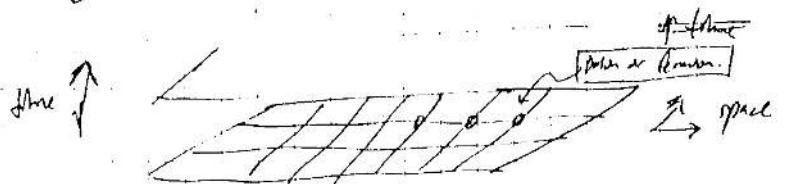


## Lecture 1 of Quantum Mechanics of Bosons and Fermions.

In this course, we're interested in the thermodynamics of bosons or fermions quantum fields. Essentially, these can be thought of as physical fields (functions) defined on all of spacetime with either a bosonic or fermionic degree of freedom living at each point:



Lake & Visser  
1707.01554

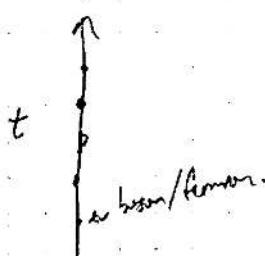
We will ultimately be constructing a path-integral representation of the thermodynamics partition function, which governs the thermodynamics of the system:

$$Z \equiv \text{Tr} [e^{-\beta \hat{H}}], \quad (\text{tr over full set of states})$$

This will allow similar derivations in QFT where one computes a vacuum transition rate:

$$\langle 0 | e^{-i \hat{H} t} | 0 \rangle.$$

(Note that  $-\beta \leftrightarrow -it$  similarly to the structure.) But first let's review some basic of thermodynamics of bosonic and fermionic systems in 0, spatial dimensions: (QM)



# Thermodynamics:

$$\beta = e^{-\frac{F}{T}}$$

$$Z(T) \equiv \text{Tr}[e^{-\beta \hat{H}}], \quad \beta \equiv \frac{1}{T}$$

$$F = -T \ln Z, \quad E = \frac{1}{2} \text{Tr} [\hat{H} e^{-\beta \hat{H}}]$$

$$S = -\frac{\delta F}{\delta T} = \ln Z + \frac{1}{T^2} \text{Tr} [\hat{H} e^{-\beta \hat{H}}] = -\frac{F}{T} + \frac{E}{T}, \\ \Rightarrow (E = F + TS)$$

## Boson: (Harmonic oscillator) (creation & annihilation operators)

$$\text{# } [\hat{a}, \hat{a}] = 0, \quad [\hat{a}^\dagger, \hat{a}^\dagger] = 0, \quad [\hat{a}, \hat{a}^\dagger] = 1$$

$$\hat{H} = \hbar \omega (\hat{a}^\dagger \hat{a} + \frac{1}{2}) = \frac{\hbar \omega}{2} (\hat{a}^\dagger \hat{a} + \hat{a} \hat{a}^\dagger)$$

(1)

$$\hat{a}^{\dagger}|n\rangle = \sqrt{n+1}|n+1\rangle, \hat{a}|n\rangle = \sqrt{n}|n-1\rangle \quad n=0, 1, 2, \dots$$

$$I = \sum_n |n\rangle \langle n|$$

# Fermionic oscillator:

$$\{\hat{a}, \hat{a}\} = 0, \{\hat{a}^{\dagger}, \hat{a}^{\dagger}\} = 0, \{\hat{a}, \hat{a}^{\dagger}\} = 1$$

vacuum state:  $\hat{a}|0\rangle \equiv 0$ .

one-particle state:  $\hat{a}^{\dagger}|1\rangle \equiv \hat{a}^{\dagger}|0\rangle$

No other states:  $\hat{a}|1\rangle = \hat{a}\hat{a}^{\dagger}|0\rangle = (1 - \hat{a}^{\dagger}\hat{a})|0\rangle = |0\rangle.$

$$\hat{a}^{\dagger}|1\rangle = \hat{a}^{\dagger}\hat{a}^{\dagger}|0\rangle = 0.$$

The Hamitonian is:

$$\underline{H} = \frac{\hbar\omega}{2} (\hat{a}^{\dagger}\hat{a} - \hat{a}\hat{a}^{\dagger}) = \hbar\omega (\hat{a}^{\dagger}\hat{a} - \frac{1}{2})$$

# Thermodynamics:

$$\begin{aligned} \text{Below: } \boxed{Z} &= \sum_{n=0}^{\infty} \langle n | e^{-\beta H} | n \rangle = \sum_{n=0}^{\infty} e^{-\beta\hbar\omega(\frac{1}{2} + n)} \\ &= \frac{e^{-\beta\hbar\omega/2}}{1 - e^{-\beta\hbar\omega}} = \boxed{\frac{1}{2} \csc\left(\frac{\pi\omega}{2T}\right)}. \end{aligned}$$

$$\boxed{F = T \ln\left(e^{\frac{\hbar\omega}{2T}} - e^{-\frac{\hbar\omega}{2T}}\right) + \frac{\hbar\omega}{2} + T \ln(1 - e^{-\beta\hbar\omega})}$$

$$\approx \begin{cases} \frac{\hbar\omega}{2} & T \ll \hbar\omega \\ -T \ln\left(\frac{T}{\hbar\omega}\right) & T \gg \hbar\omega \end{cases}$$

$$S = -\frac{\delta F}{\delta T} = \ln\left(1 - e^{-\beta\hbar\omega}\right) + \frac{\hbar\omega}{T} \frac{1}{e^{\beta\hbar\omega} - 1}$$

$$\boxed{E = F + S = \hbar\omega \left( \frac{1}{2} + \frac{1}{e^{\beta\hbar\omega} - 1} \right)}$$

(2)

# Fermions:

$$\boxed{Z} = \langle 0 | e^{-\beta H} | 0 \rangle + \langle 1 | e^{-\beta H} | 1 \rangle$$

$$= e^{+\frac{\hbar\omega}{2}} + e^{-\beta(\hbar\omega - \frac{\hbar\omega}{2})}$$

$$= [1 + e^{-\beta\hbar\omega}] e^{\frac{\beta\hbar\omega}{2}} = \boxed{2 \cosh(\frac{\beta\hbar\omega}{2})}$$

$$\boxed{F} = -T \ln(e^{\frac{\hbar\omega}{2T}} + e^{-\frac{\hbar\omega}{2T}}) = -\frac{\hbar\omega}{2} - T(1 + e^{-\beta\hbar\omega})$$

Note the minus signs from the boson case.

$$\Rightarrow S = -\ln(1 + e^{-\beta\hbar\omega}) + \frac{\hbar\omega}{T} \frac{1}{e^{\beta\hbar\omega} + 1}$$

$$\boxed{E} F + TS = \boxed{\hbar\omega \left( -\frac{1}{2} + \frac{1}{e^{\beta\hbar\omega} + 1} \right)},$$

You've seen this all before, but now we want to

(i) promote these to field calculations (with dependence on spacetime coordinates)

(ii) carry out a path-integral derivation.

Why to do this with coherent states, which is where bosons & fermions derivatives are the most similar.

### # Coherent states

~~Coherent states are eigenstates of the annihilation operator!~~

~~$\hat{a}|0\rangle = |\phi\rangle \quad \hat{a}^\dagger|\phi\rangle = \phi|0\rangle$~~

~~and for bosons/boson/fermions, these are given by:~~

~~$|\phi\rangle \equiv e^{\pm \phi a^\dagger} |0\rangle$~~

Need a quick review of how to deal with fermionic (Grassmann) variables.

Let  $\psi_i$  be ~~an~~ anticommuting variables. We define it to behave in the following way:

Andromathy:  $\psi_1\psi_2 = \psi_2\psi_1$  ( $\Rightarrow \psi^2 = 0$ )

Functions: defined by Taylor series

$$f(\psi) = f(0) + \psi f'(0) \quad \text{truncates!}$$

Differentiations:  $\frac{d}{d\psi}(c) = 0$ ,  $\frac{d}{d\psi}\psi = 1$ ; linear

(Definite) Integration:  $\int d\psi = 0$ ,  $\int d\psi \psi = 1$ ; linear

Excellently, all operators defined algebraically.

~~#~~ # Coherent States (Altland & Simons) bosons/fermions  $\boxed{\phi = +1/-1}$

These states are defined as:  $| \phi \rangle = e^{\pm \phi a^\dagger} | 0 \rangle$

These are eigenstates of the annihilation operator:  $\boxed{\hat{a} | \phi \rangle = \phi | \phi \rangle}$  in both cases

bosons:  $\hat{a} | \phi \rangle = [a, e^{\phi a^\dagger}] | 0 \rangle$   
 $= \sum_{n=0}^{\infty} \frac{[a, (a^\dagger)^n]}{n!} \phi^n | 0 \rangle$

$[a, (a^\dagger)^n] = n(a^\dagger)^{n-1}$  can easily prove this by induction;

assume:  $a(a^\dagger)^n = n(a^\dagger)^{n-1} + (a^\dagger)^n a$  (be true for  $n=1$ )

$$\Rightarrow \boxed{a(a^\dagger)^{n+1} = n(a^\dagger)^n + (a^\dagger)^n \underbrace{a a^\dagger}_{1+a^\dagger}}$$

$$\neq \boxed{(n+1)(a^\dagger)^n + (a^\dagger)^{n+1} a} \quad \checkmark$$

$$\Rightarrow \boxed{\hat{a} | \phi \rangle} = \sum_{n=1}^{\infty} \frac{\phi^n}{n!} n(a^\dagger)^{n-1} | 0 \rangle$$

$$= \phi e^{\phi a^\dagger} | 0 \rangle = \boxed{| \phi | \phi \rangle}$$

(4)

Semion:  $|\hat{a}|n\rangle = \hat{a} e^{-\hat{n}\hat{a}^\dagger}|0\rangle = \hat{a} (1 - \hat{n}\hat{a}^\dagger)|0\rangle - \hat{n} + \hat{n}\hat{a}^\dagger|0\rangle$

 $= 0 + \cancel{\hat{n}} + \eta|0\rangle \quad \downarrow \text{raise 2 lower (anti-commutes with } \hat{a} \text{ with } \eta\text{)}.$ 
 $= \eta(1 - \hat{n}\hat{a}^\dagger)|0\rangle \quad \downarrow \hat{n}^2 = 0, \text{ so can add } \hat{n}\eta.$ 
 $= |\eta|n\rangle$

Need the Glauber relation:

Completeness:  $\text{Id} = \int d(\varphi^+, \varphi) e^{-\varphi^+ \varphi} |\varphi\rangle \langle \varphi|$

overlap:  $\langle \varphi' | \varphi \rangle = \exp(\varphi'^+ \varphi)$

Measure:  $d(\varphi^+, \varphi) \equiv \frac{d\varphi^+ d\varphi}{\pi (1+\zeta)/2}$

Quick proof:

Boson:  $\langle \varphi' | \varphi \rangle = \langle 0 | e^{\frac{i}{\hbar} \varphi'^+ a} e^{\frac{i}{\hbar} \varphi a^\dagger} | 0 \rangle$  number.

Baker-Campbell-Hausdorff says  $e^{A+B} = e^{A+B+\frac{i}{\hbar}[A,B]} = [e^B e^A e^{[A,B]}]$

 $\frac{e^B e^A}{e^B e^A} = e^{A+B-\frac{i}{\hbar}[A,B]} = 1 \cancel{e^{i\hbar/2}}$

$$\begin{aligned} \Rightarrow \langle \varphi' | \varphi \rangle &= \langle 0 | e^{\frac{i}{\hbar} \varphi'^+ a} e^{\frac{i}{\hbar} \varphi a^\dagger} e^{[\varphi, \varphi^\dagger]} \varphi'^+ \varphi | 0 \rangle \\ &\quad \underbrace{\qquad}_{1+(-)\varphi^\dagger} \quad \underbrace{\qquad}_{1+(-)\varphi} \end{aligned}$$
 $= \langle 0 | 0 \rangle \exp(\varphi'^+ \varphi) \quad \checkmark$

Fermion:  $\langle \eta' | \eta \rangle = \langle 0 | (1 - \cancel{\eta \eta^\dagger})(1 - \eta \hat{a}^\dagger) | 0 \rangle$

 $= \langle 0 | 0 \rangle + \eta' \eta \langle 1 | 1 \rangle = 1 + \eta' \eta = \exp(\eta' \eta) \quad \checkmark$

Rev. of Identity:

$(|0\rangle - n|1\rangle)(\langle 0| - \eta^\dagger \langle 1|)$

fermions:  $\int d\eta^+ d\eta^- e^{-\frac{i}{\hbar} \eta^+ \eta^-} (1 - \eta \hat{a}^\dagger) |0\rangle \langle 0| (1 - \eta^\dagger \hat{a})$  number.

 $= \int d\eta^+ d\eta^- (1 - \eta^+ \eta^-) (|0\rangle \langle 0| + (\text{only one } \eta^+ \text{ or } \eta^-) + n \eta^+ |1\rangle \langle 1|)$ 
 $= \int d\eta^+ d\eta^- (|0\rangle \langle 0|) = \int d\eta^+ d\eta^- (-\eta^+ \eta^- |0\rangle \langle 0| + n \eta^+ |1\rangle \langle 1|) = [|0\rangle \langle 0| + |1\rangle \langle 1|] \quad \checkmark \quad (5)$

Bosons are actually harder:

$$\int \frac{d\phi^+ d\phi}{\pi} e^{-\phi^+ \phi} \sum_{n=0}^{\infty} \frac{\phi^n}{\sqrt{n!}} |n\rangle \langle n| \frac{\phi^{+m}}{\sqrt{m!}}$$

$\phi^+$  & the complex conjugate of  $\phi$ . So, go to radial, angular polar coordinates  
in the complex plane:  $d\phi^+ d\phi = |\phi| d|\phi| d\theta$  ( $\theta$  is the angle)

$$\begin{aligned} & \int \frac{d|\phi| d|\phi|}{\pi} \int_0^{2\pi} d\theta e^{-|\phi|^2} \sum_{n,m=0}^{\infty} \frac{|\phi|^{n+m}}{\sqrt{n! m!}} e^{i(n-m)\theta} |n\rangle \langle m| \\ &= \sum_{n=0}^{\infty} \int \frac{2|\phi| d|\phi|}{n!} |\phi|^{2n} e^{-|\phi|^2} |n\rangle \langle n| \\ &= \sum_{n=0}^{\infty} \underbrace{\int_0^{2\pi} \frac{d\theta}{n!} e^{-|\phi|^2} |\phi|^{2n}}_{=} |n\rangle \langle n| = \boxed{\sum_{n=0}^{\infty} |n\rangle \langle n|} \quad \checkmark \\ &= 1 \quad (\text{Gamma function}) \end{aligned}$$

# Path Integral: (Going to path)

$$Z = \text{Tr}[e^{-\beta \hat{H}}] = \sum_{n \geq 0} \langle n | e^{-\beta \hat{H}} | n \rangle$$

Want completeness:

$$= \int d(\phi^+, \phi) e^{-\phi^+ \phi} \sum_{n \geq 0} \underbrace{\langle n | \phi \rangle}_{\text{bra}} \underbrace{\langle \phi | e^{-\beta \hat{H}}}_{\text{operator}} |n\rangle$$

exchange (grassmann) variables:

$$\begin{aligned} &= \int d(\phi^+, \phi) e^{-\phi^+ \phi} \sum_{n \geq 0} \langle \pm \phi^+ | e^{-\beta \hat{H}} | n \rangle \langle n | \phi \rangle \\ \Rightarrow Z &= \boxed{\int d(\phi^+, \phi) e^{-\phi^+ \phi} \langle \pm \phi^+ | e^{-\beta \hat{H}} | \phi \rangle} \end{aligned}$$

\* Bosons return to same state; fermions to negative the state!

(6)

Now break up  $[0, \beta]$  into  $N$  small intervals  $\Delta t = \frac{\beta}{N}$

$$Z = \int d(\phi^+, \phi) e^{-\phi^+ \phi} \langle \pm \phi | e^{-\Delta t_N \hat{H}} e^{-\Delta t_{N-1} \hat{H}} \dots e^{-\Delta t_1 \hat{H}} |\phi \rangle$$

Identically again.

$$= \int d(\phi^+, \phi) d(\phi_N^+, \phi_N^-) e^{-\phi^+ \phi_N^-} \langle \pm \phi | e^{-\Delta t_N \hat{H}} | \phi_N^- \rangle$$

$$\langle \phi_N^- | e^{-\Delta t_{N-1} \hat{H}} \dots e^{-\Delta t_1 \hat{H}} |\phi \rangle$$

Now deal with each term, using property of coherent state:

$$e^{-\phi_i^+ \phi_i^-} \langle \phi_i^- | e^{-\Delta t \hat{H}[\hat{a}^+, \hat{a}^-]} | \phi_{i-1} \rangle$$

$$\simeq e^{-\phi_i^+ \phi_i^-} \langle \phi_i^- | \phi_{i-1} \rangle e^{-\Delta t H[\phi_i^+, \phi_{i-1}^-]}$$

$$= \exp\left\{-\phi_i^+ \phi_i^- + \phi_i^+ \phi_{i-1}^- - \Delta t H[\phi_i^+, \phi_{i-1}^-]\right\}$$

$$= \exp\left\{-\Delta t \left( \phi_i^+ \left( \frac{\phi_i^- - \phi_{i-1}^-}{\Delta t} \right) + H[\phi_i^+, \phi_{i-1}^-] \right)\right\}$$

Taking the limit of large  $N$  then gives: (and absorb the overall  $\phi^+ \phi$  variable introduced during the trace operator):

$$\Rightarrow Z = \int D\phi^+(\tau) D\phi(\tau) \exp\left\{-\int_0^\beta d\tau \left[ \phi^+(\tau) \frac{d\phi(\tau)}{d\tau} + H[\phi^+(\tau), \phi(\tau)] \right]\right\}$$

$$\phi^+(\beta) = \pm \phi(0)$$

$$\phi(\beta) = \pm \phi(0)$$

\* Legendre transform

For usual Hamiltonians, Legendre transform gives Euclidean Lagrangian

Going to fields now:

$$Z = \int D\phi^+ D\phi \exp\left\{-\int_0^\beta d\tau \int d^3x [L_E]\right\}$$

$$\phi^+(\beta) = \pm \phi(0)$$

$$\phi^-(\beta) = \mp \phi^+(\beta)$$

(7)

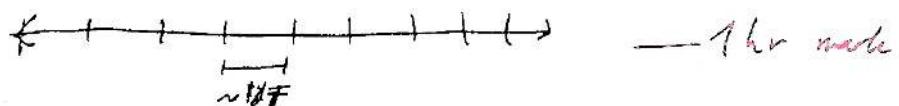
## Path Integral; Summary:

"Matsubara Frequencies"

- Compact "Time" interval ( $\beta \leftrightarrow it$ )  $\rightarrow$  discrete energies.
- Bosons periodic in imaginary time (energies  $w_n = \frac{2\pi ET}{\beta}$ )
- Fermions antiperiodic in imaginary time (energies  $w_n = \frac{2\pi ET}{\beta}(n + \frac{1}{2})$ )
- Path integral with Euclidean Lagrangian ( $t \rightarrow -it$ )  
 (Euclidean Lagrangian b/c the  $\phi^+ \frac{d\phi}{dt}$  term doesn't have a  $+i$  like in the real-time path integral.)

\* Note that "size" of compact + direction is inversely proportional to temperature  $T$ .

- high  $T$ , small + dimension
- low  $T$ , large + dimension again ( $T \rightarrow 0$  gives integral again, w/ continuous charges)



### # Free Bosonic Field

$$\mathcal{L} = \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - V(\phi) = \underbrace{-\frac{1}{2} (\partial_t \phi)^2}_{\text{Euclidean}} + \frac{1}{2} (\nabla \phi)^2 - V(\phi)$$

~~Start from this~~  $\sum_{w_n}$

Fourier of field:  $\phi(\tau, \vec{x}) = T \sum_{w_n} \frac{1}{\sqrt{v}} \sum_{\vec{k}} \tilde{\phi}(w_n, \vec{k}) e^{i w_n \tau - i \vec{k} \cdot \vec{x}}$   
 (little box)

[now we know integral]

$$\int d\phi_i d\phi_j \exp(-\frac{1}{2} \phi_i A_{ij} \phi_j) \propto [\det(A)]^{-\frac{1}{2}}$$

$$\exp(-\int \mathcal{L}_E) = \prod_n \exp\left(-\frac{1}{2v} \sum_m (w_n^2 + \vec{k}^2 + m^2) |\phi(w_n, \vec{k})|^2\right)$$

Product over  $\vec{k}$  space. The path integral then gives gaussian integrals, which can be computed as: (first for the  $\vec{k}=0$  harmonic oscillator)

$$(8) \quad Z_E^{HO} = \int d\phi e^{-S_E} = \prod_{n=-\infty}^{\infty} \frac{dx_n^2}{w_n^2} = \# T \prod_{n=-\infty}^{\infty} \frac{1}{(w_n^2 + w_n^2)^{1/2}} \quad (x)$$

Using  $\frac{\sinh \pi x}{\pi x} = \prod_{n=1}^{\infty} \left(1 + \frac{x^2}{n^2}\right)$ , can show this is: ( $\omega \in \mathbb{R}^3+m^2$ )

$$Z_{HO}^{HO} = \exp\left(-\frac{1}{T} \left[ \frac{\omega}{2} + T \ln \left( 1 - e^{-\beta \omega} \right) \right] \right) = \exp\left(-\frac{E_{HO}}{T}\right).$$

$$\Rightarrow Z^{SFT} = \prod_u Z^{HO}$$

Thus, the free energy can be written in the form (for density)

Thus, the free energy of the SFT can be written! (in 3 dimensions)

$$\lim_{V \rightarrow \infty} \frac{F^{SFT}}{V} = \int \frac{d^3 k}{(2\pi)^3} \left[ \frac{\omega_k^0}{2} + T \ln \left( 1 - e^{-\beta \omega_k} \right) \right]$$

with  $\omega_k = \sqrt{\omega^2 + k^2}$ . Note that the expression (\*) on p. 8 also tells

$$\lim_{V \rightarrow \infty} \frac{F}{V} = \lim_{V \rightarrow \infty} \left( -\frac{T}{V} \ln Z \right) = + \lim_{V \rightarrow \infty} \frac{T}{V} \sum_{u, w_n} \ln \left( \frac{e^{\beta \omega_n^0}}{2} \right) + \text{const.}$$

$$= \left[ T \sum_{u, n} \int \frac{d^3 k}{(2\pi)^3} \ln \left( \omega_n^0 + k^2 + m^2 \right) + \text{const.} \right]$$

We usually write this directly  $\oint \equiv T \sum_n \int \frac{d^3 k}{(2\pi)^3} \left( \frac{e^{\beta \omega_n^0}}{4\pi} \right)^2$   $\leftarrow$  factor to simplify  
dim-reducing expression.

We will frequently encounter sum-integrals ... (MM are back to the momenta)

## # Free Dirac Field

Minkowski Lagrangian (density):  $L_M = \bar{\Psi} (\gamma_m^\mu \partial_\mu + m) \Psi$

$$\bar{\Psi} \equiv \Psi^+ \otimes \gamma_0^0, \quad m \equiv m, \quad \text{if } 4 \times 4, \quad \{\gamma_m^\mu, \gamma_n^\nu\} = 2\eta^{\mu\nu}, \quad (\gamma_m^\mu)^+ = \gamma_n^0 \delta_m^{\mu} \gamma_n^0$$

$$\stackrel{\text{t}}{\rightarrow} (-1, +1, +1, +1), \quad (\gamma_m^\mu = i \gamma_{m, \text{other}})$$

Carrying out the Legendre transformation, we will find:

$$L_E = L_M (\tau = it) = \bar{\Psi} (\gamma_E^\mu \partial_\mu + m) \Psi$$

$$\gamma_E^0 = -i \gamma_M^0, \quad \gamma_E^i = \gamma_M^i, \quad \{\gamma_E^\mu, \gamma_E^\nu\} = 2\delta^{\mu\nu} \leftarrow \text{can raise & lower indices w/o } (-1) \text{ reversal!}$$

(9)

Fourier space for fermions is similar, we  $\{\ell\}$  to denote fermionic Matsubara component of  $\rho^0 = \frac{1}{\pi} \omega_n^f = 2\pi T(n + \frac{1}{2})$   $n \in \mathbb{Z}$ .

$$\psi(x) = \sum_{\{\ell\}} e^{i\ell \cdot x} \tilde{\psi}(\ell); \quad \tilde{\psi} \text{ is expected } (-i\ell; x)$$

This gives:  $S_E = \sum_{\{\ell\}} \tilde{\psi}(\ell) [i\ell + m] \tilde{\psi}(\ell).$

Using the Fermion Gaussian integral:

$$\int D\psi_i D\bar{\psi}_i e^{-\frac{1}{2} \sum_{\{\ell\}} \bar{\psi}_i A_{ij} \psi_j} = \det(A)$$

Hence, we arrive at:

$$\begin{aligned} Z &= \text{Tr}[e^{-S_E}] = \# \prod_{\{\ell\}} \det[i\ell + m] \quad \xrightarrow{\text{product sym. in } \ell \rightarrow -\ell} \\ &= \# \left( \prod_{\{\ell\}} \det[i\ell + m] \det[-i\ell + m] \right)^{1/2} \\ &= \# \left( \prod_{\{\ell\}} \det[(\ell^2 + m^2) \pi_{4 \times 4}] \right)^{1/2} \\ &= \# \prod_{\{\ell\}} (\ell^2 + m^2)^2 \end{aligned}$$

Therefore

$$\lim_{V \rightarrow \infty} \frac{F}{V} = \lim_{V \rightarrow \infty} \left( -\frac{T}{V} \ln(Z) \right) = -\lim_{V \rightarrow \infty} \frac{T}{V} \times 2 \sum_{\{\ell\}} \ln(\ell^2 + m^2) + \text{const.}$$

$$= -4 \sum_{\{\ell\}} \frac{1}{2} \ln(\ell^2 + m^2) + \text{const.} \quad = -4 \times \text{(bare result)} \\ \text{(but for } \sum_{\{\ell\}} \text{)}$$

\* Pressure to \* Free energy or  $-1$  of bare result (as expected)

- note that came from gaussian ~~integral~~ Gaussian

\*  $4 \times$  b/c 4 DOF of one field!

## # Manipulating bosonic + fermionic sum-integrals.

Lecture 1/2

Consider the following <sup>boson</sup> thermal sum, with  $f(p)$  a meromorphic function in the complex plane:

$$\sigma_b \equiv T \sum_{w_n} f(w_n). \quad (\star)$$

We seek a way to evaluate this expression. Introduce an auxiliary function

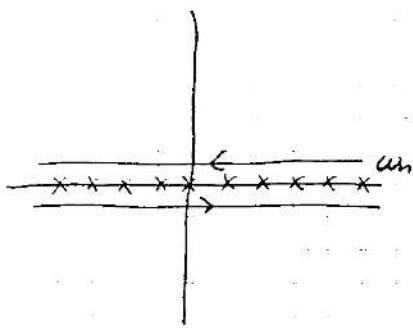
$$n_B(i\beta p) \equiv \frac{i}{\exp(i\beta p) - 1}, \quad n_B \text{ has poles.}$$

This function has poles where  $\exp(i\beta p) = 1$ ; that is at  $i\beta p = 2\pi n \quad n \in \mathbb{Z}$  Matsubara frequencies! Moreover, the points near one of these poles  $w_n + z$ : ( $p = w_n$ ).

$$n_B[i(w_n + z)] = \frac{i}{\exp(i\beta[w_n + z]) - 1} = \frac{i}{\exp(i\beta z) - 1} \\ \simeq \frac{T}{z} + O(1).$$

Hence, the residue at each pole is  $T$ . This means we can replace the sum in  $(\star)$  by the <sup>contour</sup> complex integral!

$$\sigma_b = \oint \frac{dp}{2\pi i} f(p) n_B(i\beta p) = \\ = \int_{-\infty - i0^+}^{+\infty - i0^+} \frac{dp}{2\pi i} f(p) n_B(i\beta p) \\ + \int_{+\infty + i0^+}^{-\infty + i0^+} \frac{dp}{2\pi i} f(p) n_B(i\beta p).$$



Here  $\text{sup } p \rightarrow -p$ , and use

$$\boxed{n_B(-ip) = \frac{1}{\exp(-i\beta p) - 1}} = \frac{\exp(i\beta p) - 1 + 1}{1 - \exp(i\beta p)} \\ = \boxed{-1 - n_B(ip)}$$

$$\Rightarrow \boxed{\sigma_b = \int_{-\infty - i0^+}^{+\infty - i0^+} \frac{dp}{2\pi i} \left\{ f(-p) + [f(p) + f(-p)] n_B(ip) \right\}}$$

$$= \boxed{\int_{-\infty}^{\infty} \frac{dp}{2\pi} f(p) + \int_{-\infty - i0^+}^{+\infty - i0^+} \frac{dp}{2\pi i} [f(p) + f(-p)] n_B(ip)}$$

↑  
vacuum

↑  
thermal contributions

Bosonic  
thermal  
sum

(11)

~~Another~~ Inspecting the thermal contribution, in the LHP:

$$|n_\beta(ip)|^{\rho = x - iy} = \left| \frac{1}{e^{i\beta y} e^{\beta y} - 1} \right|^{\gamma > T} \approx e^{-\beta y} \approx e^{-\beta |p|}$$

Hence, if  $f(p)$  grows slower than exponentially (slower than  $e^{\beta|p|}$ ) at large  $|p|$  — such as polynomially — the thermal contribution can be closed in the lower-half  $\ell$  plane,  $\Rightarrow$  picking up the poles and residues of  $f(p) + f(-p)$ .

$\Rightarrow$  [Physically, thermal contribution from "on-shell" particles]. 2 hr mark.

Fermionic thermal sums can actually be related back to the bosonic ones above, by being clever:

$$\sigma_f = T \sum_{\{w_n\}} f(w_n) \quad (1)$$

$$\begin{aligned} \sigma_f(T) &= T \left[ \dots + f(-3\pi T) + f(-\pi T) + f(\pi T) + \dots \right] \\ &\leftarrow T \left[ \dots + f(-3\pi T) + f(-2\pi T) + f(\pi T) + f(2\pi T) + \dots \right] \\ &\leftarrow T \left[ \dots + f(2\pi T) + f(2\pi T) + \dots \right] \\ &= T \left[ \dots + f(-3\pi T) + f(-2\pi T) + f(-\pi T) + f(0) + f(\pi T) + f(2\pi T) + \dots \right] \\ &\leftarrow T \left[ \dots + f(2\pi T) + f(0) + f(2\pi T) + \dots \right] \\ &= 2 \cdot \frac{T}{2} \left[ \dots + f(-6\pi \frac{T}{2}) + f(-4\pi \frac{T}{2}) + f(-2\pi \frac{T}{2}) + f(0) + \dots \right] \\ &\leftarrow T \left[ \dots + f(2\pi T) + f(0) + \dots \right] \end{aligned}$$

$$\Rightarrow \boxed{\sigma_f(T) = 2\sigma_b(\frac{T}{2}) - \sigma_b(T)}$$

Thus, the general fermionic sum will be derived as above w/ the auxiliary function:

$$\text{[REDACTED]} = 2n_\beta^{(T)}(ip) - n_\beta^{(T)}(ip)$$

Verbal  
2.5

(2)

This can be reduced as:

$$\begin{aligned}
 2n_B^{(\pi)}(ip) - n_B^{(\tau)}(ip) &= \frac{2}{\exp(2ip\beta) - 1} - \frac{1}{\exp(ip\beta) - 1} \\
 &= \frac{1}{\exp(ip\beta) - 1} \left[ \frac{2}{\exp(ip\beta) + 1} - 1 \right] \\
 &= \frac{1}{\exp(ip\beta) - 1} \left[ (-1) \frac{1 - \exp(ip\beta)}{\exp(ip\beta) + 1} \right] \\
 &= \boxed{-n_F^{(\tau)}(ip)}, \quad n_F(ip) \text{ fermi-dense dist.}
 \end{aligned}$$

Therefore, Fermi-Dirac thermal sums are evaluated by:

$$O_f(T) = \int_{-\infty}^{+\infty} \frac{dp}{2\pi} f(p) - \int_{-\infty - i0^+}^{+\infty + i0^+} \frac{dp}{2\pi} [f(p) + f(-p)] n_F(ip)$$

Fermi-Dirac thermal sum. (if added  $-1 \times$  boson, w/  $n_B \rightarrow n_F$ ).

Example of a thermal sum: Pressure of a scalar field.

$$\frac{F}{V} = \oint \frac{1}{p} \ln(p^2 + m^2) = \oint \frac{1}{p} \ln(\omega_n^2 + E_p)^2 \equiv \int \frac{j(E_p)}{p} \equiv \sqrt{p^2 + m^2}$$

Let's evaluate the generalized integral

$$j(E, c) \equiv T \sum \frac{1}{\omega_n (w_n + c)^2 + E^2}, \quad c \in \mathbb{C}$$

Note that, defining

~~$$j(E, c) \equiv \frac{T}{2} \sum \ln((\omega_n^2 + c)^2 + E^2)$$~~

~~$$\frac{d(j(E))}{E} \neq \frac{d(j(E))}{E} \frac{d(E^2)}{E^2} \ln((\omega_n^2 + E^2)^2) = \frac{1}{\omega_n^2 + E^2}$$~~

~~$$\frac{1}{E} \frac{dj(E, c)}{dE} = \frac{1}{\omega_n^2 + E^2} T \sum \frac{1}{\omega_n^2 + E^2}$$~~

so if we evaluate this sum to get the answer for  $j(E, 0) \equiv j(E)$ .

Taking: (In real part)

$$f(p) = \frac{1}{(p+c)^2 + E^2} = \frac{1}{2E} \left[ \frac{1}{p+c+iE} - \frac{1}{p+c-iE} \right]$$

(13)

$$(In T > 0 part) \\ f(p) + f(-p) = \frac{1}{2E} \left[ \frac{1}{p+c+iE} + \frac{1}{p-c+iE} - \frac{1}{p+c-iE} - \frac{1}{p-c-iE} \right]$$

The poles in the LHP are at  $p = \pm c - iE$ , with residue  $\frac{i}{2E}$ .

the Vacuum term:  $\int_{-\infty}^{\infty} \frac{dp}{2\pi} f(p) = \frac{1}{2\pi} (-2\pi i) \left( \frac{i}{2E} \right) = \frac{1}{2E}$

Thermal term:  $\int_{-\infty-i0^+}^{\infty-i0^+} \frac{dp}{2\pi} [f(p) + f(-p)] = \frac{1}{2\pi} (-2\pi i) \left( \frac{i}{2E} \right) \left[ n_B(i(c-iE)) + n_B(i(c+iE)) \right]$

$$\Rightarrow i(E; c) = \frac{1}{2E} [1 + n_B(E-iC) + n_B(E+iC)].$$

Now observe that

$$n_B(x) = \frac{1}{e^x - 1} = \frac{e^{-x}}{1 - e^{-x}} = \frac{d}{dx} [\ln(1 - e^{-x})]$$

Hence, we can derive

$$j(E, c) = \text{const.} + \frac{E}{2} + \frac{T}{2} \{ \ln[1 - e^{\beta(E-iC)}] + \ln[1 - e^{-\beta(E+iC)}] \}$$

Setting  $c \rightarrow 0$  reproduces the boson pressure of a single mode!

Some notes:

- \* For regularization purposes, sometimes easier to derive f of derivatives — ~~too difficult to calculate by hand~~
- Otherwise, Dimensional regularization is a great tool.

- \* Note that  $iC \rightarrow -\mu$  here looks like the result with a chemical potential. In other words, chemical potential enters as

$$[w_n \mapsto w_n - i\mu] \quad \text{chemical potential} \quad * \text{The in general}$$

~~handwritten notes~~ ~~for this part~~ ~~for this part~~

~~which is  $\mapsto (w_n, \omega_n)$  etc~~

## # Interacting Field Theory (Perturbation theory)

Lecture 2

Our time with free fields has ended. Now we want to consider a theory with a nonzero interaction term to the action:

$$S = S_0 + S_I$$

We want to expand around the free action (perturbatively)

$$e^{-\frac{F_0}{V}} = \int D\phi e^{-S_0[\phi] - S_I[\phi]} \approx \int D\phi e^{-S_0[\phi]} (1 - S_I + \frac{1}{2} S_I^2 + \dots)$$

$$\Rightarrow \boxed{F = -\frac{1}{V} \ln Z = -T \ln \left[ \int D\phi e^{-S_0[\phi]} (1 - S_I + \frac{1}{2} S_I^2 + \dots) \right]}$$

$$= F_0 - T \ln \left[ \langle 1 - S_I + \frac{1}{2} S_I^2 + \dots \rangle \right]$$

$$\uparrow \boxed{\langle \phi \rangle_0 = \frac{\int D\phi \phi e^{-S_0}}{\int D\phi e^{-S_0}}}$$

$$\approx F_0 + \left( \langle S_I \rangle_0 - \underbrace{\frac{1}{2} [\langle S_I^2 \rangle_0 - \langle S_I \rangle_0^2]}_{\text{connected.}} + \dots \right) \frac{T}{V}$$

$$\Rightarrow \boxed{F = F_0 + \langle S_I - \frac{1}{2} S_I^2 + \dots \rangle_{0,c} \frac{S_I}{V}}$$

Sum over connected components only (drop overall  $S_I$  to remove  $\frac{T}{V}$  prefactor).

To compute these connected components, we use Wick's theorem (derived by analogy to source and perturbativity)

$$\text{Write } S_0[\phi] = \sum_{x,y} \frac{1}{2} \phi_x A_{xy} \phi_y$$

~~Define~~ Introduce a source Define

$$Z[J] \equiv \int D\phi \exp \left[ \sum_{x,y} \left( -\frac{1}{2} \phi_x A_{xy} \phi_y + J_x \phi_x \right) \right] \quad \begin{aligned} &\text{Complete square} \\ &\text{shift } \phi \text{ by constant.} \end{aligned}$$

$$= \exp \left[ \sum_{x,y} \frac{1}{2} J_x A_{xy}^{-1} J_y \right] Z[0]$$

(15)

Then  $\boxed{\langle \varphi_1 \varphi_2 \dots \varphi_n \rangle_{0,c}} = \frac{1}{2[0]} \frac{\delta}{\delta J_1} \frac{\delta}{\delta J_2} \dots \frac{\delta}{\delta J_n} Z[J] \Big|_{J=0}$

$$= \frac{\delta}{\delta J_1} \frac{\delta}{\delta J_2} \dots \frac{\delta}{\delta J_n} \left( 1 + \int_{a,b} \frac{1}{2} J_a A_{ab}^{-1} J_b + \frac{1}{2} \int_{a,b} \frac{1}{2} J_a A_{ab}^{-1} J_b \int_{c,d} \frac{1}{2} J_c A_{cd}^{-1} J_d + \dots \right)$$

$$= \boxed{\sum_{\text{all comb. } c} A_{1,i}^{-1} A_{i,j}^{-1} \dots A_{n,l}^{-1} A_{l,N}^{-1}}$$

Example a scalar field theory:  $(\phi^4)$

$$f = \frac{F}{V}; \quad f_{00} = f_0 + f_1 + f_2 + \dots; \quad \boxed{L = \frac{1}{2} \partial_\mu \phi \partial^\mu \phi + \lambda \int_0^\beta d\tau \int \frac{1}{4!} \phi^4}$$

$$\boxed{f_1 = \frac{F}{V} \langle S_{\tau,0} \rangle = \cancel{\frac{F}{V}} \int_0^\beta d\tau \int_x \cancel{\frac{1}{4!}} \underbrace{\lambda \langle \phi(x) \phi(x) \phi(x) \phi(x) \rangle}_\text{Trans. invariant.}}$$

$$= \frac{\lambda}{4!} \langle \phi_0 \phi_0 \phi_0 \phi_0 \rangle_0 = \cancel{\left(\frac{3}{4!}\right)} \lambda \langle \phi_0 \phi_0 \rangle_0 \langle \phi_0 \phi_0 \rangle_0$$

$$= \boxed{\frac{1}{8} \infty}$$

$f_2$  proceeds similarly:

$$\boxed{f_2 = \left(\frac{\lambda}{4!}\right)^2 [12 \int_x \langle \phi_x \phi_0 \rangle_0^4 + 36 \langle \phi_0 \phi_0 \rangle_0^2 \int_x \langle \phi_x \phi_0 \rangle_0^2]}$$

$$= \boxed{\frac{1}{48} \infty + \frac{1}{16} \infty \infty}$$

The only combination appearing in these expressions are  $\langle \phi_x \phi_y \rangle_0$  — the propagator.  
Let's look a little more at this function.

# The (scalar) propagator:

\* remember  
 $\tau \in [0, \beta]$

$$\langle \phi_x \phi_y \rangle_0 = \boxed{\int_p \frac{e^{ip \cdot (x-y)}}{p^2 + m^2}}$$

How do the limits of this compare to  $\text{Feynman theory? Consider}$

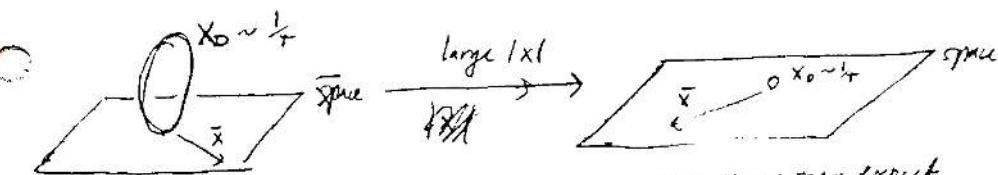
(16).  $\langle \phi_x \phi_0 \rangle$  w/o G:

UV limit:  $|x|, |x_0| \ll \frac{1}{T}, \frac{1}{m}$ . Expect dominant regime from contour representation from  $|\bar{p}|/|x| \sim 1, |\bar{p}^0|/|x| \sim 1 \Rightarrow (\bar{p}_\parallel/\bar{p}^0 \gg T, m)$  Hence, from the sum-angle perspective  $T \ll \text{small} \Rightarrow \int_{\bar{p}} \rightarrow \int_{\bar{p}} (i \sum_n + \frac{1}{\bar{p}})$

$$\Rightarrow \langle \phi_x, \phi_0 \rangle_0 \rightarrow \int_{\bar{p}} \frac{e^{i\bar{p} \cdot \bar{x}}}{{\bar{p}}^2 + m^2} \quad \text{vacuum result}$$

$$\frac{i\bar{p}}{\bar{p}^2} \sim (\text{mom})^2 \sim \frac{1}{|x|^2}$$

IR limit:  $|x| \gg \frac{1}{T}$ . In this case, ~~we take~~ the parallel tangent component is always small (at most  $\frac{1}{T}$ ). Let's take  $T$  large also, to make things clearer.



~~we expect~~ Since  $w_n \gg 1$ , only the zero mode contributes to the long-distance physics. (Can show this by doing the zero-mode computation)

$$\Rightarrow \langle \phi_x \phi_0 \rangle_0 \simeq T \int_{\bar{p}} \frac{e^{i\bar{p} \cdot \bar{x}}}{\bar{p}^2 + m^2} \sim T \cdot (\text{mom})^2 = T |\bar{x}|^{-1}$$

can now do this integral:  $z \equiv \bar{p} \cdot \bar{x} / (p x)$

$$= \frac{T}{(2\pi)^2} \int_{-1}^{+1} dz \int_0^\infty dp p^2 \frac{e^{ipx z}}{p^2 + m^2}$$

$$= \frac{T}{(2\pi)^2} \int_0^\infty \frac{dp p^2}{p^2 + m^2} \frac{e^{ipx} - e^{-ipx}}{ipx}$$

$$= \frac{T}{(2\pi)^2 x} \int_{-\infty}^\infty \frac{dp p^2 e^{ipx}}{p^2 + m^2}$$

$$\boxed{\langle \phi_x \phi_0 \rangle_0 = \left[ \frac{T e^{-mx}}{4\pi x} \right], \quad x \gg \frac{1}{T}} \quad (d=3 \text{ we would expect } \sim x^{-2})$$

\* IR physics modifies the propagator in an essential way.

\* ~~eff~~ prop  $\sim 1/\text{mass}^{d-2} = 1/\text{mass}^{d-1}$

UV: all dimensions same vacuum:  $\sim |x|^{-2}$

IR: one dim ( $x_0$ )  $\sim \frac{1}{T} \Rightarrow \sim T |x|^{-1}$

*17* (17)

## ~~WKB Approximation & Renormalization~~ # UV & IR divergences

This qualitative modification of the propagator & the IR will lead to some very interesting physics. Let us first, however, look at a few types of divergences appearing in the calculation. We have:

$$f = \int_{\vec{u}} J(m, T) + \frac{1}{8} \lambda [I(m, T)]^2 + O(\lambda^2)$$

$$(2) \quad J(m, T) = \int_{\vec{u}} \left[ \frac{E_u}{2} + T \ln(1 - e^{-\beta E_u}) \right] \quad (= \oint_{\vec{u}} \frac{1}{2} \ln(u^2 + m^2) + \text{const.})$$

$$(1) \quad I(m, T) = \frac{1}{m} \frac{d}{dm} J(m, T) = \oint_{\vec{u}} \frac{1}{2E_u} [1 + 2n_B(E_u)] \\ (= \oint_{\vec{u}} \frac{1}{B^2 + m^2}) \quad \text{can't be done for general } m,$$

Let's work at high-T to simplify the expression. We'll do it in a slightly hasty way, because the terms are not well-behaved for  $m \rightarrow 0$  (and in particular  $m/T \rightarrow 0$ ).

We will ~~first split~~: Evaluate  $I(m, T)$  by first doing  $\int_{\vec{u}}$  and then  $\sum_{wn}$ , then use:

$$J(m, T) = \int_0^m dm' m' I(m', T) + J(0, T)$$

$$J(0, T) = \frac{T^4}{2\pi^2} \int_0^\infty dx x^2 \ln[1 - \exp(-\sqrt{x^2})] = -\frac{\pi^2 T^4}{90} \quad (\text{numerical})$$

to get  $J(m, T)$ . Evaluate  $I(m, T)$ :

$$1.) \quad \oint_{\vec{u}} = \oint_{\vec{u}'} + T \int_{\vec{u}} \underbrace{\dots}_{\equiv \Phi(m, d, A)}$$

$$2.) \quad I^{(n=0)} = T \int_{\vec{u}} \frac{1}{u^2 + m^2} = T \int_{\vec{u}} \frac{1}{(u^2 + m^2)^A} \quad \text{for } A=1$$

3.) Dam reg results

$$a.) \quad d\vec{u} = \frac{\pi^{d/2}}{\Gamma(d/2)} (u^2)^{\frac{d-2}{2}} d(u^2)$$

$$b.) \quad u^2 \mapsto m^2 t \quad (\text{scale out } m):$$

$$(18) \quad \Phi(m, d, A) = \frac{m^{d-2A}}{\Gamma(d/2)} \frac{\pi^{d/2}}{(2\pi)^d} \int_0^\infty dt t^{\frac{d-2}{2}} (1+t)^{-A} \left( \frac{e^{\frac{m^2}{4t}}}{4\pi} \right)^{\frac{d}{2}}$$

It is now in the form of an Euler Beta function. Changing  $t \rightarrow 1/s - 1$ ,  $dt \mapsto -ds/s^2$  gives an Euler Beta function.

$$\Phi(m, d, A) = \frac{m^{d-2A}}{(4\pi)^{d/2} \Gamma(d/2)} \int_0^1 ds s^{A-\frac{d}{2}-1} (1-s)^{\frac{d}{2}-1} \left( \frac{e^{\gamma_E} \pi^2}{4\pi} \right)^s$$

$$= \frac{(m^2)^{\frac{d}{2}-A}}{(4\pi)^{d/2}} \frac{\Gamma(A-\frac{d}{2})}{\Gamma(A)} \left( \frac{e^{\gamma_E} \pi^2}{4\pi} \right)^s$$

Asymptotic result leads to  $\Phi(m, d, A) \approx \frac{1}{2} \Gamma(A) \Gamma(A-d)$

(4)  $d=3-2\epsilon$  and expanding in  $\epsilon$ .

Result:  $I^{(n=0)} = T \Phi(m, 3-2\epsilon, 1) = \left[ -\frac{Tm}{4\pi} + O(\epsilon) \right]$

$$\left( \Rightarrow I^{(n=0)} = -\frac{Tm^3}{12\pi} \right)$$

$\uparrow$   
odd powers of  $m$   
from  $0 \times 2$  nibbles  
modus

Residuum term:

$$I'(m, T) = T \sum_{w_n} \int \frac{1}{w_n^2 + h^2 + m^2} \approx \sum_{n=-\infty}^{\infty} \int \frac{1}{(2\pi n T)^2 + h^2 + m^2}$$

~~$$h^2 \mapsto (2\pi T)^2 h^2$$~~

$$= T^{d-1} \sum_{n=-\infty}^{\infty} \int \frac{1}{h^2 + [m/(2\pi T)]^2 + n^2}$$

~~$$= T^{d-1} \sum_{n=-\infty}^{\infty} \Omega \left( \frac{1}{(2\pi n)^2} \right)$$~~

~~$$= T (2\pi T)^{d-2} \sum_{n=-\infty}^{\infty} \int \frac{1}{h^2 + n^2 + [m/(2\pi T)]^2}$$~~

~~$$= 2T^{d-1} (2\pi)^{d-2} \sum_{n=1}^{\infty} \Omega(n, d, m, T)$$~~

(11)

$$\begin{aligned} I'(m, T) &= 2T^{d-1} (2\pi)^{d-2} \left[ \sum_{n=1}^{\infty} \frac{B_{2n} n^{d-2}}{(2\pi n)^{d-1}} \right] \frac{\Gamma(1 - \frac{1}{2})}{\Gamma(d)} \frac{1}{(4\pi)^{d/2}} \\ &= \frac{T^2}{12} + O(m^2) + O(\epsilon). \end{aligned}$$

$$I'(m, T) = 2T (2\pi T)^{d-2} \sum_{n=1}^{\infty} \sum_{\ell=0}^{\infty} \frac{[m/(2\pi T)]^{2\ell}}{[n^2 + \ell^2]^{\ell+1}} (-1)^\ell$$

$$\sum_{\ell=0}^{\infty} \int_0^{\infty} \left[ \frac{1}{u} \right]$$

$$= \overbrace{\Phi(n, d, 1)}^{\sim} \sim \frac{1}{n^{2\ell+2-d}}$$

$$= \frac{2T \left( \frac{e^{\chi E} T^3}{4\pi} \right)^\varepsilon}{(4\pi)^{d/2} (2\pi T)^{d-d}} \sum_{\ell=0}^{\infty} \left[ \frac{-m^2}{(2\pi T)^2} \right]^{\ell} \frac{\Gamma(\ell+1 - \frac{d}{2})}{\Gamma(\ell+1)} \zeta(2\ell+2-d)$$

$$\Rightarrow I'(m, T) = \frac{T^2}{12} - \frac{2m^2}{(4\pi)} \left[ \frac{1}{2\varepsilon} + \ln \left( \frac{\bar{\chi} e^{\chi E}}{4\pi T} \right) \right] + \frac{2m^4 \zeta(3)}{(4\pi)^4 T^2} + O\left(\frac{m^6}{T^4}\right) + O(\varepsilon)$$

+ zero mode

$$\Rightarrow I(m, T) = \frac{T^2}{12} - \frac{mT}{4\pi} - \frac{2m^2}{(4\pi)^2} \left[ \frac{1}{2\varepsilon} + \ln \left( \frac{\bar{\chi} e^{\chi E}}{4\pi T} \right) \right] + \dots$$

And: Recalling the  $T=0$  expression  $I_0(m) \equiv \int \frac{1}{2E_k}$ , we would

find:

$$I(m, T) = \underbrace{-\frac{m^2}{8\pi n^2} \left[ \frac{1}{\varepsilon} + \ln \frac{\bar{\chi}^2 e^{2\chi E}}{m^2} + 1 \right]}_{\text{vacuum}} + I_T(m) + \underbrace{\dots}_{\text{thermal}}$$

$$I_T(m) \approx \frac{T^2}{12} - \frac{mT}{4\pi} + \frac{2m^2}{(4\pi)^2} \left[ \ln \left( \frac{m e^{\chi E}}{4\pi T} \right) - \frac{1}{2} \right] + O\left(\frac{m^4}{T^2}\right) + O(\varepsilon)$$

And we find:

$$(20.) \quad J(m, T) = \frac{-\pi^2 T^4}{90} + \frac{m^2 T^2}{24} - \frac{m^3 T}{12\pi} - \frac{m^4}{2(4\pi)^2} \left[ \frac{1}{2\varepsilon} + \ln \left( \frac{\bar{\chi} e^{\chi E}}{4\pi T} \right) \right].$$

But then we see something funny happening with the  $\mathcal{O}(\lambda)$  term of  $f$ :

$$f_{(1)} = \frac{1}{8} \lambda [I(m, T)]^2 \Rightarrow (\text{T-dependence}) / (\frac{1}{\epsilon} \text{ divergence})$$

This doesn't look like massless renormalization, so we'll have to look into this. ~~\* UV divergence~~

Now consider part of the  $f_{(2)}$  contribution:

$$f_{(2)} \ni \frac{1}{16} \text{ (loop diagram)} = \frac{1}{16} \lambda^2 \left[ \underbrace{\frac{1}{\epsilon}}_{\text{from } I} \frac{1}{[\rho^2 + m^2]^2} \right] [I(m, T)]^2$$

$$\sim m \rightarrow 0, \rightarrow \frac{1}{\epsilon} \frac{1}{[\rho^2]^2} \sim \frac{1}{\rho^4} \text{ in the zero mode}$$

~~fix it if present.~~

Can also see this from

$$\frac{1}{\epsilon} \frac{1}{[\rho^2 + m^2]^2} = -\frac{d}{dm^2} I(m, T) = -\frac{1}{2m} \frac{d}{dm} \left( -\frac{mT}{4\pi} + \dots \right)$$

$$= \frac{1}{8\pi m} \rightarrow \infty \text{ as } m \rightarrow 0.$$

comes from  $n=0$   
Nambu-Goldstone mode

But plasma at ~~massless~~ massless particles should have well-defined thermodynamics

What's going on here?

~~\* IR divergence~~

end lecture 2

## # ~~Without~~ Renormalization: (UV)

Renormalization actually does work the same as at  $T=0$ , we need to consider coupling according to the bare mass parameter  $m_B$ :

$$m_B^2 = m_R^2 + \lambda_R \delta m_R^2$$

Then

$$\begin{aligned} \phi(T) &= \phi^{(0)}(m_B^2 + T) + \lambda_B \phi^{(1)}(m_B^2, T) \\ &= \phi^{(0)}(m_R^2, T) + \lambda_R \left[ \phi^{(1)}(m_R^2, T) + \frac{\partial \phi^{(0)}}{\partial m_R^2} \delta m_R^2 \right] + \text{term} \end{aligned} \quad \begin{matrix} \checkmark \\ \text{This term will remove the funny behavior.} \end{matrix} \quad (21)$$

One can perform the renormalization by using the on-shell scheme, by performing the usual 1-PI renormalization of the propagator

$$(1.) \quad \boxed{\langle \phi \phi \rangle_{\text{ext}} = [\langle \phi \phi \rangle_0 + \Pi]^{-1}}$$

$$= \frac{1}{P^2 - m^2 + i\Gamma} + \frac{i\Gamma}{P^2 - m^2 + i\Gamma} = \frac{1}{P^2 - m^2 + i\Gamma} \rightarrow \langle \phi \phi \rangle_{\text{ext}}^{-1} - \langle \phi \phi \rangle_0^{-1} = \Pi$$

$$\langle \phi \phi \rangle_{\text{ext}} = \langle \phi \phi \rangle_0 + \Pi \rightarrow \langle \phi \phi \rangle_{\text{ext}} = \langle \phi \phi \rangle_0 [1 + \langle \phi \phi \rangle_0 \Pi]$$

$$\rightarrow \langle \phi \phi \rangle_{\text{ext}} - \langle \phi \phi \rangle_0 = \langle \phi \phi \rangle_0 \Pi$$

$$\simeq \langle \phi \phi \rangle_0 - \langle \phi \phi \rangle_0 \Pi \langle \phi \phi \rangle_0 \quad \left[ \frac{\langle \phi \phi e^{-s_z} \rangle}{\langle e^{-s_z} \rangle} = \langle \phi \phi \rangle_0 \right] \langle \phi \phi (e^{-s_z} - 1) \rangle_0 = -\langle \phi \phi \rangle_0 \Pi \langle \phi \phi \rangle_0 \quad \text{Explain 2x!}$$

$$(2) \Rightarrow \langle \phi \phi \rangle_0 \Pi \langle \phi \phi \rangle_0 = \langle \phi S_z \phi \rangle_c + \text{HOTs} \quad \text{2x!} \quad \boxed{\langle \phi \phi \rangle_0 \Pi \langle \phi \phi \rangle_0 = \langle \phi S_z \phi \rangle_c + \text{HOTs}}$$

$$\left. \begin{aligned} & \text{Perform} \\ & \varphi = 0 \\ & \text{renorm.} \end{aligned} \right\} \quad \langle \phi \left( \frac{\lambda}{4!} \phi_x \phi_x \phi_x \phi_x \right) \phi \rangle_{0,c} = \frac{4.3\lambda}{4!} \langle \phi \phi \rangle_0 \langle \phi_x \phi_x \rangle_0 \langle \phi_x \phi_x \rangle_0 \propto m^2 \\ \Rightarrow \quad \boxed{\Pi} & = \frac{\lambda}{2} \int \frac{1}{P^2 + m^2} = \frac{\lambda}{2} I_0(m) \quad \underline{\text{Vac Adjnt}}$$

Hence,  $\langle \phi \phi \rangle_c = \int \frac{1}{P^2 + m_B^2 + \frac{\lambda}{2} I_0(m_B)} + \mathcal{O}(\lambda^2)$ . Vac Adjnt

$$\Rightarrow m_B^2 = m_B^2 + \frac{\lambda}{2} I_0(m_B) \Rightarrow m_B^2 = m_{\text{pole}}^2 - \frac{\lambda_R}{2} I_0(m_{\text{pole}})$$

We can now plug this into the free energy:  $\frac{d}{dm} f(m) = \frac{1}{2m} \frac{d}{dm} f(m)$

$$\begin{aligned} f(T) &= J(m_B, T) + \frac{1}{8} \lambda_B [I(m_B, T)]^2 \quad \leftarrow = \frac{1}{2} I_0(m_B) \cancel{f(m)} \\ &= J(m_p, T) + (m_B^2 - m_p^2) \underbrace{\frac{\partial J(m_p, T)}{\partial m^2}}_{\cancel{m_B^2}} \\ &\quad + \frac{1}{8} \lambda_B [I(m_p, T)]^2 \quad \leftarrow \cancel{(vac)(mat) \text{ cancels!}} \end{aligned}$$

$$\begin{aligned} J &= J_0 + J_T \quad \left( \begin{array}{l} J_0 = I_0 + I_T \\ I = I_0 + I_T \end{array} \right) \\ &= \left\{ J_0(m_p, T) - \frac{1}{8} \lambda_B I_0^2(m_p) \right\} + \left\{ J_T(m_p) + \frac{1}{8} \lambda_B I_T^2(m_p) \right\} \end{aligned}$$

$T=0, \text{dr}$

(finite)

$T>0$  finite

Indeed splits into vacuum & scatter parts!

$$(2) \quad \boxed{\text{+ usual } T=0 \text{ renormalization works, even if infini. results!}} \quad \text{show (vac, dr) & (finite, T) contributions}$$

$$f_{(0)} \rightarrow J^{(n=0)} \sim \boxed{m_B^3 T}$$

ratio:  $\boxed{\frac{\lambda_B T^2}{m_B^2}}$

$$f_{(1)} \rightarrow \frac{3}{4!} \lambda_B I'(0, T) J^{(n=0)}$$

$$\sim \lambda_B \left(\frac{T^2}{12}\right) \left(-\frac{m_B T}{4\pi}\right) \sim \boxed{\lambda_B m_B T^3}$$

~~$$f_{(2)} \rightarrow \frac{1}{16} \lambda_B^2 \left(I'(0, T)\right)^2 \left(T \cdot \frac{1}{(p^2 + m_B^2)^2}\right)$$~~

) ratio  $\boxed{\frac{\lambda_B T^2}{m_B^2}}$

$$f_{(2)} \rightarrow \frac{1}{16} \lambda_B^2 \left(I'(0, T)\right)^2 \left(T \cdot \frac{1}{(p^2 + m_B^2)^2}\right)$$

$$\sim \lambda_B^2 (T^2)^2 \left(\frac{p}{m_B}\right) \sim \boxed{\frac{\lambda_B^2 T^5}{m_B}}$$

So the expansion parameter is  $\lambda_B T^2 / m_B^2$ ; suggests

\* Clearly something goes wrong:  $m_B \ll \lambda_B T^2$

\* Suggests there is some new physics at scale  $\lambda_B T^3$ .

## #~~IR divergences~~ # Renormalization (IR)

Now we turn to the IR divergences, which hint at new & interesting physics from the  $T \geq 0$ .

~~Note that~~ these correspond to the limit  $m_B \rightarrow 0$ , and seem to be connected to the zero Matsubara mode, which has no "mass term" (bubble  $\omega_n \propto T$ ,  $n \neq 0$ ).

Looking back at what we have above, we should see that the  $n=0$  mode is the issue; and it contributes to odd powers of  $m$ :

$$\begin{aligned}
 C. \quad I'(0, T) &\Rightarrow f(0) \propto T^{(n=0)} = \frac{\lambda_B T}{T + \frac{m_B^2}{\lambda_B}} \sim m_B^3 T \quad \text{ratio} \sim \frac{\lambda_B T^2}{m_B^2} \\
 &= \frac{T}{12} \\
 I^{(n=0)} &= -\frac{m_T^2}{4\pi} \\
 f_{(1)} &\Rightarrow \frac{3}{4} \lambda_B \cdot I'(0, T); \quad I^{(n=0)} \sim (+\lambda_B m_B T^3) \quad \text{ratio} \sim \frac{\lambda_B T^2}{m_B^2} \\
 f_{(2)} &\Rightarrow \frac{9}{4} \lambda_B^2 [I'(0, T)]^2 \sim (+\lambda_B^2 T^5) \quad \text{(check on p. 22.5)}
 \end{aligned}$$

This ratio is not small if  $m_B^2 \ll \lambda_B T^2$  — suggests that there is some new physics at the scale  $\lambda_B T^2$ .

We can deal with these difficulties in two ways:

- 1.) Diagrammatic Renormalization
- 2.) EFT for the  $n=0$  mathe mode.

Let's do them in turn.

## Renormalization:  $\text{---} = n=0$  mode,  $\text{—} = n \neq 0$  mode

$$\begin{aligned}
 \text{bad } f(0) &= \text{---} \\
 \text{bad } f_{(1)} &= \text{---} \\
 \text{bad } f_{(2)} &= \text{---} \\
 &\quad \left. \right\} \text{redundant!} \\
 &\quad \text{---}; \quad \text{"mg" or "davy"} \\
 &\quad \text{---}; \quad \text{diagram}
 \end{aligned}$$

Combinations:

(23)

$$\begin{aligned}
 f(T) &= \left\langle S_x - \frac{1}{2} S_x^2 + \dots + \frac{(-1)^{N+1}}{N!} S_x^{(N)} \right\rangle_{c, \text{ drop } f_x} \\
 &\Rightarrow \frac{(-1)^{N+1}}{N!} \left( \frac{\lambda_B}{4!} \right)^N \left\langle \underbrace{\phi \phi \phi \phi}_{2(N-1)} \underbrace{\phi \phi \phi \phi}_{2(N-2)} \dots \underbrace{\phi \phi \phi \phi}_{2} \dots \right\rangle_{c_1} \\
 &= \frac{(-1)^{N+1}}{N!} \left( \frac{\lambda_B}{4!} \right)^N \binom{N}{2} \underbrace{\frac{[2(N-1)][2(N-2)] \dots [2]}{2^{N-1}(N-1)!}}_{\text{non-zero mode}} I'(0, T) \\
 &\quad \boxed{\cdot T \int_{\vec{p}} \left( \frac{1}{p^2 + m_B^2} \right)^N} \quad \text{(non-zero mode)}
 \end{aligned}$$

Look at the zero mode part for small  $N$ :

$$N=1: \int \frac{1}{p^2 + m_B^2} = -\frac{m_B}{4\pi} = -\frac{d}{dm_B} \left( +\frac{m_B^3}{6\pi} \right)$$

$$N=2: \int \frac{1}{(p^2 + m_B^2)^2} = -\frac{d^2}{dm_B^2} \left( -\frac{m_B}{4\pi} \right) = \left( -\frac{d}{dm_B} \right) \left( -\frac{d}{dm_B} \right) \left( +\frac{m_B^3}{6\pi} \right)$$

$$\begin{aligned}
 \Rightarrow \text{generally: } \boxed{\int \frac{1}{(p^2 + m_B^2)^N}} &= -\frac{1}{N-1} \frac{d}{dm_B} \int \frac{1}{p^2 + m_B^2} \\
 &= \frac{(-1)(-1) \dots (-1)}{(N-1)} \left( \frac{d}{dm_B} \right)^{N-1} \int \frac{1}{p^2 + m_B^2} \\
 &= \boxed{\frac{(-1)^N}{(N-1)!} \left( \frac{d}{dm_B} \right)^{N-1} \left( +\frac{m_B^3}{6\pi} \right)}
 \end{aligned}$$

Plugging A M R:

$$f_{(N)} \Rightarrow -\frac{T}{2} \frac{1}{N!} \left( \frac{\lambda_B T^2}{2 \cdot 4!} \right)^N \left( \frac{d}{dm_B} \right)^{N-1} \left( +\frac{m_B^3}{6\pi} \right)$$

Taylor series

$$\begin{aligned}
 \Rightarrow f_{(N)} &\sum_{n=0}^{\infty} -\frac{1}{n!} \left( \frac{\lambda_B T^2}{4!} \right)^n \left( \frac{d}{dm_B} \right)^n \left( +\frac{m_B^3}{12\pi} T \right) \\
 &= \boxed{-\frac{T}{12\pi} \left( m_B^2 + \frac{\lambda_B T^2}{4!} \right)^{3/2}}
 \end{aligned}$$

- 1 hr

(24) \*  $m_B \rightarrow 0$  can be taken. \*  $\lambda_B^{3/2}$  contribution to this case!

\* Induced divergences modify the order & the weak-coupling expansion!

$$\boxed{f(T) = -\frac{\pi^2 T^4}{90} + \frac{\lambda_B T^4}{4! \times 48} - \frac{T}{12\pi} \left[ \frac{\lambda_B T^2}{4!} \right]^{3/2} + \mathcal{O}(\lambda_B^2 T^4)}$$

$$= \boxed{-\frac{\pi^2 T^4}{90\pi} \left[ 1 - \frac{5}{64} \frac{\lambda_B}{\pi^2} + \frac{5}{32} \left( \frac{\lambda_B}{\pi^2} \right)^{3/2} + \mathcal{O}(\lambda_B^2) \right]}$$

massless limit.

## EFT method (Better for computing/understanding physics.)

(1.) Determine T-dependent pole mass in the  $m_B \rightarrow 0$  limit.

Effective thermal mass  $m_{\text{eff}}^2$

(2.) Argue in  $\lambda_n \ll 1$  limit,  $m_{\text{eff}}$  only needed for  $n=0$  mode.

(3.) Rewrite Lagrangian ~~for~~ (for  $m_B^2 = 0$ ) & the term:

$$\mathcal{L}_E = \underbrace{\frac{1}{2} (\partial_\mu \phi)^2}_{\mathcal{L}_0} + \underbrace{\frac{1}{2} m_{\text{eff}}^2 \phi_{n=0}^2}_{\mathcal{L}_I} + \underbrace{\frac{1}{4!} \lambda_B \phi^4}_{\mathcal{L}_I} - \underbrace{\frac{1}{2} m_{\text{eff}}^2 \phi_{n=0}^2}_{\mathcal{L}_I}$$

This leads to same answer as above. Consider

$$m_{\text{eff}}^2 = \lim_{m_B^2 \rightarrow 0} [m_B^2 + 3\lambda_B I(m_B, T)] = \frac{\lambda_B T^2}{4!} + \mathcal{O}(\lambda_B^2)$$

$\xrightarrow{f_p \rightarrow f}$  version of above

\* For  $n \neq 0$ ,  $m_{\text{eff}}^2 \ll \omega_n^2$ , ~~modifies~~ ~~modifies~~ mode significantly!

(So needs only need it for the  $n=0$  mode)

$$\Rightarrow \boxed{f(T) = \int \ln(p^2) \frac{1}{2} + T \int \frac{1}{p} \ln(p^2 + m_{\text{eff}}^2) - \text{const.}}$$

$$= J(0, T) + J^{(n=0)}(m_{\text{eff}}, T) = \boxed{-\frac{\pi^2 T^4}{90} - \frac{m_{\text{eff}}^3 T}{12\pi}}$$

$$f_{(1)}(T) = \frac{3}{4!} \lambda_B \langle \phi_0, \langle \phi_1, \phi_0 \rangle, \langle \phi_0, \phi_0 \rangle \rangle - \frac{1}{2} m_{\text{eff}}^2 \langle \phi_0, \phi_0 \rangle$$

$$= \frac{3}{4!} \lambda_B \left[ I(0, T) + I^{(n=0)}(m_{\text{eff}}, T) \right]^2 - \frac{1}{2} m_{\text{eff}}^2 I^{(n=0)}(m_{\text{eff}}, T)$$

(25)

$$\Rightarrow f_{(1)}(T) = \frac{3}{4!} \lambda_B \left[ \frac{T^4}{144} - \frac{m_{\text{eff}} T^4}{384\pi^2} + \frac{m_{\text{eff}}^2 T^2}{16\pi^2} \right] + \frac{1}{2} m_{\text{eff}}^2 \left( \frac{m_{\text{eff}} T}{4\pi} \right)$$

$$\Rightarrow f(T) = -\frac{\pi^2 T^4}{90} + \frac{3}{4!} \lambda_R \frac{T^4}{144} - \frac{m_{\text{eff}}^3 T}{12\pi}$$

agrees w/ previous expression.

$$\underbrace{\frac{\lambda_R T^2}{4!}}_{\frac{3}{8\pi} \frac{\lambda_R m_{\text{eff}}^2 T^2}{4!}}$$

removes the unwanted term!

(one of the propagators  
more IR diverging terms)

\* This method works even at higher orders.

\* IR divergences (1) get canceled by the reorganization  
(note that, in EFT styles need to compute  $m_{\text{eff}}^2$  to higher  
orders as well.)

## # QCD

Transition to more advanced part of the course now, in real QCD.

\* Want derive Path-Integral w/ gauge fields. There are subtleties, related  
to fixing a gauge in the derivation (and  $A_0$ ), but in the end, the path  
integral ultimately becomes the same as expected  $\mathcal{Z}_E = \mathcal{Z}_M (\vec{x} \rightarrow i\tau)$

$$S_E = \mathcal{Z}_{QCD} = C \int \prod_{\text{per.}} dA_0^a dA_\mu^a \prod_{\text{anti.}} \bar{c}^a c^a \int d\bar{\psi} d\psi e^{-S_{QCD}}$$

$$S_{QCD} = \int_0^\beta d\tau \int d^3k \left[ \frac{1}{4} F_{\mu\nu}^a F^{a\mu\nu} + \cancel{(\text{higher order terms})} + \frac{1}{2} g f^{abc} \bar{c}^a \gamma^\mu c^b \bar{c}^c \right] \quad (\text{glue})$$

$$+ \frac{1}{2g} \partial_\mu A_\mu^a \partial_\nu A_\nu^a + \partial_\mu \bar{c}^a \partial_\nu c^a + g f^{abc} \partial_\mu \bar{c}^a \gamma^\mu c^b \bar{c}^c \quad (\text{gauge})$$

$$+ \bar{\psi}_a (\not{D}_{ab} + m \not{\gamma}_5 \delta_{ab}) \psi_b \quad (\text{fermion})$$

$$P_{\mu\nu}^{\text{tot}} = P_{\mu\nu}^4$$

These lead to the propagators:

$$\langle \tilde{A}_\mu^a(p) \tilde{A}_\nu^b(q) \rangle = S^{ab} \delta(p+q) \left[ \frac{S_{\mu\nu} - P_\mu P_\nu / P^2}{P^2} + \frac{S_{\mu\nu} - P_\mu P_\nu / P^2}{P^2} \right]$$

(26)

$$\tilde{S}_x^a e^{i(p+q) \cdot x} = \beta S_{m+q_n, 0} (2\pi)^d S^{(d)}(\vec{p} + \vec{q})$$

Gluon contribution:

→ ~~RECP~~

$$1/p_{\mu\nu} = P^2 \frac{P^\perp}{p_{\mu\nu}} + \frac{P^2}{5} \frac{P^{\parallel}}{p_{\mu\nu}}$$

$\nwarrow$   $P^\perp = S_{\mu\nu} - 1 = d$        $P^{\parallel} = 1$

$$S^{aa} = N_c^2 - 1 = d_A \text{ copies of this:}$$

$$\rightarrow f(t) \Big|_{\text{gluon}} = d_A \left\{ d \cdot \frac{1}{2} \oint_P f(\ln(P)) + \text{const} \right\} + \frac{1}{2} \oint_P f(\ln(\frac{1}{5} P^2) - \text{const}) \}$$

Ans:

$$= d_A \left\{ (d+1) J(0, T) - \frac{1}{2} \oint_P f(\ln(\frac{1}{5})) \right\} \quad \begin{matrix} \text{so in chiral} \\ \text{(no scales)} \end{matrix}$$

C ~~Check that each ghost field has point in spacetime here.~~

~~ghosts~~

The ghosts have a gauge integral (no  $\frac{1}{2}$  const, but  $-1$ ), but are

~~$$f(t) \Big|_{\text{ghosts}} = 1 - \oint_P f(\ln(t))$$~~

bosons  $\Rightarrow \oint_P$ :

$$\text{C} \quad \text{ghosts: } f(t) \Big|_{\text{gh}} = d_A \left\{ -2 \cdot \frac{1}{2} \oint_P f(\ln(P^2)) + \text{const} \right\}$$

$$= -2 d_A J(0, T)$$

so that  $f \Big|_{\text{gauge}} = f \Big|_{\text{gh}} + f \Big|_{\text{gh}} = d_A \underbrace{(d-1)}_{\# \text{ physical modes.}} J(0, T)$

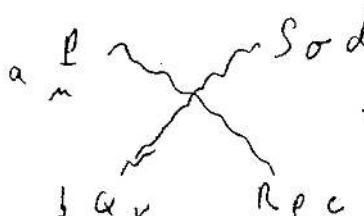
(1)

(26.5)

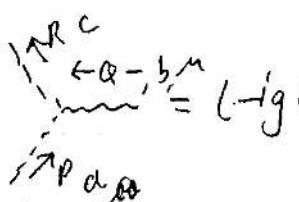
$$\langle \tilde{c}(l) \tilde{c}(\alpha) \rangle = \delta^{ab} f(l-\alpha) \frac{1}{l^2}, \quad = \text{massive } \xrightarrow{\alpha \rightarrow -l}$$

$$\langle \tilde{\psi}_A(l) \tilde{\psi}_B(\alpha) \rangle = S_{AB} \delta(l-\alpha) \frac{-il+m}{l^2+m^2}. \quad \xrightarrow{\alpha \rightarrow l}$$

$$a \overset{P_1}{\underset{Q_V, b}{\swarrow \nearrow}} \overset{R, C}{+} = ig f^{abc} \left\{ \delta_{m\rho} (l_\nu - R_\nu) + \text{cycle } [(P, Q, R), (l_\nu, \nu, \rho)] \right\}$$



$$= g^2 f^{cab} f^{ecd} (\delta_{m\rho} \delta_{n\nu} - \delta_{m\nu} \delta_{n\rho}) + \dots \text{cycle } (P, Q, R, S) \\ (m, n, \rho, \nu) \\ (a, b, c, d).$$



$$(-ig f^{abc} P_m), \quad \overset{R \wedge B}{\underset{P, A}{\swarrow \nearrow}} \underset{Q, n}{\underset{P, A}{\swarrow \nearrow}} = -ig T_{AB}^a \delta^m_n$$

## Nonabelian blackbody radiation:

$$J(0, T) = \frac{1}{2} \oint [ln(P^2) + \text{const.}] = \left[ -\frac{\pi^2 T^4}{90} \right]$$

$$\begin{aligned} \tilde{J}(0, T) &= \frac{1}{2} \oint [ln(P^2) + \text{const.}] = \left\{ 2 \cdot J(0, \frac{T}{2}) - J(0, T) \right. \\ &\quad \left. = \left( \frac{2}{2^4} - 1 \right) \left( \frac{\pi^2 T^4}{90} \right) = \frac{7}{8} \frac{\pi^2 T^4}{90} \right\} \end{aligned}$$

The hadronic propagator can be written as:

$$P_{mn} = P^2 \delta_{mn} - (1 - \frac{1}{3}) P_m P_n = P^2 \delta_{mn} - P_m P_n$$

Gluon:  $f(T) = (N_c - 1) \left\{ d \times J(0, T) + 1 \times J(0, T) \right\}$

# gluon propagator

Ghosts:  $f(T) = (N_c - 1) \left\{ d \times J(0, T) - \frac{1}{V} \ln Z_{gh} \right\} - \frac{1}{V} \ln Z_{gh}$  (p. 26.5)

$$= -\frac{1}{V} \ln (T^2 P^2) = -2(N_c - 1) J(0, T) \quad (2)$$

$$\text{Fermions: } f(T) = -4 \cdot \underbrace{N_c N_f}_{\# \text{ ferm}} \hat{f}(0, T)$$

$$\Rightarrow f(T) = -\frac{\pi^2 T^4}{90} \left[ 2(N_c^2 - 1) + \frac{7}{2} N_f N_c \right]$$

\* Note that first removed unphysical DOF of the gluas.

End  
lecture 3

### # Thermal gluon mass

Want to go through the calculation in detail. As in the scalar theory, we want:

Feynman gauge  $\xi = 1$

$$\frac{\langle A_m^a(u) A_\nu^b(Q) e^{-S_I} \rangle_0}{\langle e^{-S_I} \rangle_0} = S^{ab} \delta(u+Q) \left[ \frac{\delta_{\mu\nu}}{u^2} - \frac{T m(u)}{u^4} + O(g^4) \right]$$

Doing lots of contractions & manipulations, shown in detail in Lake & Veltman, one finds: (Using the fact that  $u^2 \geq 0$  from the bare gluon mass = 0)  $\Rightarrow$  ignore  $k^2$  term of internal propagator

$$\begin{aligned} \langle \tilde{A}_m^a(u) \tilde{A}_\nu^b(Q) (-S_I) \rangle_{0,c} &= \text{cloud} = -g^2 N_c d \tilde{I}_T(0) \left[ \frac{S^{ab} \delta(u+Q)}{(u^2)^2} \right] \\ &= \frac{g^2 N_c}{2} \left[ \frac{S^{ab} \delta(u+Q)}{(u^2)^2} \right] \left\{ -3 \delta_{\mu 0} \delta_{\nu 0} + 7 \delta_{\mu i} \delta_{\nu i} \right\} \frac{T^2}{12} \end{aligned}$$

$$\text{cloud} = -\frac{g^2 N_c}{2} \left[ \text{---} \right] (-\delta_{\mu 0} \delta_{\nu 0} + \delta_{\mu i} \delta_{\nu i}) \frac{T^2}{12}$$

$I_T(0)$  &

$\tilde{I}_T(0)$

$$\text{cloud} = -g^2 N_f \left[ \text{---} \right] \left\{ \delta_{\mu 0} \delta_{\nu 0} + 0 \times \delta_{\mu i} \delta_{\nu i} \right\} \frac{T^2}{6}$$

approx.

$$\text{Total: } -\frac{\delta^{ab} \delta_{\mu 0} \delta_{\nu 0} \delta(u+Q)}{T (u^2)^2} \boxed{\frac{g^2 T^2 \left( \frac{N_c}{3} + \frac{N_f}{6} \right)}{M_E^2}}$$

(3) \* Actually, not

taking the  $T \rightarrow 0$  limit will show even more interesting structure later.

absolutely gauge invariant at this order.

In rest frame, must be  $T$

for coordinate  $tt = (0, t)$ ; W.F says T-time component

$$\Rightarrow \langle \tilde{A}_m^a(u) \tilde{A}_\nu^b(Q) \rangle \stackrel{u \neq 0}{\approx} \frac{\delta^{ab} \delta_{\mu 0} \delta(u+Q)}{u^2 + \delta_{\mu 0} \delta_{\nu 0} M_E^2}$$

Rebyle mass parameter

(1) from A. Atel exponentially screened in order, for non-limit.

(2) \* "Color-magnetic" terms  $\cancel{A}_\mu$  do not get screened at this order

## # Red Thermal Loops:

Let's consider the ~~other~~ self energy again, but let's not restrict to the ~~lowest~~ Matsubara mode — the reason is that as  $T$  gets smaller, more of

I'll now show a sketch of how to compute the  $O(g^2)$  corrections to the  $f(T)$  of pQCD at high  $T$ . The main steps are:

1.) Do the Color, Lorentz contractions to ~~reduc~~ the exp.  
(the  $I$  won't show).

$$2) \text{use } D_{\mu\nu} = \cancel{D}_{\mu\nu} \cancel{D}_{\alpha\beta} = \frac{1}{2} [(L - u)^2 - p^2 - k^2]$$

To introduce more propagators and cancel some propagators.

3.) Use Momentum shifts to reduce to small set of master integrals, etc.

4.) Compute (or look up) these master integrals. As ~~QCD~~ pointed out,  
it is always possible  $\square$

2-loop free energy of QCD  $\rightarrow$  leads to fully factorizable integrals — great!  $2 \text{ loops}$

$$\frac{1}{2} \text{ (cloud)} = -g^2 \frac{N_c}{4} d I_T(0) \oint \frac{\delta^{ad} \delta^{mn}}{u} \frac{1}{u^2}$$

$$= \frac{g^2}{4} N_c (N_c^2 - 1) d(d+1) [I_T(0)]^2$$

$$\frac{1}{12} \text{ (cloud)} = -\frac{g^2}{4} N_c d \oint \frac{\delta^{ad}}{u, L} \delta^{mn} \frac{u^2 + (u-p)^2 + L^2}{u^2 L^2 (u-p)^2} = \frac{g^2}{4} N_c d \oint \frac{\delta^{ad}}{u, L} \delta^{mn} [I_T(0)]^2$$

$$= -\frac{g^2}{4} N_c (N_c^2 - 1) d \times 3 [I_T(0)]^2$$

$$\frac{1}{2} \text{ (cloud)} = \frac{g^2}{4} N_c \delta^{ad} \oint \frac{\delta^{mn}}{L, u} \frac{L^2 + (u-L)^2 - u^2}{u^2 L^2 (u-L)^2}$$

$$= \frac{g^2}{4} N_c (N_c^2 - 1) [I_T(0)]^2$$

$$-\frac{1}{2} \text{ (cloud)} = -g^2 N_c \delta^{ad} \frac{d-1}{2} \oint \frac{L^2 + (u-L)^2 - u^2}{u, L, 0} = -\frac{g^2}{2} N_c \delta^{ad} (d-1) \left\{ 2 I_T(0) \tilde{I}_T(0) - [I_T(0)]^2 \right\}$$

(29)

$$\text{Using } d=3, \quad I_+(0) = \frac{T^2}{12}, \quad \tilde{I}_r(0) = 2 I_{\frac{3}{2}}(0) - I_+(0) = \left[ \frac{2}{2^2} - 1 \right] \frac{T^2}{12} = \frac{-T^2}{24}$$

one finds:

$$f_{(1)}(T)|_{\text{act}} = -\frac{\pi^2 T^4}{90} (N_c^2 - 1) \left( -\frac{5}{2} \frac{g^2}{4\pi^2} \right) \left( N_c + \frac{5}{4} N_f \right).$$

With the thermal mass from above, this means now the free energy density is:

$$f(T)|_{\text{act}} = -\frac{\pi^2 T^4}{45} (N_c^2 - 1) \left\{ 1 + \frac{7}{4} \frac{N_c N_f}{N_c^2 - 1} - \frac{5}{4} \left( N_c + \frac{5}{4} N_f \right) \frac{\alpha_s}{\pi} \right. \\ \left. + 30 \left( \frac{N_c}{3} + \frac{N_f}{6} \right)^{3/2} \left( \frac{\alpha_s}{\pi} \right)^{3/2} + \mathcal{O}(\alpha_s^2) \right\}$$

- \* First relative correction is negative — Interactions decrease the pressure.
- \* Shows the power of (1)-(4) above. Very little work for  $\mathcal{O}(\alpha_s^{3/2})$  result! (Also known by this now thermal mass idea.)
- \* Well return to the general EFT description next time.

Next today, however, will turn to the quark sector, and the asymptotic limit at high  $m_s$ ,  $T=0$ .

# EFTs of light-temperature QCD

## Lecture 4

Saw above that there's interesting IR physics in general QFTs. What happens in QCD specifically, and what's the low-energy EFT?<sup>2</sup> (Looked like the  $n=0$  gluon Matsubara mode at least ~~approx~~ becomes massless.)

$$\text{horizon: } \oint_{\Gamma} f(w_n, \bar{p}) = \int_{-\infty - i0^+}^{\infty + i0^+} \frac{dw}{2\pi} [f(w, \bar{p}) + f(-w, \bar{p})] \{1 + 2n_B(iw)\}$$

$$\text{fermion: } \oint_{\Gamma} f(w_n, \bar{p}) = \int_{-\infty - i0^+}^{\infty + i0^+} \frac{dw}{2\pi} [f(w, \bar{p}) + f(-w, \bar{p})] \{1 - 2n_F(iw)\}$$

In general,  $f$  is product of propagators and additional structures from vertices.

~~approximate~~ 
$$f(w, \bar{p}) \sim \frac{1}{w^2 + E_p^2} \quad \text{simplest case.}$$

What dominates in these expressions? (Keeping in mind  $w \sim T$  for  $n \neq 0$ , for bosons)

$$\text{boson: } \text{LHS, zero mode: } \left[ \frac{T}{E_p^2} \right]$$

$$\begin{aligned} \text{RHS, pole enhancement: } & \frac{1}{2} \left( \frac{-2\pi i}{2\pi} \right) \frac{2}{-2iE_p} \left[ 1 + 2n_B(E_p) \right] \\ \text{near } w = -iE_p: & = \frac{1}{E_p} \left( \frac{1}{2} + \frac{1}{e^{E_p/T} - 1} \right) \xrightarrow{\text{high-T}} \\ & \approx \frac{1}{E_p} \left( \frac{1}{2} + \frac{1}{E_p/T + E_p^2/(2T^2)} + \dots \right) \\ & \approx \left[ \frac{T}{E_p^2} \right] \checkmark \text{ agree.} \end{aligned}$$

$$\text{fermion: LHS, no zero mode: } \left( \frac{T}{\pi T} \right)^2 \sim \left[ \frac{1}{\pi^2 T} \right]$$

$$\begin{aligned} \text{RHS, similarly: } & \frac{1}{E_p} \left( \frac{1}{2} - \frac{1}{e^{E_p/T} + 1} \right) \approx \frac{1}{E_p} \left( \frac{1}{2} - \frac{1}{2 + E_p/T} + \dots \right) \\ & = \left[ \frac{1}{\pi^2 T} \right] \checkmark \end{aligned}$$

What's the dimensionless scale associated with these loop integrals?

- Adding loop adds factor  $g^2$  (vertices)  $\int g^2$

Add Ward prop-gives  $T$  in Matsubara sum.  $\int g^2$

- Where does  $w$  ~~approx~~ can give  $E_p^{-\frac{1}{2}}$   $\rightarrow$  in  $\frac{1}{w}$  add in  $\bar{p}$  integral

- Fermion can only give  $T^{-\frac{1}{2}}$ .

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Vance, we expect: bosons:  $\lambda_B \sim \frac{g^2 T}{m}$  (and this will be my diagram)  
 fermions:  $\lambda_B \sim \frac{g^2 T}{T} \sim g^2$

### Conclusions:

- 1.) Fermions are always perturbative at high  $T$ .
- 2.) Boson zero modes have poor convergence, w/  $m \rightarrow 0$ .

↳ In fact,  $m_{\text{eff}}^2 \sim g^2 T^2$  indicates

$$\lambda_{B,0} \sim \frac{g^2 T}{\sqrt{g^2 T^2}} \sim g.$$

⇒ odd powers of  $g$  in weak-coupling expansion.

- 3.) Special gluonic modes have  $m_{\text{eff}} \lesssim g^2 T$

$$\Rightarrow \lambda_{B, \text{glue}} \sim \frac{g^2 T}{g^2 T} \sim O(1) \quad \text{non-perturbative}$$

This agrees w/ diagrammatic analysis of 't Hooft PLB 96, 289 (1980)  
 ("t Hooft problem").

So we have scale hierarchy at high- $T$   $[g^2 T \ll g T \ll T]$

\* Idea of EFT. — Given full theory w/ 1PI Green Functions  $F_n$

- Identify IR DOFs. Compute GFs & EFT  $\tilde{F}_n$   
 ↳ approximation of full theory to extract operators

- For momenta & energy small ( $p, w \lesssim g M$  & heavy scale)  
 ↳ small coupling

$$\frac{\delta F_n}{F_n} = \frac{|F_n - \tilde{F}_n|}{F_n} \leq O(g^\epsilon) \text{ for some } \epsilon > 0.$$

If mass & coupling in low-energy theory are suitably tuned. (matching)

In hot QCD: (Kapusta, Laine, Reinhardt, Schröder 0211321v2)

- 1.) DOFs: nonbroken O modes (w/ gluary  $A_0, A_i$ ) These are irrelevant.  
 So Dimensionally Reduced EFT of  $d=3-2\epsilon$  dimensions.

(B5) 2.) Symmetries a) Spatial translations & rotations (medium breaks horizontality)

(B6)

Symmetries b.)  $A_0 \rightarrow -A_0$  (if no quark chiral. pacts)

c.) Gauge boson mass split: b/c fields are  $\tau$  independent:

$$A_m' = U A_m U^{-1} + i g U \partial_m U^{-1}$$

$$\hookrightarrow \begin{cases} \bar{A}_i' = U \bar{A}_i U^{-1} + i g U \partial_i U^{-1} \\ \bar{A}_0' = U \bar{A}_0 U^{-1} \end{cases} \begin{matrix} \leftarrow \text{gauge fields} \\ \leftarrow \text{adjoint rep scalar.} \end{matrix} \begin{matrix} \leftarrow \text{can't have} \\ \text{mass!} \end{matrix}$$

Lagrangian (Full) for  $\tau$  independent fields gives:

$$F_{0i}^a = \partial_j A_i^{aj} - D_i^{ab} A_0^b$$

$$\rightarrow L_E = \frac{1}{4} F_{ij}^a F_{ij}^a + \frac{1}{2} (D_i^{ab} A_0^b) (D_i^{ac} A_0^c)$$

can rewrite as  $\boxed{L_{\text{eff}}^{(0)} = \frac{1}{4} \bar{F}_{ij}^a \bar{F}_{ij}^a + \text{Tr} \{ [\bar{D}_i, \bar{A}_0] [\bar{D}_i, \bar{A}_0] \}}$

Plugged in remaining terms at the time:

$$d_m = 2: \text{Tr} [A_0^2]$$

$$d_m = 4: \text{Tr} [A_0^4], (\text{Tr} [A_0^2])^2 \quad \begin{matrix} \text{actually} \\ \text{connected to } N_c = 2, 3 \end{matrix}$$

$$d_m = 6: \text{Tr} \{ [\bar{D}_i, \bar{F}_{ij}] [\bar{D}_k, \bar{F}_{kj}] \} + \dots$$

$\Rightarrow$  Electrostatic QCD:

$$S_{\text{eff}} = \frac{1}{T} \int_x \left\{ \frac{1}{4} \bar{F}_{ij}^a \bar{F}_{ij}^a + \text{Tr} \{ [\bar{D}_i, \bar{A}_0] [\bar{D}_i, \bar{A}_0] \} \right.$$

$$\left. + \bar{m}^2 \text{Tr} [\bar{A}_0^2] + \bar{\lambda}_1^{(1)} (\text{Tr} [\bar{A}_0^2])^2 + \bar{\lambda}_1^{(2)} \text{Tr} [\bar{A}_0^4] \right. \\ \left. + \dots \right\} \quad (+ i \bar{g} \text{Tr} [\bar{A}_0^3] \text{ if } m_2 > 0)$$

Matching gives:  $\bar{m}^2 = g^2 T^{-2} \left( \frac{N_c}{3} + \frac{N_f}{6} \right) + \mathcal{O}(g^4 T^{-2})$

Start @  $\mathcal{O}(g^4)$

$$\bar{\lambda}_1^{(1)} = \frac{g^4}{4\pi^2} + \dots, \quad \bar{\lambda}_1^{(2)} = \frac{g^4}{12\pi^2} (N_c - N_f) + \mathcal{O}(g^6) \dots$$

in pQCD pressure:  $\bar{g}^2 = g^2 + \mathcal{O}(g^4), \quad \bar{g} = \sum_{i=1}^{N_f} \frac{m_i g_i^3}{3\pi^2} + \mathcal{O}(g^5)$

(Analyticity reflected operators suggest,  $\frac{S\bar{T}}{F} \sim \left( \frac{\bar{m}}{T} \right)^3 \sim g^3$ ). (c.f. p. 123 LBV).

(up to the order shown)

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Can even go further and integrate out the massive  $A_0$  field. (If we are interested in long-wavelength physics  $1/X > 1/m$ .)

$$\Rightarrow \text{Magnetohydro QCD. } S_{\text{eff}} = \frac{1}{4} \int \left\{ \frac{1}{4} F_{ij}^a F_{ij}^a + \text{HOOps} \right\}$$

\* 2 state parameter  $\rightarrow$  the gauge coupling.

Stark  $\propto \partial(g)$  Rescaling the fields to remove the  $\frac{1}{T}$  gives  $\tilde{g}^2 T$  as the coupling.  $A \mapsto A T^{1/2}$

$$\Rightarrow \text{All dimensionless quantities, } \propto [\tilde{g}^2 T]^{1/2} \times (\text{non-perturbative #})$$

$$= [(\partial A)^2 + g A \partial A + g^2 A A A A] T$$

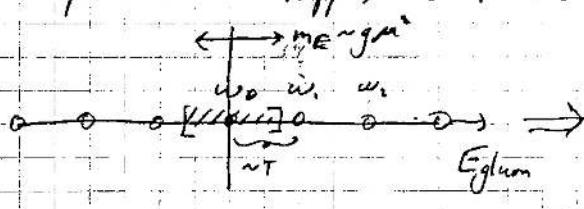
Note that ~~WILTED~~ if  $p_T > 0$ , then the essentially all still holds

@ high  $T$  ← NEXT PAGES (MORE) HERE  $\nwarrow m_B > 0$  &

# High- $m$ , EFT<sub>2</sub> \* low or 0  $T$ . \*

@ high  $m$ , high  $T$ , the essentially all still holds. (Just a few more operators, since  $A_0 \rightarrow -A_0$  no longer a symmetry.)

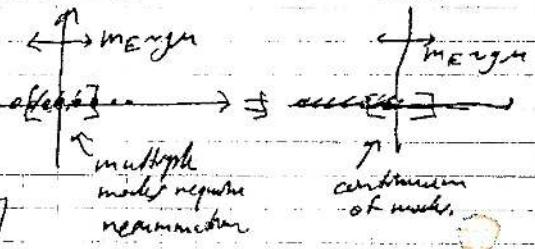
However, note what happens as  $T \rightarrow 0$ :



1603.00750  
Kurkela + Vuorinen.

$T \rightarrow 0$

$T=0$



@ high  $m$  and small  $T$  (or  $T=0$ ), can't assume only one Matsubara mode!



$\Rightarrow$  Thermal mass is not enough; require  $\Pi(K, \frac{k_0}{k_1}) \sim \Pi_{\text{HTL}}(k_0) + k_0^2 \dots$

Require more general self-energy than just

the  $\Pi^{uv}|_{k_0=0}$ , with non-trivial  $\frac{k_0}{k_1}$  dependence.

Hard-thermal loop

"Hard-thermal loop" self energy, since the loops integrated are dominated by the ~~most~~ highest p. w. "hardest" momenta.

$$\Rightarrow \Pi^{uv}(k_0, k) = \int \frac{d^2 Q}{(2\pi)^2} g^2 N_F S \sum_a \frac{\text{Tr}(\gamma^a \gamma^v (\not{Q} + \not{Q}) \gamma^u)}{Q^2 (\not{P} + \not{Q})^2}$$

$$\star \text{ Might show more details here} \quad \sim \not{Q}^3 \cdot \frac{1}{\not{Q}} \frac{1}{\not{Q}} \rightarrow \not{Q} \text{ at } T=0.$$

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~~High  $m_p$ ,  $T=0$  - argue (cold Quark matter) (X marked on p. 37)~~

In the limit,  $\frac{g}{\sqrt{2}Q} \rightarrow S$  (4d integral again, which is much simpler!)

Note that chemical potential enters as:

$$Z = \exp \left[ -\beta (\hat{H} - \mu \hat{N}) \right], \text{ but } \hat{N} = \int_{-\infty}^{\infty} \text{conserved current, } 0^\text{ component, which} \\ \text{contributes}$$

~~Such a component normally couples the gauge field  $A^0$ .~~

~~Matrix?~~

$$\begin{matrix} & \partial_0 + g A_0 & & & \\ \partial_0 + g A_0 & \xrightarrow{\text{Eucl.}} & \xrightarrow{\text{Eucl.}} & & \\ & \downarrow \text{chem. pol} & & & \\ & \partial_0 + m & & \xrightarrow{-i\partial_T + m} & -i(\partial_T + i\omega) \end{matrix}$$

~~Take the Feynman gauge field in the Euclidean Lagrangian~~

$$\mathcal{L}_E \rightarrow \boxed{\partial_\mu \bar{\psi} \gamma^\mu \psi} \leftrightarrow (\bar{\psi} \gamma_\mu \psi) - (\bar{\psi} \gamma_E^\mu \psi) m$$

$$\boxed{\partial_0 + g A_0 \rightarrow \partial_0 - m}$$

$$i\omega_n \mapsto i\omega_n - m = i(\omega_n + i\mu); \text{ imaginary shift to frequency!}$$

The reason is that ~~\*~~ no more sum-integrals

~~\*... but contour integrals with necessary b/c of  $(Q^0 + i\mu)$ .~~

$$\text{E.g.: } \tilde{I}_n = \int \frac{1}{(\vec{p})^n} = \int \frac{1}{[(p^0 + i\mu)^2 + \vec{p}^2]^n} = \lim_{T \rightarrow 0} \oint \frac{1}{[(p^0 + i\mu)^2 + \vec{p}^2]^n}$$

Normal approach here is to do the  $\vec{p}$  integral first ~~at  $T=0$~~  and then ~~do the~~  $p^0$  integral last;

Two possible approaches here:

- 1.) Proceed as we did ~~before~~ in the ~~first~~ case & do the  $\vec{p}$  integral first; and also keep  $T>0$  for its part.

$$\begin{aligned}
 & T \sum_{p^0} \int \frac{1}{[\bar{p}^2 + (p^0 + i\mu)^2]^k} = T \sum_{n=0}^{\infty} \text{Res}(2\pi T) \int_{\gamma}^{d-2k} \frac{1}{\bar{p}^2 + [(n + \frac{1}{2})^2 + \frac{i\mu}{2\pi T}]^2 + (\bar{p})^2} dk \\
 & \quad \text{scale out } T \\
 & \quad \text{approx. by } \text{Res}(p + i\mu, d, \mu) \\
 & \quad \frac{1}{(\frac{1}{2} + \frac{i\mu}{2\pi T})^2 + (\frac{1}{2} + \frac{i\mu}{2\pi T})^2} = T \sum_{n=1}^{\infty} (2\pi T)^{d-2k} \left[ \int_{\gamma}^{\frac{1}{2}} \frac{1}{\bar{p}^2 + (n + \frac{1}{2} + \frac{i\mu}{2\pi T})^2} dk + \int_{\gamma}^{\frac{1}{2}} \frac{1}{\bar{p}^2 + (n + \frac{1}{2} - \frac{i\mu}{2\pi T})^2} dk \right] \\
 & \quad \left( \frac{1}{2} - \frac{i\mu}{2\pi T} \right)^2 \\
 & \quad = T \frac{(2\pi T)^{d-2k}}{(4\pi)^{d/2}} \frac{\Gamma(k - \frac{d}{2})}{\Gamma(k)} \left( \frac{e^{i\pi/4} \pi^2}{4\pi} \right)^k \times \left[ \sum_{n=1}^{\infty} \frac{1}{(n + \frac{1}{2} + \frac{i\mu}{2\pi T})^{2k-d}} + \sum_{n=1}^{\infty} \frac{1}{(n + \frac{1}{2} - \frac{i\mu}{2\pi T})^{2k-d}} \right]
 \end{aligned}$$

"Hurwitz zeta function"  $\rightarrow L_2(2k-d, \frac{1}{2} + \frac{i\mu}{2\pi T}) + \text{Res}_1(2k-d, \frac{1}{2} - \frac{i\mu}{2\pi T})$

\* \* \* 2008-08-14 next 42 pages

(2.) Other option is to just directly perform the integrals at  $T=0$

- no  $\sum_n$  necessary!

- no taking  $T \rightarrow 0$  limit at the end (note the  $\frac{1}{2\pi T} \rightarrow \infty$ ).

$$\int_{\bar{p}} \int_{p^0} \frac{dp^0}{2\pi} \frac{1}{[(p^0 + i\mu)^2 + \bar{p}^2]^n} \leftarrow \frac{1}{[(p^0 + i\mu + i\bar{p})][p^0 + i\mu - i\bar{p}]]^n}$$

$$\text{Goes as } \int_{p^0} \partial(p^0 - \mu) \left( \frac{d}{dp^0} \right)^{n-1} \frac{1}{(p^0 + i\mu + i\bar{p})^n}$$

$$\begin{cases} p^0 \\ i(p-\mu) \end{cases}$$

$$\star i(p-\mu) + i(-p-\mu)$$

\* Inspectively there has results cancellations,

one finds that the results don't agree for  $n > 2$

\* [22.08.14 479] examines why this is

$\bar{p}$  related to non-integrable divergence @  $p^0 = 0$ ,  $p \neq 0$  for  $n > 2$

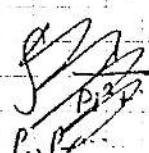
\* \* \* See Milne (Schauder) ~~Integration Terms~~  $\rightarrow$  derivatives of delta functions (homology terms), arising from  $n_F \rightarrow 0$

from or mostly to the residue theorem?

$$\frac{(d^{n-1})}{(dp^0)} \left[ \frac{n_F[i\beta(p^0 - i\mu)]}{(p^0 + i\bar{p})^n} \right]$$

~~#11~~ Cutting rules!: At  $T=0$ , these residue calculations are asked by the Boltzmann character from 0912.1856, proven in 1609.04339

$N$ -loop  
Given an  $n$ -loop integral  $F(\{P_u\}, n)$

 Given an arbitrary 1PI  $N$ -loop  $n$ -point Feynman graph  $F(\{P_u\}, n)$ , where  $P_u$ ,  $u=1, 2, \dots, n$ , are the external momenta. Then

$$F(\{P_u\}, n) = \underbrace{F_{0\text{-cut}}(\{P_u\})}_{\text{Vacuum}} + \underbrace{F_{1\text{-cut}}(\{P_u\}, n)}_{1 \text{"cut" internal ferm. line}} + \dots + \underbrace{F_{n\text{-cut}}(\{P_u\}, n)}_{n \text{"cut" internal ferm. lines.}}$$

A "cut" means:

- 1.) Remove the cut prop. from the original graph.
- 2.) Evaluate the resulting  $N-j$  loop,  $n+2j$ -point amplitude in vacuum.
- 3.) Set the cut momenta  $Q_i$  on-shell; i.e. usually  $q_0^i = iE_i$ .
- 4.) Integrate the resulting expression over the cut 3d momenta w/ weight:  $-\delta(\mu - E_i)/(2E_i)$

E.g.:   $= \text{vac}[\theta] \otimes \circ - \int \frac{\delta(\mu - E_p)}{2E_p} \left[ \begin{array}{c} \text{Feynman diagram with a cut} \\ p_0 \rightarrow iE_p \end{array} \right]$   
 $+ \int \frac{\delta(\mu - E_p)}{2E_p} \int \frac{\delta(\mu - E_q)}{2E_q} \left[ \begin{array}{c} \text{Feynman diagram with two cuts} \\ p_0 \rightarrow iE_p \\ q_0 \rightarrow iE_q \end{array} \right]$

\* Subtleties w/ higher-order propagators & intermediate divergences.

~~#12 Thermal Loops~~, #13 EFTs at High  $T$  and  $\mu$ . (Remarkably simple.)

\* D<sub>IR</sub>) for calculating perturbative pressure corrections

\* Cut rule

\* Soft momenta

\* Try your best to factorize

$$\lim_{T \rightarrow 0} \tilde{I}_k(\mu, T) = -\left(\frac{e^{\delta E \pi^2}}{4\pi}\right)^{\varepsilon} \frac{i\mu}{2\pi} \frac{\Gamma(u - \frac{d}{2})}{(4\pi)^{d/2} \Gamma(u)(1+d-2u)} \\ \left[ (i\mu)^{d-2k} - (-i\mu)^{d-2k} \right]$$

Other way ( $T=0$  & resum):

$$I_k(\mu) \xrightarrow[\text{po. Art.}]{} -\left(\frac{e^{\delta E \pi^2}}{4\pi}\right)^{\varepsilon} \frac{\mu}{\Gamma(\frac{d}{2})} \frac{\Gamma(u - \frac{d}{2})}{(4\pi)^{d/2} \Gamma(u)(1+d-2u)} \mu^{1-2k}$$

These only agree for  $k=1$ .

Zero-T limit of  $\tilde{I}_k(n, u)$  appears on p. (32)

More details of HTL derivations

$$\frac{K}{Q} \frac{Q}{Q} \frac{Q}{Q}$$

$$\Pi^{\mu\nu}(k^0, \vec{k}) = g^2 \frac{N_f}{2} S^{\mu\nu} \int \frac{\text{Tr}[R \gamma^\mu (k \bar{Q} Q) \gamma^\nu]}{P_{\bar{Q}Q}^2 (k \bar{Q} Q)^2}$$

(may be off by sign)

$$= -g^2 N_f \int \frac{S^{\mu\nu}}{2p_n^2} \left[ -u^2 + \frac{2p_n^2}{Q^2} + 2K^{\mu\nu} u^2 - \frac{-4p^2 p^2}{Q^2 (u - Q)^2} \right]$$

Focus on special components: & shift  $Q \rightarrow \cancel{Q} + k$  one term:

$$\Pi^{ij}(u) = -g^2 N_f \int \frac{2S^{ij}}{P_{\bar{Q}Q}^2} \left[ \frac{2S^{ij}}{Q^2} + \frac{-u^2 S^{ij} + 2u^i u^j - 4p^i p^j}{Q^2 (u - Q)^2} \right] \\ T \sum_{\{p_n\}} \in w_n + i\mu$$

Denoting  $E_1 = |\vec{p}_1|$ ,  $E_2 = |\vec{p}_2 - \vec{u}|$ , (energetics from here are dimensional)

$$T \sum_{\{p_n\}} \frac{1}{(w_n + i\mu)^2 + E_1^2} \approx \frac{1}{2E_1} [1 - n_F(E_1 - \mu) - n_F(E_2 + \mu)]$$

other, need  $S$  functions

$$S = T \sum_{\{p_n\}} \frac{1}{[\tilde{p}_n^2 + E_1^2][(\vec{u}_n - \vec{p}_n)^2 + E_2^2]} \\ = T \sum_{\{p_n\}} \left[ T \sum_{\{v_n\}} \beta S(\tilde{v}_n + \vec{u}_n - \vec{p}_n) \right] \frac{1}{[\tilde{p}_n^2 + E_1^2][\tilde{v}_n^2 + E_2^2]} \\ \int_0^\beta d\tau e^{i\tau(\tilde{v}_n + \vec{u}_n - \vec{p}_n)} \\ = \int_0^\beta d\tau e^{i\vec{u}_n \cdot \tau} \left\{ T \sum_{\{p_n\}} \underbrace{\frac{e^{-i\vec{p}_n \cdot \tau}}{\tilde{p}_n^2 + E_1^2}}_{(1)} \right\} \left\{ T \sum_{\{v_n\}} \underbrace{\frac{e^{i\tilde{v}_n \cdot \tau}}{\tilde{v}_n^2 + E_2^2}}_{(2)} \right\}$$

$$(2) = \frac{1}{2E_2} [n_F(E_2 - \mu) e^{(\beta - \tau)E_2 - \beta\mu} - n_F(E_2 + \mu) e^{-\tau E_2}]$$

$$(1) = \frac{1}{2E_1} [n_F(E_1 + \mu) e^{(\beta - \tau)E_1 + \beta\mu} - n_F(E_1 - \mu) e^{-\tau E_1}]$$

good exercise.

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$$\text{Then } \mathcal{G} = \int_0^\beta d\tau e^{ik_n \tau} \left\{ \dots n_F(E_1 + m) n_F(E_2 - m) e^{i\tau(\Delta E_s + \beta(-))} \dots \right\}$$

$\textcircled{T} \quad \frac{1}{4E_1 E_2}$

$$\rightarrow \frac{1}{ik_n + \Delta E_s}$$

$$\Rightarrow \mathcal{G} = \frac{1}{4E_1 E_2} \left\{ \frac{1}{ik_n - E_1 - E_2} [n_F(E_1 + m) + n_F(E_2 - m) - 1] \right.$$

$+ \dots \text{ similar} \}$

Now use the fact that  $E_1 = p$ ,  $E_2 = |p - k| \approx p - k_i \frac{\partial}{\partial p_i} |p|$

Vacuum part drops out  $\propto k^0, \bar{v}$

(vac  $\delta E \propto K^2 |p^\perp(\vec{k})|$ )

; manipulate integrals a lot, integrate by parts on  ~~$\int dp \not\propto \vec{p} \rightarrow 0$~~

$$\int_p dp n'_F(p \pm m) = -(d-1) \int_p \frac{1}{p} n(p \pm m)$$

$$\int d\Omega_v v_i v_j = \frac{\delta_{ij}}{8}$$

$$\boxed{\pi_{ij}^{(F)}(k) = g^2 N_F (d-1) \int_p \frac{1}{p} [n_F(p+m) + n_F(p-m)]}$$

$$\int d\Omega_v \frac{v_i v_j i k_n}{ik_n - k \cdot v}$$

Once you compute the other diagrams (glue)

$$\Rightarrow \pi_{ij}(k) = m_E^{-2} \int d\Omega_v \frac{v_i v_j i k_n}{ik_n - k \cdot v} + \text{subleading.}$$

$$m_E^{-2} = g^2 (d-1) \int_p \frac{1}{p} \left\{ N_C [n_F(p+m) + n_F(p-m)] + (d-1) N_C n_B(p) \right\}$$

$$\stackrel{d=3}{=} g^2 \left[ N_F \left( \frac{T^2}{6} + \frac{m^2}{2\pi^2} \right) + \frac{N_C T^2}{3} \right].$$

The last expression turns out to be:

$$\text{II}^{uv}(k) = m_e^2 \int_0^\infty (\delta^{u0}\delta^{v0} - \frac{i k^0}{k \cdot V} V^u V^v) s ds$$

$\uparrow$   
 $V = (-i, \vec{v})$

The LCM result must be written as a sum of two projectors:

$$P_T^{uv}(k) = \delta^{ui}\delta^{vj}(s_{ij} - \hat{h}_i\hat{h}_j)$$

$$P_L^{uv}(k) = \cancel{\delta^{u0}\delta^{v0}} \cancel{(k^0)} \underbrace{P_T^{uv}(k) - P_T^{uv}(k)}_{\Delta \text{ odd transv}}$$

$\Delta \text{ odd transv}$

They have the properties:

$$P_{T/L}^{uv}(k) h_m = 0 \quad (\Delta \text{-dim transv})$$

$$P_T^{uv}(k) h_i = 0 \quad (\Delta \text{-dim transv})$$

& they are orthogonal  $P_T^{ua} P_L^{av} = 0$ .

Then defining:  $\boxed{\text{II}_{ij}(k) = P_{ij}^T(k) \text{II}_T(k) + P_{ij}^L(k) \text{II}_L(k)}$

Linear algebra then gives:

$$\text{II}_T(k) = \boxed{\frac{m_e^2}{d-1} \left[ \frac{k^2}{k^2} + \frac{k^2}{k^2} \right] = \frac{m_e^2}{d-1} \left[ \frac{k_n^2}{k^2} + \frac{k^2}{k^2} \right]}$$

$$\text{II}_L(k) = \boxed{\frac{m_e^2 k^2}{k^2} (1 - L) = \frac{m_e^2 k^2}{k^2} (1 - L)}$$

$$L = \boxed{\int_0^\infty \frac{i k_n^0}{i k_n - k \cdot V} \frac{d=3}{2k} \frac{i k_n^0}{2k} \ln \left( \frac{i k_n + k}{i k_n - k} \right)}$$

*gauge constant?*

$$\Rightarrow \text{II}^{uv}(k) = m_e^2 \int_0^\infty (\delta^{u0}\delta^{v0} - \frac{i k^0}{k \cdot V} V^u V^v) s ds \quad V = (-i, \vec{v}).$$

(37-4)

There are also similar higher-point functions, e.g.:

$$\text{Diagram} + \text{Diagram} = \frac{m^2}{p \cdot Q} = m_E^2 \int V^\mu V^\nu V^\phi \left[ \frac{iQ\alpha}{p \cdot V \cdot Q \cdot V} \right] \frac{iR^\phi}{p \cdot V \cdot R \cdot V}$$

+ glue loop      gauge invariant

Very importantly, they generalized what happens:

$\mathcal{O}(m_E^{-1})$  when momenta  
soft - same  
is free vertex.

$$P^\mu S^{\text{tree}}(L, Q, R) = \Pi^\mu \epsilon(R) - \Pi^\nu \epsilon(Q)$$

## HTL structure:

\* Can I derive this better?

$$\Pi_{ab}^{uv}(k^0, k) = m_E^2 \int (\delta^{u0} \delta^{v0} - \frac{i k^0}{k \cdot v} V^u V^v) S_{ab}$$

$$@ T=0, \quad \begin{matrix} \nearrow & \downarrow \\ \hat{v} & \hat{v} \\ \uparrow & \downarrow \\ \text{normalized to 1.} \end{matrix}$$

$$m_E^2 = \frac{g^2 m_e^2}{2 \pi^2}$$

$$\text{Also Debye mass: } (\Pi_{ab}^{uv}(0, k) = m_E^2 \delta^{u0} \delta^{v0} S_{ab}).$$

$\Rightarrow$  Non-trivial functional dependence:  $-\frac{i k^0}{2k} \ln\left(\frac{i k^0 + k}{i k^0 - k}\right)$ .

$\Rightarrow$  Also similar HTL contributions for N-point gluon function.

\* At high T, also gluon contributions which are needed for dynamics in the soft long wavelength limit.

Can also write the HTL self-energy as the sum of two 4-d transverse components:

1 component,      d-1 components

$$\Pi_{HTL}^{uv} = \underbrace{\Pi_{L,0}^{uv}(k)}_{\text{3d trans}} + \underbrace{\Pi_{T,0}^{uv}(k)}_{\text{3d trans}}$$

$$\left( \begin{array}{l} \Pi_{L,0}^{uv}(k) = S^{u0} \delta^{v0} (\delta^{ij} - \frac{k^i k^j}{k^2}) \\ \text{3d trans to } k \end{array} \right)$$

$$\left( \begin{array}{l} \Pi_{T,0}^{uv}(k) = \overbrace{\Pi^{uv}(k)}^{\text{4d trans}} - \overbrace{\Pi^{uv}(k)}^{\text{3d long, to } k^i (k^i)} \\ \text{3d long, to } k^i (k^i) \end{array} \right)$$

$$\text{Mink: } \left[ \begin{array}{l} \Pi_T^{uv} = \frac{m_E^2}{2} \left[ \frac{k^2}{k^2} + \left( 1 - \frac{k^2}{k^2} \right) L(k) \right], \quad \Pi_L^{uv} = m_E^2 \left( 1 - \frac{k^2}{k^2} \right) L(k); \quad L(k) = \frac{k^0}{2k} \ln\left(\frac{k^0 + k}{k^0 - k}\right) \end{array} \right]$$

HTL Lagrangian does exist, but is nonlocal: (Mink here)

$$S_{HTL} = \frac{1}{2} \int d^4x \partial_\mu V^\alpha \partial_\mu V^\beta F_{\alpha\beta} + \frac{m_E^2}{2} \int \text{Tr} \left[ \left( \frac{1}{V \cdot \partial} V^\alpha \right) \left( \frac{1}{V \cdot \partial} V^\beta F_{\alpha\beta} \right) \right]$$

$V = (1, \vec{v})$ , D covariant derivative in adjoint representation.

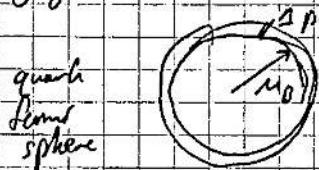
$\exists$  local description, but additional DOF are classical fields, rather than quantum fields, so estimating relative accuracy is difficult.

At shortly  $T=0$ , however, this theory is still perturbative.

$$\sim \int d^4x \int \frac{\Gamma(U, P, R)^2 d^4P}{[P^2 + \Pi(P)] [N^2 + \Pi(N)]} \sim \mathcal{O}(\alpha_s m_E^{-2})$$

$\Rightarrow$  Perturbative connection to SE  $\Pi \sim \mathcal{O}(m_E^{-2})$ !

- \* Very importantly, @ high  $\mu_B$  &  $T=0$ , no gluons w/ tail enhancement, ~~so no g scale~~ which means no ultrisoft  $g^2 \mu$  scale with nonperturbative physics! \* No zero-mode enhancement nearby  $g_c$ .
- \* There is quark parity, leading to new bosonic DOF, but this is a negligible pressure contribution at high  $T/\mu_B$



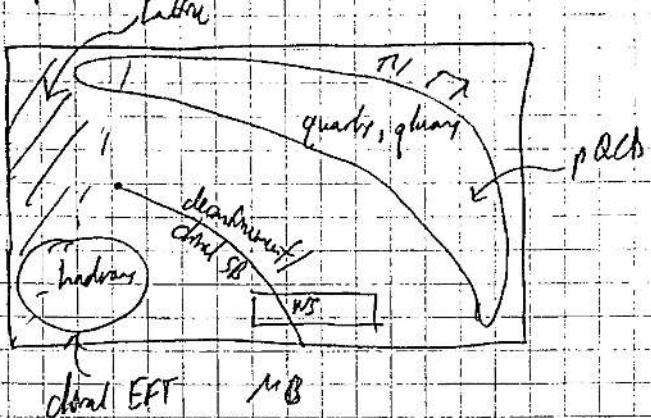
Parity  $P = (4\pi m_B^2 \cdot 1) \cdot 1 \leftarrow$  phase space  
 $\sim m_B^2 B^2 \ll m_B^4$  negligible pressure

See review by Afford + Neemod Phys. 80, 1455 (2008).

- \* Then next time we'll focus on the pressure of cold Quark Matter to N3LO, and discuss the EFT structure etc.

## # N3LO pressure of cold Quark Matter

Why do we care? Pressure of dense QCD matter @ high  $n_{\text{c}}$  unknown  
from first principles:



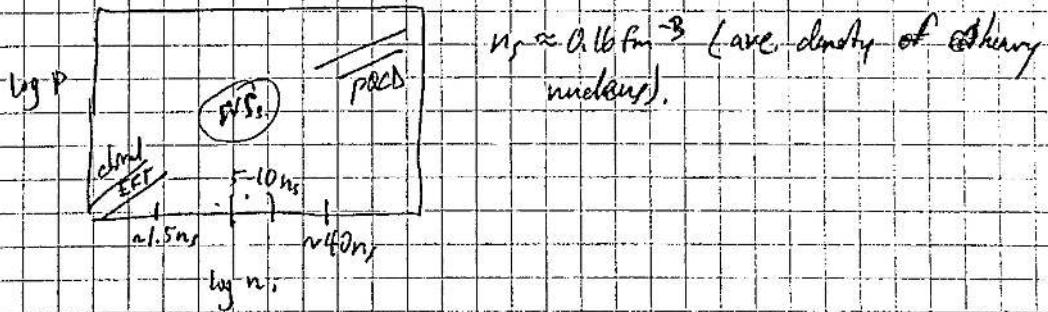
Q: What is the behavior of dense QCD matter?

Q: Do neutron stars probe high enough densities to add cold nuclear matter in their cores?

Approach of Kurkela + APJ 789 127 (2014); Annals TG 46 PRL 120, 17, (2018)

to combine chiral EFT, NS observations, and pQCD to constrain the NS-matter EOS + look for evidence of deconfined behavior in dense matter.

Nat. Phys. 16, 9, (2020)  
PRX 12, 1, (2022) ...



Recent work by Kurkela & Kurkela PRL 128, 20 (2022) and Gorda et al (APJ) 2204.11877 show that the presence of the pQCD constant has impact on NS-EOS inference. The reason is that there are very general conditions of thermodynamic stability, causality, and consistency w/ the NS-matter EOS.

Observe:

(40)

Causality means

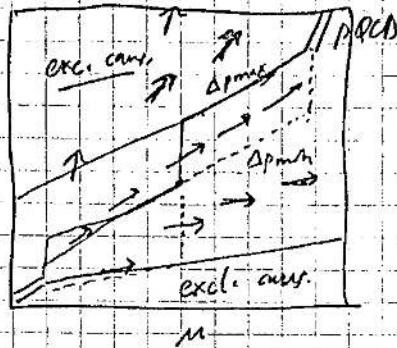
$$C^{-2} = \frac{d\ln n}{n} \frac{dn}{dm} > 1$$

$$\Rightarrow \partial^t C(u) > \frac{n}{m}$$

Schulky means:

$$\frac{\partial^2 n}{\partial m^2} > 0 \Rightarrow n''(m) > 0$$

already satisfied by causality.



consistency  $\rightarrow$  area under  
the curve is fixed

$$\int n'(u) du = P_{\text{QCD}} - P_{\text{CEFT}}$$

$$m_{\text{CEFT}} = \Delta p$$

If  $A_p \neq [A_{\text{push}}, A_{\text{max}}]$  inconsistent with CEFT + QCD

\* Of course, fundamental pQCD calculations are also just lots of sum to do, and interesting in principle.  $\approx$ .

unprimed  
Pressure of a cold Quark-matter nucleus comes from two scales.

$$P = p_0^h + \alpha_s^2 p_1^h + \alpha_s^2 p_2^h + \alpha_s^3 p_3^h \quad P \sim m_B$$

$$+ \alpha_s^2 p_2^s + \alpha_s^2 p_3^s \quad P \sim m_E \text{ and } m_B$$

$$+ \alpha_s^3 p_3^m \quad \text{Interaction between scales.}$$

Starting at  $O(\alpha_s^2)$ , there are IR singularities, leading to the need for renormalization.

$$\frac{m}{m} \sim \frac{d^4 p}{(p^2)^2} \sim \int \frac{\pi^2 \delta(m, u)}{(p^2)^2} \sim \frac{d^4 p}{p^4} \sim \log \text{diluent}$$

We thus expect some sort of  $m_E^4 \log(\cdot)$  terms in the pressure, with some ratio of scales appearing in the logarithm. What scales?  $m_B$ ? Expect  $\frac{m_B}{m_E}$ ?

~~$$\frac{m}{m} \sim \frac{d^4 p}{(p^2)^2} \sim \int \frac{\pi^2 \delta(m, u)}{(p^2)^2} \sim \frac{d^4 p}{p^4} \sim \log \text{diluent}$$~~

so expect there to be a  $m_E^4 \log\left(\frac{m_B}{m_E}\right)$  term. This is different from  $m_B^4$  -

T because more than just the  $\Theta(n=0)$  mode is contributing here.

(41)

The O-mole at high  $T$  was going  $\pi_E^{3/2}$  contribution,  $\pi_{IR}^{1/2}$  enhanced.  
 This is only by  $(g)$  enhanced. Because of this by substitution, actually easy to  
 derive the coefficient of this logarithm, this is because it arises from between  
 the two [UV of soft theory ( $\text{HTL}$ ) = IR of hard theory.]

Explicitly:

$$P \rightarrow \underbrace{\text{HTL}}_{\text{HTL}} + \underbrace{\text{IR}}_{\text{IR}} + \underbrace{\text{IR}}_{\text{IR}} + \dots = \text{Pres}(\Pi_1)$$

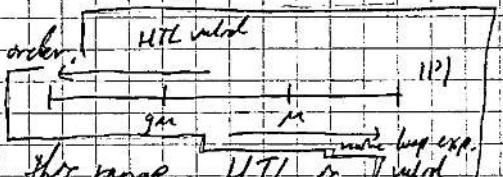
$$= [\text{Pres}(\Pi_1) + \text{Pres}(\Pi_{IR})] + \text{Pres}(\Pi_{IR})$$

IR rate! can  
loop expand!

$$\rightarrow \boxed{\text{Pres}(\Pi_1) = \text{Pres}(\Pi_{HTL}) + [\text{loop}(\Pi_1) - \text{Pres}(\Pi_{HTL})]}$$

$\rightarrow 0 \rightarrow 0$  Reg at  $T=0$

$$= \boxed{\underbrace{\text{HTL}}_{\text{HTL}} + \underbrace{\text{IR}}_{\text{IR}}} \text{ off this order.}$$



But now consider  $|\vec{p}| \ll |\vec{P}| \ll \mu$ . In this range, HTL is more loop exp.  
 enough, b/c  $|\vec{P}| \ll \mu$ ; but a loop expansion works b/c  $g \gg |\vec{P}| \gg g_m$ .

$$\Rightarrow \cancel{\text{HTL}} \cancel{\frac{d^4 p}{(2\pi)^4} \frac{d^4 P}{(2\pi)^4} \frac{d^4 T}{(2\pi)^4}}$$

vacuum subtracted

$$\Rightarrow P \rightarrow -\frac{da}{2} \int \frac{d^4 p}{(2\pi)^4} \left[ 2 \ln \left( \frac{|\vec{p}|^2 + \frac{\mu^2}{p^2}}{1} \right) + \ln \left( 1 + \frac{\mu^2}{p^2} \right) \right]$$

$$\approx -\frac{da}{2} \int \frac{d^4 p}{(2\pi)^4} \left[ \frac{\mu^2}{p^2} + \frac{\frac{\mu^2}{2}}{\frac{p^2}{2}} \right]$$

$$= -\frac{da}{(4\pi)^2} \left[ \frac{m^2}{m^2 - g_m^2} - \frac{4}{m^2} \ln \left( \frac{m}{g_m} \right) \right] + \frac{da}{(4\pi)^2} m^2 \ln \left( \frac{A}{g} \right)$$

Here,  $m^2 = \frac{m_E^2}{2}$ . By the asymptotic off-mass-shell mass: Transverse pole given

$$by \gamma^2 = k^2 + m^2$$

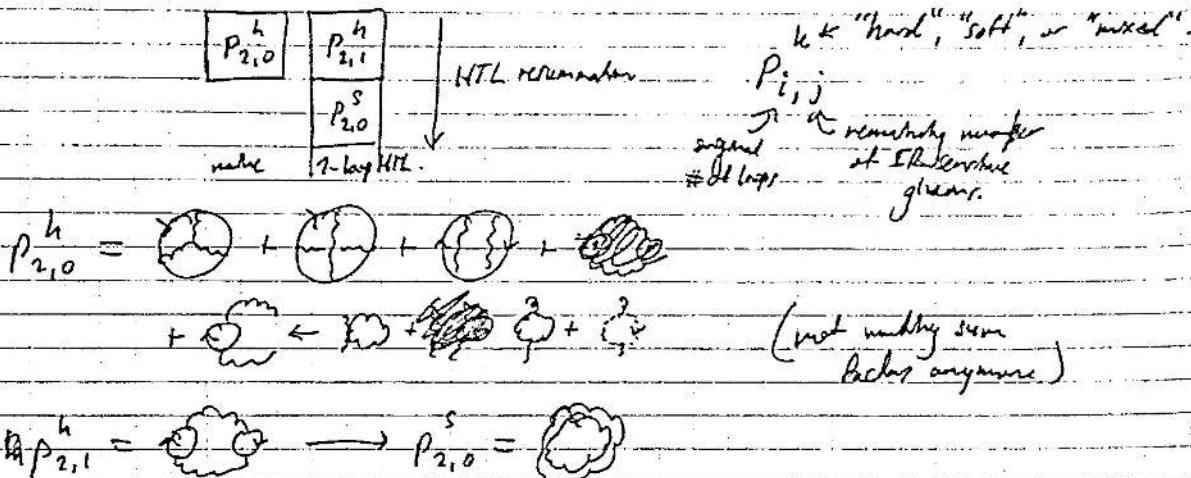
\* It agrees with the full diagrammatic calculations from Freedman & Melleson (1977) PRD 16, 100 papers!

(Multiple papers.)

(42)

to approach was demonstrated in Gorda + PRL 121 08020 (2018).

Note that we've perturbed soft and hard diagrams that are related to each other (the soft "expands into" the hard one): We can group these diagrams as:



The "easiest" way to compute the full  $O(\alpha_s^2)$  pressure of cold QM is to compute all of these terms ~~separately~~<sup>one-by-one</sup>, then add them together.

\* Since  $P_{\text{loop}}(\Pi_{\text{HTL}})$  is scale-independent to  $\Lambda$ -theory, there is no double-counting!

Schematically, the results appear as:

~~$$P_{2,0}^h = \frac{d_4 m_E^4}{(8\pi)^2} \left( \frac{m_E}{\Lambda_h} \right)^{-2\varepsilon} \left( \frac{P_{-1}^{WZLO}}{\varepsilon} + P_0^{WZLO} \right)$$~~

$$P_{2,0}^S = \frac{d_4 m^4}{(8\pi)^2} \left( \frac{m}{\Lambda_h} \right)^{-2\varepsilon} \left( \frac{P_{-1}^{WZLO}}{\varepsilon} + P_0^{WZLO} \right)$$

An "interaction scale"  
(but still  $\Lambda$ -theory scale)

~~$$P_{2,1}^h = \frac{d_4 m^4}{(8\pi)^2} \left( \frac{m}{\Lambda_h} \right)^{-2\varepsilon} \left( \frac{P_{-1}^{WZLO}}{\varepsilon} + P_0^{WZLO} \right)$$~~

↑ Frequency cancel, but leave behind  
 $P_{-1} \log \left( \frac{m_E}{m} \right)$ .

$P_{2,0}^h$  made after renormalization

↑ polar angle,  $\phi = \tan^{-1} \left( \frac{p}{p_0} \right)$ .

~~at WZLO the terms are the same~~

$$(43) \quad P_{2,0}^S \sim \int d\Omega d\phi \left[ S(d-1) \log \left( P^2 + \Pi_r(\phi) \right) \right] + \log \left( P^2 + \Pi_r(\phi) \right)$$

$$\left[ \text{at WZLO level} \right] = \int_P d\Omega d\phi \left[ \log \left( P^2 + \Pi_r(\phi) \right) \right]$$

The result  $\sigma$  of the sum:

$$\left\langle \frac{1}{\varepsilon} \left[ \text{Tr} \{ \Pi^2 \log(\Pi) \} \right] + \left[ \dots \right] \right\rangle_{\phi}$$

+ ...

Angular average over 4d  
polar angle

$$\langle \phi \rangle_{\phi} = \frac{\int d\Omega \phi^2 d\Omega}{S_0 \pi^2 d\Omega^2}$$

$$\rightarrow \frac{d_4 m_E^4}{(8\pi)^2} \left[ + \frac{1}{2\varepsilon} - \log\left(\frac{m_E}{\Lambda_h}\right) - \frac{3}{4} + \frac{\pi^2}{6} - \frac{8}{4} - \frac{5\log(2)}{6} + \frac{4\ln^2 2}{3} \right]$$

$\delta = -0.85638\dots$ , arises from  $\langle \text{Tr} \{ \Pi^2 \log(\Pi) \} \rangle_{\phi}$  integral.

The  $N3LO$  pressure has a similar structure, but the fully IR- $\sigma$  now the two-loop HTL diagrams!

$$P_{3,0}^S = \text{Diagram 1} + \text{Diagram 2} + \text{Diagram 3} \quad (\text{not really symmetric}).$$

and now there's a more complicated "pressure table":

	$P_{3,0}^h$	$P_{3,1}^h$	$P_{3,2}^h$	
hard	$P_{3,0}^h$	$P_{3,1}^h$	$P_{3,2}^h$	
mixed	$P_{3,0}^m$	$P_{3,1}^m$	$P_{3,2}^m$	
soft	$P_{3,0}^s$	$P_{3,1}^s$	$P_{3,2}^s$	
Name	1-loop HTL	2-loop HTL		
			HTL reconnection	For example:
				$P_{5,0}^h \Rightarrow$
				$P_{3,2}^h \Rightarrow$
				$P_{3,1}^m \Rightarrow$
				$P_{3,0}^s \Rightarrow$
				"2-loop HTL diagrams"
				" $\Pi_{\text{IR}}^{(1\text{loop})} = K^2 \Pi(\phi)$ " power corrections
				" $\Pi_{\text{IR}}^{(2\text{loop})} = K^2 \Pi(\phi)$ " 2-loop corrections
	$P_{3,1}^h \Rightarrow$			
	$P_{3,0}^h \Rightarrow$			

to go beyond LO HTL self-energy - need 2-loop corrections and power corrections

[Gorda, Pihlström, Vuorinen, Pankka, Säppi, Vuorinen PRD 127 16 (2021)  
computed soft 2-loop HTL pressure]

(44)

$$P_{3,0}^n = \frac{g^2 N_c \alpha_s g^2 m_E^4}{(2\pi)^6} \left(\frac{m_E}{\lambda_n}\right)^{-4E} \left[ \frac{P_{-2}}{(2\varepsilon)^2} + \frac{P_{-1}}{2\varepsilon} + P_0 \right]$$

$$P_{-2} = \frac{11}{6} \int_D \text{Tr} \{ \Pi_0(\vec{u})^2 \} = \frac{11}{6} \left( \frac{\pi^2}{4} \right).$$

$$P_{-1} = \int_D \left\{ \frac{19+11\pi^2}{72} - \frac{11}{6} \text{Tr} \{ \Pi_0(\vec{u})^2 \ln [\Pi_0(\vec{u})] \} \right\}$$

~~$\int_0^{\pi/2} dx \sin(2x) \text{Tr} \{ 8V_{3g}^{12} (\vec{u} \sin x, \vec{p} \cos x) \}$~~

$$\int_0^{\pi/2} dx \sin(2x) \text{Tr} \left\{ 8V_{3g}^{12} (\vec{u} \sin x, \vec{p} \cos x) \right\} / 24 [1 + \sin(2x) \vec{u} \cdot \vec{p}]$$

$$= 11.6840(15)$$

From  topology, and bare propagators,  
HTL vertices.

$$P_0 = 17.150(7)$$

(See 2111.11944 for a readable summary of the full calculation)

The self-energy corrections  $\Pi_{\text{HTL},0}^{(2)}$  and  $\Pi_{\text{HTL}}^{(\text{pert})}$  were first computed in QED by Mannel, Soto, Stelzer PRD 94 025017 (2016); Cagnacci, Mannel, Soto PLB 780, 308 (2018); Cagnacci, Landry, Soto PLB 801, 135193 (2020) (order + PRD, 036012; L031501 (2023), Ekelstadt)

Gluon self-energy recently computed by Ekelstadt 2302.04894 @ LHC T and now Gorda, Prabhatan, Sippi, Sippinen 2304.09187  
Ekelstadt 2302.04894, 2304.09255

Interestingly, new angular structures appear, as does a new soft scale that controls a different combination of  $T$  and  $m$  than  $m_E$  does:

$$\Pi_{\text{HTL}}^{(2)}(n) = \# \int \frac{d^3x}{4\pi} \left\{ V^{\mu\nu} V^{\rho\sigma} \left[ \frac{(n^\rho)^2}{L(n^\mu n^\nu)} - \frac{2n^\rho}{n^\mu n^\nu} \right] + [V^{\mu\nu} n^\rho + n^\mu V^{\rho\nu}] \frac{n^\rho}{n^\mu n^\nu} \right\}$$

(see Ekelstadt 2302.04894 for a derivation) using kinetic theory.

$$(45) \quad m_E^2 \sim \left( \frac{T^2}{3} + \frac{\mu^2}{\pi^2} \right), \quad \Pi_{\text{HTL}}^{(2)} \sim \left( T^2 + \frac{\mu^2}{\pi^2} \right) \quad \text{in QED}$$

(see Bush + 2304.09187)

For the  $T=0$  pressure contribution  $P_m$  or anything