partial v_{\perp}} + k_{\parallel} v_{\perp} \frac{\partial f_{0}}{\partial v_{\parallel}} \right] \sin \chi E_{y}
+ \left[\omega \frac{\partial f_{0}}{\partial v_{\parallel}} + \left(k_{\perp} v_{\parallel} \frac{\partial f_{0}}{\partial v_{\perp}} - k_{\perp} v_{\perp} \frac{\partial f_{0}}{\partial v_{\parallel}} \right) \cos \chi \right] E_{z} \right) \times
\exp \left[i \mu \sin \chi \right] \exp \left[-i \mu \sin \theta \right] \exp \left[i \left(k_{\parallel} v_{\parallel} - \omega \right) \left(t' - t \right) \right] \tag{2.0.28}$$

This expression is quite cumbersome so we will introduce some new variables P and Q

$$P = \left(\omega - k_{\parallel} v_{\parallel}\right) \frac{\partial f_0}{\partial v_{\perp}} + k_{\parallel} v_{\perp} \frac{\partial f_0}{\partial v_{\parallel}}$$
 (2.0.29)

$$Q = k_{\perp} v_{\parallel} \frac{\partial f_0}{\partial v_{\perp}} - k_{\perp} v_{\perp} \frac{\partial f_0}{\partial v_{\parallel}}$$
 (2.0.30)

Making these substitutions we can write our expression for f_1 with our rearranged integrand

$$f_{1} = \frac{-e}{m\omega} \int_{-\infty}^{t} \left(P \cos \chi E_{x} + P \sin \chi E_{y} + \left[\omega \frac{\partial f_{0}}{\partial v_{\parallel}} + Q \cos \chi \right] E_{z} \right) \times \exp \left[i\mu \sin \chi \right] \exp \left[-i\mu \sin \theta \right] \exp \left[i\left(k_{\parallel}v_{\parallel} - \omega \right) \left(t' - t \right) \right] dt'$$
(2.0.31)

To perform this time integration we can expand these terms using the Jacobi-Anger expansion

$$e^{iz\sin\theta} = \sum_{n=-\infty}^{\infty} J_n(z) e^{in\theta}$$
 (2.0.32)

where J_n is the Bessel function of the first kind. We can do some manipulation to get 2 more useful identities

$$\cos\theta e^{iz\sin\theta} = \frac{1}{iz}\frac{\partial}{\partial\theta}e^{iz\sin\theta} = \sum_{n=-\infty}^{\infty} \frac{nJ_n(z)}{z}e^{in\theta}$$
 (2.0.33)

$$\sin \theta e^{iz\sin \theta} = \frac{1}{i} \frac{\partial}{\partial z} e^{iz\sin \theta} = -i \sum_{n=-\infty}^{\infty} J'_n(z) e^{in\theta}$$
 (2.0.34)

Using these identities we can replace $\cos \chi e^{i\mu\sin\chi}$, $\sin \chi e^{i\mu\sin\chi}$, and $e^{i\mu\sin\chi}$ terms

$$f_{1} = \frac{-e}{m\omega} \int_{-\infty}^{t} \sum_{n=-\infty}^{\infty} \left(\frac{nJ_{n}\left(\mu\right)}{\mu} PE_{x} - iJ_{n}'\left(\mu\right) PE_{y} + \left[\omega \frac{\partial f_{0}}{\partial v_{\parallel}} J_{n}\left(\mu\right) + \frac{nJ_{n}\left(\mu\right)}{\mu} Q\right] E_{z} \right) \times$$

$$\exp\left[-i\mu\sin\theta\right]\exp\left[in\theta\right]\exp\left[i\left(n\Omega+k_{\parallel}v_{\parallel}-\omega\right)\left(t'-t\right)\right]dt'$$
(2.0.35)

where we have split $\exp[in\chi] = \exp[in\theta] \exp[in\Omega(t'-t)]$ and combined some terms. From now on we'll drop the argument of the Bessel functions for brevity. We can rewrite the coefficient for E_z as

$$\omega \frac{\partial f_0}{\partial v_{\parallel}} J_n + \frac{nJ_n}{\mu} \left(k_{\perp} v_{\parallel} \frac{\partial f_0}{\partial v_{\perp}} - k_{\perp} v_{\perp} \frac{\partial f_0}{\partial v_{\parallel}} \right)$$

$$= \left[\frac{n\Omega v_{\parallel}}{v_{\perp}} \frac{\partial f_0}{\partial v_{\perp}} + (\omega - n\Omega) \frac{\partial f_0}{\partial v_{\parallel}} \right] J_n \qquad (2.0.36)$$

We can define a new simpler variable R to replace Q

$$R = \frac{n\Omega v_{\parallel}}{v_{\perp}} \frac{\partial f_0}{\partial v_{\perp}} + (\omega - n\Omega) \frac{\partial f_0}{\partial v_{\parallel}}$$
 (2.0.37)

allowing us to write

$$f_1 = \frac{-e}{m\omega} \int_{-\infty}^{t} \sum_{n=-\infty}^{\infty} \left(\frac{nJ_n}{\mu} PE_x - iJ_n' PE_y + J_n RE_z \right) \times$$

$$\exp\left[-i\mu\sin\theta\right]\exp\left[in\theta\right]\exp\left[i\left(n\Omega+k_{\parallel}v_{\parallel}-\omega\right)\left(t'-t\right)\right]dt'$$
(2.0.38)

We now the only time dependence in our integrand comes from the final complex exponential term, so we can write

$$f_1 = \frac{-e}{m\omega} \sum_{n=-\infty}^{\infty} \left(\frac{nJ_n}{\mu} PE_x - iJ'_n PE_y + J_n RE_z \right) \times$$

$$\exp\left[-i\mu\sin\theta\right] \int_{-\infty}^{t} \exp\left[in\theta\right] \exp\left[i\left(n\Omega + k_{\parallel}v_{\parallel} - \omega\right)(t'-t)\right] dt' \qquad (2.0.39)$$

This is an easy integral

$$\int_{-\infty}^{t} \exp\left[i\left(n\Omega + k_{\parallel}v_{\parallel} - \omega\right)(t' - t)\right]dt' = \frac{i}{\omega - n\Omega - k_{\parallel}v_{\parallel}}$$
(2.0.40)

where we make some arguments using delta functions about how the complex exponential evaluates to 0 at $-\infty$. So we get our simplest form for f_1

$$f_{1} = \frac{-e}{m\omega} \sum_{n=-\infty}^{\infty} \left(\frac{nJ_{n}}{\mu} PE_{x} - iJ'_{n} PE_{y} + J_{n} RE_{z} \right) \times$$

$$\exp\left[-i\mu \sin \theta \right] \exp\left[in\theta \right] \frac{i}{\omega - n\Omega - k_{\parallel} v_{\parallel}}$$
(2.0.41)

Now we need to take the first velocity moment f_1 to get our currents

$$j_i = e \int v_i f_1 d^3 \mathbf{v} \tag{2.0.42}$$

We already saw $v_x = v_{\perp} \cos \theta$, $v_y = v_{\parallel} \sin \theta$ and $v_z = v_{\parallel}$ so we get these expressions for the currents

$$j_{x} = \frac{-e^{2}v_{\perp}}{m\omega} \sum_{n=-\infty}^{\infty} \left(\frac{nJ_{n}}{\mu} PE_{x} - iJ'_{n} PE_{y} + J_{n} RE_{z} \right) \times$$

$$\cos\theta \exp\left[-i\mu \sin\theta\right] \exp\left[in\theta\right] \frac{i}{\omega - n\Omega - k_{\parallel}v_{\parallel}} \qquad (2.0.43)$$

$$j_{y} = \frac{-e^{2}v_{\perp}}{m\omega} \sum_{n=-\infty}^{\infty} \left(\frac{nJ_{n}}{\mu} PE_{x} - iJ'_{n} PE_{y} + J_{n} RE_{z} \right) \times$$

$$\sin\theta \exp\left[-i\mu \sin\theta\right] \exp\left[in\theta\right] \frac{i}{\omega - n\Omega - k_{\parallel}v_{\parallel}} \qquad (2.0.44)$$

$$j_{z} = \frac{-e^{2}v_{\parallel}}{m\omega} \sum_{n=-\infty}^{\infty} \left(\frac{nJ_{n}}{\mu} PE_{x} - iJ'_{n} PE_{y} + J_{n} RE_{z} \right) \times$$

$$\exp\left[-i\mu \sin\theta\right] \exp\left[in\theta\right] \frac{i}{\omega - n\Omega - k_{\parallel}v_{\parallel}} \qquad (2.0.45)$$

which are all very similar apart from factors of v_{\perp} or v_{\parallel} and more expressions $\propto \exp\left[-i\mu\sin\theta\right]$ we can replace using (2.0.32), (2.0.33) and (2.0.34) where we use the expression $e^{-i\mu\sin\theta} = e^{i\mu\sin(-\theta)}$. Doing these replacements we get

$$j_{x} = \frac{-e^{2}v_{\perp}}{m\omega} \sum_{n=-\infty}^{\infty} \left(\frac{nJ_{n}}{\mu}PE_{x} - iJ'_{n}PE_{y} + J_{n}RE_{z}\right) \times$$

$$\sum_{m=-\infty}^{\infty} \frac{mJ_{m}}{\mu} \exp\left[i\left(n-m\right)\theta\right] \frac{i}{\omega - n\Omega - k_{\parallel}v_{\parallel}} \qquad (2.0.46)$$

$$j_{y} = \frac{-e^{2}v_{\perp}}{m\omega} \sum_{n=-\infty}^{\infty} \left(\frac{nJ_{n}}{\mu}PE_{x} - iJ'_{n}PE_{y} + J_{n}RE_{z}\right) \times$$

$$\sum_{m=-\infty}^{\infty} -iJ'_{m} \exp\left[i\left(n-m\right)\theta\right] \frac{i}{\omega - n\Omega - k_{\parallel}v_{\parallel}} \qquad (2.0.47)$$

$$j_{z} = \frac{-e^{2}v_{\parallel}}{m\omega} \sum_{n=-\infty}^{\infty} \left(\frac{nJ_{n}}{\mu}PE_{x} - iJ'_{n}PE_{y} + J_{n}RE_{z}\right) \times$$

$$\sum_{m=-\infty}^{\infty} J_m \exp\left[i\left(n-m\right)\theta\right] \frac{i}{\omega - n\Omega - k_{\parallel}v_{\parallel}}$$
 (2.0.48)

We can replace $\exp [i (n - m) \theta] = 2\pi \delta (n - m)$ where δ is the Dirac delta. To get rid of these terms we'll just integrate over gyrophase θ now which just sets n = m and we pick up a factor of 2pi. This collapses the double sum to a single sum, allowing us to write an expression which almost fits on a single line!

$$\int_{0}^{2\pi} j_{x} d\theta = \frac{-2\pi i e^{2} v_{\perp}}{m\omega} \sum_{n=-\infty}^{\infty} \frac{\left(\frac{nJ_{n}}{\mu}\right)^{2} P E_{x} - i \frac{nJ_{n}J'_{n}}{\mu} P E_{y} + \frac{nJ_{n}^{2}}{\mu} R E_{z}}{\omega - n\Omega - k_{\parallel}v_{\parallel}} \quad (2.0.49)$$

$$\int_{0}^{2\pi} j_{y} d\theta = \frac{2\pi e^{2} v_{\perp}}{m\omega} \sum_{n=-\infty}^{\infty} \frac{\frac{nJ_{n}J_{n}'}{\mu} PE_{x} - i \left(J_{n}'\right)^{2} PE_{y} + J_{n}J_{n}'RE_{z}}{\omega - n\Omega - k_{\parallel}v_{\parallel}}$$
(2.0.50)

$$\int_{0}^{2\pi} j_{z} d\theta = \frac{-2i\pi e^{2}v_{\parallel}}{m\omega} \sum_{n=-\infty}^{\infty} \frac{\frac{nJ_{n}^{2}}{\mu}PE_{x} - iJ_{n}J_{n}'PE_{y} + J_{n}^{2}RE_{z}}{\omega - n\Omega - k_{\parallel}v_{\parallel}}$$
(2.0.51)

We can now use the definition of the dielectric tensor (2.0.7) to write

$$\epsilon_{ij} = \delta_{ij} + \sum_{s} \frac{e_s^2}{\omega^2 \epsilon_0 m_s} \sum_{n = -\infty}^{\infty} \int_0^{\infty} \int_0^{\infty} \frac{2\pi S_{ij}}{\omega - n\Omega - k_{\parallel} v_{\parallel}} dv_{\perp} dv_{\perp} \qquad (2.0.52)$$

where

$$S_{ij} = \begin{bmatrix} v_{\perp} \left(\frac{nJ_n}{\mu}\right)^2 P & iv_{\perp} \frac{nJ_nJ'_n}{\mu} P & v_{\perp} \frac{nJ_n^2}{\mu} R \\ -iv_{\perp} \frac{nJ_nJ'_n}{\mu} P & v_{\perp} \left(J'_n\right)^2 P & -iv_{\perp} J_nJ'_nR \\ v_{\parallel} \frac{nJ_n^2}{\mu} P & iv_{\perp} J_nJ'_nP & v_{\perp} J_n^2R \end{bmatrix}$$

$$P = \left(\omega - k_{\parallel}v_{\parallel}\right) \frac{\partial f_{0s}}{\partial v_{\perp}} + k_{\parallel}v_{\perp} \frac{\partial f_0}{\partial v_{\parallel}}$$

$$R = \frac{n\Omega v_{\parallel}}{v_{\perp}} \frac{\partial f_{0s}}{\partial v_{\perp}} + (\omega - n\Omega) \frac{\partial f_0}{\partial v_{\parallel}}$$

and the arguments of the Bessel functions are $\mu = \frac{k_{\perp}v_{\perp}}{\Omega}$. We now use the fact the equilibrium distribution function f_{0s} is a Maxwellian

$$f_{0s} = \frac{n_s}{\pi^{\frac{3}{2}} v_{Ts}^3} \exp\left[-\frac{v_{\perp}^2 + v_{\parallel}^2}{v_{Ts}^2}\right]$$
 (2.0.54)

where $v_{Ts} = \sqrt{\frac{2kT_s}{2m_s}}$ is the thermal velocity.

3 Application to EBWs

The previous section was completely general. Now we'll apply these formulas to EBWs.

3.0.1 Electrostatic Approximation

We can break down the wave electric field \mathbf{E}_k into a transverse component \mathbf{E}_{kT} and a longitudinal component \mathbf{E}_{kL}

4 Appendix

4.1 Complex Exponential Quantities

Based on section 7.4 in [1]. A useful mathematical notation is to represent an oscillating physical quantity as a phasor, with the understanding the actual physical quantity is the real part, i.e.

$$\psi(t) = \mathcal{R} \left[\tilde{\psi} e^{i\mathbf{k} \cdot \mathbf{x} - \mathbf{i}\omega \mathbf{t}} \right] = \frac{1}{2} \left(\psi + \psi^* \right)$$
 (4.1.1)

For linear relationships taking the real is done implicitly as it has no effect. However, for non-linear relationships we must do this to get the correct answer. For example, the product of two oscillating quantities $\psi(t) = \tilde{\psi}e^{i(\mathbf{k}\cdot\mathbf{x}-\omega t)}$ and $\chi(t) = \tilde{\chi}e^{i(\mathbf{k}\cdot\mathbf{x}-\omega t)}$ must be written as

$$\psi(t)\chi(t) = \mathcal{R}\left[\tilde{\psi}e^{i(\mathbf{k}\cdot\mathbf{x}-\omega t)}\right] \times \mathcal{R}\left[\tilde{\chi}e^{i(\mathbf{k}\cdot\mathbf{x}-\omega t)}\right]$$
(4.1.2)

Expanding this out, assuming $\mathbf{k} = \mathbf{k}_R + i\mathbf{k}_I$ and $\omega = \omega_R + i\omega_I$

$$\psi(t)\chi(t) = \frac{1}{4} \left(\tilde{\psi}e^{i(\mathbf{k}\cdot\mathbf{x}-\omega t)} + \tilde{\psi}^*e^{i(\mathbf{k}^*\cdot\mathbf{x}-\omega^*t)} \right) \left(\tilde{\chi}e^{i(\mathbf{k}\cdot\mathbf{x}-\omega t)} + \tilde{\chi}^*e^{i(\mathbf{k}^*\cdot\mathbf{x}-\omega^*t)} \right)$$

$$= \frac{1}{4} \left(\tilde{\psi}\tilde{\chi}e^{2i(\mathbf{k}\cdot\mathbf{x}-\omega t)} + \tilde{\psi}\tilde{\chi}^*e^{i[(\mathbf{k}-\mathbf{k}^*)\cdot\mathbf{x}-(\omega-\omega^*)t]} + \tilde{\psi}^*\tilde{\chi}^*e^{i[(\mathbf{k}-\mathbf{k}^*)\cdot\mathbf{x}-(\omega-\omega^*)t]} + \tilde{\psi}^*\tilde{\chi}^*e^{2i(\mathbf{k}\cdot\mathbf{x}-\omega t)} \right)$$

$$(4.1.3)$$

As $-i(\omega - \omega^*) = 2\omega_I$ and $i(\mathbf{k} - \mathbf{k}^*) = -2\mathbf{k}_R$ we can rewrite this as

$$\psi(t)\chi(t) = \frac{1}{4} \left[\tilde{\psi}\tilde{\chi} + \tilde{\psi}^*\tilde{\chi}^* \right] e^{2i(\mathbf{k}\cdot\mathbf{x} - \omega t)} + \frac{1}{4} \left[\tilde{\psi}^*\tilde{\chi} + \tilde{\psi}\tilde{\chi}^* \right] e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})}$$
(4.1.4)

The first term is oscillatory so if we time average over a period of the oscillation they will vanish, while the second term will remain as it is non-oscillatory. Assuming ω_I and \mathbf{k}_R are very small and so $e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})}$ is constant over a single wave period, the time averaged quantity $\langle \psi(t)\chi(t)\rangle$ is

$$\langle \psi(t)\chi(t)\rangle = \frac{1}{4} \left(\tilde{\psi}\tilde{\chi}^* + \tilde{\psi}^*\tilde{\chi}\right) e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})}$$
(4.1.5)

The term inside the brackets can be rewritten as the real part of a complex number using (4.1.1)

$$\langle \psi(t)\chi(t)\rangle = \frac{1}{2}\mathcal{R}\left[\tilde{\psi}\tilde{\chi}^*\right]e^{2(\omega_I t - \mathbf{k}_I \cdot \mathbf{x})}$$
 (4.1.6)

4.2 Vector Identities

$$\mathbf{4.2.1} \quad \|\mathbf{N} \times \mathbf{e}\|^2 = \mathbf{e}_k^* \cdot \left(N^2 \underline{\underline{I}} - \mathbf{N} \otimes \mathbf{N}\right) \cdot \mathbf{e}_k$$

Start at the end and work backwards

$$\mathbf{e}_k^* \cdot \left(N^2 \underline{I} - \mathbf{N} \otimes \mathbf{N}\right) \cdot \mathbf{e}_k =$$

$$\begin{pmatrix} e_x & e_y & e_z \end{pmatrix} \begin{pmatrix} N_y^2 + N_z^2 & -N_x N_y & -N_x N_z \\ -N_x N_y & N_x^2 + N_z^2 & -N_y N_z \\ -N_x N_z & -N_y N_z & N_y^2 + N_z^2 \end{pmatrix} \begin{pmatrix} e_x \\ e_y \\ e_z \end{pmatrix}$$

$$= (N_y^2 + N_z^2) e_x^2 + (N_x^2 + N_z^2) e_y^2 + (N_x^2 + N_y^2) e_z^2 - 2N_x N_y e_x e_y - 2N_x N_z e_x e_z - 2N_y N_z e_y e_z$$

$$= (N_y e_z - N_z e_y)^2 + (N_z e_x - N_x e_z)^2 + (N_x e_y - N_y e_x)^2$$
(4.2.1)

These terms are the components of a cross product, hence we see this is equal to $\|\mathbf{N} \times \mathbf{e}\|^2$.

$$\mathbf{4.2.2} \quad \mathbf{N} - \mathcal{R}\left[\left(\mathbf{e}_k^* \cdot \mathbf{N}\right) \mathbf{e}_k\right] = -\tfrac{1}{2} \mathbf{e}_k^* \cdot \tfrac{\partial \left(\mathbf{N} \otimes \mathbf{N} - N^2 \frac{I}{\Xi}\right)}{\partial \mathbf{N}} \cdot \mathbf{e}_k$$

4.3 Bessel Functions

4.3.1 Bessel's Differential Equation

$$J_{-n}(x) = (-1)^n J_n(x)$$
(4.3.1)

$$J_n(x) = \frac{1}{\pi} \int_0^{\pi} \cos(n\theta - x\sin\theta) d\theta = \frac{1}{\pi} \int_0^{\pi} \cos(x\sin\theta - n\theta) d\theta \qquad (4.3.2)$$

4.3.2 Bessel Function Recursion Relations

$$J_{n+1}(x) + J_{n-1}(x) = \frac{2n}{x} J_n(x)$$
(4.3.3)

$$J'_n(x) = \frac{1}{2} \left(J_{n+1}(x) - J_{n-1}(x) \right) \tag{4.3.4}$$

$$I_{n+1}(x) - I_{n-1}(x) = -\frac{2n}{x}I_n(x)$$
(4.3.5)

$$I'_n(x) = \frac{1}{2} \left(I_{n+1}(x) + I_{n-1}(x) \right) \tag{4.3.6}$$

4.3.3 Bessel Function Integrals

$$\int_0^\infty \frac{J_n(x)}{x} dx = \frac{1}{n} \tag{4.3.7}$$

$$2\int_{0}^{\infty} J_{n}^{2}\left(x\sqrt{2b}\right)xe^{-x^{2}}dx = I_{n}(b)e^{-b}$$
(4.3.8)

$$4\int_{0}^{\infty} J_{n}\left(w\sqrt{2b}\right) J_{n}'\left(x\sqrt{2b}\right) x^{2} e^{-x^{2}} dx = \sqrt{2b} \left(I_{n}'(b) - I_{n}(b)\right) e^{-b}$$
 (4.3.9)

$$4\int_0^\infty \left(J_n'\left(x\sqrt{2b}\right)\right)^2 x^3 e^{-x^2} dx = \left[\frac{n^2}{b}I_n(b) + 2b\left(I_n(b) - I_n'(b)\right)\right] e^{-b}$$
(4.3.10)

First integral is from []. (4.3.9) and (4.3.10) are derived from derivatives of (4.3.8). We get (4.3.8) from Weber's second integral with $p = \sqrt{2b}$ and q = 1.

$$\int_{0}^{\infty} J_n^2(px) x e^{-q^2 x^2} dx = \frac{1}{2q^2} e^{-\frac{p^2}{2q^2}} I_n\left(\frac{p^2}{2q^2}\right)$$
 (4.3.11)

Taking a derivative with respect to b

$$\frac{\partial}{\partial b} \left[2 \int_0^\infty J_n^2 \left(\sqrt{2b} x \right) x e^{-x^2} dx \right] = 2 \int_0^\infty 2 J_n \left(\sqrt{2b} x \right) J_n' \left(\sqrt{2b} x \right) \frac{x}{\sqrt{2b}} x e^{-x^2} dx$$

$$\implies 4 \int_0^\infty J_n \left(w \sqrt{2b} \right) J_n' \left(x \sqrt{2b} \right) x^2 e^{-x^2} dx = \sqrt{2b} \frac{\partial}{\partial b} \left(e^{-b} I_n(b) \right)$$

$$\sqrt{2b} \left(I_n'(b) - I_n(b) \right) e^{-b} \tag{4.3.12}$$

Taking another derivative

$$\frac{\partial}{\partial b} \left[4 \int_0^\infty J_n \left(w \sqrt{2b} \right) J_n' \left(x \sqrt{2b} \right) x^2 e^{-x^2} dx \right]$$

$$J \tag{4.3.13}$$

4.4 Plasma Dispersion Function

$$Z(\zeta) = \frac{1}{\sqrt{\pi}} \int_{-\infty}^{\infty} \frac{e^{-x^2}}{x - \zeta} dx \tag{4.4.1}$$

$$Z'(\zeta) = 1 + \zeta Z(\zeta) = \frac{1}{\sqrt{\pi}} \int_{-\infty}^{\infty} \frac{xe^{-x^2}}{x - \zeta} dx$$
 (4.4.2)

$$\zeta Z'(\zeta) = \zeta \left(1 + \zeta Z(\zeta)\right) = \frac{1}{\sqrt{\pi}} \int_{-\infty}^{\infty} \frac{x^2 e^{-x^2}}{x - \zeta} dx \tag{4.4.3}$$

References

 $^{1}\mathrm{P.~M.}$ Bellan, Fundamentals~of~plasma~physics (Cambridge University Press, 2008).

 $^{^2\}mathrm{R.}$ Fitzpatrick, Plasma~physics:~an~introduction (Crc Press, 2014).