

Cosmology

Pre-class reading 7

- 1) For mode with $k_1 < k_{eq}$, the modes were outside the horizon during radiation era and grew as a^2 . When the Universe becomes matter dominated, growth slowed to a . After the mode enters the horizon during the matter era, the growth still goes as a . For mode with $k_2 > k_{eq}$, before horizon entry, modes grow as a^2 . After horizon entry (still during radiation era), the growth becomes logarithmic in a or almost constant. This mode grows again as a once the Universe becomes matter dominated. The ratio of linear growth factors of the k_2 mode with respect to the k_1 mode is $(a_{k_2}/a_{eq})^2$ or $(k_{eq}/k)^2 < 1$
- 2) Accurately normalizing the power spectrum requires a method that doesn't depend on non-linear growth (variance of galaxy distribution measured $\sim 8/h$ Mpc will trace the perturbations when they may have become non-linear) and that doesn't depend on assuming galaxies tracing mass distribution. Thus, σ_8 is calculated as $\sigma_m(8/h \text{ Mpc})$ from the initial power spectrum evolved to the present time according to linear theory
- 3) From the spherical (non-linear) model, we see that at $t=t_{collapse}$ ($\theta = 2\pi$), the density contrast diverges. This corresponds to a density contrast of 1.69 if we extrapolate the linear model to $t_{collapse}$. So, the reason this number is so low is just math? Dark matter halos are significantly more dense compared to the average density of the Universe because they underwent non-linear collapse.
- 4) For halos that deviate even slightly from spherical symmetry, shells start crossing and halo ultimately virializes- balancing out potential and kinetic energies. By applying virial theorem and energy conservation, we find that virialized DM halo has a density contrast ~ 178 times greater than the background density
- 5) The virial radius is the radius when the halo has virialized i.e its total energy is given by $-(1/2)*3/5GM^2/R_{vir}$. This is set by energy conservation between energy at turn-around and energy after virialization. This is not a good definition I think because it grows with time as more and more dark matter from outside the virial radius gets accreted?

