

# **IB Electromagnetism**

Ishan Nath, Lent 2023

Based on Lectures by Prof. Gordon Ogilvie

January 25, 2023

## Contents

<b>1</b>	<b>Introduction</b>	<b>2</b>
1.1	Charges and Currents . . . . .	2
1.2	Fields and Forces . . . . .	3
1.3	Maxwell's equations . . . . .	3
1.4	Units . . . . .	4
<b>2</b>	<b>Electrostatics</b>	<b>5</b>
2.1	Gauss' Law . . . . .	5
2.1.1	Spherical Symmetry . . . . .	5
2.1.2	Cylindrical Symmetry . . . . .	6
2.1.3	Planar Symmetry . . . . .	7
2.1.4	Surface Charge and Discontinuity . . . . .	8
	<b>Index</b>	<b>9</b>

# 1 Introduction

## 1.1 Charges and Currents

*Electric charge* is a physical property of elementary particles. It is:

- Positive, negative or zero.
- Quantized (an integer multiple of the *elementary charge*  $e$ ).
- Conserved (even if particles are created or destroyed).

By convention, the electron has charge  $-e$ , the proton has charge  $+e$ , and the neutron has charge 0.

On macroscopic scales, the number of particles is so large that charge can be considered to have continuous *electric charge density*  $\rho(\mathbf{x}, t)$ . The total charge in a volume  $V$  is then

$$Q = \int_V \rho \, dV.$$

The *electric current density*  $\mathbf{J}(\mathbf{x}, t)$  is the flux of electric charge per unit area. The current flowing through a surface  $S$  is

$$I = \int_S \mathbf{J} \cdot d\mathbf{S}.$$

Consider a time-independent volume  $V$  with boundary  $S$ . Since charge is conserved, we have

$$\begin{aligned} \frac{dQ}{dt} &= -I, \\ \frac{d}{dt} \int_V \rho \, dV + \int_S \mathbf{J} \cdot d\mathbf{S} &= 0, \\ \int_V \left( \frac{\partial \rho}{\partial t} + \nabla \cdot \mathbf{J} \right) dV &= 0. \end{aligned}$$

Since this is true for any  $V$ , we must have

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \mathbf{J} = 0.$$

This *equation of charge conservation* has the typical form of a conservation law.

The discrete charge distribution of a single particle of charge  $q_i$ , and position vector  $\mathbf{x}_i(t)$  is

$$\begin{aligned}\rho &= q_i \delta(\mathbf{x} - \mathbf{x}_i(t)), \\ \mathbf{J} &= q_i \dot{\mathbf{x}}_i \delta(\mathbf{x} - \mathbf{x}_i(t)).\end{aligned}$$

For  $N$  particles, it is

$$\begin{aligned}\rho &= \sum_{i=1}^N q_i \delta(\mathbf{x} - \mathbf{x}_i(t)), \\ \mathbf{J} &= \sum_{i=1}^N q_i \dot{\mathbf{x}}_i \delta(\mathbf{x} - \mathbf{x}_i(t)).\end{aligned}$$

We can verify that these distributions satisfy the charge conservation equation.

## 1.2 Fields and Forces

Electromagnetism is a *field theory*. Charged particles interact not directly, but by generating fields around them that are experienced by other charged particles.

In general, we have two time-dependent vector fields: the *electric field*  $\mathbf{E}(\mathbf{x}, t)$ , and the *magnetic field*  $\mathbf{B}(\mathbf{x}, t)$ .

The *Lorentz force* on a particle of charge  $q$  and velocity  $\mathbf{v}$  is

$$\mathbf{F} = q(\mathbf{E} + \mathbf{v} \times \mathbf{B}).$$

## 1.3 Maxwell's equations

In this course we will explore some consequences of *Maxwell's equations*

$$\begin{aligned}\nabla \cdot \mathbf{E} &= \frac{\rho}{\epsilon_0}, & \nabla \cdot \mathbf{B} &= 0, \\ \nabla \times \mathbf{E} &= -\frac{\partial \mathbf{B}}{\partial t}, & \nabla \times \mathbf{B} &= \mu_0 \left( \mathbf{J} + \epsilon_0 \frac{\partial \mathbf{E}}{\partial t} \right).\end{aligned}$$

Some properties of Maxwell's equations are:

- They are coupled linear PDE's in space and time.
- They involve two positive constants:  $\epsilon_0$  (vacuum permittivity), and  $\mu_0$  (vacuum permeability).
- Charges ( $\rho$ ) and currents ( $\mathbf{J}$ ) are the sources of the electromagnetic fields.

- Each equation has an equivalent integral form, related via the divergence theorem of Stokes' theorem.
- These are the vacuum equations that apply on microscopic scales (or in a vacuum). A related macroscopic version applies in media (for examples air).
- The equations are consistent with each other and with charge conservation. For example,  $\nabla \cdot (M3) = \frac{\partial}{\partial t}(M2)$ , and

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \mathbf{J} = \frac{\partial}{\partial t}(\epsilon_0 \nabla \cdot \mathbf{E}) + \nabla \cdot \left( -\epsilon_0 \frac{\partial \mathbf{E}}{\partial t} + \frac{1}{\mu_0} \nabla \times \mathbf{B} \right) = 0.$$

## 1.4 Units

The SI unit of electric charge is the coulomb (C). The elementary charge is (exactly)

$$e = 1.602\,176\,634 \times 10^{-19} \text{ C}.$$

The SI unit of electric current is the ampere, or amp (A), equal to  $1 \text{ C s}^{-1}$ .

The SI base units needed in electromagnetism are:

second (s)

metre (m)

kilogram (kg)

ampere (A)

From the Lorentz force law, we can see that the units of  $\mathbf{E}$  and  $\mathbf{B}$  must be

$$\text{kg m s}^{-3} \text{ A}^{-1} \text{ and } \text{kg s}^{-2} \text{ A}^{-1}.$$

The latter is also called the tesla (T). From Maxwell's equations, we can work out the units of  $\epsilon_0$  and  $\mu_0$ . The experimentally determined values are

$$\begin{aligned} \epsilon_0 &= 8.854 \dots \times 10^{-12} \text{ kg}^{-1} \text{ m}^{-3} \text{ s}^4 \text{ A}^2 \\ \mu_0 &= 1.256 \dots \times 10^{-6} \text{ kg m s}^{-2} \text{ A}^{-2} \\ &\approx 4\pi \times 10^{-7} \text{ kg m s}^{-2} \text{ A}^{-2}. \end{aligned}$$

The speed of light is (exactly)

$$c = \frac{1}{\sqrt{\mu_0 \epsilon_0}} = 299\,792\,458 \text{ m s}^{-1} \approx 3 \times 10^8 \text{ m s}^{-1}.$$

## 2 Electrostatics

In a time-independent situation, Maxwell's equations reduce to

$$\begin{aligned}\nabla \cdot \mathbf{E} &= \frac{\rho}{\epsilon_0}, & \nabla \times \mathbf{E} &= \mathbf{0}, \\ \nabla \cdot \mathbf{B} &= 0, & \nabla \times \mathbf{B} &= \mu_0 \mathbf{J}.\end{aligned}$$

Since  $\mathbf{E}$  and  $\mathbf{B}$  are decoupled, we can study them separately.

*Electrostatics* is the study of the electric field generated by a stationary charge distribution

$$\nabla \cdot \mathbf{E} = \frac{\rho}{\epsilon_0}, \quad \nabla \times \mathbf{E} = \mathbf{0}.$$

### 2.1 Gauss' Law

Consider a closed surface  $S$  enclosing a volume  $V$ . Integrating (M1) over  $V$  and using the divergence theorem, we obtain *Gauss' law*

$$\int_S \mathbf{E} \cdot d\mathbf{S} = \frac{Q}{\epsilon_0},$$

where  $Q = \int_V \rho dV$  is the total charge in  $V$ .

Gauss' law is the integral version of (M1) and is valid generally. This says that the electric flux of a closed surface is proportional to the total charge enclosed.

In special situations, we can use Gauss' law together with symmetry to deduce  $\mathbf{E}$  from  $\rho$ . By choosing the *Gaussian surface*  $S$  appropriately.

#### 2.1.1 Spherical Symmetry

Consider a spherically symmetric charge distribution,  $\rho(r)$  in spherical polar coordinates, with total charge  $Q$  contained within an outer radius  $R$ .

To have spherical symmetry, the electric field should have the form

$$\mathbf{E} = E(r)\mathbf{e}_r.$$

This will satisfy (M3'), as required. To find  $E(r)$ , we apply Gauss' law to a sphere of radius  $r$ . If  $r > R$ , then

$$\int_S \mathbf{E} \cdot d\mathbf{S} = E(r) \int_S \mathbf{e}_r \cdot d\mathbf{S} = E(r) \int_S dS = E(r) 4\pi r^2 = \frac{Q}{\epsilon_0}.$$

Thus, outside of the sphere of radius  $R$ ,

$$\mathbf{E} = \frac{Q}{4\pi\epsilon_0 r^2} \mathbf{e}_r.$$

So the external electric field of a spherically symmetric body depends only on the total charge.

The Lorentz force on a particle of charge  $q$  in  $r > R$  is

$$\mathbf{F} = q\mathbf{E} = \frac{Qq}{4\pi\epsilon_0 r^2} \mathbf{e}_r.$$

This is the *Coulomb force* between charged particles. The force is repulsive if the charges have the same sign ( $Qq > 0$ ) and attractive if they have opposite signs ( $Qq < 0$ ).

If we take the limit  $R \rightarrow 0$ , we obtain the electric field of a *point charge*  $Q$ , corresponding to

$$\rho = Q\delta(\mathbf{x}).$$

There is a close analogy between the Coulomb force and the gravitational force between massive particles,

$$\mathbf{F} = -\frac{GMm}{r^2} \mathbf{e}_r.$$

Both involve an inverse-square law, and the product of the charges/masses. However,

- While gravity is always attractive, electric forces can be repulsive or attractive.
- Gravity is very much weaker than the Coulomb force, e.g. for two protons the ratio of the electric to gravitational forces is

$$\frac{e^2}{4\pi\epsilon_0 G m_p^2} \approx 10^{36}.$$

On the atomic scale, gravity is irrelevant. But positive and negative charges balance so accurately that on the planetary scale, gravity is dominant.

### 2.1.2 Cylindrical Symmetry

Consider a cylindrically symmetric charge distribution  $\rho(r)$  in cylindrical polar coordinates, with total charge  $\lambda$  per unit length, contained within an outer radius  $R$ .

To have cylindrical symmetry,

$$\mathbf{E} = E(r)\mathbf{e}_r.$$

To find  $E(r)$  we apply Gauss' law to a cylinder of radius  $r$  and arbitrary length  $L$ . Again, we consider  $r > R$ . Then, since only the curved part of the cylinder contributes to the flux,

$$\int_S \mathbf{E} \cdot d\mathbf{S} = E(r) \int_S \mathbf{e}_r \cdot d\mathbf{S} = E(r) \int_S dS = E(r)2\pi rL = \frac{\lambda L}{\epsilon_0}.$$

Thus, we get

$$\mathbf{E} = \frac{\lambda}{2\pi\epsilon_0 r} \mathbf{e}_r.$$

In the limit  $R \rightarrow 0$ , we obtain the electric field of a *line charge*  $\lambda$  per unit length, corresponding to

$$\rho = \lambda\delta(x)\delta(y).$$

### 2.1.3 Planar Symmetry

We consider a planar charge distribution  $\rho(z)$  in Cartesian coordinates, with total charge  $\sigma$  per unit area, contained within a region  $-d < z < d$  of thickness  $2d$ . We assume reflectional symmetry, so  $\rho(z)$  is even.

To have planar symmetry, we need

$$\mathbf{E} = E(z)\mathbf{e}_z,$$

which will satisfy (M3'). Reflectional symmetry implies  $E(-z) = -E(z)$ . To find  $E(z)$  for  $z > 0$ , apply Gauss' law to a "Gaussian pillbox" of height  $2z$  and arbitrary area  $A$ . If  $z > d$ , then

$$\int_S \mathbf{E} \cdot d\mathbf{S} = E(z)A - E(-z)A = 2E(z)A = \frac{\sigma A}{\epsilon_0}.$$

Thus,

$$\mathbf{E} = \begin{cases} \frac{\sigma}{2\epsilon_0} \mathbf{e}_z & z > d, \\ -\frac{\sigma}{2\epsilon_0} \mathbf{e}_z & z \leq -d. \end{cases}$$

In the limit  $d \rightarrow 0$ , we obtain the electric field of a *surface charge*  $\sigma$  per unit area, corresponding to

$$\rho = \sigma\delta(z).$$



### 2.1.4 Surface Charge and Discontinuity

Let  $\mathbf{n}$  be a unit vector normal to the charged surface, pointing from region 1 to region 2. In our example,  $\mathbf{n} = \mathbf{e}_z$ .

The discontinuity in  $\mathbf{E}$  is given by

$$[\mathbf{n} \cdot \mathbf{E}] = \frac{\sigma}{\epsilon_0},$$

where  $\sigma$  is the surface charge density, and

$$[X] = X_2 - X_1$$

denotes a discontinuity. The tangential components are continuous (they are both 0), so

$$[\mathbf{n} \times \mathbf{E}] = \mathbf{0}.$$

# Index

- Coulomb force, 6
- electric charge, 2
- electric charge density, 2
- electric field, 3
- electrostatics, 5
- equation of charge conservation, 2
- field theory, 3
- Gauss' law, 5
- Gaussian surface, 5
- line charge, 7
- Lorentz force, 3
- magnetic field, 3
- Maxwell's equations, 3
- point charge, 6
- surface charge, 7
- vacuum permeability, 3
- vacuum permittivity, 3