

Lecture 4



Example: A s/c is orbiting Earth on an orbit with an eccentricity of 0.2. The radius of periaresis is 1000 km altitude. What is the speed at periaresis, the radius of apoaesis & the speed @ apoaesis?

$$r_p = 6378 + 1000 = 7378 \text{ km}$$

$$\text{Speed at } r_p: \Sigma = \frac{v^2}{2} - \frac{\mu}{r} = -\frac{\mu}{2a}$$

$$r_p = a(1-e) \Rightarrow a = r_p / (1-e) = 9222.5 \text{ km}$$

$$\frac{v_p^2}{2} - \frac{\mu}{r_p} = -\frac{\mu}{2a} \Rightarrow v_p = \sqrt{\frac{2\mu}{r_p} - \frac{\mu}{a}}$$

$$\text{Earth} \Rightarrow \mu = 3.986 \times 10^5 \text{ km}^3/\text{s}^2 \Rightarrow v_p = 8.052 \text{ km/s}$$

$$\text{Radius of Apoaesis: } r_a = a(1+e) \Rightarrow r_a = 11067 \text{ km}$$

$$\text{Speed at apoaesis: from energy eqn: } v_a = \sqrt{\frac{2\mu}{r_a} - \frac{\mu}{a}} \Rightarrow v_a = 5.368 \text{ km/s}$$

Example: A s/c is orbiting Earth. At periaresis, it has an altitude of 1500 km and a velocity of 8.5 km/s. What is the eccentricity of the orbit? What is the flight path angle & speed of the s/c when its altitude is 3000 km and $v < 180^\circ$?

$$\text{Eccentricity: } r_p = a(1-e)$$

$$r_p = 1500 + 6378 = 7878 \text{ km}$$

$$\Sigma = \frac{v_p^2}{2} - \frac{\mu}{r_p} = -\frac{\mu}{2a}$$

$$v_p = 8.5 \text{ km/s}$$

$$\Rightarrow a = \frac{-\mu}{2} \left[\frac{v_p^2}{2} - \frac{\mu}{r_p} \right]^{-1} \Rightarrow a = 27,543.6 \text{ km}$$

$$1-e = r_p/a \Rightarrow e = 1 - r_p/a \Rightarrow e = 0.719$$

$$\text{Velocity @ } r = 3000 + 6378 = 9378 \text{ km}$$

$$\Sigma = \frac{v^2}{2} - \frac{\mu}{r} = -\frac{\mu}{2a}$$

$$v = \sqrt{\frac{2\mu}{r} - \frac{\mu}{a}} \Rightarrow v = 8.399 \text{ km/s}$$

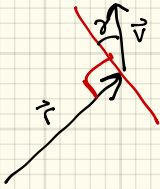
Flight Path Angle: γ

$$h = r v \cos \gamma$$

From the previous part, we know r, v at this location

$$r = 9378 \text{ km}$$

$$v = 8.379 \text{ km/s}$$

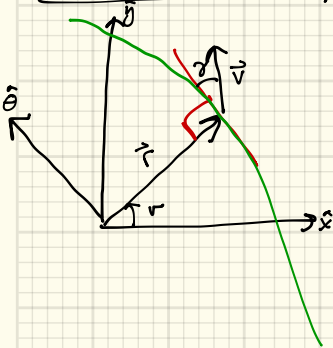


$$h = r v \cos \gamma = r_p v_p \cos \gamma_p, \quad \gamma_p = 0$$

$$= r_p v_p$$

$$\cos \gamma = \frac{r_p v_p}{r v} \Rightarrow \boxed{\gamma = 31.77^\circ} \quad \text{b/c } v < 180^\circ, \gamma > 0$$

Period of an Ellipse: time required to traverse the orbit



From Kinematics:

$$\vec{r} = r \hat{r}$$

$$\dot{\vec{r}} = \dot{r} \hat{r} + r \dot{\hat{r}}$$

$$\hat{r} = \cos \nu \hat{x} + \sin \nu \hat{y}$$

$$\hat{\theta} = -\sin \nu \hat{x} + \cos \nu \hat{y}$$

$$\dot{\hat{r}} = -\dot{\nu} \sin \nu \hat{x} + \dot{\nu} \cos \nu \hat{y}$$

$$\dot{\hat{r}} = \dot{\nu} \hat{\theta}$$

$$\dot{\vec{r}} = \dot{r} \hat{r} + r \dot{\nu} \hat{\theta}$$

$$h = r v \cos \gamma$$

$$v \cos \gamma = r \dot{\nu} \quad (\text{Component of the velocity in the } \hat{\theta} \text{ direction})$$

$$\Rightarrow h = r^2 \dot{\nu} = r^2 \frac{d\nu}{dt} \Rightarrow dt = \frac{r^2}{h} d\nu \quad r = f(\nu) \text{ then the trajectory eqn}$$

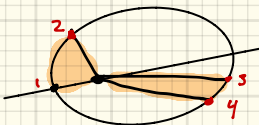
$$r = \frac{p}{1 + \cos \nu}$$

$$\text{Area of a circle} = \pi r^2$$

$$\text{Portion of a circle} = \frac{\nu}{2\pi} \pi r^2 \quad (\nu \text{ in rad here})$$

$$dA = \frac{1}{2} r^2 d\nu$$

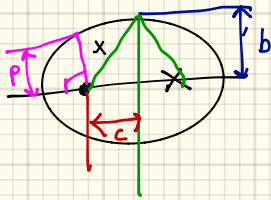
$$dt = \frac{2}{h} dA \quad \leftarrow \text{Proves Kepler's 2nd law: equal areas in equal times}$$



Suppose the 2 shaded areas are the same size, then it will take the spacecraft the same amount of time to go from $r_1 \rightarrow r_2$ and from $r_3 \rightarrow r_4$

The area of an ellipse is: $A = \pi ab$ (b = semi-minor axis)

$$\text{Integrating } dt = \frac{2}{h} dA \Rightarrow TP = \frac{2\pi ab}{h}$$



$$e = \frac{c}{a}$$

$$\text{length of string} = 2a = 2a(1+e)$$

$$= 2x + 2c = 2x + 2ea$$

$$2a(1+e) = 2x + 2ea$$

$$a + ea = x + ea \Rightarrow x = a$$

$$a^2 = b^2 + c^2 \Rightarrow b = \sqrt{a^2 - c^2} = \sqrt{a^2 - e^2 a^2} = \sqrt{a^2(1 - e^2)} = \sqrt{ap}$$

$$h = \sqrt{ap}$$

$$TP = \frac{2\pi a \sqrt{ap}}{\sqrt{ap}} \Rightarrow \boxed{TP = 2\pi \sqrt{\frac{a^3}{h}}}$$

Velocity of a Circular orbit:

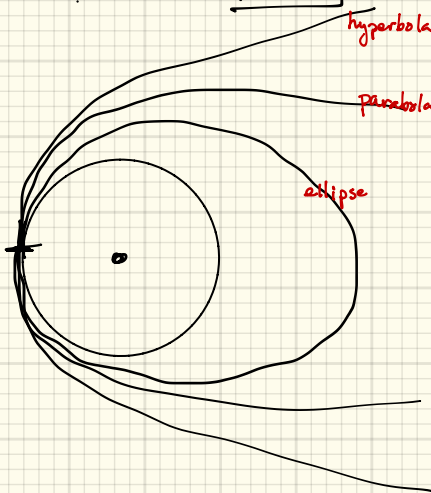
$$r = \text{constant} \Rightarrow v = \text{constant}$$

$$\Sigma = \frac{v^2}{2} - \frac{\mu}{r} = -\frac{\mu}{2a}$$

$$a = r \text{ (for circular orbit)}$$

$$\frac{v^2}{2} - \frac{\mu}{r} = -\frac{\mu}{2r}$$

$$\Rightarrow \boxed{v_c = \sqrt{\frac{\mu}{r}}}$$



Parabola: $E = 0$

The escape speed is the speed of a parabolic orbit at a given radius (r).

$$E = \frac{V_{esc}^2}{2} - \frac{M}{r} = \frac{V_{\infty}^2}{2} - \frac{M}{r_{\infty}} = 0$$

$$V_{esc} = \sqrt{\frac{2M}{r}}$$

Hyperbola: The hyperbola has some non-zero velocity at infinity. The hyperbolic excess speed is this velocity.

hyperbolic excess speed = Speed at $r = \infty$

$$E = \frac{V_{\infty}^2}{2} - \frac{M}{r_{\infty}} = -\frac{M}{2a} \quad \text{b/c } r_{\infty} = \infty$$

$$\Rightarrow V_{\infty} = \sqrt{-\frac{M}{a}}$$

Note $a < 0$ for hyperbola