Many examples of normal forms for stochastic or non-autonomous differential equations

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Instructions

- Download and install the computer algebra package $Reduce^1$ via http://www.reduce-algebra.com
- Navigate to folder Examples within folder StoNormForm.
- For each example of interest, start-up *Reduce* and enter the command in_tex "filename.tex"\$ where filename is the root name of the example (as listed in the following table of contents).

The results involve convolutions over the non-autonomous/stochastic factors (Roberts 2008, 2015b, (12) and §19.2, respectively):

$$e^{\mu t} \star v := \begin{cases} \int_{-\infty}^t e^{\mu(t-\tau)} v(\tau) d\tau & \Re \mu < 0, \\ \int_t^\infty e^{\mu(t-\tau)} v(\tau) d\tau & \Re \mu > 0, \end{cases}$$

which has the crucial differential property that

$$\frac{d}{dt}e^{\mu t}\star v = \mu(e^{\mu t}\star v) - (\operatorname{sgn} \Re \mu)v.$$

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¹ We use the computer algebra package *Reduce* because it is both free and perhaps the fastest general purpose computer algebra system (Fateman 2003, e.g.).

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1 ratodes: Simple rational ODEs

A simple system of fast/slow odes in rational functions is

$$\dot{x} = -\frac{xy}{1+z}, \quad \dot{y} = -\frac{y}{1+2y} + x^2, \quad \dot{z} = 2\frac{z}{1+3x}.$$
 (1)

Use x(1) to denote variable x, y(1) to denote variable y, and z(1) to denote z. Multiply each ODE by the denominator for the ODE and shift the nonlinear d/dt terms to the right-hand side.

Start by loading the procedure.

1 in_tex "../stoNormForm.tex"\$

Execute the construction of a normal form for this system.

2 stonormalform(
3 {-x(1)*y(1)-z(1)*df(x(1),t)},

The procedure embeds the system as the $\varepsilon = 1$ version of the family

$$\dot{x}_1 = \varepsilon \left(-\frac{\mathrm{d} x_1}{\mathrm{d} t} z_1 - x_1 y_1 \right)$$

$$\dot{y}_1 = \varepsilon \left(-2 \frac{\mathrm{d} y_1}{\mathrm{d} t} y_1 + 2x_1^2 y_1 + x_1^2 \right) - y_1$$

$$\dot{z}_1 = -3\varepsilon \frac{\mathrm{d} z_1}{\mathrm{d} t} x_1 + 2z_1$$

Time dependent coordinate transform

$$z_1 = 6\varepsilon^2 X_1 Y_1 Z_1 + Z_1$$

$$y_1 = \varepsilon^2 (2X_1^4 - 4X_1^2 Y_1^2 + 6Y_1^3) + \varepsilon (X_1^2 - 2Y_1^2) + Y_1$$

$$x_1 = \varepsilon^2 (2X_1^3 Y_1 - 1/2X_1 Y_1^2 + X_1 Y_1 Z_1) + \varepsilon X_1 Y_1 + X_1$$

Result normal form DEs

$$\dot{Z}_1 = -54\varepsilon^3 X_1^3 Z_1 + 18\varepsilon^2 X_1^2 Z_1 - 6\varepsilon X_1 Z_1 + 2Z_1
\dot{Y}_1 = 8\varepsilon^3 X_1^4 Y_1 + 4\varepsilon^2 X_1^2 Y_1 + 2\varepsilon X_1^2 Y_1 - Y_1
\dot{X}_1 = \varepsilon^3 \left(-2X_1^5 - 2X_1 Y_1^2 Z_1 \right) - \varepsilon^2 X_1^3$$

2 futureNoise: Future noise in the transform

An interesting pair of fast/slow SDEs derived from stochastic advection/dispersion is

$$\dot{x} = -\sigma y w(t)$$
 and $\dot{y} = -y + \sigma x w(t)$, (2)

where lowercase w(t) denotes the formal derivative dW/dt of a Stratonovich Wiener process $W(t,\omega)$. Parameter σ controls the strength of the noise. In stochastic advection/dispersion parameter σ represents the lateral wavenumber of the concentration profile.

Start by loading the procedure.

Execute the construction of a normal form for this system.

Being linear in x, y the nonlinear parameter ε does not appear in the analysis and results. Consequently, the procedure analyses the system as prescribed (since given w changed to σw). The interest in this example is the noise and the noise-noise interactions. As usual, the noise-noise interactions are truncated to errors $\mathcal{O}(\sigma^3)$.

Time dependent coordinate transform

$$y_1 = \sigma e^{-1t} \star w_1 X_1 + Y_1$$

 $x_1 = \sigma e^t \star w_1 Y_1 + X_1$

Result normal form DEs

$$\dot{Y}_1 = \sigma^2 e^t \star w_1 \, w_1 Y_1 - Y_1
\dot{X}_1 = -\sigma^2 e^{-1t} \star w_1 \, w_1 X_1$$

The interesting aspect of this example is the explicit presence of non-Markovian, future time integrals, anticipation integrals, in the convolutions $e^t \star w_1$. These appear in both the coordinate transform, and the evolution *off* the stochastic slow manifold. But, as guaranteed by theory, they do not appear on the stochastic slow manifold.

Further, this example could go to higher order noise-noise interactions very quickly, that is, to higher orders in σ . However, I do not compute such higher order terms in this code.

3 othersFail: Other methodologies fail

Consider, for small bifurcation parameter ϵ , the system

slow mode
$$\dot{x} = \epsilon x + x^3 - (1 - \sigma w)xy$$
, fast mode $\dot{y} = -y + x^2 + y^2 + \sigma yw$.

Deterministically, there is a bifurcation to two equilibria for small $\epsilon > 0$. The noise w affects this bifurcation somehow.

Why is this tricky? Cross-sectional averaging is simply projection onto the slow space y=0 which predicts instability of subcritical bifurcation $\dot{x}=\epsilon x+x^3$. Whereas adiabatic approximation, singular perturbation, and multiple scales set $\dot{y}=0$ whence $y\approx x^2$ and thus predict only the linear growth of $\dot{x}=\epsilon x$. Our normal form transforms get the deterministic dynamics correctly. But what happens for stochastic dynamics?

Start by loading the procedure.

Execute the construction of a normal form for this system. Multiply a cubic terms in the x SDE in order to count orders of approximation best (since the right-hand side is multiplied by small). Multiply the bifurcation parameter by small in order to make it scale with ε^2 .

```
16 stonormalform(
17 {small*epsilon*x(1)+small*x(1)^3
18 -x(1)*y(1)*(1-small*w(1))},
19 {-y(1)+x(1)^2+y(1)^2+y(1)*w(1)},
20 {},
21 5)$
22 end;
```

With the above artifices, the procedure analyses the following system which reduce to the given one for $\varepsilon = 1$:

$$\dot{x}_1 = \sigma \varepsilon w_1 x_1 y_1 + \varepsilon^2 \left(x_1^3 + x_1 \epsilon \right) - \varepsilon x_1 y_1$$
$$\dot{y}_1 = \sigma w_1 y_1 + \varepsilon \left(x_1^2 + y_1^2 \right) - y_1$$

Time dependent coordinate transform This transform is quite complicated, due to the noise, and involve fast-time future and history integrals. The deterministic terms at the end.

$$y_{1} = \sigma \varepsilon^{3} \left(-4e^{t} \star e^{t} \star w_{1} X_{1}^{2} Y_{1}^{2} + 4e^{-1t} \star e^{-1t} \star w_{1} X_{1}^{4} - 2e^{-1t} \star e^{-1t} \star w_{1} X_{1}^{2} \epsilon + 2e^{2t} \star w_{1} X_{1}^{2} Y_{1}^{2} - 10e^{t} \star w_{1} X_{1}^{2} Y_{1}^{2} - 3e^{t} \star w_{1} Y_{1}^{4} + e^{-1t} \star w_{1} X_{1}^{4} + 3e^{-1t} \star w_{1} X_{1}^{2} Y_{1}^{2} - 2e^{-1t} \star w_{1} X_{1}^{2} \epsilon) + \sigma \varepsilon^{2} \left(2e^{t} \star w_{1} Y_{1}^{3} - 2e^{-1t} \star w_{1} X_{1}^{2} Y_{1} \right) + \sigma \varepsilon \left(-e^{t} \star w_{1} Y_{1}^{2} + e^{-1t} \star w_{1} X_{1}^{2} \right) + \varepsilon^{3} \left(X_{1}^{4} - 7X_{1}^{2} Y_{1}^{2} - 2X_{1}^{2} \epsilon - Y_{1}^{4} \right) + \varepsilon^{2} Y_{1}^{3} + \varepsilon \left(X_{1}^{2} - Y_{1}^{2} \right) + Y_{1}$$

$$x_{1} = \sigma \varepsilon^{3} \left(-e^{3t} \star w_{1} X_{1} Y_{1}^{3} + e^{2t} \star w_{1} X_{1} Y_{1}^{3} + 3e^{t} \star w_{1} X_{1}^{3} Y_{1} \right) + \sigma \varepsilon^{2} \left(e^{2t} \star w_{1} X_{1} Y_{1}^{2} - e^{t} \star w_{1} X_{1} Y_{1}^{2} + e^{-1t} \star w_{1} X_{1}^{3} \right) + 2\varepsilon^{3} X_{1}^{3} Y_{1} + \varepsilon X_{1}^{3} Y_{1} + X_{1}^{3}$$

Result normal form DEs

$$\begin{split} \dot{Y}_1 &= \sigma^2 \varepsilon^4 \left(8 \mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_1 \ w_1 X_1^4 Y_1 - 4 \mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_1 \ w_1 X_1^2 Y_1 \epsilon + \right. \\ & \left. 6 \mathrm{e}^t \star w_1 \ w_1 X_1^4 Y_1 + 22 \mathrm{e}^{-1t} \star w_1 \ w_1 X_1^4 Y_1 - 4 \mathrm{e}^{-1t} \star w_1 \ w_1 X_1^2 Y_1 \epsilon \right) + \\ & \left. 2 \sigma^2 \varepsilon^2 \mathrm{e}^{-1t} \star w_1 \ w_1 X_1^2 Y_1 + \sigma \varepsilon^4 \left(22 w_1 X_1^4 Y_1 - 4 w_1 X_1^2 Y_1 \epsilon \right) + \right. \\ & \left. 2 \sigma \varepsilon^2 w_1 X_1^2 Y_1 + \sigma w_1 Y_1 + \varepsilon^4 \left(6 X_1^4 Y_1 - 4 X_1^2 Y_1 \epsilon \right) + 4 \varepsilon^2 X_1^2 Y_1 - Y_1 \right. \\ & \dot{X}_1 &= -3 \sigma^2 \varepsilon^4 \mathrm{e}^{-1t} \star w_1 \ w_1 X_1^5 - 2 \sigma \varepsilon^4 w_1 X_1^5 + \varepsilon^4 \left(-X_1^5 + 2X_1^3 \epsilon \right) + \varepsilon^2 X_1 \epsilon \right. \end{split}$$

- As expected, $Y_1 = 0$ is the stochastic slow manifold, and is exponentially attractive (almost always) in some domain about the origin.
- The slow X_1 evolution is independent of Y_1 . Deterministically $(\sigma = 0)$, we predict a bifurcation to $X_1 \approx \pm \epsilon^{1/4}$. The noise appears to modify this slightly.
- The time-dependent coordinate transform maps these predictions back into the xy-plane.

4 offdiagonal: Levy area contraction: off-diagonal example

Pavliotis & Stuart (2008) assert the following system of five coupled SDEs are interesting for various parameters a and for small ϵ .

$$dx_1 = \epsilon y_1 \, dt \,,$$

$$dx_2 = \epsilon y_2 dt,$$

$$dx_3 = \epsilon (x_1 y_2 - x_2 y_1) dt,$$

$$dy_1 = (-y_1 - ay_2) dt + dW_1,$$

$$dy_2 = (+ay_1 - y_2) dt + dW_2.$$

This stochastic system has two noise sources. We treat $W_i(t,\omega)$ as Stratonovich Wiener processes. Use $\mathbf{x}(\mathbf{i})$ to denote variable x_i , $\mathbf{y}(\mathbf{i})$ to denote variable y_i , and $\mathbf{w}(\mathbf{i})$ to denote noise dW_i/dt .

Start by loading the procedure.

```
23 in_tex "../stoNormForm.tex"$
```

It is convenient to factor written results on the two given parameters ϵ, a :

```
24 factor epsilon,a;
```

Execute the construction of a normal form for this system. A coding is to specify the system as given: specify the slow SDEs via a three component list; and the fast stable SDEs via a two component list.

```
25 stonormalform(
26
       {epsilon*y(1),
27
        epsilon*y(2),
        epsilon*(x(1)*y(2)-x(2)*y(1))},
28
       {-y(1)-a*y(2)+w(1),}
29
30
        -y(2)+a*y(1)+w(2),
       {},
31
32
       4)$
33 end;
```

Now the approach can only analyse systems which are linearly diagonalised, but this system has two off-diagonal terms in the \vec{y} -SDEs (terms that cause oscillations in \vec{y} with frequency a as \vec{y} decays in magnitude like e^{-t}). In order to make some sort of progress, the procedure is brutal with such off-diagonal terms. Anything linear and off-diagonal is multiplied by the parameter small and so is treated as asymptotically small. When it does so, it gives the warning message

```
34 ***** Warning ****
35 Off diagonal linear terms in y- or z- equations
36 assumed small. Answers are rubbish if not
37 asymptotically appropriate.
```

As the message says, the results may consequently be rubbish.

Here then, the procedure analyses the following system which reduce to the given one for $\varepsilon = 1$:

$$\dot{x}_1 = \epsilon \varepsilon y_1$$
$$\dot{x}_2 = \epsilon \varepsilon y_2$$

$$\dot{x}_3 = \epsilon \varepsilon \left(-x_2 y_1 + x_1 y_2 \right)$$
$$\dot{y}_1 = -a \varepsilon y_2 + \sigma w_1 - y_1$$
$$\dot{y}_2 = a \varepsilon y_1 + \sigma w_2 - y_2$$

That is, the code treats the frequency parameter a as small, and so the results are appropriate only for small a, as well as only for small ϵ .

If one really needs to analyse non-small a, then more sophisticated code has to be developed.

Time dependent coordinate transform This transform is quite complicated, due to the noise, and involve fast-time future and history integrals.

$$\begin{split} y_1 &= -a^2 \sigma \varepsilon^2 \mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_1 - a \sigma \varepsilon \mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_2 + \\ & \sigma \mathrm{e}^{-1t} \star w_1 + Y_1 \\ y_2 &= -a^2 \sigma \varepsilon^2 \mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_2 + a \sigma \varepsilon \mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_1 + \\ & \sigma \mathrm{e}^{-1t} \star w_2 + Y_2 \\ x_1 &= a \epsilon \sigma \varepsilon^2 \left(\mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_2 + \mathrm{e}^{-1t} \star w_2 \right) + a \epsilon \varepsilon^2 Y_2 - \epsilon \sigma \varepsilon \mathrm{e}^{-1t} \star w_1 - \\ & \epsilon \varepsilon Y_1 + X_1 \\ x_2 &= a \epsilon \sigma \varepsilon^2 \left(-\mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_1 - \mathrm{e}^{-1t} \star w_1 \right) - a \epsilon \varepsilon^2 Y_1 - \\ & \epsilon \sigma \varepsilon \mathrm{e}^{-1t} \star w_2 - \epsilon \varepsilon Y_2 + X_2 \\ x_3 &= a \epsilon \sigma \varepsilon^2 \left(-\mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_2 X_2 - \mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_1 X_1 - \\ & \mathrm{e}^{-1t} \star w_2 X_2 - \mathrm{e}^{-1t} \star w_1 X_1 \right) + a \epsilon \varepsilon^2 \left(-X_2 Y_2 - X_1 Y_1 \right) + \\ & \epsilon^2 \sigma \varepsilon^2 \left(\mathrm{e}^t \star w_2 Y_1 - \mathrm{e}^{1t} \star w_1 Y_2 \right) + \epsilon \sigma \varepsilon \left(-\mathrm{e}^{-1t} \star w_2 X_1 + \\ & \mathrm{e}^{-1t} \star w_1 X_2 \right) + \epsilon \varepsilon \left(X_2 Y_1 - X_1 Y_2 \right) + X_3 \end{split}$$

Result normal form DEs

$$\begin{split} \dot{Y}_1 &= -a\varepsilon Y_2 - Y_1 \\ \dot{Y}_2 &= a\varepsilon Y_1 - Y_2 \\ \dot{X}_1 &= -a^2 \epsilon \sigma \varepsilon^3 w_1 - a\epsilon \sigma \varepsilon^2 w_2 + \epsilon \sigma \varepsilon w_1 \\ \dot{X}_2 &= -a^2 \epsilon \sigma \varepsilon^3 w_2 + a\epsilon \sigma \varepsilon^2 w_1 + \epsilon \sigma \varepsilon w_2 \\ \dot{X}_3 &= a^2 \epsilon \sigma \varepsilon^3 \left(-w_2 X_1 + w_1 X_2 \right) + a\epsilon^2 \sigma^2 \varepsilon^3 \left(\mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_2 \ w_2 + \mathrm{e}^{-1t} \star w_1 \ w_1 + \mathrm{e}^{-1t} \star w_2 \ w_2 + \mathrm{e}^{-1t} \star w_1 \ w_1 \right) + \\ &= a\epsilon \sigma \varepsilon^2 \left(w_2 X_2 + w_1 X_1 \right) + \epsilon^2 \sigma^2 \varepsilon^2 \left(\mathrm{e}^{-1t} \star w_2 \ w_1 - \mathrm{e}^{-1t} \star w_1 \ w_2 \right) + \\ &= \epsilon \sigma \varepsilon \left(w_2 X_1 - w_1 X_2 \right) \end{split}$$

- As expected, $\vec{Y} = 0$ is the stochastic slow manifold, and is exponentially attractive (almost always) in some domain about the origin.
- The slow \vec{X} evolution is independent of \vec{Y} : X_1, X_2 undergo a correlated 'slow' random walk; whereas X_3 is dominantly some multiplicative random walk.
- The time-dependent coordinate transform maps these predictions back into the \vec{x}, \vec{y} -space.

5 jordanForm: the Jordan form of position-momentum variables

Suppose x(t) is the spatial position of some particle, and you want to analyse the 'mechanical' system of SDEs

$$\ddot{x} = -xy$$
 and $\dot{y} = -2y + x^2 + \dot{x}^2 + \sigma w(t)$,

where w(t) denotes the formal derivative dW/dt of a Stratonovich Wiener process $W(t,\omega)$, or some other time-dependent forcing, called noise. Introduce position and velocity variables $x_1=x$ and $x_2=\dot{x}$, and also $y_1=y$ to convert to the system of three coupled first-order SDEs

$$\begin{split} \dot{x}_1 &= x_2 \,, \\ \dot{x}_2 &= -x_1 y_1 \,, \\ \dot{y}_1 &= -2 y_1 + x_1^2 + x_2^2 + \sigma w(t) \,. \end{split}$$

Start by loading the procedure.

```
38 in_tex "../stoNormForm.tex"$
```

Execute the construction of a normal form for this system.

Why divide x(2) by small? A possible coding is to specify the system as given, but recall that the slow SDEs are always multiplied by small thus changing the first SDE to $\dot{x}_1 = \varepsilon x_2$ and hence changing the relation between position and velocity—this would be OK if x_2 was viewed as momentum and the particle had large mass. But what if really do we want x_2 to be velocity. Fortunately, the coded iteration scheme works for systems with linear part in Jordan form, but one has to code the system as follows. Namely, divide the off-diagonal term of the Jordan form by small to cancel out the procedure's brutal multiplication by small.

Then the coded procedure reports that it analyses the following system which not only reduces to the given one for $\varepsilon = 1$, but also preserves the physical relation between position x_1 and velocity x_2 :

$$\dot{x}_1 = x_2$$
 $\dot{x}_2 = -\varepsilon x_1 y_1$
 $\dot{y}_1 = \sigma w_y + \varepsilon (x_2^2 + x_1^2) - 2y_1$

Further, here ε counts the order of nonlinearity so truncating to errors $\mathcal{O}(\varepsilon^3)$ is the same as truncating to errors $\mathcal{O}(|(\vec{x},y)|^4)$.

The cost of preserving the physical relation between position x_1 and velocity x_2 is that more iterations are needed in the construction.

Time dependent coordinate transform This transform is quite complicated, due to the noise, and involve fast-time future and history integrals.

$$y_{1} = \sigma e^{-2t} \star w_{y} + \varepsilon \left(3/4X_{2}^{2} - 1/2X_{2}X_{1} + 1/2X_{1}^{2} \right) + Y_{1}$$

$$x_{1} = \sigma \varepsilon \left(-1/4e^{-2t} \star w_{y} X_{2} - 1/4e^{-2t} \star w_{y} X_{1} \right) + \varepsilon \left(-1/4X_{2}Y_{1} - 1/4X_{1}Y_{1} \right) + X_{1}$$

$$x_{2} = \sigma \varepsilon \left(1/4e^{-2t} \star w_{y} X_{2} + 1/2e^{-2t} \star w_{y} X_{1} \right) + \varepsilon \left(1/4X_{2}Y_{1} + 1/2X_{1}Y_{1} \right) + X_{2}$$

Result normal form DEs

$$\begin{split} \dot{Y}_1 &= \varepsilon^2 \left(1/2 X_2^2 Y_1 + 1/2 X_2 X_1 Y_1 - 1/2 X_1^2 Y_1 \right) - 2 Y_1 \\ \dot{X}_1 &= \sigma^2 \varepsilon^2 \left(-3/64 \mathrm{e}^{-2t} \star w_y \, w_y X_2 - 3/32 \mathrm{e}^{-2t} \star w_y \, w_y X_1 \right) + \\ &\quad \sigma \varepsilon \left(1/4 w_y X_2 + 1/4 w_y X_1 \right) + X_2 \\ \dot{X}_2 &= \sigma^2 \varepsilon^2 \left(3/32 \mathrm{e}^{-2t} \star w_y \, w_y X_2 + 1/8 \mathrm{e}^{-2t} \star w_y \, w_y X_1 \right) + \sigma \varepsilon \left(- \\ &\quad 1/4 w_y X_2 - 1/2 w_y X_1 \right) + \varepsilon^2 \left(-3/4 X_2^2 X_1 + 1/2 X_2 X_1^2 - 1/2 X_1^3 \right) \end{split}$$

- As expected, $Y_1 = 0$ is the stochastic slow manifold, and is exponentially attractive (almost always) in some domain about the origin.
- As expected, the slow \vec{X} evolution is independent of Y_1 : X_2 is approximately a 'velocity' variable for 'position' X_1 , and shows some nonlinear noise affected dynamics.
- The time-dependent coordinate transform maps these predictions back into the \vec{x}, y -space. Observe that \vec{X} are not precisely the physical position-veloxity \vec{x} , but instead are affected by nonlinearity, and the noise, and their interaction.

6 slow0sc: Radek's slow oscillation with fast noise

Consider Radek's system

$$\dot{x} = -\epsilon xz$$
, $\dot{y} = +\epsilon yz$ and $\dot{z} = -(z-1) + \sigma w(t)$.

In this linear system, x,y oscillate with 'frequency' ϵz . But z(t) is an Ornstein–Uhlenbeck process with mean one. What are the dynamics?

Transform to our standard form via

$$x = x_1$$
, $y = x_2$ and $z = 1 + y_1$.

Then start by loading the procedure.

46 in_tex "../stoNormForm.tex"\$

Execute the construction of a normal form for this system.

```
47 factor x;

48 stonormalform(

49 { -x(2)*(1+y(1)),

50 x(1)*(1+y(1)) },

51 { -y(1)+w(1) },

52 {},

53 4)$

54 end;
```

With the above input the procedure analyses the following system:

$$\dot{x}_1 = x_2 \varepsilon (-y_1 - 1)$$
$$\dot{x}_2 = x_1 \varepsilon (y_1 + 1)$$
$$\dot{y}_1 = \sigma w_1 - y_1$$

This is precisely the original system, but with variables changed as above, and with parameter $\varepsilon = \epsilon$ (here we use the procedure's multiplication by ε to incorporate Radek's ϵ).

Time dependent coordinate transform

$$y_{1} = \sigma e^{-1t} \star w_{1} + Y_{1}$$

$$x_{1} = -\sigma \varepsilon^{2} e^{-1t} \star w_{1} X_{1} Y_{1} + \sigma \varepsilon e^{-1t} \star w_{1} X_{2} - 1/2 \varepsilon^{2} X_{1} Y_{1}^{2} + \varepsilon X_{2} Y_{1} + X_{1}$$

$$X_{1}$$

$$x_{2} = -\sigma \varepsilon^{2} e^{-1t} \star w_{1} X_{2} Y_{1} - \sigma \varepsilon e^{-1t} \star w_{1} X_{1} - 1/2 \varepsilon^{2} X_{2} Y_{1}^{2} - \varepsilon X_{1} Y_{1} + X_{2}$$

Result normal form DEs In such linear systems, the following normal form is straightforward.

$$\begin{aligned} \dot{Y}_1 &= -Y_1 \\ \dot{X}_1 &= -\sigma \varepsilon w_1 X_2 - \varepsilon X_2 \\ \dot{X}_2 &= \sigma \varepsilon w_1 X_1 + \varepsilon X_1 \end{aligned}$$

- As expected, $Y_1 = 0$ is the emergent stochastic slow manifold.
- The slow \vec{X} evolution clearly oscillates in (X_1, X_2) , $X_j \propto e^{i\theta}$, with phase angle $\theta = \varepsilon(t + \sigma W(t, \omega))$, recalling $W = \int w \, dt$. This phase grows linearly with a superposed random walk.
- The time-dependent coordinate transform maps these predictions back into the \vec{x}, y_1 -plane, and thence to the original xyz-space.
- In this system, higher-order terms in ε only affect the coordinate transform, they do not change the evolution of \vec{X} .

7 linearHyper: simple linear hyperbolic noisy system

The procedure also analyses hyperbolic systems, and recovers the classic stochastic/non-autonomous results guaranteed by the Hartman–Grobman Theorem. Consider the following linear SDEs with one stable variable, and one unstable variable:

$$\dot{y}_1 = -y_1 + \sigma w_1 z_1$$
$$\dot{z}_1 = z_1 + \sigma w_1 y_1$$

Start by loading the procedure.

Execute the construction of a normal form for this system: the parameter σ is automatically inserted by the procedure.

```
56 stonormalform(
57 {},
58 { -y(1)+z(1)*w(1) },
59 { +z(1)+y(1)*w(1) },
60 3 )$
61 end;
```

Time dependent coordinate transform This simply mixes Y, Z a little depending upon the noise.

$$z_1 = -\sigma e^{2t} \star w_1 Y_1 + Z_1$$

 $y_1 = \sigma e^{-2t} \star w_1 Z_1 + Y_1$

Result normal form DEs In such linear systems the normal form is straightforward, as follows.

$$\dot{Z}_1 = \sigma^2 e^{-2t} \star w_1 \, w_1 Z_1 + Z_1$$
$$\dot{Y}_1 = -\sigma^2 e^{2t} \star w_1 \, w_1 Y_1 - Y_1$$

The Y, Z variables are decoupled. Their evolution retains effects from noise-noise interactions: Z from the past history; and Y from future anticipation.

8 foliateHyper: Duan's hyperbolic system for foliation

To illustrate a stochastic/non-autonomous Hartman–Grobman Theorem, Sun et al. (2011) used the following simple hyperbolic system with one stable variable, and one unstable variable:

$$\dot{y}_1 = -y_1 + \sigma w_1 y_1$$
$$\dot{z}_1 = z_1 + y_1^2 + \sigma w_1 z_1$$

The stable y-dynamics is simply an Ornstein-Uhlenbeck process, independent of z(t). The unstable z-dynamics is similar, but with a quadratic forcing by the stable variable y. Let's unfold this effect.

Start by loading the procedure.

```
62 in_tex "../stoNormForm.tex"$
```

Execute the construction of a normal form for this system: the parameter σ is automatically inserted by the procedure.

```
63 stonormalform(
64 {},
65 { -y(1)+y(1)*w(1) },
66 { +z(1)+y(1)^2+z(1)*w(1) },
67 9 )$
68 end;
```

In the procedure, the y_1^2 term is automatically multiplied by ε , and so, in the results, ε counts the order of nonlinearity of each term. We analyse to high-order, errors $\mathcal{O}(\varepsilon^9, \sigma^3)$, because the results are simple.

Time dependent coordinate transform To decouple the stochastic dynamics, we just need to stochastically 'bend' the z-variable. This bending forms a stochastic foliation of the system.

$$z_1 = -1/3\sigma\varepsilon e^{3t} \star w_1 Y_1^2 - 1/3\varepsilon Y_1^2 + Z_1$$

 $u_1 = Y_1$

Result normal form DEs The normal form dynamics is linear and decoupled, as per Hartman–Grobman, namely

$$\dot{Z}_1 = \sigma w_1 Z_1 + Z_1$$
 $\dot{Y}_1 = \sigma w_1 Y_1 - Y_1$

9 monahanFive: Monahan's five examples

Monahan & Culina (2011) discuss stochastic averaging and give several examples in the body and an appendix, of which we analyse five. They really need this approach as "a large separation often does not exist in atmosphere or ocean dynamics" between the fast and slow time scales.

9.1 Example four: 'three' time scales

Monahan & Culina (2011) comment that this, their fourth example, a linear system, has three time scales. But I do not see these time scales, I only see varying strength interactions. They consider

$$\frac{dx}{dt} = -x + \frac{a}{\sqrt{\tau}}y$$
 and $\frac{dy}{dt} = \frac{1}{\sqrt{\tau}}x - \frac{1}{\tau}y + \frac{b}{\sqrt{\tau}}\dot{W}$.

Let's rescale time, $t = \tau t'$ so that $d/dt = \frac{1}{\tau}d/dt'$ and $\dot{W} = \frac{1}{\sqrt{\tau}}dW/dt'$. Then, dropping dashes, the SDE system is

$$\frac{dx}{dt} = -\tau x + a\sqrt{\tau}y$$
 and $\frac{dy}{dt} = \sqrt{\tau}x - y + b\dot{W}$.

Start by loading the procedure.

69 in_tex "../stoNormForm.tex"\$

Execute the construction of a normal form for this system. Using the default inbuilt parametrisation of noise by sigma to represent parameter b, and using small in the x-SDE so that it counts the number of small $\sqrt{\tau}$, code these as the following.

The procedure reports that it analyses the following family

$$\dot{x}_1 = \sqrt{\tau} y_1 \varepsilon a - \varepsilon^2 \tau x_1$$
 $\dot{y}_1 = w_1 \sigma - y_1 + \sqrt{\tau} \varepsilon x_1$

in which we indeed see ε only in the grouping $\varepsilon\sqrt{\tau}$.

Time dependent coordinate transform This is linear as the system is linear.

$$y_1 = -e^{-1t} \star e^{-1t} \star w_1 \ \sigma \varepsilon^2 \tau a - e^{-1t} \star w_1 \ \sigma \varepsilon^2 \tau a + e^{-1t} \star w_1 \ \sigma + Y_1 + \sqrt{\tau} \varepsilon X_1$$
$$x_1 = -\sqrt{\tau} e^{-1t} \star w_1 \ \sigma \varepsilon a - \sqrt{\tau} Y_1 \varepsilon a + X_1$$

Result normal form DEs The normal form dynamics is linear and decoupled, as per Hartman–Grobman, namely

$$\dot{Y}_1 = -Y_1 \varepsilon^2 \tau a - Y_1$$

$$\dot{X}_1 = w_1 \sigma \varepsilon^3 \tau \sqrt{\tau} \left(-2a^2 + a \right) + \sqrt{\tau} w_1 \sigma \varepsilon a + \varepsilon^2 \tau \left(X_1 a - X_1 \right)$$

Monahan & Culina (2011) derive the last two terms in the X-equation, but not the first as it is too small for their averaging analysis. They comment that a>1 is some sort of difficulty, presumably because X grows when a>1: but here we have no problem with a>1, especially as the decay rate to the stochastic slow manifold, the Y-SDE, is $(1+\tau a)$ which gets stronger with increasing parameter a.

9.2 Example one: simple rational nonlinear

With 'small' scale-separation parameter τ , Monahan & Culina (2011) first consider the example

$$\frac{dx}{dt} = -x + \Sigma(x)y$$
 and $\frac{dy}{dt} = -\frac{1}{\tau}y + \frac{1}{\sqrt{\tau}}\dot{W}$,

for general smooth functions $\Sigma(x)$. Rescale time, $t=\tau t'$ so that $d/dt=\frac{1}{\tau}d/dt'$ and $W=\frac{1}{\sqrt{\tau}}dW/dt'$. Then, dropping dashes, the SDE is

$$\frac{dx}{dt} = -\tau x + \tau \Sigma(x)y$$
 and $\frac{dy}{dt} = -y + \dot{W}$.

The stonormalform procedure is already loaded. Write a message saying we are now analysing the next system.

```
77 write "**** Example One of Monahan (2011) ****";
```

Execute the construction of a normal form for this system. But let's restrict the general function $\Sigma(x)$ to the rational form $\Sigma(x) := (a_0 + a_1x + a_2x^2)/(1 + b_1x + b_2x^2)$. Code this form as the following (after multiplying through by the denominator).

```
78 factor df;
79 operator a; defindex a(down);
80 operator b; defindex b(down);
  stonormalform(
       {-tau*x(1)*(1+b(1)*x(1)+b(2)*x(1)^2)}
82
         -df(x(1),t)*(b(1)*x(1)+b(2)*x(1)^2)
83
         +tau*y(1)*(a(0)+a(1)*x(1)+a(2)*x(1)^2)},
84
       \{ -y(1)+w(1) \},
85
       {},
       3)$
87
88 %end; % optionally end examples here
```

The procedure reports that it analyses the following family

$$\dot{x}_1 = \frac{\frac{\mathrm{d}\,x_1}{\mathrm{d}\,t}\,\varepsilon\left(-b_2x_1^2 - b_1x_1\right) + y_1\varepsilon\tau\left(a_2x_1^2 + a_1x_1 + a_0\right) + \varepsilon\tau\left(-b_2x_1^3 - b_1x_1^2 - x_1\right)$$

$$\dot{y}_1 = w_1\sigma - y_1$$

so evaluate the results at $\varepsilon = 1$ to compare with the modelling of Monahan & Culina (2011).

Time dependent coordinate transform

$$y_1 = e^{-1t} \star w_1 \sigma + Y_1$$

$$x_1 = e^{-1t} \star w_1 \sigma \varepsilon \tau \left(-a_2 X_1^2 - a_1 X_1 - a_0 \right) + Y_1 \varepsilon \tau \left(-a_2 X_1^2 - a_1 X_1 - a_0 \right) + X_1$$

Result normal form DEs

$$\begin{split} \dot{Y}_1 &= -Y_1 \\ \dot{X}_1 &= w_1 \sigma \varepsilon^2 \tau^2 \big(a_2 b_2 X_1^4 - a_2 X_1^2 + 2 a_1 b_2 X_1^3 + a_1 b_1 X_1^2 + 3 a_0 b_2 X_1^2 + \\ & 2 a_0 b_1 X_1 + a_0 \big) + w_1 \sigma \varepsilon^2 \tau \big(-a_2 b_2 X_1^4 - a_2 b_1 X_1^3 - a_1 b_2 X_1^3 - \\ & a_1 b_1 X_1^2 - a_0 b_2 X_1^2 - a_0 b_1 X_1 \big) + w_1 \sigma \varepsilon \tau \big(a_2 X_1^2 + a_1 X_1 + a_0 \big) + \\ & \varepsilon^2 \tau \big(b_2^2 X_1^5 + 2 b_2 b_1 X_1^4 + b_2 X_1^3 + b_1^2 X_1^3 + b_1 X_1^2 \big) + \varepsilon \tau \big(- \\ & b_2 X_1^3 - b_1 X_1^2 - X_1 \big) \end{split}$$

Monahan & Culina (2011) derive some of this X equation. The other terms here are higher order terms that become significant at finite parameter values. For example, the next correction to their analysis, $w_1\tau^2(-3a_4X_1^4-2a_3X_1^3-a_2X_1^2+a_0)$, is probably derivable as $\tau^2(\Sigma-x\Sigma')\dot{W}$ (when rescaled).

9.3 Example three: several fast stable modes

Monahan & Culina (2011) third considered the example

$$\frac{dx}{dt} = -x + \Sigma(x) \|\vec{y}\|$$
 and $\frac{d\vec{y}}{dt} = -\frac{1}{\tau}\vec{y} + \sqrt{\frac{2}{\tau}}\dot{\vec{W}}$,

for general smooth functions $\Sigma(x)$, and 'small' scale separation parameter τ . As before, rescale time, $t=\tau t'$ so that $d/dt=\frac{1}{\tau}d/dt'$ and $\dot{W}=\frac{1}{\sqrt{\tau}}dW/dt'$. Here I also cheat: they have $\|\vec{y}\|$ in the slow equation; but $\|\vec{y}\|$ is not a smooth multinomial and so my generic procedure cannot apply; instead I replace $\|\vec{y}\|$ with $\|\vec{y}\|^2$ which has the same symmetry but is multinomial. Then, upon the rescaling of time, and dropping dashes, the SDE is

$$\frac{dx}{dt} = -\tau x + \tau \Sigma(x) \|\vec{y}\|^2$$
 and $\frac{d\vec{y}}{dt} = -\vec{y} + \sigma \dot{\vec{W}}$.

The stonormalform procedure is already loaded. Write a message saying we are now analysing the next system.

```
89 write "**** Example Three of Monahan (2011) ****";
```

Restrict analysis to the general quartic $\Sigma(x) := a_0 + a_1 x + \dots + a_4 x^4$, and so code the system as the following (the generic program automatically inserts the σ in the noise). Currently restrict to just a two component \vec{y} as I do not see any reason for any more, and Monahan & Culina (2011) do not appear to specify.

```
90
  stonormalform(
       \{-tau*x(1)+tau*(y(1)^2+y(2)^2)
91
         *(a(0)+a(1)*x(1)+a(2)*x(1)^2
92
         +a(3)*x(1)^3+a(4)*x(1)^4)},
93
       \{ -y(1)+w(1),
94
         -y(2)+w(2) },
95
       {},
96
       3)$
97
98 %end; % optionally end examples here
```

The procedure reports that it analyses the following family

$$\dot{x}_1 = y_2^2 \varepsilon \tau \left(a_4 x_1^4 + a_3 x_1^3 + a_2 x_1^2 + a_1 x_1 + a_0 \right) + y_1^2 \varepsilon \tau \left(a_4 x_1^4 + a_3 x_1^3 + a_2 x_1^2 + a_1 x_1 + a_0 \right) - \varepsilon \tau x_1$$

$$\dot{y}_1 = w_1 \sigma - y_1$$

$$\dot{y}_2 = w_2 \sigma - y_2$$

in which we see ε only in the grouping $\varepsilon\tau$, so truncating to errors $\mathcal{O}(\varepsilon^3)$ is the same as to errors $\mathcal{O}(\tau^3)$.

Time dependent coordinate transform

$$\begin{split} y_1 &= \mathrm{e}^{-1t} \star w_1 \, \sigma + Y_1 \\ y_2 &= \mathrm{e}^{-1t} \star w_2 \, \sigma + Y_2 \\ x_1 &= \mathrm{e}^t \star w_2 \, Y_2 \sigma \varepsilon \tau \left(-a_4 X_1^4 - a_3 X_1^3 - a_2 X_1^2 - a_1 X_1 - a_0 \right) + \\ &= \mathrm{e}^{-1t} \star w_2 \, Y_2 \sigma \varepsilon \tau \left(-a_4 X_1^4 - a_3 X_1^3 - a_2 X_1^2 - a_1 X_1 - a_0 \right) + \\ &= \mathrm{e}^t \star w_1 \, Y_1 \sigma \varepsilon \tau \left(-a_4 X_1^4 - a_3 X_1^3 - a_2 X_1^2 - a_1 X_1 - a_0 \right) + \\ &= \mathrm{e}^{-1t} \star w_1 \, Y_1 \sigma \varepsilon \tau \left(-a_4 X_1^4 - a_3 X_1^3 - a_2 X_1^2 - a_1 X_1 - a_0 \right) + Y_2^2 \varepsilon \tau \left(-a_4 X_1^4 - 1/2 a_3 X_1^3 - 1/2 a_2 X_1^2 - 1/2 a_1 X_1 - 1/2 a_0 \right) + Y_1^2 \varepsilon \tau \left(-a_4 X_1^4 - 1/2 a_3 X_1^3 - 1/2 a_2 X_1^2 - 1/2 a_1 X_1 - 1/2 a_0 \right) + X_1 \end{split}$$

The complicated form of x_1 only reflects the transient effects of the decaying \vec{Y} : once $\vec{Y} \to 0$, then $x_1 = X_1$.

Result normal form DEs

$$\begin{split} \dot{Y}_1 &= -Y_1 \\ \dot{Y}_2 &= -Y_2 \\ \dot{X}_1 &= \mathrm{e}^{-1t} \star w_2 \, w_2 \sigma^2 \varepsilon^2 \tau^2 \big(-3/2 a_4 X_1^4 - a_3 X_1^3 - 1/2 a_2 X_1^2 + 1/2 a_0 \big) + \\ &\mathrm{e}^{-1t} \star w_2 \, w_2 \sigma^2 \varepsilon \tau \big(a_4 X_1^4 + a_3 X_1^3 + a_2 X_1^2 + a_1 X_1 + a_0 \big) + \\ &\mathrm{e}^{-1t} \star w_1 \, w_1 \sigma^2 \varepsilon^2 \tau^2 \big(-3/2 a_4 X_1^4 - a_3 X_1^3 - 1/2 a_2 X_1^2 + 1/2 a_0 \big) + \\ &\mathrm{e}^{-1t} \star w_1 \, w_1 \sigma^2 \varepsilon \tau \big(a_4 X_1^4 + a_3 X_1^3 + a_2 X_1^2 + a_1 X_1 + a_0 \big) - \varepsilon \tau X_1 \end{split}$$

These show the decay of \vec{Y} , and that the irreducible noise-noise interactions are the only modifications to the slow decay of X_1 , and hence of x_1 .

9.4 Example two: irregular slow manifold

Monahan & Culina (2011) second consider the example SDEs

$$\frac{dx}{dt} = x - x^3 + \Sigma(x)y$$
 and $\frac{dy}{dt} = -\frac{1}{x\tau}y + \frac{1}{\sqrt{\tau}}\dot{W}$,

for general smooth functions $\Sigma(x)$, and 'small' scale separation parameter τ . Since the y-dynamics are exponentially unstable for negative x, we restrict attention to x>0. Even for positive x the system is singular as $x\to 0$ so the slow manifold is irregular in some sense (although 'singular' in a good way in that the scale

separation between fast and slow becomes infinite). Let's be more sophisticated in rescaling time: let's choose the new fast time t' so that $dt = x\tau dt'$; that is, $t' = \int (x\tau)^{-1} dt$ which would not be explicitly known until after a solution x(t') is found. I presume that the noise then transforms as $\dot{W} = \frac{1}{\sqrt{x\tau}} dW/dt'$ (needs checking). Then, dropping dashes, the SDEs are

$$\frac{dx}{dt} = \tau \left[x^2 - x^4 + x\Sigma(x)y \right]$$
 and $\frac{dy}{dt} = -y + \sqrt{x}\dot{W}$.

The \sqrt{x} is a problem in my generic procedure as it requires multinomial systems, so transform to $x = x_1^2$ (not the usual $x = x_1$) so that $2x_1dx_1 = dx$. Then the SDE system becomes

$$\frac{dx_1}{dt} = \frac{1}{2}\tau \left[x_1^3 - x_1^7 + x_1 \Sigma(x_1^2) y \right] \quad \text{and} \quad \frac{dy}{dt} = -y + x_1 \dot{W}.$$

The stonormalform procedure is already loaded. Write a message saying we are now analysing the next system.

```
99 write "**** Example Two of Monahan (2011) ****";
```

Restricting to the general linear $\Sigma(x) := a_0 + a_1 x$, code the SDE system as the following (remember $\mathbf{x}(1) = x_1 = \sqrt{x}$).

```
100 stonormalform(
101 { 1/2*tau*( x(1)^3 -x(1)^7
102 +x(1)*(a(0)+a(1)*x(1)^2)*y(1) ) },
103 { -y(1) +x(1)*w(1) },
104 {},
105 3 )$
106 %end; % optionally end examples here
```

The procedure reports that it analyses the following family

$$\dot{x}_1 = y_1 \varepsilon \tau \left(1/2a_1 x_1^3 + 1/2a_0 x_1 \right) + \varepsilon \tau \left(-1/2x_1^7 + 1/2x_1^3 \right)$$

$$\dot{y}_1 = w_1 \sigma x_1 - y_1$$

Again, usefully, the artificial ε only occurs in the combination $\varepsilon\tau$ and so just counts the number of factors of τ in each term. That is, errors $\mathcal{O}(\varepsilon^3)$ is the same as errors $\mathcal{O}(\tau^3)$.

Time dependent coordinate transform Straightforwardly

$$y_1 = e^{-1t} \star e^{-1t} \star w_1 \ \sigma \varepsilon \tau \left(\frac{1}{2} X_1^7 - \frac{1}{2} X_1^3 \right) + e^{-1t} \star w_1 \ \sigma X_1 + Y_1$$

$$x_1 = e^{-1t} \star w_1 \ \sigma \varepsilon \tau \left(-\frac{1}{2} a_1 X_1^4 - \frac{1}{2} a_0 X_1^2 \right) + Y_1 \varepsilon \tau \left(-\frac{1}{2} a_1 X_1^3 - \frac{1}{2} a_0 X_1 \right) + X_1$$

Result normal form DEs As expected, $Y_1 = 0$ is a stochastic slow manifold, that is almost surely emergent in some domain.

$$\begin{split} \dot{Y}_1 &= \mathrm{e}^t \star w_1 \, w_1 Y_1 \sigma^2 \varepsilon^2 \tau^2 \left(1/4 a_1^2 X_1^6 + 1/2 a_1 a_0 X_1^4 + 1/4 a_0^2 X_1^2 \right) + \\ & \mathrm{e}^{-1t} \star w_1 \, w_1 Y_1 \sigma^2 \varepsilon^2 \tau^2 \left(3/4 a_1^2 X_1^6 + a_1 a_0 X_1^4 + 1/4 a_0^2 X_1^2 \right) + \\ & w_1 Y_1 \sigma \varepsilon^2 \tau^2 \left(-a_1 X_1^9 - 3/2 a_0 X_1^7 + 1/2 a_0 X_1^3 \right) + w_1 Y_1 \sigma \varepsilon \tau \left(-1/2 a_1 X_1^3 - 1/2 a_0 X_1 \right) - Y_1 \end{split}$$

$$\dot{X}_{1} = e^{-1t} \star w_{1} w_{1} \sigma^{2} \varepsilon^{2} \tau^{2} \left(-\frac{1}{4} a_{1}^{2} X_{1}^{7} - \frac{1}{2} a_{1} a_{0} X_{1}^{5} - \frac{1}{4} a_{0}^{2} X_{1}^{3} \right) + w_{1} \sigma \varepsilon^{2} \tau^{2} \left(a_{1} X_{1}^{10} + \frac{3}{2} a_{0} X_{1}^{8} - \frac{1}{2} a_{0} X_{1}^{4} \right) + w_{1} \sigma \varepsilon \tau \left(\frac{1}{2} a_{1} X_{1}^{4} + \frac{1}{2} a_{0} X_{1}^{2} \right) + \varepsilon \tau \left(-\frac{1}{2} X_{1}^{7} + \frac{1}{2} X_{1}^{3} \right)$$

Using just the leading order terms for \dot{X}_1 , the terms linear in τ , and recalling $X_1 \approx x_1 = \sqrt{x}$, the last SDE gives the model

$$\frac{dx}{dt'} \approx \tau \left[x^2 - x^4 + x^{3/2} \Sigma(x) \sigma \frac{dW}{dt'} \right].$$

But recall that $dt' = dt/(x\tau)$ (although one should be more careful as $X_1 \approx \sqrt{x}$, not exact equality) and similarly $dW/dt' = \sqrt{x\tau}\dot{W}$ so that this model becomes

$$\frac{dx}{dt} \approx x - x^3 + \sqrt{\tau} x \Sigma(x) \sigma \frac{dW}{dt}.$$

This agrees with the Stratonovich part of (A28) by Monahan & Culina (2011). But again, the above derivation has the systematic higher order corrections that are needed for finite scale separation τ .

9.5 Idealised Stommel-like model of meridional overturning circulation

Monahan & Culina (2011) also analyse the Idealised Stommel-like model, for small scale-separation parameter τ ,

$$\frac{dx}{dt} = \mu - |y - x|x + \sigma_A \dot{W}_1,$$

$$\frac{dy}{dt} = +\frac{1}{\tau} (1 - y) - |y - x|y + \sqrt{\frac{2}{\tau}} \sigma_M \dot{W}_2.$$

The mod-functions do not fit into my generic computer algebra so replace them with squares to preserve the symmetry. As before, rescale time, $t = \tau t'$ so that $d/dt = \frac{1}{\tau}d/dt'$ and $\dot{W}_j = \frac{1}{\sqrt{\tau}}dW_j/dt'$. Since for small τ , the fast variable y is strongly attracted to one, change the reference point for y by setting $y = 1 + y_1(t)$. Then the SDEs becomes akin to

$$\frac{dx}{dt'} = \epsilon^2 \left[\mu - (1 + y_1 - x)^2 x \right] + \epsilon \sigma_A \frac{dW_1}{dt'},$$

$$\frac{dy_1}{dt'} = -y_1 - \epsilon^2 (1 + y_1 - x)^2 (1 + y_1) + \sqrt{2} \sigma_M \frac{dW_2}{dt'}$$

The stonormalform procedure is already loaded. Write a message saying we are now analysing the next system.

107 write "**** Stommel-like model of Monahan (2011) ****"; 108 factor rho;

Let $\rho := \sigma_A/(\sqrt{2}\sigma_M)$, use the inbuilt $\sigma := \sqrt{2}\sigma_M$, and invoke small to correctly count the number of small $\sqrt{\tau}$ s in the analysis. Code the above dynamics as the following.

The procedure reports that it analyses the following family, an expended version of the prescribed system,

$$\dot{x}_1 = \sqrt{\tau} w_1 \rho \sigma \varepsilon - y_1^2 \varepsilon^2 \tau x_1 + y_1 \varepsilon^2 \tau \left(2x_1^2 - 2x_1 \right) + \varepsilon^2 \tau \left(-x_1^3 + 2x_1^2 - x_1 + \mu \right)$$

$$\dot{y}_1 = w_2 \sigma - y_1^3 \varepsilon^2 \tau + y_1^2 \varepsilon^2 \tau \left(2x_1 - 3 \right) + y_1 \varepsilon^2 \tau \left(-x_1^2 + 4x_1 - 3 \right) - y_1 + \varepsilon^2 \tau \left(-x_1^2 + 2x_1 - 1 \right)$$

Again the artificial ε only occurs in the combination $\varepsilon^2 \tau$ and so just counts the number of factors of τ in each term. That is, errors $\mathcal{O}(\varepsilon^4)$ is the same as errors $\mathcal{O}(\tau^2)$.

Time dependent coordinate transform Straightforward but complicated in detail:

$$y_{1} = e^{-1t} \star e^{-1t} \star w_{2} \ \sigma \varepsilon^{2} \tau \left(-X_{1}^{2} + 4X_{1} - 3\right) + 3/2 e^{t} \star w_{2} Y_{1}^{2} \sigma \varepsilon^{2} \tau + 3/2 e^{-1t} \star w_{2} Y_{1}^{2} \sigma \varepsilon^{2} \tau + e^{-1t} \star w_{2} Y_{1} \sigma \varepsilon^{2} \tau \left(-4X_{1} + 6\right) + e^{-1t} \star w_{2} \sigma + 1/2 Y_{1}^{3} \varepsilon^{2} \tau + Y_{1}^{2} \varepsilon^{2} \tau \left(-2X_{1} + 3\right) + Y_{1} + \varepsilon^{2} \tau \left(-X_{1}^{2} + 2X_{1} - 1\right)$$

$$x_{1} = e^{t} \star w_{2} Y_{1} \sigma \varepsilon^{2} \tau X_{1} + e^{-1t} \star w_{2} Y_{1} \sigma \varepsilon^{2} \tau X_{1} + e^{-1t} \star w_{2} \sigma \varepsilon^{2} \tau \left(-2X_{1}^{2} + 2X_{1}\right) + 1/2 Y_{1}^{2} \varepsilon^{2} \tau X_{1} + Y_{1} \varepsilon^{2} \tau \left(-2X_{1}^{2} + 2X_{1}\right) + X_{1}$$

Result normal form DEs As expected, $Y_1 = 0$ is a stochastic slow manifold which is almost surely emergent:

$$\begin{split} \dot{Y}_1 &= -3\mathrm{e}^{-1t} \star w_2 \, w_2 Y_1 \sigma^2 \varepsilon^2 \tau + 4 \sqrt{\tau} \mathrm{e}^{-1t} \star w_2 \, w_1 Y_1 \rho \sigma^2 \varepsilon^3 \tau \, + \\ & w_2 Y_1 \sigma \varepsilon^2 \tau \left(4 X_1 - 6 \right) + Y_1 \varepsilon^2 \tau \left(- X_1^2 + 4 X_1 - 3 \right) - Y_1 \\ \dot{X}_1 &= -\mathrm{e}^{-1t} \star w_2 \, w_2 \sigma^2 \varepsilon^2 \tau X_1 + \mathrm{e}^{-1t} \star w_2 \, w_1 \rho \sigma^2 \varepsilon^3 \tau^{3/2} \left(4 X_1 - 2 \right) + \\ & w_2 \sigma \varepsilon^2 \tau \left(2 X_1^2 - 2 X_1 \right) + \sqrt{\tau} w_1 \rho \sigma \varepsilon + \varepsilon^2 \tau \left(- X_1^3 + 2 X_1^2 - X_1 + \mu \right) \end{split}$$

Deterministically, this model has multiple equilibria for small μ , but only one equilibria for $\mu > 4/27$, at finite amplitude. The noise \dot{W}_1 causes transitions between such multiple equilibria, and the multiplicative noise \dot{W}_2 contributes as well. But the same order of smallness is the *first* term in the X_1 SDE above which is a quadratic noise-noise interaction that has a mean drift effect which should enhance the stability of the small x equilibrium.

10 majdaTriad: Majda's two triad models

Majda et al. (2002) investigated averaging in two 3D SDE systems. Let's compare with their stochastic normal form.

10.1 Multiplicative triad model

The multiplicative triad model of Majda et al. (2002) consists of three modes, v_1 , v_2 and v_3 . These evolve in time according to

$$\frac{dv_1}{dt} = b_1 v_2 v_3$$
, $\frac{dv_2}{dt} = b_2 v_1 v_3$, $\frac{dv_3}{dt} = -v_3 + b_3 v_1 v_2 + \sigma \dot{W}$,

where b_j and σ are some constants, and the noise forces the third mode. Here I have already scaled the equations so that the rate of decay of the third mode is one. Thus on long time scales we expect the third mode to be essentially negligible and the system to be modelled by the relatively slow evolution of the first two modes.

Start by loading the procedure.

```
116 in_tex "../stoNormForm.tex"$
```

The system uses parameters b_j so define

117 operator b; defindex b(down);

Execute the construction of a normal form for this system using $x_j = v_j$ and $y_1 = v_3$.

```
118 factor yy;

119 stonormalform(

120 { b(1)*x(2)*y(1),

121 b(2)*x(1)*y(1) },

122 { -y(1)+b(3)*x(1)*x(2)+w(3) },

123 {},

124 4 )$

125 %end; % optionally end examples here
```

The procedure reports that it analyses the following family

$$\dot{x}_1 = \varepsilon b_1 x_2 y_1$$
 $\dot{x}_2 = \varepsilon b_2 x_1 y_1$ $\dot{y}_1 = w_3 \sigma + \varepsilon b_3 x_2 x_1 - y_1$

Here, ε counts the order of nonlinearity so that errors $\mathcal{O}(\varepsilon^4)$ are errors $\mathcal{O}(|\vec{v}|^5 + \sigma^5)$ (due to the noise driving fluctuations of size σ).

Time dependent coordinate transform Straightforwardly,

$$y_{1} = e^{-1t} \star e^{-1t} \star w_{3} \ \sigma \varepsilon^{2} \left(-b_{3}b_{2}X_{1}^{2} - b_{3}b_{1}X_{2}^{2}\right) + e^{-1t} \star w_{3} \ \sigma \varepsilon^{2} \left(-b_{3}b_{2}X_{1}^{2} - b_{3}b_{1}X_{2}^{2}\right) + e^{-1t} \star w_{3} \ \sigma \varepsilon^{2} \left(-b_{3}b_{2}X_{1}^{2} - b_{3}b_{1}X_{2}^{2}\right) + e^{-1t} \star w_{3} \ \sigma + Y_{1} + \varepsilon b_{3}X_{2}X_{1}$$

$$x_{1} = e^{-1t} \star w_{3} Y_{1} \sigma \varepsilon^{2} b_{2}b_{1}X_{1} - e^{-1t} \star w_{3} \ \sigma \varepsilon b_{1}X_{2} + 1/2Y_{1}^{2} \varepsilon^{2} b_{2}b_{1}X_{1} - Y_{1}\varepsilon b_{1}X_{2} + X_{1}$$

$$x_{2} = e^{-1t} \star w_{3} Y_{1} \sigma \varepsilon^{2} b_{2}b_{1}X_{2} - e^{-1t} \star w_{3} \ \sigma \varepsilon b_{2}X_{1} + 1/2Y_{1}^{2} \varepsilon^{2} b_{2}b_{1}X_{2} - Y_{1}\varepsilon b_{2}X_{1} + X_{2}$$

Result normal form DEs As expected, $Y_1 = 0$ is the emergent (almost always) stochastic slow manifold.

$$\dot{Y}_1 = 4w_3Y_1\sigma\varepsilon^3b_3b_2b_1X_2X_1 + Y_1\varepsilon^2(-b_3b_2X_1^2 - b_3b_1X_2^2) - Y_1$$

$$\dot{X}_1 = w_3 \sigma \varepsilon^3 \left(-2b_3 b_2 b_1 X_2 X_1^2 - 2b_3 b_1^2 X_2^3 \right) + w_3 \sigma \varepsilon b_1 X_2 + \varepsilon^2 b_3 b_1 X_2^2 X_1$$

$$\dot{X}_2 = w_3 \sigma \varepsilon^3 \left(-2b_3 b_2^2 X_1^3 - 2b_3 b_2 b_1 X_2^2 X_1 \right) + w_3 \sigma \varepsilon b_2 X_1 + \varepsilon^2 b_3 b_2 X_2 X_1^2$$

Majda et al. (2002) predicts, their equation (52), the two leading order terms in the deterministic part and the linear noise part. I suspect their first term in each equation is an Ito version of my Stratonovich modelling. All higher order terms are missed by their averaging, but easily constructed here by increasing the argument 4 to the procedure.

10.2 Additive triad model

The additive triad model of Majda et al. (2002) consists of three modes, v_1 , v_2 and v_3 , as before. However, these now evolve in time according to

$$\begin{split} \frac{dv_1}{dt} &= b_1 v_2 v_3 \,, \\ \frac{dv_2}{dt} &= -v_2 + b_2 v_1 v_3 + \sigma_2 \dot{W}_2 \,, \\ \frac{dv_3}{dt} &= -v_3 + b_3 v_1 v_2 + \sigma_3 \dot{W}_3 \,, \end{split}$$

where b_j and σ_j are some constants, and there is independent stochastic forcing of the second and third modes. Here I have already scaled the equations so that the rate of decay of *both* the second and third mode is one.² Thus on long time scales we expect the second and third modes to be essentially negligible and the system to be modelled by the relatively slow evolution of the first mode. This section constructs the stochastic normal form of its centre manifold model as the basis for a model over long time scales with new noise processes.

The procedure stonormalform is already loaded. Write a message saying we are now analysing the next system.

126 write "**** Additive Triad system of Majda (2002) ****";

Execute the construction of a normal form for this system using $x_1 = v_1$, $y_i = v_{i+1}$, and $b_{i1}\sigma = \sigma_i$.

² In contrast, Majda et al. (2002) set the two modes to have different decay rates. Do not expect much difference in using the same decay rate, it is just more convenient that the memory convolutions are then identical for the two modes rather than being different. Having the decay rates the same is also closer to my expected application to spatial problems.

The procedure reports that it analyses the following family

$$\dot{x}_1 = \varepsilon b_1 y_2 y_1
\dot{y}_1 = \sigma b_{21} w_2 + \varepsilon b_2 x_1 y_2 - y_1
\dot{y}_2 = \sigma b_{31} w_3 + \varepsilon b_3 x_1 y_1 - y_2$$

Here, ε counts the order of nonlinearity so that the errors $\mathcal{O}(\varepsilon^3)$ are errors $\mathcal{O}(|\vec{v}|^4 + \sigma^4)$ (due to the noise driving fluctuations of size σ).

Time dependent coordinate transform Straightforwardly,

$$y_{1} = Y_{1} + \sigma \varepsilon b_{31} b_{2} e^{-1t} \star e^{-1t} \star w_{3} X_{1} + \sigma b_{21} e^{-1t} \star w_{2}$$

$$y_{2} = Y_{2} + \sigma \varepsilon b_{21} b_{3} e^{-1t} \star e^{-1t} \star w_{2} X_{1} + \sigma b_{31} e^{-1t} \star w_{3}$$

$$x_{1} = -1/2 Y_{2} Y_{1} \varepsilon b_{1} + Y_{2} \sigma \varepsilon \left(-1/2 b_{21} b_{1} e^{t} \star w_{2} - 1/2 b_{21} b_{1} e^{-1t} \star w_{2}\right) + Y_{1} \sigma \varepsilon \left(-1/2 b_{31} b_{1} e^{t} \star w_{3} - 1/2 b_{31} b_{1} e^{-1t} \star w_{3}\right) + X_{1}$$

Result normal form DEs As expected, $Y_1 = Y_2 = 0$ is the emergent (almost always) stochastic slow manifold. Unusually, on this slow manifold $x_1 = X_1$ exactly (to at least the next few orders).

$$\begin{split} \dot{Y}_1 &= Y_2 \sigma^2 \varepsilon^2 \big(-1/2b_{31}b_{21}b_2b_1 \mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_3 \ w_2 - \\ & 1/2b_{31}b_{21}b_2b_1 \mathrm{e}^{-1t} \star w_3 \ w_2 - 1/2b_{31}b_{21}b_2b_1 \mathrm{e}^{-1t} \star w_2 \ w_3 \big) + \\ & Y_2 \varepsilon b_2 X_1 + Y_1 \sigma^2 \varepsilon^2 \big(-1/2b_{31}^2b_2b_1 \mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_3 \ w_3 - \\ & 1/2b_{31}^2b_2b_1 \mathrm{e}^{-1t} \star w_3 \ w_3 \big) - Y_1 \\ \dot{Y}_2 &= Y_2 \sigma^2 \varepsilon^2 \big(-1/2b_{21}^2b_3b_1 \mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_2 \ w_2 - \\ & 1/2b_{21}^2b_3b_1 \mathrm{e}^{-1t} \star w_2 \ w_2 \big) - Y_2 + Y_1 \sigma^2 \varepsilon^2 \big(- \\ & 1/2b_{31}b_{21}b_3b_1 \mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_2 \ w_3 - 1/2b_{31}b_{21}b_3b_1 \mathrm{e}^{-1t} \star w_3 \ w_2 - \\ & 1/2b_{31}b_{21}b_3b_1 \mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_2 \ w_3 \big) + Y_1 \varepsilon b_3 X_1 \\ \dot{X}_1 &= \sigma^2 \varepsilon^2 \big(1/2b_{31}^2b_2b_1 \mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_3 \ w_3 X_1 + \\ & 1/2b_{31}^2b_2b_1 \mathrm{e}^{-1t} \star w_3 \ w_3 X_1 + 1/2b_{21}^2b_3b_1 \mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_2 \ w_2 X_1 + \\ & 1/2b_{21}^2b_3b_1 \mathrm{e}^{-1t} \star w_2 \ w_2 X_1 \big) + \sigma^2 \varepsilon \big(1/2b_{31}b_{21}b_1 \mathrm{e}^{-1t} \star w_3 \ w_2 + \\ & 1/2b_{31}b_2b_1 \mathrm{e}^{-1t} \star w_2 \ w_3 \big) \end{split}$$

The only terms in the model for \dot{X}_1 are the quadratic noise-noise interaction terms. Majda et al. (2002) recognise the last, σ^2 term, but not the first, $X_1\sigma^2$ term. They represent the last as a mean drift and independent noise (the mean drift comes from the Ito representation of the above Stratonovich noise-noise interaction).

11 nonautoTwo: Potzsche and Rasmussen non-autonomous examples

Potzsche & Rasmussen (2006) establish Taylor approximations of various integral manifolds of non-autonomous systems. They give two examples.

11.1 Lorenz near the pitchfork bifurcation

Example 5.1 of Potzsche & Rasmussen (2006) is

$$\begin{split} \dot{x}_1 &= \sigma_{\epsilon}(x_2 - x_1), \\ \dot{x}_2 &= \rho_{\epsilon}x_1 - x_2 - x_1x_3, \\ \dot{x}_3 &= -\beta_{\epsilon}x_3 + x_1x_2. \end{split}$$

where parameters are $\sigma_{\epsilon} = \sigma_0 + \epsilon \sigma(t)$, $\rho_{\epsilon} = 1 + \rho_0 + \epsilon \rho(t)$ and $\beta_{\epsilon} = \beta_0 + \epsilon \beta(t)$. When there is no parametric fluctuations, $\epsilon = 0$, there is a pitchfork bifurcation as ρ_0 crosses zero so they set $\rho_0 = 0$. For the general procedure we must set σ_0 and β_0 to some definite values, here $\sigma_0 = \beta_0 = 1$.

```
134 s0:=beta0:=1;
135 s1:=s0/(s0+1);
```

To analyses dynamics at this pitchfork bifurcation in the presence of fluctuations, Potzsche & Rasmussen (2006) [p.449] take a linear transform of the system to variables

$$\vec{y} = \begin{bmatrix} -\sigma_0 & 0 & 1\\ 1 & 0 & 1\\ 0 & 1 & 0 \end{bmatrix} \vec{x} \,.$$

In the following coding I use $x(1) = y_3$, $y(1) = y_1$ and $y(2) = y_2$; there are no unstable modes.

There is a notational glitch in that the procedure uses σ for the size of the non-autonomous effects, whereas they use σ as a parameter to the Lorenz system. Herein let s denote their σ , and let σ denote their ϵ . Then their fluctuations $\epsilon \rho(t)$ are represented in the coded input by $\mathbf{w(rho)}$ whereas in the output it is represented by σw_{ρ} , and similarly for the other non-autonomous quantities.

Start by loading the procedure.

```
136 in_tex "../stoNormForm.tex"$
```

Execute the construction of a normal form for the system.

```
137 factor yy,y,xx,x;
138 stonormalform(
       \{ s0*s1*y(1)*y(2) -s1*x(1)*y(2) +s1*x(1)*w(rho) \}
139
140
            +(w(s)-w(rho)/(s0+1))*y(1)},
       \{-(s0+1)*y(1)+s1*y(1)*y(2)-x(1)*y(2)/(s0+1)
141
            +w(rho)/(s0+1)*x(1)
142
            -(w(s)+w(rho)/(s0+1))*y(1),
143
         -beta0*y(2)-s0*y(1)^2+(1-s0)*x(1)*y(1)+x(1)^2
144
145
            -w(beta)*y(2)},
       {},
146
       3)$
147
148 %end; % optionally end examples here
```

The procedure reports that it analyses the following family

$$\dot{x}_1 = -1/2x_1y_2\varepsilon + 1/2x_1\sigma w_\rho + 1/2y_2y_1\varepsilon + y_1\sigma(-1/2w_\rho + w_s)
\dot{y}_1 = -1/2x_1y_2\varepsilon + 1/2x_1\sigma w_\rho + 1/2y_2y_1\varepsilon + y_1\sigma(-1/2w_\rho - w_s) - 2y_1
\dot{y}_2 = x_1^2\varepsilon - y_2\sigma w_\beta - y_2 - y_1^2\varepsilon$$

Time dependent coordinate transform Straightforwardly,

$$y_{1} = X_{1}Y_{2}\sigma\varepsilon\left(-\frac{1}{2}e^{-1t}\star w_{\beta} + e^{-1t}\star w_{\rho} - \frac{1}{4}e^{-2t}\star w_{\rho} + \frac{1}{2}e^{-1t}\star w_{s}\right) - \frac{1}{2}X_{1}Y_{2}\varepsilon + \frac{1}{2}X_{1}\sigma e^{-2t}\star w_{\rho} + \frac{1}{2}e^{2t}\star w_{\beta} - \frac{1}{4}e^{2t}\star w_{\rho} + \frac{1}{2}e^{2t}\star w_{\beta} - \frac{1}{2}e^{t}\star w_{\beta} + \frac{1}{2}e^{t}\star w_{\beta} - \frac{1}{$$

Result normal form DEs As expected, $Y_1 = Y_2 = 0$ is the emergent slow manifold, albeit hideously complicated SDEs due to the nonlinear interaction of the three non-autonomous effects.

```
\dot{Y}_1 = X_1^2 Y_1 \sigma^2 \varepsilon^2 (1/2 e^{2t} \star e^{2t} \star w_\rho \ w_\rho + 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star w_\rho \ w_\rho - 1/4 e^{-2t} \star e^{-2t} \star w_\rho \ w_\rho - 1/4 e^
                                                                                                                    1/2e^{-2t} \star e^{-2t} \star w_{\rho} w_{s} - e^{2t} \star e^{2t} \star w_{s} w_{\rho} + 1/12e^{2t} \star w_{\beta} w_{\rho} +
                                                                                                                        1/2e^{-1t}\star w_{\beta}w_{\beta} + 1/3e^{-1t}\star w_{\beta}w_{\rho} - 1/2e^{-1t}\star w_{\beta}w_{s} -
                                                                                                                        3/4e^{2t}\star w_{\rho}w_{\beta} + 13/48e^{2t}\star w_{\rho}w_{\rho} + 3/8e^{2t}\star w_{\rho}w_{s} +
                                                                                                                        1/3e^t \star w_\rho w_\beta - 2/3e^t \star w_\rho w_\rho - 1/3e^t \star w_\rho w_s +
                                                                                                                        1/2e^{-1t}\star w_{\rho} w_{\beta} + 1/3e^{-1t}\star w_{\rho} w_{\rho} - 1/2e^{-1t}\star w_{\rho} w_{s} -
                                                                                                                        1/6e^{-2t} \star w_{\rho} w_{\beta} + 7/12e^{-2t} \star w_{\rho} w_{\rho} - 7/12e^{-2t} \star w_{\rho} w_{s} +
                                                                                                                        3/2e^{2t}\star w_s w_{\beta} - 3/8e^{2t}\star w_s w_{\rho} - 3/4e^{2t}\star w_s w_s - e^t\star w_s w_{\beta} +
                                                                                                                    2e^{t}\star w_{s} w_{\rho} + e^{t}\star w_{s} w_{s} + X_{1}^{2}Y_{1}\sigma\varepsilon^{2}(-1/2w_{\beta} - 3/4w_{\rho} - 3/4w_{\rho})
                                                                                                                    1/4w_s) + 1/2X_1^2Y_1\varepsilon^2 + X_1Y_2^2\sigma^2\varepsilon^2( - 1/8e^{2t}\star e^{2t}\star w_\rho w_\rho +
                                                                                                                        1/4e^{2t} \star e^{2t} \star w_s \ w_\rho - 1/12e^{2t} \star w_\beta \ w_\rho + 1/4e^t \star w_\beta \ w_\beta - 1/4e^t \star w_\beta
                                                                                                                        1/2e^{t}\star w_{\beta}w_{\rho} - 1/4e^{t}\star w_{\beta}w_{s} + 1/4e^{-1t}\star w_{\beta}w_{\beta} +
                                                                                                                        1/6e^{-1t}\star w_{\beta}w_{\rho} - 1/4e^{-1t}\star w_{\beta}w_{s} + 1/96e^{2t}\star w_{\rho}w_{\rho} +
                                                                                                                        1/6e^t \star w_\rho w_\beta - 1/3e^t \star w_\rho w_\rho - 1/6e^t \star w_\rho w_s - 1/6e^t \star w_\rho w_\rho - 1/6e^t \star w_
                                                                                                                        1/2e^{-1t}\star w_{\rho}w_{\beta} - 1/3e^{-1t}\star w_{\rho}w_{\rho} + 1/2e^{-1t}\star w_{\rho}w_{s} -
                                                                                                                    1/12e^{-2t}\star w_{\rho}w_{\beta} + 1/96e^{-2t}\star w_{\rho}w_{\rho} - 1/6e^{-2t}\star w_{\rho}w_{s} - 1/6e^{-2t}\star w_{\rho}
                                                                                                                        1/6e^{2t} \star w_s w_\rho - 1/4e^t \star w_s w_\beta + 1/2e^{1t} \star w_s w_\rho + 1/4e^t \star w_s w_s - 1/4e^t \star w_s w_\rho + 1/4e^t \star w_\phi + 1/4e
                                                                                                                        1/4e^{-1t}\star w_s w_\beta - 1/6e^{-1t}\star w_s w_\rho + 1/4e^{-1t}\star w_s w_s + 1/4e^{-1t}\star w_
                                                                                                                    3/4X_1Y_2^2\sigma\varepsilon^2w_{\rho} - 1/2X_1Y_2^2\varepsilon^2 + Y_1\sigma^2(1/4e^{2t}\star w_{\rho}w_{\rho} - 1/2V_1Y_2^2\varepsilon^2)
                                                                                                                        1/2e^{2t} \star w_s w_o + Y_1 \sigma (-1/2w_o - w_s) - 2Y_1
\dot{Y}_2 = X_1^2 Y_2 \sigma^2 \varepsilon^2 (e^t \star w_\beta w_\beta + e^t \star w_\beta w_\rho - 1/3 e^{-1t} \star w_\beta w_\rho + 1/3 e^{-1t} \star w_\phi + 1/3 e^{-1t} 
                                                                                                                    e^{-1t} \star w_{\beta} w_{s} + 2/3e^{t} \star w_{\rho} w_{\beta} + 2/3e^{t} \star w_{\rho} w_{\rho} + 2/3e^{-1t} \star w_{\rho} w_{\rho} - 2/3e^{t} \star w_{\rho} w_{\rho} + 2/3e^{t} \star w_{\rho} w
                                                                                                                        2e^{-1t}\star w_{\rho} w_s + 1/3e^{-2t}\star w_{\rho} w_{\beta} - 13/24e^{-2t}\star w_{\rho} w_{\rho} +
                                                                                                                        5/6e^{-2t} \star w_{\rho} w_{s} - e^{t} \star w_{s} w_{\beta} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} - e^{t} \star w_{s} w_{\rho} + 1/3e^{-1t} \star w_{s} w_{\rho} + 1/3e^{
                                                                                                                        e^{-1t} \star w_s w_s + X_1^2 Y_2 \sigma \varepsilon^2 (-w_\beta - 1/2w_\rho + w_s) + X_1^2 Y_2 \varepsilon^2 - W_\beta w_s + W_\beta w_s 
                                                                                                                        Y_2 \sigma w_\beta - Y_2
```

$$\begin{split} \dot{X}_1 &= X_1^3 \sigma^2 \varepsilon^2 \big(-1/4 \mathrm{e}^{-2t} \star \mathrm{e}^{-2t} \star w_\rho \ w_\rho + 1/2 \mathrm{e}^{-2t} \star \mathrm{e}^{-2t} \star w_\rho \ w_s - \\ & 1/2 \mathrm{e}^{-1t} \star w_\beta \ w_\beta - 1/3 \mathrm{e}^{-1t} \star w_\beta \ w_\rho + 1/2 \mathrm{e}^{-1t} \star w_\beta \ w_s + \\ & 1/4 \mathrm{e}^{-2t} \star w_\beta \ w_\rho - 1/2 \mathrm{e}^{-2t} \star w_\beta \ w_s - 1/2 \mathrm{e}^{-1t} \star w_\rho \ w_\beta - \\ & 1/3 \mathrm{e}^{-1t} \star w_\rho \ w_\rho + 1/2 \mathrm{e}^{-1t} \star w_\rho \ w_s - 1/12 \mathrm{e}^{-2t} \star w_\rho \ w_\beta - \\ & 13/48 \mathrm{e}^{-2t} \star w_\rho \ w_\rho + 3/8 \mathrm{e}^{-2t} \star w_\rho \ w_s - 1/8 \mathrm{e}^{-2t} \star w_s \ w_\rho + \\ & 1/4 \mathrm{e}^{-2t} \star w_s \ w_s \big) + X_1^3 \sigma \varepsilon^2 \big(1/2 w_\beta + 3/4 w_\rho - 1/4 w_s \big) - 1/2 X_1^3 \varepsilon^2 + \\ & X_1 \sigma^2 \big(-1/4 \mathrm{e}^{-2t} \star w_\rho \ w_\rho + 1/2 \mathrm{e}^{-2t} \star w_\rho \ w_s \big) + 1/2 X_1 \sigma w_\rho \end{split}$$

In their analysis Potzsche & Rasmussen (2006) explicitly report the last and third-to-last terms above, for these choices of σ_0 and β_0 , to deduce their model (5.3) which here is

$$\dot{X} \approx \frac{1}{2}\sigma w_{\rho}X - \frac{1}{2}X^3.$$

Nice agreement.

11.2 Fluctuating kdV example

Potzsche & Rasmussen (2006) [Example 5.4] seek travelling wave solutions, u(x-ct) with wave speed c, of a modified KdV equation. This leads to the following system

$$\dot{x}_1 = x_2$$
, $\dot{x}_2 = x_3$, $\dot{x}_3 = c^2 x_2 - a(t) x_1^2 x_2$.

This analysis is for wave speed $c^2 = 1$. A transform to diagonalise the linear part into slow variable x, stable y and unstable z is then that $x_1 = x + y + z$, $x_2 = z - y$ and $x_3 = z + y$. Using w(a) to denote the variable coefficient a(t), it is represented in this output by σw_a .

The procedure **stonormalform** is already loaded. Write a message saying we are now analysing the next system.

```
149 write "**** Fluctuating kdV example of P&R (2006) ****"; 150 factor zz,z;
```

Execute the construction of a normal form for this system.

The procedure reports that it analyses the following family

$$\begin{split} \dot{x}_1 &= -x_1^2 y_1 \sigma w_a + x_1^2 z_1 \sigma w_a - 2 x_1 y_1^2 \sigma w_a + 2 x_1 z_1^2 \sigma w_a - y_1^3 \sigma w_a - \\ & y_1^2 z_1 \sigma w_a + y_1 z_1^2 \sigma w_a + z_1^3 \sigma w_a \\ \dot{y}_1 &= 1/2 x_1^2 y_1 \sigma w_a - 1/2 x_1^2 z_1 \sigma w_a + x_1 y_1^2 \sigma w_a - x_1 z_1^2 \sigma w_a + \\ & 1/2 y_1^3 \sigma w_a + 1/2 y_1^2 z_1 \sigma w_a - 1/2 y_1 z_1^2 \sigma w_a - y_1 - 1/2 z_1^3 \sigma w_a \\ \dot{z}_1 &= 1/2 x_1^2 y_1 \sigma w_a - 1/2 x_1^2 z_1 \sigma w_a + x_1 y_1^2 \sigma w_a - x_1 z_1^2 \sigma w_a + \\ & 1/2 y_1^3 \sigma w_a + 1/2 y_1^2 z_1 \sigma w_a - 1/2 y_1 z_1^2 \sigma w_a - 1/2 z_1^3 \sigma w_a + z_1 \end{split}$$

Time dependent coordinate transform Straightforwardly,

$$\begin{split} z_1 &= -1/2X_1^2Y_1\sigma\mathrm{e}^{2t}\star w_a - X_1Y_1^2\sigma\mathrm{e}^{3t}\star w_a - X_1Z_1^2\sigma\mathrm{e}^{-1t}\star w_a - \\ & 1/2Y_1^3\sigma\mathrm{e}^{4t}\star w_a - 1/2Y_1^2Z_1\sigma\mathrm{e}^{2t}\star w_a - 1/2Z_1^3\sigma\mathrm{e}^{-2t}\star w_a + Z_1 \\ y_1 &= -1/2X_1^2Z_1\sigma\mathrm{e}^{-2t}\star w_a - X_1Y_1^2\sigma\mathrm{e}^t\star w_a - X_1Z_1^2\sigma\mathrm{e}^{-3t}\star w_a - \\ & 1/2Y_1^3\sigma\mathrm{e}^{2t}\star w_a - 1/2Y_1Z_1^2\sigma\mathrm{e}^{-2t}\star w_a + Y_1 - 1/2Z_1^3\sigma\mathrm{e}^{-4t}\star w_a \\ x_1 &= X_1^2Y_1\sigma\mathrm{e}^t\star w_a + X_1^2Z_1\sigma\mathrm{e}^{-1t}\star w_a + 2X_1Y_1^2\sigma\mathrm{e}^{2t}\star w_a + \\ & 2X_1Z_1^2\sigma\mathrm{e}^{-2t}\star w_a + X_1 + Y_1^3\sigma\mathrm{e}^{3t}\star w_a + Y_1^2Z_1\sigma\mathrm{e}^t\star w_a + \\ & Y_1Z_1^2\sigma\mathrm{e}^{-1t}\star w_a + Z_1^3\sigma\mathrm{e}^{-3t}\star w_a \end{split}$$

Result normal form DEs

$$\begin{split} \dot{Z}_1 &= -1/4X_1^4 Z_1 \sigma^2 \mathrm{e}^{-2t} \star w_a \, w_a + X_1^2 Y_1 Z_1^2 \sigma^2 \left(3/4 \mathrm{e}^{2t} \star w_a \, w_a - \mathrm{e}^{t} \star w_a \, w_a + 5/4 \mathrm{e}^{-2t} \star w_a \, w_a - 2 \mathrm{e}^{-3t} \star w_a \, w_a \right) - 1/2 X_1^2 Z_1 \sigma w_a + \\ & \quad Y_1^2 Z_1^3 \sigma^2 \left(3/4 \mathrm{e}^{2t} \star w_a \, w_a - \mathrm{e}^{t} \star w_a \, w_a - 3/4 \mathrm{e}^{-2t} \star w_a \, w_a + \\ & \quad + 2^{-3t} \star w_a \, w_a - 3/4 \mathrm{e}^{-4t} \star w_a \, w_a \right) - 1/2 Y_1 Z_1^2 \sigma w_a + Z_1 \end{split}$$

$$\dot{Y}_1 = 1/4 X_1^4 Y_1 \sigma^2 \mathrm{e}^{2t} \star w_a \, w_a + X_1^2 Y_1^2 Z_1 \sigma^2 \left(2 \mathrm{e}^{3t} \star w_a \, w_a - \frac{1}{2} (2 \mathrm{e}^{3t} \star w_a \, w_a + \frac{1}{2} (2 \mathrm{e}^{3t} \star w_a \, w_a$$

Since $Z_1 = 0$ is invariant, putting $Z_1 = 0$ into the coordinate transform gives the centre-stable manifold. Then the expression for z_1 in the above coordinate transform leads to the same convolutions as those of Potzsche & Rasmussen (2006) [pp.453–4]. Conversely, since $Y_1 = 0$ is invariant, putting $Y_1 = 0$ gives the centre-unstable manifold and the expression for y_1 above leads to the same convolutions as those of Potzsche & Rasmussen (2006). Presumably the distortions of the other variables have a higher order influence on this nice agreement.

The X_1 -evolution is zero on either of these invariant manifolds.

12 noisyMMH: noisy Michaelis-Menten-Henri chemical kinetics

The Michaelis-Menten-Henri system, in non-dimensional form, is

$$\dot{x} = \epsilon[-x + (x + \kappa - \lambda)y], \qquad \dot{y} = x - (x + \kappa)y.$$

The usual approach is singular perturbation theory, but let's not obfuscate with singular limits, and instead use the clarity of being regular. A manifold of equilibria occur at $y = x/(x + \kappa)$ and $\epsilon = 0$ (also if $\epsilon \neq 0$ and $\lambda = 0$ but we do not consider this case). Let's explore dynamics based at arbitrary point on this equilibrium manifold: substitute $x(t) = x_0 + x_1(t)$ and $y(t) = x_0/(x_0 + \kappa) + y_1(t)$. Also, to get the decay rate of y_1 to be a simple number, stretch

time by the factor $x_0 + \kappa$: that is, $(x_0 + \kappa)dt = d\tau$ where τ is the time of the analysis so that

$$\frac{1}{x_0 + \kappa} \dot{x}_1(t) = x_1'(\tau), \quad \frac{1}{x_0 + \kappa} \dot{y}_1(t) = y_1'(\tau).$$

This does mean that we have to be careful interpreting the results. Hence derive

$$x_1' = \frac{\epsilon}{x_0 + \kappa} \left[-x_0 - x_1 + (x_0 + x_1 + \kappa - \lambda)(x_0/(x_0 + \kappa) + y_1) \right],$$

$$y_1' = -y_1 + \frac{1}{x_0 + \kappa} \left[-x_1 y_1 + x_0 + x_1 - \frac{(x_0 + x_1 + \kappa)x_0}{x_0 + \kappa} \right].$$

As a prototypical example, let's investigate the simplest stochastic effects on this MM system of additive noises $w_1(\tau)$ and $w_2(\tau)$. The additive noise transforms to a multiplicative noise on the slow manifold, so it is important to remember that *all* analysis and results are in the *Stratonovich* interpretation.

The analysis here is *strong*, *pathwise*. The transformations here only rely on the 'noise' being measurable, so the results also apply to deterministic non-autonomous forcing. The analysis may also apply to to non-Brownian noise provided the appropriate interpretation is used (e.g., the Marcus interpretation). That is, as long as standard rules of integral calculus are valid.

Start by loading the procedure.

```
158 in_tex "../stoNormForm.tex"$
```

Set some variables for simplicity.

```
159 let rho*x0=>1-rho*kappa; % define rho=1/(x0+kappa)
160 kappa:=1; lam:=1/2; % set for simplicity
```

Execute the construction of a normal form for the system.

```
161 factor epsilon;
162 stonormalform(
        { epsilon*rho*( -x0-x(1)
163
          +(x0+x(1)+kappa-lam)*(x0*rho+y(1)))
164
          +w(1) },
165
        {-y(1)+rho*(-x(1)*y(1)+x0+x(1))}
166
          -(x0+x(1)+kappa)*x0*rho)
167
          +w(2) },
168
169
        {},
170
        3)$
171 end;
```

The procedure reports that it analyses the following family

$$\dot{x}_1 = \epsilon \varepsilon \left(x_1 y_1 \rho - x_1 \rho^2 - 1/2 y_1 \rho + y_1 + 1/2 \rho^2 - 1/2 \rho \right) + \sigma w_1$$

$$\dot{y}_1 = \sigma w_2 + \varepsilon \left(-x_1 y_1 \rho + x_1 \rho^2 \right) - y_1$$

Time dependent coordinate transform The algorithm constructs a coordinate transform to variables (X_1, Y_1) , including terms quadratic in σ , that to errors $\mathcal{O}(\sigma^2, \varepsilon^2)$, is the following. The coordinate transform depends upon both the past and the future via convolutions $e^{-1t}\star$ and $e^t\star$, respectively. The following expressions are complicated because stochastic effects interact through nonlinearity in a combinatorial explosion of ways. We almost certainly do not need all these terms. I subsequently explain why the blue terms are the ones describing the emergent stochastic slow manifold and the evolution thereon. Further, remember that the dominant terms are towards the end of each expression.

$$y_{1} = \sigma \varepsilon \left(-e^{-1t} \star e^{-1t} \star w_{2} X_{1} \rho - e^{-1t} \star w_{1} \rho^{2} \right) + \sigma e^{-1t} \star w_{2} + \varepsilon X_{1} \rho^{2} + Y_{1}$$

$$x_{1} = \epsilon \sigma \varepsilon \left(-e^{-1t} \star w_{2} X_{1} \rho + 1/2 e^{-1t} \star w_{2} \rho - e^{-1t} \star w_{2} - e^{t} \star w_{1} Y_{1} \rho \right) + \epsilon \varepsilon \left(-X_{1} Y_{1} \rho + 1/2 Y_{1} \rho - Y_{1} \right) + X_{1}$$

These new coordinates (X_1, Y_1) are non-Markovian in relation to (x_1, y_1) , in some sense, but the non-Markovian nature is exponentially decaying away from the current time. The construction of a non-autonomous stochastic slow manifold has to look to the future and the past in order to find out what variations are going to stay bounded for all time.

Result normal form DEs In the (X_1, Y_1) coordinates, the stochastic system satisfies the following Stratonovich system, to errors $\mathcal{O}(\sigma^3, \varepsilon^3)$.

$$\begin{split} Y_1' &= \epsilon \sigma^2 \varepsilon^2 \big(\mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_2 \ w_1 Y_1 \rho^2 + 2 \mathrm{e}^{-1t} \star w_2 \ w_1 Y_1 \rho^2 \big) + \\ &\quad \epsilon \sigma \varepsilon^2 \big(2 w_2 X_1 Y_1 \rho^2 - w_2 Y_1 \rho^2 + 2 w_2 Y_1 \rho - w_1 Y_1 \rho^3 \big) + \epsilon \varepsilon^2 \big(- X_1 Y_1 \rho^3 + 1/2 Y_1 \rho^3 - Y_1 \rho^2 \big) - \varepsilon X_1 Y_1 \rho - Y_1 \\ X_1' &= 1/2 \epsilon^2 \sigma \varepsilon^2 w_2 \rho^2 + \epsilon \sigma^2 \varepsilon^2 \big(- \mathrm{e}^{-1t} \star \mathrm{e}^{-1t} \star w_2 \ w_1 X_1 \rho^2 - \\ &\quad 2 \mathrm{e}^{-1t} \star w_2 \ w_1 X_1 \rho^2 + 1/2 \mathrm{e}^{-1t} \star w_2 \ w_1 \rho^2 - \mathrm{e}^{-1t} \star w_2 \ w_1 \rho - \\ &\quad - \mathrm{e}^{-1t} \star w_1 \ w_1 \rho^3 \big) + \epsilon \sigma^2 \varepsilon \mathrm{e}^{-1t} \star w_2 \ w_1 \rho + \epsilon \sigma \varepsilon^2 \big(- w_2 X_1^2 \rho^2 + \\ &\quad 1/2 w_2 X_1 \rho^2 - w_2 X_1 \rho - w_1 X_1 \rho^3 + 1/2 w_1 \rho^3 - w_1 \rho^2 \big) + \\ &\quad \epsilon \sigma \varepsilon \big(w_2 X_1 \rho - 1/2 w_2 \rho + w_2 \big) + \epsilon \varepsilon^2 \big(X_1^2 \rho^3 - 1/2 X_1 \rho^3 + \\ &\quad X_1 \rho^2 \big) + \epsilon \varepsilon \big(- X_1 \rho^2 + 1/2 \rho^2 - 1/2 \rho \big) + \sigma w_1 \end{split}$$

Discussion

- In the Y_1' SDE, by construction, every term is $\propto Y_1$, and, further, the leading term gives $Y_1' \approx -Y_1$. Hence, $Y_1 \approx \mathcal{O}(e^{-\tau}) = \mathcal{O}(e^{-\int (x_0 + \kappa)dt})$ as time increases. Consequently, by continuity, in some finite domain about $(x_0, y_0), Y_1 \to 0$ to form the emergent stochastic slow manifold $Y_1 = 0$.
- The local shape of the slow manifold is thus given by substituting $Y_1 = 0$ into the expressions for (x_1, y_1) . Thus the slow manifold is locally parametrised by X_1, ϵ, σ . Now the variation in X_1 is the Taylor series for the variation in x_0

(as they are both describing the same slow manifold). So all we need is to set $X_1 = 0$ and look at the shape of the slow manifold in terms of x_0, ϵ, σ , that is, the blue terms.

- Setting $X_1 = Y_1 = 0$ gives $y_1 \approx -\sigma e^{-1t} \star w_1 \rho^2 + \sigma e^{-1t} \star w_2$, that is, since $\rho = 1/(x + \kappa) = 1/(x + 1)$,

$$y = y_0 + y_1 \approx \frac{x_0}{x_0 + 1} - \frac{\sigma e^{-1t} \star w_1}{(x+1)^2} + \sigma e^{-1t} \star w_2$$
.

Dominantly, the slow manifold jitters up/down in y due to the recent history of noise w_2 , but also is affected from the recent history of the noise w_1 in x.

- Setting $X_1 = Y_1 = 0$ gives $x_1 \approx \epsilon \sigma \varepsilon (1/2e^{-1t} \star w_2 \rho - e^{-1t} \star w_2)$, that is,

$$x = x_0 + x_1 \approx x_0 - \epsilon \sigma \frac{2x_0 + 1}{2(x_0 + 1)} e^{-1t} \star w_2$$
.

The noise w_2 in y generates a history dependent slip between x and the relevant x_0 !

This slip may be seen to be due to the slope of isochrons transversal to the slow manifold—a slope not detected in Singular Perturbation Analysis.

This stochastic-MM example also shows the general property that although the *existence* of a slow manifold has future dependence, here via $e^t \star$ convolutions, the slow manifold itself and the evolution thereon depends only upon the history, here via $e^{-1t} \star$ convolutions.

• Now for the x-evolution on the stochastic slow manifold. Consider $X(t) = x_0 + x_1(t)$, so that $X' = x_1'$. Recall that on the slow manifold, $Y_1 = 0$ and $x \approx x_0 - \epsilon \sigma \frac{2x+1}{2(x+1)} e^{-1t} \star w_2$, $X_1 = 0$, so also putting $X_1 = 0$ and $x_0 = X$ give the evolution for the slow variable X, namely the global slow evolution is

$$X' \approx -\epsilon \frac{X}{2(X+1)^2} + \epsilon^2 \frac{X(2X+1)}{4(X+1)^5} + \sigma w_1 + \epsilon \sigma \frac{2X+1}{2(X+1)} w_2.$$

The stochastic slow variable X is not quite the same as the physical x. This coordinate transform lacks any convolutions in time. That lack is part of the art of the construction.

If, instead, one wants the slow variable to be precisely x, as many implicitly assume they can, then convolutions must occur in x'). We may see this by constructing a nonlinear coordinate transform that maintains, when $Y_1 = 0$, that $x = x_0$, precisely. It is straightforward to modify the algorithm to do so. The generic consequence is that terms *linear* in the noise appear in the evolution x' that have fast-time history convolutions. That is, the consequence is that undesirable fast-time history integrals occur in the evolution of the supposedly slow variable x.

Noise-noise interactions However, effects which are quadratic in the noise, due to noise-noise interactions, generally involve convolutions that cannot be removed from the evolution of the slow variable, as seen in expressions for X'_1 . Here, the lowest order example is the term

$$+\epsilon\sigma^2 e^{-1t}\star w_2 w_1/(x+1)$$

which could be included in the retained terms of X_1' . We argued (Chao & Roberts 1996, §4) that such terms 'bring up' new information from the fluctuations on the fast-time microscale, and hence cause noise effects in the slow model that are independent of slow-scale sampling of w_1 and w_2 . We argued that such terms, when one only samples them on the long-times of the slow manifold, should thus be replaced by a new noise, namely $e^{-\beta t} \star w_2 w_1 \sim \frac{1}{2\sqrt{\beta}} w_3$ when all w_j are formally 'the derivatives' of independent Wiener processes.

The above results are for one example of a stochastic MM system. Almost all other stochastic MM systems would have the same issues.

13 heatXchanger: Local analysis of heat exchanger

Roberts (2015a) provides novel theoretical support for the method of multiple scales in spatio-temporal systems, and then extends this important method. Perhaps the simplest example is the heat exchanger: the non-autonomous slow manifold analysis that is at the heart of the novel methodology is determined here. An internal technical report³ considered the scenario where hot fluid enters one pipe from the right having temperature field a(x,t), and cold fluid enters the other pipe from the left with temperature field b(x,t). Non-dimensional governing PDEs are

$$\frac{\partial a}{\partial t} = + \frac{\partial a}{\partial x} + \frac{1}{2}(b-a), \quad \frac{\partial b}{\partial t} = - \frac{\partial b}{\partial x} + \frac{1}{2}(a-b).$$

Transform to mean and difference fields:

$$c(x,t) := \tfrac{1}{2}(a+b), \quad d(x,t) := \tfrac{1}{2}(a-b), \quad \text{i.e., } a = c+d, \quad b = c-d.$$

The mean and difference of the PDEs gives the equivalent PDE system

$$\frac{\partial c}{\partial t} = \frac{\partial d}{\partial x}, \quad \frac{\partial d}{\partial t} = -d + \frac{\partial c}{\partial x}.$$

In this form we readily see that the difference field d tends to decay exponentially quickly, but that interaction between gradients of the mean and difference fields generates other effects.

 $^{^3\,\}mathrm{AJR}$ (2012), Derive boundary conditions for heat exchanger modelling using near boundary dynamics, <code>dbefhem.pdf</code>

The approach is to expand the fields in their local spatial structure based around a station x = X. Expand advection-exchange in a heat exchanger in powers of $(x - X)^n/n!$.

$$c(x,t) = c_0(X,t) + c_1(X,t)(x-X) + c_2(X,t)\frac{1}{2}(x-X)^2 + c_3(X,t)\frac{1}{6}(x-X)^3 + c_4(X,x,t)\frac{1}{24}(x-X)^4,$$

$$d(x,t) = d_0(X,t) + d_1(X,t)(x-X) + d_2(X,t)\frac{1}{2}(x-X)^2 + d_3(X,t)\frac{1}{6}(x-X)^3 + d_4(X,x,t)\frac{1}{24}(x-X)^4,$$

With Taylor Remainder Theorem closing the problem in terms of unknown functions which here are represented by the non-autonomous forcing w_i . Variables $y(j) = d_{j-1}$ and $x(j) = c_{j-1}$. Also $w(1) = d_{4X}\eta_x$ and $w(2) = c_{4X}\xi_x$ and evaluate at intensity $\sigma = 5$.

Start by loading the procedure.

```
172 in_tex "../stoNormForm.tex"$
```

Execute the construction of a normal form for the system.

```
173 stonormalform(
        {y(2),y(3),y(4),y(5),w(1)},
174
        {-y(1)+x(2)}
175
176
        ,-y(2)+x(3)
        ,-y(3)+x(4)
177
178
        ,-y(4)+x(5)
        ,-y(5)+w(2),
179
        { },
180
181
        99);
182 %end; % optional finish here
```

Here, since the results are exact, we can notionally carry out analysis to high-order, here coded 99th order. Alternatively, as in the next section, one could divide by small all the y(j) terms in the x(i) equations, and all the x(i) terms in the y(j) equations to analyse the original system without the algorithm's artifice of small/ ε .

Specified dynamical system The above embeds the odes as the following.

$$\dot{x}_1 = \varepsilon y_2
\dot{x}_2 = \varepsilon y_3
\dot{x}_3 = \varepsilon y_4
\dot{x}_4 = \varepsilon y_5
\dot{x}_5 = \sigma w_1
\dot{y}_1 = \varepsilon x_2 - y_1
\dot{y}_2 = \varepsilon x_3 - y_2
\dot{y}_3 = \varepsilon x_4 - y_3$$

$$\dot{y}_4 = \varepsilon x_5 - y_4$$

$$\dot{y}_5 = \sigma w_2 - y_5$$

Time dependent coordinate transform

$$y_{1} = \sigma \varepsilon^{4} \left(e^{-1t} \star e^{-1t} \star e^{-1t} \star w_{2} + 2e^{-1t} \star e^{-1t} \star w_{2} + 3e^{-1t} \star w_{2} \right) - \varepsilon^{3} X_{4} + \varepsilon X_{2} + Y_{1}$$

$$y_{2} = \sigma \varepsilon^{3} \left(e^{-1t} \star e^{-1t} \star w_{1} + 2e^{-1t} \star w_{1} \right) - \varepsilon^{3} X_{5} + \varepsilon X_{3} + Y_{2}$$

$$y_{3} = \sigma \varepsilon^{2} \left(- e^{-1t} \star e^{-1t} \star w_{2} - e^{-1t} \star w_{2} \right) + \varepsilon X_{4} + Y_{3}$$

$$y_{4} = -\sigma \varepsilon e^{-1t} \star w_{1} + \varepsilon X_{5} + Y_{4}$$

$$y_{5} = \sigma e^{-1t} \star w_{2} + Y_{5}$$

$$x_{1} = \sigma \varepsilon^{4} \left(- e^{-1t} \star e^{-1t} \star w_{1} - 3e^{-1t} \star w_{1} \right) + \varepsilon^{3} Y_{4} - \varepsilon Y_{2} + X_{1}$$

$$x_{2} = \sigma \varepsilon^{3} \left(e^{-1t} \star e^{-1t} \star w_{2} + 2e^{-1t} \star w_{2} \right) + \varepsilon^{3} Y_{5} - \varepsilon Y_{3} + X_{2}$$

$$x_{3} = \sigma \varepsilon^{2} e^{-1t} \star w_{1} - \varepsilon Y_{4} + X_{3}$$

$$x_{4} = -\sigma \varepsilon e^{-1t} \star w_{2} - \varepsilon Y_{5} + X_{4}$$

$$x_{5} = X_{5}$$

Result normal form DEs

$$\dot{Y}_1 = \varepsilon^4 Y_5 - \varepsilon^2 Y_3 - Y_1$$

$$\dot{Y}_2 = -\varepsilon^2 Y_4 - Y_2$$

$$\dot{Y}_3 = -\varepsilon^2 Y_5 - Y_3$$

$$\dot{Y}_4 = -Y_4$$

$$\dot{Y}_5 = -Y_5$$

$$\dot{X}_1 = 3\sigma \varepsilon^4 w_1 - \varepsilon^4 X_5 + \varepsilon^2 X_3$$

$$\dot{X}_2 = -2\sigma \varepsilon^3 w_2 + \varepsilon^2 X_4$$

$$\dot{X}_3 = -\sigma \varepsilon^2 w_1 + \varepsilon^2 X_5$$

$$\dot{X}_4 = \sigma \varepsilon w_2$$

$$\dot{X}_5 = \sigma w_1$$

Clearly, the emergent slow manifold is $\vec{Y} = 0$. Then the slow evolution of \vec{X} leads to the emergent dynamics being described by

$$\frac{\partial c}{\partial t} \approx \frac{\partial^2 c}{\partial x^2} - \frac{\partial^4 c}{\partial x^4} + 3\sigma w_1.$$

13.1 Near the boundary

This is for the case of boundary conditions $c + pd = \operatorname{cd}_0(t)$ at x = 0 for some parameter p. Computer algebra finds boundary conditions on the fields that reduce the dynamics near the boundary to the following with $x(1) = c_1$, $x(2) = c_3$, $y(1) = d_0$, $y(2) = d_2$ and $w(1) = d_{3X}\eta_x$ with $\sigma = 4$. Curiously, there is no dependence upon parameter p in these dynamics.

The procedure **stonormalform** is already loaded. Write a message saying we are now analysing the next system.

```
183 write "**** Near the boundary ****";
```

Execute the construction of a normal form for this system

Again, the following results are exact.

Specified dynamical system

$$\begin{aligned}
\dot{x}_1 &= y_2 \\
\dot{x}_2 &= \sigma w_1 \\
\dot{y}_1 &= x_1 - y_1 \\
\dot{y}_2 &= x_2 - y_2
\end{aligned}$$

Time dependent coordinate transform

$$y_{1} = \sigma(e^{-1t} \star e^{-1t} \star w_{1} + 2e^{-1t} \star w_{1}) - X_{2} + X_{1} + Y_{1}$$

$$y_{2} = -\sigma e^{-1t} \star w_{1} + X_{2} + Y_{2}$$

$$x_{1} = \sigma e^{-1t} \star w_{1} + X_{1} - Y_{2}$$

$$x_{2} = X_{2}$$

Result normal form DEs

$$\begin{aligned} \dot{Y}_1 &= -Y_2 - Y_1 \\ \dot{Y}_2 &= -Y_2 \\ \dot{X}_1 &= -\sigma w_1 + X_2 \\ \dot{X}_2 &= \sigma w_1 \end{aligned}$$

13.2 Heat exchanger with quadratic reaction

Expand advection-reaction-exchange in a heat exchanger in powers of $(x-X)^n/n!$. The reaction is some quadratic that should generate Burgers' equation model. With Taylor Remainder Theorem closing the problem in terms of unknown functions which here are represented by the non-autonomous forcing w_i . Note that $y(j) = d_{j-1}$ and $x(j) = c_{j-1}$. Also $w(1) = 3d_{2x}$ and $w(2) = 3c_{2x}$ and evaluate at intensity $\sigma = 1$.

The procedure **stonormalform** is already loaded. Write a message saying we are now analysing the next system.

190 write "**** with quadratic reaction ****";

Execute the construction of a normal form for this system

```
191 stonormalform(
        {y(2)-x(1)*y(1)},
192
          y(3)-x(1)*y(2)-x(2)*y(1),
193
          small*w(1)-x(1)*y(3)-2*x(2)*y(2)-x(3)*y(1) },
194
        \{-y(1)+x(2)-(x(1)^2+y(1)^2)/2,
195
          -y(2)+x(3)-x(1)*x(2)-y(1)*y(2),
196
          -y(3)+small*w(2)-x(2)^2-x(1)*x(3)-y(2)^2-y(1)*y(3)},
197
       { },
198
199
       3);
200 end;
```

We could divide the off-diagonal linear terms by small (and remove the multiplication of forcing w), and the algorithm still converges, albeit in more iterations. The resulting asymptotic expressions then do not assume that x derivatives are 'successively smaller'. The following uses the default scaling which corresponds to successively smaller x-derivatives provided I also multiply the forcing by small.

Specified dynamical system

$$\dot{x}_1 = \varepsilon \left(-x_1 y_1 + y_2 \right)
\dot{x}_2 = \varepsilon \left(-x_2 y_1 - x_1 y_2 + y_3 \right)
\dot{x}_3 = \sigma \varepsilon w_1 + \varepsilon \left(-x_3 y_1 - 2x_2 y_2 - x_1 y_3 \right)
\dot{y}_1 = \varepsilon \left(x_2 - 1/2x_1^2 - 1/2y_1^2 \right) - y_1
\dot{y}_2 = \varepsilon \left(x_3 - x_2 x_1 - y_2 y_1 \right) - y_2
\dot{y}_3 = \sigma \varepsilon w_2 + \varepsilon \left(-x_3 x_1 - x_2^2 - y_3 y_1 - y_2^2 \right) - y_3$$

Time dependent coordinate transform

$$y_{1} = \varepsilon (X_{2} - 1/2X_{1}^{2} + 1/2Y_{1}^{2}) + Y_{1}$$

$$y_{2} = \varepsilon (X_{3} - X_{2}X_{1} + Y_{2}Y_{1}) + Y_{2}$$

$$y_{3} = \sigma \varepsilon e^{-1t} \star w_{2} + \varepsilon (-X_{3}X_{1} - X_{2}^{2} + Y_{3}Y_{1} + Y_{2}^{2}) + Y_{3}$$

$$x_{1} = \varepsilon (X_{1}Y_{1} - Y_{2}) + X_{1}$$

$$x_{2} = \varepsilon (X_{2}Y_{1} + X_{1}Y_{2} - Y_{3}) + X_{2}$$

$$x_{3} = \varepsilon (X_{3}Y_{1} + 2X_{2}Y_{2} + X_{1}Y_{3}) + X_{3}$$

Result normal form DEs

$$\dot{Y}_1 = \varepsilon^2 \left(-\frac{1}{2} X_1^2 Y_1 + 2X_1 Y_2 - Y_3 \right) - Y_1
\dot{Y}_2 = \varepsilon^2 \left(-\frac{X_2 X_1 Y_1 + 2X_2 Y_2 - \frac{1}{2} X_1^2 Y_2 + 2X_1 Y_3}{2X_1 Y_1 + \varepsilon^2 \left(-\frac{X_3 X_1 Y_1 - X_3 Y_2 - X_2^2 Y_1 - 2X_2 X_1 Y_2 + X_2 Y_3 - \frac{1}{2} X_1^2 Y_3 \right) - Y_3}
\dot{X}_1 = \varepsilon^2 \left(X_3 - 2X_2 X_1 + \frac{1}{2} X_1^3 \right)$$

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$$\dot{X}_2 = \sigma \varepsilon^2 w_2 + \varepsilon^2 \left(-2X_3 X_1 - 2X_2^2 + 3/2X_2 X_1^2 \right)$$

$$\dot{X}_3 = -\sigma \varepsilon^2 w_2 X_1 + \sigma \varepsilon w_1 + \varepsilon^2 \left(-3X_3 X_2 + 3/2X_3 X_1^2 + 3X_2^2 X_1 \right)$$

Hmmm, looks like this generates the slowly varying model that

$$\frac{\partial C}{\partial t} \approx \frac{\partial^2 C}{\partial x^2} - 2C \frac{\partial C}{\partial x} + \frac{1}{2}C^3.$$

Interestingly there is an extra factor of two in the nonlinear advection, and a net cubic reaction.

References

Chao, X. & Roberts, A. J. (1996), 'On the low-dimensional modelling of Stratonovich stochastic differential equations', *Physica A* **225**, 62–80.

Fateman, R. (2003), 'Comparing the speed of programs for sparse polynomial multiplication', *ACM SIGSAM Bulletin* **37**(1), 4–15. http://www.cs.berkeley.edu/~fateman/papers/fastmult.pdf

Majda, A., Timofeyev, I. & Vanden-Eijnden, E. (2002), 'A priori tests of a stochastic mode reduction strategy', *Physica D* **170**, 206–252.

Monahan, A. H. & Culina, J. (2011), 'Stochastic averaging of idealized climate models', *Journal of Climate* **24**(12), 3068–3088.

Pavliotis, G. A. & Stuart, A. M. (2008), Multiscale methods: averaging and homogenization, Vol. 53 of Texts in Applied Mathematics, Springer.

Potzsche, C. & Rasmussen, M. (2006), 'Taylor approximation of integral manifolds', *Journal of Dynamics and Differential Equations* **18**, 427–460.

Roberts, A. J. (2008), 'Normal form transforms separate slow and fast modes in stochastic dynamical systems', *Physica A* **387**, 12–38.

Roberts, A. J. (2015a), 'Macroscale, slowly varying, models emerge from the microscale dynamics in long thin domains', *IMA Journal of Applied Mathematics* **80**(5), 1492–1518.

Roberts, A. J. (2015b), Model emergent dynamics in complex systems, SIAM, Philadelphia.

http://bookstore.siam.org/mm20/

Sun, X., Kan, X. & Duan, J. (2011), Approximation of invariant foliations for stochastic dynamical systems, Technical report, Illinois Institute of Technology.