

149(1) : Equivalence of ECE Metric and Gravitational Metric  
for the Precessing Ellipse.

In the XY plane defined by:

$$dz^2 = 0 \quad (1)$$

The ECE metric is:

$$d\tau^2 = dt^2 - \frac{1}{c^2} (dr^2 + r^2 d\phi^2) \quad (2)$$

and the gravitational metric is:

$$d\tau^2 = x dt^2 - \frac{1}{c^2} \left( \frac{dr^2}{x} + r^2 d\phi^2 \right) \quad (3)$$

where

$$x = 1 - \frac{r_0}{r} = 1 - \frac{2mG}{c^2 r} \quad (4)$$

The two metrics are the same if:

$$dt^2 - \frac{1}{c^2} dr^2 = x dt^2 - \frac{1}{xc^2} dr^2 \quad (5)$$

$$\text{i.e. } (1-x) dt^2 = \frac{1}{c^2} dr^2 \left( 1 - \frac{1}{x} \right) \quad (6)$$

$$\frac{1}{c^2} \left( \frac{dr}{dt} \right)^2 = - \left( 1 - \frac{2mG}{c^2 r} \right) \quad (7)$$

$$\text{or } \frac{1}{c^2} \left( \frac{dr}{dt} \right)^2 + 1 = \frac{2mG}{c^2 r}, \quad (8)$$

$$\left( \frac{dr}{dt} \right)^2 = \left( \frac{a^2}{1+a^2} \right) v^2, \quad (9)$$

$$a = \frac{y \times r}{d} \sin(\gamma \phi) \quad (10)$$

$$v = |v| = \left| \frac{dr}{dt} \right| \quad (11)$$

$$dr \cdot dv = dr^2 + r^2 d\phi^2 \quad (12)$$

Therefore:

$$\frac{1}{c^2} \left( \frac{a^2}{1+a^2} \right) v^2 + 1 = \frac{2mg}{c^2 r} - (13)$$

i.e.

$$\frac{1}{2} \left( v^2 \left( \frac{a^2}{1+a^2} \right) + c^2 \right) = \frac{mg}{r} = -V - (14)$$

Multiply both sides by the mass of the test particle  $m$ ,

To find:

$$\frac{1}{2} \left( \frac{a^2}{1+a^2} \right) mv^2 + \frac{1}{2} mc^2 = \frac{m \frac{mg}{r}}{r} = -U - (15)$$

where

$$v^2 = \left( \frac{dr}{dt} \right)^2 + r^2 \left( \frac{d\phi}{dt} \right)^2 - (16)$$

Here  $U$  is the potential energy:

$$U = mV - (17)$$

where  $V$  is the gravitational potential

Denote the kinetic energy by:

$$T = \frac{1}{2} \left( \frac{a^2}{1+a^2} \right) mv^2 + \frac{1}{2} mc^2 - (18)$$

To find that

$$H = T + U = 0 - (19)$$

here  $H$  is the Hamiltonian.

This means that the total energy of the orbit

is zero.

By transforming the gravitational metric (3) into the ECE Minkowski metric (2) the gravitational potential energy has been transformed into pure kinetic energy, defined by eq. (18). Thus eq. (2) is a pure kinetic description of the orbit, which is observed experimentally to be the precessing ellipse:

$$r = \frac{d}{1 + \epsilon \cos(\gamma\phi)} \quad (20)$$

In the limit:

$$\gamma \rightarrow 1 \quad (21)$$

the ellipse stops precessing, and the perihelia stops advancing.

In the limit

$$\epsilon \rightarrow 0 \quad (22)$$

the orbit approaches the circular orbit, i.e.

$$a \rightarrow \infty \quad (23)$$

and

$$U + \frac{1}{2}mc^2 \rightarrow 0 \quad (24)$$

This means that in the limit of a circular orbit:

$$\frac{mg}{r} \rightarrow \frac{1}{2}c^2, \quad (25)$$

$$\boxed{r \rightarrow \frac{2mg}{c^2} = r_0} \quad (26)$$

In the standard literature  $r_0$  is incorrectly referred to as the Schwarzschild radius. However, temporary scholarship shows conclusively that

4) Schwarzschild did not infer the metric due to incorrectly attributed to him. In ECE physics this is known as the gravitational metric, eq. (3). Both eqs. (2) and (3) are solutions of the ECE orbital equations of UFT III.

## Results

For the precessing elliptical orbit, (the relativistic Keplerian orbit):

$$H = T + u = 0 \quad -(27)$$

where

$$T = \frac{m}{2} \left( \frac{a^2}{1+a^2} \right) \left( \left( \frac{dr}{dt} \right)^2 + r^2 \left( \frac{d\phi}{dt} \right)^2 \right) + \frac{1}{2} mc^2, \quad -(28)$$

$$u = -\frac{mMb}{r}. \quad -(29)$$

The total linear velocity in a plane ( $x_1$ ) is:

$$\sqrt{v^2} = \left( \frac{dr}{dt} \right)^2 + r^2 \left( \frac{d\phi}{dt} \right)^2 \quad -(30)$$

$$\text{and } a = \frac{y \times r}{\partial d} \sin(y\phi) \quad -(31)$$

## Numerical Calculation

Use: Mass of earth  $m = 5.98 \times 10^{24}$  kg

Mass of sun  $M = 1.99 \times 10^{30}$  kg

$$G = 6.67 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}$$

5) Earth to sun distance  $r = 1.50 \times 10^{11} \text{ m}$   
 Speed of light  $c \sim 3 \times 10^8 \text{ m s}^{-1}$   
 Orbital velocity  $v$  of earth around sun  $= 2.98 \times 10^4 \text{ m s}^{-1}$

Use close data in the equation:

$$\frac{1}{2} \left( \frac{a^2}{1+a^2} \right) mv^2 + \frac{1}{2} mc^2 = \frac{mMg}{r} \quad (32)$$

$$\text{Assume: } \frac{a^2}{1+a^2} \sim 1 \quad (33)$$

in a rough approximation. Thus, if  $r$  is the  
earth to sun distance to be calculated  
 $\frac{mMg}{r} = 7.937 \times 10^{44} / r$  joules

$$\text{and } \frac{1}{2} mc^2 = 2.691 \times 10^{33} \text{ joules.}$$

$$\frac{1}{2} mv^2 = 2.655 \times 10^{33} \text{ joules.} \quad 44$$

$$\text{So: } 2.655 \times 10^{33} \left( \frac{a^2}{1+a^2} \right) + 2.691 \times 10^{33} = \frac{7.937 \times 10^{44}}{r} \quad (34)$$

$$\text{So using eq. (33)} \quad r \sim \frac{2mG}{c^2} = 2.95 \times 10^3 \text{ m} \quad (35)$$

To an excellent approximation.

Eq. (32) is:

$$! \frac{m\mathbf{h}^2}{r} - \frac{1}{2} mc^2 = \frac{1}{2} \left( \frac{a^2}{1+a^2} \right) m v^2 - (36)$$

where

$$f(\epsilon, \phi) = \frac{a^2}{1+a^2} = \frac{\left( \frac{y \epsilon r \sin(y\phi)}{d} \right)^2}{1 + \left( \frac{y \epsilon r \sin(y\phi)}{d} \right)^2} - (37)$$

$$\text{It is seen that } f_r \sim 2.95 \times 10^{-3} \text{ m} - (38)$$

then

$$f(\epsilon, \phi) \sim 1 - (39)$$

self consistently.

### Graphical Work

It may be interesting to graph eq. (37)  
for various  $\epsilon$ ,  $d$  and  $y$  of process 25 ellipse.

## 7) Overall Conclusion

The experimentally observed orbit is eq. (20), and is described straightforwardly by the metric (2) of Minkowski spacetime using  $d\tau/dt$  from eq. (20). There are problems with the received opinion (3) because when

$$r_0 = r - (40)$$

there is a singularity. The latter is seen severely misinterpreted as "Big Bang". Data show that the latter does not exist.

This note shows that to much wanted, but deeply flawed, eq. (3) is equivalent to the observational eq. (2) only if:

$$r \leq r_0 - (41)$$

This analysis reveals another severe limitation that eq. (3) is well known by now that Einstein field equation was in correct connection. Schwarzschild did not derive eq. (3), which is merely a meaningless solution of a incorrect equation. It produces the observed function (20) only in the limit (41). So eq. (3) is not generally applicable. Eq. (2) on the other hand is generally applicable to all observed orbits.

149(2) : Equations of Motion from the Minkowski Metric, ECE  
Equation of Motion.

The Minkowski metric in cylindrical polar coordinates is:

$$ds^2 = c^2 d\tau^2 = c^2 dt^2 - dr \cdot dr \quad (1)$$

where

$$\begin{aligned} dr \cdot dr &= dr^2 + r^2 d\phi^2 + dz^2 \\ &= dx^2 + dy^2 + dz^2 \end{aligned} \quad (2)$$

The total linear velocity is defined as:

$$v = \frac{dr}{dt} \quad (3)$$

Therefore:  $c^2 d\tau^2 = c^2 dt^2 - v^2 dt^2 \quad (4)$

$$= (c^2 - v^2) dt^2 \quad (5)$$

so  $d\tau^2 = \left(1 - \frac{v^2}{c^2}\right) dt^2 \quad (6)$

The infinitesimal of proper time is:

$$d\tau = \left(1 - \frac{v^2}{c^2}\right)^{1/2} dt \quad (7)$$

and

$$dt = \gamma d\tau \quad (8)$$

where  $\gamma = \left(1 - \frac{v^2}{c^2}\right)^{-1/2} \quad (9)$

If we restrict attention to the  $xy$  plane:

$$dz = 0 \quad (10)$$

The rest energy is defined as:

$$E_0 = mc^2 = m \left( \frac{ds}{d\tau} \right)^2 = mg_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \quad (11)$$

where  $S$  is the action.

Thus:

$$2) E_0 = mc^2 \left( \frac{dt}{d\tau} \right)^2 - m \left( \frac{dr}{d\tau} \right)^2 - mr^2 \left( \frac{d\phi}{d\tau} \right)^2 - (12)$$

The Lagrange equation is:

$$\frac{d}{d\tau} \left( \frac{\partial E_0}{\partial \dot{x}^\mu} \right) = \frac{\partial E_0}{\partial x^\mu} = 0 - (13)$$

so

$$\frac{d}{d\tau} \left( mr^2 \frac{d\phi}{d\tau} \right) = 0 - (14)$$

$$\frac{d}{d\tau} \left( m \frac{dr}{d\tau} \right) = 0 - (15)$$

$$\frac{d}{d\tau} \left( mc^2 \frac{dt}{d\tau} \right) = 0 - (16)$$

The constants of motion from eqns (14) to (16) are:

$$E = mc^2 \frac{dt}{d\tau} = \gamma mc^2 - (17)$$

$$L = mr^2 \frac{d\phi}{d\tau} = \gamma mr^2 \frac{d\phi}{dt} - (18)$$

$$P = m \frac{dr}{d\tau} = \gamma m \frac{dr}{dt} - (19)$$

$$\text{So: } \frac{P^2}{2m} = \frac{E^2}{mc^2} - mc^2 - \frac{L^2}{mr^2} - (20)$$

$$\boxed{(Y+1)\tau = (Y+1)(Y-1)mc^2 = \frac{P_m^2}{2m} + \frac{L^2}{mr^2}} - (21)$$

where

$$E = T + E_0 - (22)$$

Here  $E$  is the total energy,  $T$  is the relativistic energy,  $L$  is the relativistic angular momentum, and  $P$  is the relativistic momentum. The orbit is purely hyperbolic, and there is no concept of potential energy or attractive

∴ from Eq. (21) "equivalent to the description:

$$\frac{dr}{dt} \cdot \frac{d\phi}{dt} = c^2 (dt^2 - d\tau^2) \quad (23)$$

$$= v^2 dt^2$$

i.e.

$$mv^2 = m \left( \left( \frac{dr}{dt} \right)^2 + r^2 \left( \frac{d\phi}{dt} \right)^2 \right) \quad (24)$$

where

$$\frac{d\phi}{dt} = \frac{d\phi}{dr} \frac{dr}{dt} \quad (25)$$

so

$$mv^2 = m \left( \frac{dr}{dt} \right)^2 \left( 1 + r^2 \left( \frac{d\phi}{dr} \right)^2 \right) \quad (26)$$

For a precessing elliptical orbit, it is found by experimental observation that:

$$\frac{d\phi}{dr} = \frac{d}{y \epsilon r \sin(\gamma\phi)} \quad (27)$$

so

$$mv^2 = m \left( \frac{dr}{dt} \right) \left( 1 + \left( \frac{d}{y \epsilon r \sin(\gamma\phi)} \right)^2 \right) \quad (28)$$

In the non-relativistic limit:

$$(y^2 - 1)mc^2 \rightarrow \left( \left( 1 - \frac{v^2}{c^2} \right)^{-1} - 1 \right) mc^2$$

$$\sim mv^2 \quad (29)$$

so the non-relativistic limit of eq. (21) is

$$4) \quad m v^2 = \frac{p^2}{m} + \frac{L^2}{mr^3} \quad - (30)$$

So

$$\boxed{T = \frac{1}{2}mv^2 = \frac{p^2}{2m} + \frac{L^2}{2mr^3}} \quad - (31)$$

In this limit  $v = \left(1 - \frac{v^2}{c^2}\right)^{-1/2} \rightarrow 1 \quad - (32)$

and  $L \rightarrow mr^2 \frac{d\phi}{dt}, \quad - (33)$

$$p \rightarrow m \frac{dr}{dt} \quad - (34)$$

In the conventional non-relativistic treatment of orbits (e.g. Maria and Thonta) the second term on the right hand side of eq. (31) is called the centripetal kinetic energy. However, it is part of the Newtonian treatment, therefore the conventional treatment is well known. The ECE treatment fully describes the orbit without using the concept of potential energy or the derived concept of centrifugal force.

$$F_C := - \frac{\partial U_C}{\partial r} = \frac{L^2}{mr^3} = mr \left( \frac{d\phi}{dt} \right)^2 \quad - (35)$$

5) Eq. (20) is:

$$m \left( \frac{dr}{d\tau} \right)^2 = \frac{p^2}{m} = \frac{E^2}{mc^2} - mc^2 - \frac{L^2}{mr^2} \quad -(36)$$

and can be compared w/ the well known result from  
 the gravitational metric:

$$m \left( \frac{dr}{d\tau} \right)^2 = \frac{E^2}{mc^2} - mc^2 - \frac{L^2}{mr^2} + \frac{2mMg}{r} + \frac{ML^2}{c^2 mr^3} \quad -(37)$$

Both eqs. (36) and (37) descend of same proceeding  
 ellipse def'd by eq. (27), but eq. (36) is  
 preferred by Planck's Razor, it is simpler and  
 much clearer. So the last two terms of eq.  
 (37) can be removed by using the limit:

$$r \rightarrow \infty, \quad -(38)$$

in which case eq. (37) reduces to the purely kinetic  
 eq. (36). Conventionally the last two terms of  
 eq. (37) are the "relativistically corrected potential  
 energy of attraction". The concept of attraction  
 and "corrected" identification (35) shows clearly  
 and "unless eq. (31)  
 energy is kinetic energy".

all is kinetic

For a circular orbit:

$$\frac{dr}{d\tau} = 0 \quad -(39)$$

$$\therefore \text{so } (\gamma^2 - 1)mc^2 = \frac{L^2}{mr^2} - (40)$$

In the limit:

$$v \ll c - (41)$$

eq. (40) becomes:

$$\boxed{\frac{1}{2}mv^2 = \frac{L^2}{2mr^2}} - (42)$$

i.e. the two components of the kinetic energy are the same in a circular orbit.

Note carefully that in the description, the orbit (of any observable type) is always described by the simplest possible solution of the orbital equation, the Minkowski metric. There is no concept of attractive and repulsive force, and no concept of potential energy. The orbit is the geodesic or metric. Light orbits in a null geodesic:

$$ds^2 = 0 - (43)$$

$$\text{i.e. } v = c, - (44)$$

so the orbit of light from eq. (26) is:

$$\boxed{m \left( \frac{dr}{dt} \right)^2 \left( 1 + r^2 \left( \frac{d\phi}{dr} \right)^2 \right) = mc^2 = E_0} - (45)$$

7) In the received opinion, originating w/ Kepler's idea of "force", eq. (4) is:

$$\underline{F} = \underline{0} = \underline{m g} + m r \left( \frac{d\phi}{dt} \right)^2 \underline{k} \quad -(4b)$$

so if attractive, Newtonian, inward  $\underline{mg}$  is balanced exactly by the outward, non-Newtonian, repulsive  $F_c \underline{k}$ . Newton's first law  $\rightarrow$  that  $m$  travels in "a straight line" unless acted upon by "force". In the ECE treatment  $m$  travels in an orbit which has kinetic energy only. The "straight line" of Newton is replaced by the orbit, which is metric.

Newtonian metric is:

$$ds^2 = dx^2 + dy^2 + dz^2 \quad ] - (47)$$

$$= dr^2 + r^2 d\phi^2 + dz^2$$

and as is well known, exists only in three dimensional space, not spacetime.

If the quantity  $d\phi/dt$  vanishes, then in eq. (38), in the non-relativistic limit:

$$m v^2 = m \left( \frac{dr}{dt} \right)^2 \quad - (48)$$

and the particle  $m$  has a velocity:

$$v = \frac{dr}{dt} \quad - (49)$$

and an acceleration  $\frac{d^2r}{dt^2}$  from the Lagrange equation.

8) In general orbit, the total kinetic energy in  
the limit  $v \ll c$  is:

$$T = \frac{1}{2}mv^2 = \frac{1}{2}m\left(\frac{dr}{dt}\right)^2 + \frac{L^2}{2mr^2} \quad -(50)$$

$$\text{so } \frac{dT}{dt} = m \frac{d^2r}{dt^2} + \frac{d}{dt}\left(\frac{L^2}{2mr^2}\right)$$

$$\rightarrow m \frac{d^2r}{dt^2} \quad -(51)$$

$$L \rightarrow 0, \quad -(52)$$

as  
so the particle  $m$  falls towards  $M$  at with  
an acceleration:

$$\boxed{\frac{d^2r}{dt^2} = \frac{1}{m} \frac{dT}{dt}} \quad -(53)$$

In the received opinion (second Newton law)  
the change of kinetic energy with time is described  
as "force", which is defined in the Newton's second  
law as mass multiplied by acceleration. In  
ECE description, which is simpler and clearer,  
 $F$  is defined purely by the metric itself.

49(3): Re Conventional Description of Orbits from GR  
Gravitational Metric.

The gravitational metric is:

$$ds^2 = c^2 d\tau^2 = \left(1 - \frac{r_s}{r}\right) c^2 dt^2 - \left(1 - \frac{r_s}{r}\right)^{-1} dr^2 - r^2 d\phi^2 - dz^2 \quad (1)$$

in cylindrical polar coordinates, where:

$$r_s = \frac{2GM}{c^2 r} \quad (2)$$

It is claimed incorrectly that it was derived by Schwarzschild, for the Einstein field equation. The latter is now known to be incorrect. However, eqn (1) is formally a possible solution of the ECE Orbital Theorem, one of an infinite set of possible solutions of the Orbital Theorem. So formally the following method derives from the ECE Orbital Theorem. In fact, received opinion, it is described as "the relativistic Kepler problem".

Define:

$$T := \frac{1}{2} mc^2 = \frac{1}{2} m \left( \frac{ds}{d\tau} \right)^2 = \frac{m}{2} g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \quad (3)$$

where  $S$  is the action and:

$$c^2 d\tau^2 = g_{\mu\nu} dx^\mu dx^\nu \quad (4)$$

so:

$$g_{00} = 1 - \frac{r_s}{r}, \quad g_{11} = -\left(1 - \frac{r_s}{r}\right)^{-1}, \quad g_{22} = -1, \quad g_{33} = -1, \\ dx^0 = c dt, \quad dx^1 = dr, \quad dx^2 = r d\phi, \quad dx^3 = dz. \quad (6)$$

So:

$$\frac{dT}{m} = c^2 = g_{00} \left( \frac{dx^0}{d\tau} \right)^2 + g_{11} \left( \frac{dx^1}{d\tau} \right)^2 + g_{22} \left( \frac{dx^2}{d\tau} \right)^2 + g_{33} \left( \frac{dx^3}{d\tau} \right)^2 \quad (7) \\ = \left(1 - \frac{r_s}{r}\right) c^2 \left( \frac{dt}{d\tau} \right)^2 - \left(1 - \frac{r_s}{r}\right)^{-1} \left( \frac{dr}{d\tau} \right)^2 - r^2 \left( \frac{d\phi}{d\tau} \right)^2 - \left( \frac{dz}{d\tau} \right)^2$$

The Lagrange equation is:

$$\frac{d}{d\tau} \left( \frac{\partial T}{\partial \dot{x}^n} \right) = \frac{\partial T}{\partial x^n} - (8)$$

$$\frac{d}{d\tau} \left( m r^2 \frac{d\phi}{d\tau} \right) = 0 - (9)$$

$$\frac{d}{d\tau} \left( m \left(1 - \frac{r_s}{r}\right) \frac{dt}{d\tau} \right) = 0 - (10)$$

$$\frac{d}{d\tau} \left( m \left(1 - \frac{r_s}{r}\right)^{-1} \frac{dr}{d\tau} \right) = 0 - (11)$$

Eqs (a) to (11) give the constants of motion:

$$E = mc^2 \left( 1 - \frac{r_s}{r} \right) \frac{dt}{d\tau} - (12)$$

$$L = m c^2 \frac{d\phi}{d\tau} - (13)$$

$$P = m \left(1 - \frac{r_s}{r}\right)^{-1} \frac{dr}{d\tau} - (14)$$

These are the total energy  $E$ , the momentum  $P$  and the angular momentum  $L$ .

In the plane:

$$dz = 0 - (15)$$

we have:

$$\begin{aligned} \left(1 - \frac{r_s}{r}\right) T &= mc^2 \left(1 - \frac{r_s}{r}\right)^2 \left(\frac{dt}{d\tau}\right)^2 - \frac{m}{2} \left(\frac{dr}{d\tau}\right)^2 - m \left(1 - \frac{r_s}{r}\right) r^2 \left(\frac{d\phi}{d\tau}\right)^2 \\ &= \frac{E^2}{mc^2} - m \left(\frac{dr}{d\tau}\right)^2 - \left(1 - \frac{r_s}{r}\right) \frac{L^2}{mr^2} - (16) \end{aligned}$$

Therefore:  $m \left(\frac{dr}{d\tau}\right)^2 = \frac{E^2}{mc^2} - \left(1 - \frac{r_s}{r}\right) T - \left(1 - \frac{r_s}{r}\right) \frac{L^2}{mr^2}$

$$3) \text{ i.e. } \frac{1}{2} m \left( \frac{dr}{d\tau} \right)^2 = \frac{1}{2} \frac{E^2}{mc^2} - \frac{1}{2} \left( 1 - \frac{rs}{r} \right) \left( mc^2 - \frac{L^2}{mr^2} \right) \\ = \left( \frac{E^2}{2mc^2} - \frac{1}{2} mc^2 \right) + \left( \frac{mMG}{r} - \frac{L^2}{2mr^2} + \frac{6ML^2}{mc^2 r^2} \right) - (17)$$

This is eq. (30) of note 149(2). It gives the same precessing elliptical orbit as the ECE Mihoski theory.

$$m \left( \frac{dr}{d\tau} \right)^2 = (\gamma - 1) mc^2 - \frac{L^2}{mr^2} - (18)$$

$$m \left( \frac{dr}{d\tau} \right)^2 = (\gamma - 1) mc^2 - \frac{L^2}{2mr^2} - (19)$$

$$\text{or } \frac{1}{2} m \left( \frac{dr}{d\tau} \right)^2 = \frac{1}{2} (\gamma - 1) mc^2 - \frac{L^2}{2mr^2}$$

This is because the Mihoski and gravitational metrics are soft solution of the same equation, the ECE Orbital Theorem of ~~UFT~~ <sup>HT</sup> for spherically symmetric spacetime. So the two solutions (17) and (19) contain the same information - it precesses elliptical orbit. However, eq. (19) can be generalized to any orbit, using observation, and but eq. (17) cannot, it always gives a precessing elliptical orbit. Eq. (17) is one that uses the Nottie constant which does not appear in eq. (19). Eq. (17)

reduces to eq. (19) when:

$$r \rightarrow \infty - (20)$$

$$\frac{rs}{r} \rightarrow 0 - (21)$$

4) and

$$E \rightarrow \gamma m c^2 = mc^2 \frac{dt}{d\tau} \quad (22)$$
$$P \rightarrow \gamma m \frac{dr}{dt} = m \frac{dr}{d\tau} \quad (23)$$

As shown in note 14a(2), the concepts of force and potential energy are not needed. Neither the first nor the second law of Newton are used.

The Einsteinian description is not used because the covariant must be unsymmetric, not symmetric as used by Einstein. Dark matter is not used. The concepts in these notes are radically new, but the calculations are relatively straightforward.

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1) 149(4): The Equations of Orbit

Start with the fully relativistic definition:

$$v^2 = \left(\frac{dr}{dt}\right)^2 + c^2 \left(\frac{d\phi}{dt}\right)^2 - (1)$$

and use the fact that

$$p^2 = \gamma^2 m^2 \left(\frac{dr}{dt}\right)^2 - (2)$$

is a constant of motion:

$$\gamma^2 \left(\frac{dr}{dt}\right)^2 = \left(\frac{p}{m}\right)^2 := A - (3)$$

So

$$\gamma^2 v^2 = A + \gamma^2 r^2 \left(\frac{d\phi}{dt}\right)^2 - (4)$$

$$= A + \gamma^2 r^2 \left(\frac{dp}{dr}\right)^2 \left(\frac{dr}{dt}\right)^2$$

$$\boxed{\gamma^2 v^2 = A \left(1 + r^2 \left(\frac{d\phi}{dr}\right)^2\right)} - (5)$$

The equation of orbits is therefore:

$$1 + r^2 \left(\frac{d\phi}{dr}\right)^2 = \frac{\gamma^2 v^2}{A} - (6)$$

i.e.

$$\boxed{\frac{d\phi}{dr} = \frac{1}{r} \left(\frac{\gamma^2 v^2}{A} - 1\right)^{1/2}} - (7)$$

Precessing Ellipse

By observation:

$$2) \frac{d\phi}{dr} = \frac{d}{y^2 \sin(y\phi)} \cdot \frac{1}{r} - (8)$$

$$\therefore = \frac{b}{r}$$

$$\text{so } b = \left( \frac{y^2 v^2}{A} - 1 \right)^{1/2} - (9)$$

$$\underline{\text{Whirlpool Galaxy}}$$

$$\frac{dr}{d\phi} = b_1 r - (10)$$

$$\frac{dr}{d\phi} = r \exp(b_1 \phi(t)) - (11)$$

$$\text{where } r(t) = \exp(b_1 \phi(t)) - (12)$$

$$\text{so } \frac{1}{b_1} = \left( \frac{y^2 v^2}{A} - 1 \right)^{1/2} - (13)$$

$$\text{i.e. } 1 + \frac{1}{b_1^2} = \frac{y^2}{A} v^2 - (14)$$

$$v = \frac{1}{r} \left( A \left( 1 + \frac{1}{b_1^2} \right) \right)^{1/2} - (15)$$

In the limit  $r \rightarrow 1$ ,  $v = \underline{\text{constant}}$

Similarly, for precessing ellipse:

$$v = \frac{1}{r} \left( A \left( 1 + b^2 \right) \right)^{1/2} - (16)$$

$$v = \frac{A^{1/2}}{r} \cdot \left( 1 + \left( \frac{d}{y^2 \sin(y\phi)} \right)^2 \right)^{1/2} - (16)$$

$$3) \text{ For the ellipse: } \left. \begin{array}{l} y \rightarrow 1 \\ y \rightarrow 1 \end{array} \right\} - (17)$$

$$\text{so } v^2 = A \left( 1 + r^2 \left( \frac{d\phi}{dr} \right)^2 \right) - (18)$$

$$= \left( \frac{2\pi}{\tau} \right)^2 a^3 \left( \frac{2}{r} - \frac{1}{a} \right) - (19)$$

from Kepler's equation of the ellipse:

$$\frac{(x + a\epsilon)^2}{a^2} + \frac{y^2}{b^2} = 1 - (20)$$

and where  $2\pi t/\tau$  is the mean anomaly.

1) 149(5), Another Form of the Equation of Orbits.

In note 149(2) it was shown that a purely kinetic description of an orbit is, from the Minkowski metric:

$$(1 - \frac{v^2}{c^2}) m c^2 = \frac{p^2}{m} + \frac{L^2}{m r^2}, \quad (1)$$

where:

$$p = m \frac{dr}{dt} = \gamma m v r, \quad (2)$$

$$L = m r^2 \frac{d\phi}{dt} = \gamma m r^2 \frac{d\phi}{dt} \quad (3)$$

In the non-relativistic limit, eq. (1) becomes:

$$T = \frac{1}{2} m v^2 = \frac{m}{2} \left( \left( \frac{dr}{dt} \right)^2 + r^2 \left( \frac{d\phi}{dt} \right)^2 \right) \quad (4)$$

$$\boxed{T = \frac{1}{2} m ( \dot{r}^2 + r^2 \dot{\phi}^2 )} \quad (5)$$

In the purely kinetic theory of orbits there is no potential energy, all is kinetic energy. Therefore the Hamiltonian and Lagrangian are:

$$H = L = T \quad (6)$$

How is it possible to describe an orbit without the traditional:

$$U = -\frac{m M G}{r} \quad (7)$$

$$F = -\frac{\partial U}{\partial r} = -\frac{m M G}{r^2} \quad (8)$$

of the inverse square law?

To answer this question set up the lagrangian:

$$L = \frac{1}{2} m ( \dot{r}^2 + r^2 \dot{\phi}^2 ) \quad (9)$$

instead of the traditional:

$$L = T - U \quad (10)$$

The two La Lagrange equations are:

$$\frac{\partial L}{\partial \dot{\phi}} = \frac{d}{dt} \frac{\partial L}{\partial \dot{\phi}} \quad - (11)$$

$$\frac{\partial L}{\partial r} = \frac{d}{dt} \frac{\partial L}{\partial \dot{r}} \quad - (12)$$

with Q Lagrangian (9). Eq (11) gives:

$$L = m r^2 \dot{\phi}^2 = \text{constant} \quad - (13)$$

and eq (12) gives:

$$m \ddot{r} = m r^2 \dot{\phi}^2 \quad - (14)$$

$$mr = \boxed{\frac{L^2}{m^2}} \quad - (15)$$

"Or conjugate

To the traditional point of view

force of repulsion is

$$F_C = \frac{L^2}{m^2} \quad - (16)$$

"and the Newtonian force of attraction is:

$$F_A = m \ddot{r} \quad - (17)$$

so

$$F_A = F_C \quad - (18)$$

It is important

if there is no potential energy. it is a complete description  
to note that eq (15) is a place X Y. This can be  
of the stable orbit in any kind.

3) Eq. (5) is:

$$T = \frac{p^2}{2m} + \frac{L^2}{2mr^2} - (19)$$

value in the non-relativistic limit:

$$\sqrt{p^2/c^2} \ll c \quad \rightarrow (20)$$

we have

$$p = mv \quad \rightarrow (21)$$

$$L = mr^2\dot{\phi} \quad \rightarrow (22)$$

$$\text{Therefore } T = \frac{p^2}{2m} + \frac{L^2}{2mr^2} = \left( \frac{T}{r} \right)^2 - \frac{p^2}{2m} - (23)$$

so in eq. (15):

$$mr\ddot{r} = \frac{L^2}{mr^3} - \frac{2}{r} \left( T - \frac{p^2}{2m} \right) - (24)$$

$$\text{In eq. (24)}: \dot{r}^2 = \left( \frac{d\phi}{dt} \right)^2 = \left( \frac{d\phi}{dr} \right)^2 \left( \frac{dr}{dt} \right)^2 - (25)$$

$$\text{i.e.} \quad \left( \frac{d\phi}{dr} \right)^2 = \dot{\phi}^2 / \dot{r}^2 - (26)$$

$$= \frac{2}{m(\dot{r}\dot{r})^2} \left( T - \frac{p^2}{2m} \right)$$

$$\frac{d\phi}{dr} = \frac{1}{\dot{r}\dot{r}} \left( \frac{2}{m} \left( T - \frac{p^2}{2m} \right) \right)^{1/2} - (27)$$

$$\frac{d\phi}{dr} = \frac{L}{mr^2\dot{r}} - (28)$$

4) In the non-relativistic limit the orbit is observed in astronomy to be an ellipse:

$$r = \frac{d}{1 + e \cos \phi} \quad - (29)$$

$$\text{so } \frac{dr}{d\phi} = \left( \frac{e \sin \phi}{d} \right) r^2 \quad - (30)$$

From eqs. (28) and (30):

$$\frac{d\phi}{dr} = \frac{L}{mr^2 i} = \left( \frac{d}{e \sin \phi} \right) \cdot \frac{1}{r^2} \quad - (31)$$

$$\text{so } L = m \left( \frac{d}{e \sin \phi} \right) \frac{dr}{dt} = mr^2 \frac{d\phi}{dt}. \quad - (32)$$

$$= L \sqrt{\frac{1}{r^2(1-e^2)} + \frac{2e}{r(1-e^2)}} \quad - (33)$$

Now we use  $\frac{dr}{dt} = \frac{dr}{d\phi} \frac{d\phi}{dt}$  which becomes, self-consistently,

$$\text{in eq. (32), } \frac{dr}{d\phi} = \frac{e \sin \phi}{d} \quad - (34)$$

Eq. (28) is:

$$\frac{dr}{d\phi} = \left( \frac{m}{L} \right) r^2 \left( \frac{dr}{dt} \right). \quad - (35)$$

The areal velocity of any orbit is:

$$\frac{dA}{dt} = \frac{L}{2m} = \frac{1}{2} r^2 \dot{\phi} \quad - (36)$$

$$\text{so } \frac{dr}{d\phi} = \frac{1}{2} \frac{dr}{dt} \frac{dt}{dA} \quad r^2 = \frac{1}{2} \frac{dr}{dA} \dot{\phi}^2 \quad - (37)$$

and for any orbit:

5)

$$\frac{dA}{dr} = \frac{\dot{r}^2}{2} \frac{d\phi}{dr} - (38)$$

This is smaller form of the general equation of orbits.

For an ellipse:

$$\begin{aligned} \frac{dA}{dr} &= \frac{1}{2} \left( \frac{d}{e \sin \phi} \right) \frac{1}{r^2} - (39) \\ &= \frac{1}{2} \left( \frac{d}{e \sin \phi} \right) \end{aligned}$$

$$- (40)$$

so

$$\sin \phi = \frac{d}{2e} \frac{dr}{dA} - (40)$$

Summary

1) The purely kinetic description of all orbits in the XY plane is:

$$\begin{aligned} m\ddot{r} &= \frac{L^2}{mr^3} - (41) \\ \frac{dr}{d\phi} &= \frac{mr^2}{L} \left( \frac{dr}{dt} \right) - (42) \\ \frac{dA}{dr} &= \frac{r^2}{2} \frac{d\phi}{dr} - (43) \\ H = \frac{L}{r} &= T - (44) \end{aligned}$$

2) The caventional description

$$F(r) = -\frac{\partial U}{\partial r} = m\ddot{r} - \frac{L^2}{mr^3} - (45)$$

$$H = T + U - (46)$$

$$L = T - U - (47)$$

$$U = -\frac{mM_0}{r} - (48)$$

6) Therefore it's conventional description:

$$m\ddot{r} = \frac{L^2}{mr^3} - \frac{\partial U}{\partial r} = (49)$$

$$\therefore \ddot{r} = \frac{L^2}{mr^3} - \frac{mMG}{r^2} = (50)$$

then we can write the above as  $m\ddot{r} = \frac{L^2}{mr^3} - \frac{mMG}{r^2}$  therefore

$$(49) + (50) = 0 \Rightarrow \frac{L^2}{mr^3} = \frac{mMG}{r^2}$$

The conventional equivalence principle is, therefore:

$$F = mg = m\ddot{r} = \frac{L^2}{mr^3} = -\frac{mMG}{r^2} \quad (51)$$

In the purely kinetic view:

$$\ddot{r} = \frac{L^2}{mr^3} = -\frac{mMG}{r^2} \quad (52)$$

i.e.

$$\frac{d^2r}{dt^2} = -\frac{mG}{r^2} \quad (53)$$

Eq. (53) is the purely kinetic interpretation of the inverse square law:

$$\text{observation of} \quad \frac{d^2r}{dt^2} = \frac{d}{dt} \left( \frac{1}{2} m \left( \frac{dr}{dt} \right)^2 \right) = -\frac{mG}{r^2} \quad (54)$$

$$\frac{d^2r}{dt^2} = \frac{d}{dt} \left( \frac{1}{2} m \left( \frac{dr}{dt} \right)^2 \right) = -\frac{mG}{r^2} \quad (55)$$

where  $T_N = \frac{1}{2} m \left( \frac{dr}{dt} \right)^2 \quad (56)$

is the purely central, a Newtonian, kinetic energy.

## Interpretation

Eq. (41) means that in any orbit in the XY plane,

$$\frac{d\bar{T}_N}{dt} = \frac{\mathbf{L}^2}{mr^3} = F_2 - (57)$$

The change of central kinetic energy with time is counter-balanced by a kinetic  $\mathbf{L}^2/(mr^3)$  deriving from angular momentum. Eq. (57) is a property of spacetime itself, as always in relativity.

The conventional description is:

$$\frac{d\bar{T}_N}{dt} = \frac{\mathbf{L}^2}{mr^3} - \frac{\partial U}{\partial r} = 0 - (58)$$

and artificially introduces the concepts of

$$F_A = - \frac{\partial U}{\partial r} - (59)$$

i.e. the concepts of force and potential energy.

The kinetic description is simpler and preferred by Ockham's Razor. Note that the introduction of  $U$  into the Lagrangian makes no difference to eq. (41) or (43), i.e. to:

$$dA = \frac{1}{2} r^2 d\phi - (60)$$

$$\frac{dA}{dt} = \frac{\mathbf{L}}{mr^2} - (61)$$

which is Kepler's second law (1609). In other words the absence of potential  $U$  makes no

8)

difference to Kepler's Second Law, which was derived from the elliptical orbit of Mars in 1609. Kepler's Second Law is true for all types of orbits, and is independent of any choice of  $u$ . Therefore all orbits are given by:

$$\boxed{\frac{dA}{dr} = \frac{1}{2} r^2 \frac{d\phi}{dr}} \quad - (62)$$

and the concepts of force and potential energy are not needed.

This is another statement of the ECE Principle of Orbits.

149(6): Effect of Inverse Square Term

As shown in previous notes the Minkowski metric:

$$ds^2 = c^2 dt^2 - dr^2 - r^2 d\phi^2 \quad (1)$$

gives:

$$(v^2 - 1)mc^2 = \frac{p_r^2}{m} + \frac{L^2}{mr^2} \quad (2)$$

where

$$T = (v - 1)mc^2 \quad (3)$$

is the relativistic kinetic energy. The relativistic momentum is

$$\underline{p} = \gamma m \frac{d\underline{r}}{dt} \quad (4)$$

where

$$d\underline{r} \cdot d\underline{r} = dr^2 + r^2 d\phi^2 \quad (5)$$

In eq. (2):

$$p_r := \gamma m \frac{dr}{dt} \quad (6)$$

$$L := \gamma m r^2 \frac{d\phi}{dt} \quad (7)$$

and are constraints of motion.

If it is assumed that  $r$  and  $\phi$  are related by the conical sector equation:

$$r = \frac{\kappa}{1 + \epsilon \cos \phi} \quad (8)$$

then

$$\frac{dr}{d\phi} = \left( \frac{\epsilon \sin \phi}{\kappa} \right) r^2 \quad (9)$$

Eq. (9) introduces a constraint on the free Lichouski metric. For the circle:

$$\epsilon = 0 \rightarrow (10)$$

$$\text{so: } dr = 0 \quad - (11)$$

and the metric become:

$$ds^2 \cdot dr^2 = r^2 d\phi^2 \quad - (12)$$

$$ds^2 = c^2 d\tau^2 = c^2 dt^2 - r^2 d\phi^2 \quad - (13)$$

$$ds^2 = c^2 d\tau^2 = c^2 (dt^2 - r^2 d\phi^2) \quad - (14)$$

$$\text{i.e. } r^2 d\phi^2 = c^2 (dt^2 - r^2 d\phi^2)$$

$$= c^2 dt^2$$

$$= c^2 dt^2 \quad - (15)$$

$$\text{i.e. } \frac{d\phi}{dt} = \frac{c}{r} = \omega \quad - (16)$$

In the free metric (1) there is no relation between  $dr$  and  $d\phi$ . Therefore as a result of any type constraints the free Minkowski metric.

For the ellipse (a):

$$P_1 := 8m \frac{dr}{dt} = \left( \frac{\epsilon \sin \phi}{a} \right) L \quad - (17)$$

$$\text{so } \frac{\epsilon \sin \phi}{a} = \frac{P_1}{L} \quad - (18)$$

is a constant of motion. The orbital equation is:

$$\frac{d\phi}{dr} = \frac{1}{r^2} \left( \frac{L}{P_1} \right) \quad - (19)$$

The non-relativistic limit of eq. (2) is

$$T = \frac{1}{2} m v^2 = \frac{1}{2} m \left( \left( \frac{dr}{dt} \right)^2 + r^2 \left( \frac{d\phi}{dt} \right)^2 \right) \quad - (20)$$

This kinetic energy is stored from the work:

$$W_{12} = \int_1^2 \underline{F} \cdot d\underline{r} = T_2 - T_1 \quad (21)$$

in which  $d\left(\frac{1}{2}mv^2\right) = \underline{F} \cdot d\underline{r}$  — (22)  
where  $\underline{F}$  is the force. Usually, the potential energy  
is introduced as

$$W_{12} = \int_1^2 \underline{F} \cdot d\underline{r} = T_2 - T_1 = U_1 - U_2 \quad (23)$$

so the Hamiltonian is:

$$H = T_1 + U_1 = T_2 + U_2 \quad (24)$$

$$U_2 = T_1 = 0 \quad (25)$$

$$U_1 = T_2 \quad (26)$$

The Hamiltonian and Lagrangian are:

$$H = T + T_2 \quad (27)$$

$$L = T - T_2 \quad (28)$$

If the inverse square attraction is introduced:

$$T = \frac{1}{2}m(r^2 + r^2\dot{\phi}^2) \quad (29)$$

$$T_2 = U_1 = -\frac{k}{r} \quad (30)$$

$$k = mMB \quad (31)$$

where  $R = mMB$ .

Note carefully that this is a purely kinetic theory. Eq. (23) shows clearly that the idea of "potential energy" is interchangeable with kinetic energy.

+ ) The lagrangian is therefore:

$$L = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\phi}^2) - \frac{k}{r} \quad (32)$$

The Lagrange equation is:

$$\frac{\partial L}{\partial r} = \frac{d}{dt} \frac{\partial L}{\partial \dot{r}} \quad (33)$$

giving:  $m (\ddot{r} - r \dot{\phi}^2) = - \frac{k}{r^2} \quad (34)$

which is the same as:

$$\frac{d^2 u}{d\phi^2} + u = - \frac{m}{L^2} \frac{1}{u^2} F(u) \quad (35)$$

where

$$F(u) = - \frac{\partial T_2}{\partial r} \quad (36)$$

$$\begin{aligned} L &= m r^2 \dot{\phi} \\ &= \text{constant of motion} \end{aligned} \quad (37)$$

The solution of eq. (35), i.e. the Lagrange equation of motion (32), is the ellipse (8).

This is seen from the fact that:

$$F(u) = - \frac{k}{r^2} = - k u^2 \quad (38)$$

so

$$\frac{d^2 u}{d\phi^2} + u = \frac{mk}{L^2}, \quad (39)$$

$$u = \frac{1}{d} (1 + \epsilon \cos \phi) \quad (40)$$

5)

So

$$\frac{d^2 u}{d\phi^2} = -\frac{\epsilon \cos \phi}{d} - (41)$$

and

$$\frac{d^2 u}{d\phi^2} + u = \frac{1}{d} = \frac{mk}{L^2} - (42)$$

i.e.

$$d = \frac{L^2}{mk} - (43)$$

### Discussion

The object:

$$T_2 = U_1 = -\frac{k}{r} - (44)$$

constrains the free Minkowski metric to produce an elliptical or conical section function of  $r$  on  $\phi$ , i.e. it relates  $dr$  and  $d\phi$  by eq. (9).  
 In the free metric this relation does not exist. So  
 to free metric:

$$\frac{k}{r} \rightarrow 0, r \rightarrow \infty - (45)$$

producing  $\frac{d^2 u}{d\phi^2} + u = 0 - (46)$

and  $d \rightarrow \infty - (47)$

in the conical section equation (8). In the free metric:

$$H = \mathcal{L} = T - (48)$$

and

$$m\dot{r} = m\dot{\phi}^2 - (49)$$

i.e.

$$F_N = F_C - (50)$$

where

$$F_N = m\ddot{r} \quad (51)$$

is the central, a Newtonian, force, and  $F_C$  is the centrifugal force.

Otherwise:

$$m\ddot{r} = mr\dot{\phi}^2 - \frac{k}{r^2} \quad (51)$$

i.e.

$$F_N = F_C + F_2 \quad (52)$$

where

$$F_2 = -\frac{\partial T_2}{\partial r} \quad (53)$$

is the change of kinetic energy w.r.t distance. Therefore  
the balance of forces in the elliptical orbit is:

$$F_C = F_N + \frac{\partial T_2}{\partial r} = F_N - F_2 \quad (54)$$

where

$$-\frac{\partial T_2}{\partial r} = F_2 = \frac{mMg}{r^2} \quad (55)$$

is determined by observation. There is no  
mathematical way of deriving eq. (55), and in  
a whirlpool galaxy, this equation no longer  
holds at all.

Note that both  $F_N$  and  $F_2$  are inward  
towards  $M$ , and  $F_C$  is outward.

# 1) 149(7): Relation of Minkowski and Gravitational Metrics.

As shown in note 149(2) the Minkowski metric:

$$ds^2 = c^2 dt^2 - dx \cdot dx \quad (1)$$

is equivalent to the Einstein energy equation:

$$E^2 = c^2 p^2 + m^2 c^4 \quad (2)$$

where

$$p^2 = p_i^2 + \frac{L^2}{m r^2} \quad (3)$$

In this equation the following constants of motion are

defined:  $E = mc^2 \frac{dt}{d\tau} = \gamma mc^2 \quad (4)$

$$p_i = m \frac{dx}{d\tau} = \gamma m \frac{dx}{dt} \quad (5)$$

$$L = m r^2 \frac{d\phi}{d\tau} = \gamma m r^2 \frac{d\phi}{dt} \quad (6)$$

$$and \quad ds^2 = dx^2 + r^2 d\phi^2 + dz^2 \quad (7)$$

The total energy for eq. (2) is:

$$E = \gamma mc^2 \quad (8)$$

and the relativistic kinetic energy is:

$$T = (\gamma - 1)mc^2 \quad (9)$$

The relativistic linear momentum is:

$$p_i = |\underline{p}_i| = \gamma m \frac{dx}{dt} \quad (10)$$

and the total relativistic linear momentum is:

$$\underline{p} = \gamma m \underline{v} \quad (11)$$

$$= \gamma m \left( p_i^2 + \frac{L^2}{m r^2} \right)^{1/2} \quad (12)$$

2) In the non-relativistic limit:  
 $\sqrt{1 - \frac{v^2}{c^2}} \approx 1 - \frac{v^2}{2c^2} \quad (13)$

Eq. (9) becomes:  $T = \frac{1}{2}mv^2 \quad (14)$

The introduction of an inverse square law of attraction is equivalent to defining the Hamiltonian:

 $H = (\gamma - 1)mc^2 + U \quad (15)$ 

where  $U = -\frac{mMG}{r} \quad (16)$

A Hamiltonian of the type (15) was used by Sommerfeld in a slightly different context in his semi-classical description of the hydrogen atom. It was shown by Sommerfeld that a Hamiltonian of type (15) produces a precessing ellipse. Because this is no need for the Schwarzschild metric. Eq. (15) is:

$$H = T + U \quad (17)$$

so the effect of  $U$  is:

$$\boxed{T \rightarrow T - \frac{k}{r}} \quad (18)$$

in the Minkowski metric (1). It is shown as follows that the Hamiltonian (17) is obtained from the metric:

$$ds^2 = c^2 dt^2 = \left(1 - \frac{r_0}{r}\right)c^2 dt^2 - dr \cdot dr \quad (19)$$

where  $r_0 = \frac{2mG}{c^2} \quad (20)$

∴ The constants of motion of this metric are:

$$E = mc^2 \left(1 - \frac{r_0}{r}\right) \frac{dt}{d\tau} = \gamma mc^2 \left(1 - \frac{r_0}{r}\right) - (21)$$

$$P_1 = m \frac{dr}{d\tau} = \gamma m \frac{dr}{dt} - (22)$$

$$L = mc^2 \frac{d\phi}{d\tau} = \gamma mc^2 \frac{d\phi}{dt} - (23)$$

With these definitions:

$$\frac{1}{2} \left(1 - \frac{r_0}{r}\right) \frac{P_1^2}{m} = \frac{E^2}{2mc^2} - \frac{1}{2} \left(1 - \frac{r_0}{r}\right) \left(mc^2 + \frac{L^2}{mr^2}\right) - (24)$$

$$\therefore \frac{1}{2} \left(1 - \frac{r_0}{r}\right) m \left(\frac{dr}{d\tau}\right)^2 = \frac{E^2}{2mc^2} - \frac{1}{2} \left(1 - \frac{r_0}{r}\right) \left(mc^2 - \frac{L^2}{mr^2}\right) - (25)$$

For comparison, the standard gravitational metric is:

$$\frac{1}{2} m \left(\frac{dr}{d\tau}\right)^2 = \frac{E^2}{2mc^2} - \frac{1}{2} \left(1 - \frac{r_0}{r}\right) \left(mc^2 - \frac{L^2}{mr^2}\right) - (26)$$

In fact metrics the right hand side is: - (27)

$$RHS = \left(\frac{E^2}{2mc^2} - \frac{1}{2} mc^2\right) + \left(\frac{nm\hbar}{r} - \frac{L^2}{2mr^2} + \frac{6mL^2}{mc^2 r^2}\right)$$

This gives a processsing ellipse as is well known.  
i.e. the metric (19) gives a processsing ellipse by:

$$dt^2 \rightarrow \left(1 - \frac{r_0}{r}\right) dt^2 - (28)$$

verified experimentally as the gravitational redshift.

4)

 $\tilde{L}$  the limit:

$$r_0 \ll r \rightarrow (29)$$

eq. (27) becomes:

$$\text{RHS} \rightarrow \frac{1}{2} mc^2 (\gamma^2 - 1) + \frac{mMg}{r} - \frac{L^2}{2mr^2} \rightarrow (30)$$

 $\tilde{L}$  the limit:

$$v \ll c \rightarrow (31)$$

this becomes:

$$\begin{aligned} \text{RHS} &\rightarrow mc^2 (\gamma - 1) + \frac{mMg}{r} - \frac{L^2}{2mr^2} \\ &= \frac{1}{2} mv^2 + \frac{mMg}{r} + \frac{L^2}{2mr^2}, \end{aligned} \rightarrow (32)$$

As

$$r \rightarrow \infty \rightarrow (33)$$

eq. (25) therefore becomes:

$$\frac{1}{2} mv^2 = \frac{1}{2} m (\ddot{r}^2 + r^2 \dot{\phi}^2) \rightarrow (34)$$

The Lagrangian used by Sommerfeld is:

$$L = mc^2 (\gamma - 1) + \frac{mMg}{r} \rightarrow (35)$$

and the Hamiltonian used by Sommerfeld is:

$$H = mc^2 (\gamma - 1) - \frac{mMg}{r} \rightarrow (36)$$

These are approximations of the complete

5) Lagrangian:

$$L = \left( \frac{E^2}{2mc^2} - \frac{1}{2} mc^2 \right) + \left( \frac{mMG}{r} + \frac{GL^2}{mc^2 r^3} \right) \quad (37)$$

and the complete Hamiltonian:

$$H = \left( \frac{E^2}{2mc^2} - \frac{1}{2} mc^2 \right) - \left( \frac{mMG}{r} + \frac{GL^2}{mc^2 r^3} \right) \quad (38)$$

for the metric (19). This metric is:

$$ds^2 = c^2 d\tau^2 = c^2 dt^2 - \frac{dr \cdot dr}{\left( \frac{2mG}{c^2 r} \right) dt^2}$$

- (39)

This metric gives a precessing ellipse:

$$r = \frac{d}{1 + \epsilon \cos(y\phi)} \quad (40)$$

and in the Sommerfeld approximation, also gives Dirac's equation of the hydrogen atom. Using eq. (40) and the principle of orbits, the metric (39) is:

$$ds^2 = c^2 d\tau^2 = c^2 dt^2 - \left( 1 + \frac{r^2 \omega^2}{c^2} \right) r^2 d\phi^2$$

where  $x = \frac{y \epsilon}{d} \sin(y\phi)$ . - (41)

Eq. (41) is a Misner-Shapiro metric with:

$$dr = x r^2 d\phi \quad (42)$$

6)

## Conclusion

The usual gravitational metric:

$$ds^2 = c^2 d\tau^2 = c^2 dt^2 \left(1 - \frac{r_0}{r}\right) - \left(1 - \frac{r_0}{r}\right)^{-1} dr^2 - r^2 d\phi^2 - dz^2 \quad (44)$$

is not needed for the description of a precessing ellipse. The most economical solution of the problem is the metric (41), which is a Minkowski metric with the functional relation (43) of the precessing ellipse. The metrics (41) and (39) are the same. In the Sommerfeld approximation (36) they give the same solution for the orbital of the electron in a H atom as the Dirac equation.

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1) 149(8): The Kinetic Nature of General Relativity

The lagrangian in general relativity is purely kinetic:

$$L = T = \frac{1}{2} mc^2 = \frac{1}{2} m g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \quad - (1)$$

In the new metric of note 149(7):

$$L = \frac{1}{2} mc^2 = \frac{1}{2} mc^2 \left(1 - \frac{c_0}{r}\right) \left(\frac{dt}{d\tau}\right)^2 - \frac{m}{2} \left(\left(\frac{dr}{d\tau}\right)^2 + r^2 \left(\frac{d\phi}{d\tau}\right)^2\right) \quad - (2)$$

$$\therefore \frac{1}{2} mc^2 \left(\left(1 - \frac{c_0}{r}\right) \left(\frac{dt}{d\tau}\right)^2 - 1\right) = \frac{m}{2} \left(\left(\frac{dr}{d\tau}\right)^2 + r^2 \left(\frac{d\phi}{d\tau}\right)^2\right) \quad - (3)$$

In the non-relativistic limit this means that

$$\boxed{\frac{1}{2} mv^2 \rightarrow \frac{1}{2} mv^2 - \frac{nmG}{r}} \quad - (4)$$

because

$$\frac{m}{2} \left(\left(\frac{dr}{d\tau}\right)^2 + r^2 \left(\frac{d\phi}{d\tau}\right)^2\right) \rightarrow \frac{1}{2} mv^2 \quad - (5)$$

and

$$\begin{aligned} \frac{1}{2} mc^2 \left(\left(1 - \frac{c_0}{r}\right) \left(\frac{dt}{d\tau}\right)^2 - 1\right) \\ = \frac{1}{2} mc^2 (\gamma^2 - 1) - \frac{1}{2} mc^2 \frac{c_0}{r} \left(\frac{dt}{d\tau}\right)^2 \\ \rightarrow \frac{1}{2} mv^2 - \frac{nmG}{r} \end{aligned} \quad - (6)$$

This is a clear way of showing that in a  
local analysis, there is no concept of potential  
energy, because the lagrangian is  $T$ .

The usual analysis is obtained by multiplying eq. (2) by  $(1 - \frac{r_0}{r})$ :

$$\frac{1}{2} mc^2 \left(1 - \frac{r_0}{r}\right) = \frac{1}{2} mc^2 \left(1 - \frac{r_0}{r}\right)^2 \left(\frac{dt}{d\tau}\right)^2 - \left(1 - \frac{r_0}{r}\right) \frac{m}{2} \left(\left(\frac{dr}{d\tau}\right)^2 + r^2 \left(\frac{d\phi}{d\tau}\right)^2\right),$$

and re-arranging:

$$\begin{aligned} \frac{1}{2} \left(1 - \frac{r_0}{r}\right) m \left(\frac{dr}{d\tau}\right)^2 &= \frac{1}{2} mc^2 \left(1 - \frac{r_0}{r}\right)^2 \left(\frac{dt}{d\tau}\right)^2 - \frac{1}{2} mc^2 \left(1 - \frac{r_0}{r}\right) \\ &\quad - \left(1 - \frac{r_0}{r}\right)^2 \frac{m}{2} r^2 \left(\frac{d\phi}{d\tau}\right)^2 \\ &= \frac{E^2}{2mc^2} - \frac{1}{2} mc^2 + \left(\frac{mr^2}{r} - \frac{L^2}{2mr^2} + \frac{6ML^2}{mc^2 r^3}\right) \end{aligned}$$

as in note 149(7). Eq (8) means -(8)

$$J = T = \frac{1}{2} mv^2 \rightarrow \frac{1}{2} mv^2 + \frac{mr^2}{r} - (9)$$

is the non-relativistic limit, and this again is purely kinetic.

Inverse square attraction in the metric of note 149(7) means:

$$\boxed{dt^2 \rightarrow \left(1 - \frac{2GM}{c^2 r}\right) dt^2} - (10)$$