Model Equations - Axisymmetric Atmosphere Simulation

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September 25, 2021

1 Original Equations from [1]

Equations of motion for zonal, meridional, and vertical winds:

$$\partial_t u = -\nabla \cdot \begin{pmatrix} vu \\ wu \end{pmatrix} + fv + \frac{uv \tan \varphi}{a} + \partial_z (\nu \partial_z u)$$

$$\partial_t v = -\nabla \cdot \begin{pmatrix} vv \\ wv \end{pmatrix} - fu - \frac{u^2 \tan \varphi}{a} - \frac{1}{a} \partial_\varphi \Phi + \partial_z (\nu \partial_z v)$$

$$\partial_t w = -\nabla \cdot \begin{pmatrix} vw \\ ww \end{pmatrix} - \partial_z \Phi + \frac{g\Theta}{\Theta_0}$$

First law of thermodynamics

$$\partial_t \Theta = -\nabla \cdot \begin{pmatrix} v\Theta \\ w\Theta \end{pmatrix} - (\Theta - \Theta_E)\tau^{-1} + \partial_z (\nu \partial_z \Theta)$$

Continuity equation:

$$0 = -\nabla \cdot \begin{pmatrix} v \\ w \end{pmatrix}$$

where u,v,w denote the zonally averaged winds in zonal, meridional, and vertical directions; z,φ denote the spherical coordinates; $\nabla = \begin{pmatrix} (a\cos\varphi)^{-1}\partial_\varphi(\cos\varphi) \\ \partial_z \end{pmatrix}$ is the gradient operator in the meridional-vertical plane; $f=2\Omega\sin(\varphi)$ is the Coriolis parameter, a is Earth's radius, $\Phi=gz$ denotes the geopotential, ν viscosity, and g the gravitational acceleration. Θ is the potential temperature field, τ is a constant radiative damping time. Heating/cooling is implemented as a relaxation towards the equilibrium temperature field Θ_E , defined as:

$$\frac{\Theta_E}{\Theta_0} = 1 - \frac{2}{3} \Delta_H P_2(\sin \varphi) + \Delta_v \left(\frac{z}{H} - \frac{1}{2}\right)$$

 Θ_0 is the average of Θ_E ; Δ_H is a nondimensional parameter representing the fractional change in potential temperature from equator to pole in radiative equilibrium; Δ_v is a nondimensional parameter representing the fractional change in potential temperature from top to bottom in radiative equilibrium; $P_2(x) = \frac{1}{2}(3x^2 - 1)$ is the second Legendre polynomial.

1.1 Boundary Conditions

• at z = H:

$$- w = 0$$

$$- \partial_z u = \partial_z v = \partial_z \Theta = 0$$

• at z = 0:

$$- w = 0$$

$$-\partial_z\Theta=0$$

$$-v\partial_z u = Cu$$

 $-v\partial_z v = Cv$; C is a constant drag coefficient.

1.2 Model Parameters

- $\nu = 25, 10, 5, 2.5, 1, 0.5 \frac{m^2}{s}$
- $\Omega = \frac{2\pi}{8.64 \times 10^6 s}$
- $a = 6.4 \times 10^6 m$
- $q = 9.8ms^{-2}$
- $H = 8.0 \times 10^3 m$
- $\Delta_H = \frac{1}{3}$
- $\Delta_v = \frac{1}{8}$
- $C = 0.005 ms^{-1}$
- $\tau = 20 \text{ days}$

2 Simplified Equations

By plugging the definition of ∇ , the continuity equation becomes:

$$0 = \frac{1}{a} \tan \varphi \cdot v - \frac{1}{a} \partial_{\varphi} v - \partial_z w$$

Doing the same for the equations of motion and the first law of thermodynamics, assuming constant viscosity ν and plugging in the continuity equation results in the following reformulations:

$$\partial_t u = -\frac{1}{a} v \partial_\varphi u + f v + \frac{u v \tan \varphi}{a} - w \partial_z u + \nu \partial_z^2 u$$

$$\partial_t v = -\frac{1}{a} v \partial_\varphi v - f u - \frac{u^2 \tan \varphi}{a} - w \partial_z v - \frac{1}{a} \partial_\varphi \Phi + \nu \partial_z^2 v$$

$$\partial_t w = -\frac{1}{a} v \partial_\varphi w - w \partial_z w - \partial_z \Phi + g \frac{\Theta}{\Theta_0}$$

$$\partial_t \Theta = -\frac{1}{a} v \partial_\varphi \Theta - w \partial_z \Theta - (\Theta - \Theta_E) \tau^{-1} + \nu \partial_z^2 \Theta$$

3 Discretizing the Equations

3.1 Space

Making use of the staggered grid finite difference approaches used in [2, 3], I suggest the following discretizations for the spatial derivatives:

Equations of Motion:

$$\begin{split} \partial_t u^k_{j+\frac{1}{2}} &= -\frac{1}{a} v^k_{j+\frac{1}{2}} \frac{u^k_{j+\frac{3}{2}} - u^k_{j-\frac{1}{2}}}{2\Delta \varphi_j} + f_j v^k_{j+\frac{1}{2}} + \frac{u^k_{j+\frac{1}{2}} v^k_{j+\frac{1}{2}} \tan \varphi_j}{a} \\ &- w^k_{j+\frac{1}{2}} \frac{u^{k+1}_{j+\frac{1}{2}} - u^{k-1}_{j+\frac{1}{2}}}{2\Delta z_k} + \nu \frac{u^{k+1}_{j+\frac{1}{2}} - 2u^k_{j+\frac{1}{2}} + u^{k-1}_{j+\frac{1}{2}}}{4\Delta z_k^2} \end{split}$$

$$\begin{split} \partial_t v_{j+\frac{1}{2}}^k &= -\frac{1}{a} v_{j+\frac{1}{2}}^k \frac{v_{j+\frac{3}{2}}^k - v_{j-\frac{1}{2}}^k}{2\Delta \varphi_j} - f_j u_{j+\frac{1}{2}}^k - \frac{\left(u_{j+\frac{1}{2}}^k\right)^2 \tan \varphi_j}{a} - w_{j+\frac{1}{2}}^k \frac{v_{j+\frac{3}{2}}^{k+1} - v_{j-\frac{1}{2}}^{k-1}}{2\Delta z_k} \\ &- \frac{1}{a} \frac{\Phi_{j+1}^k - \Phi_{j-1}^k}{2\Delta \varphi_j} + \nu \frac{v_{j+\frac{1}{2}}^{k+1} - 2v_{j+\frac{1}{2}}^k + v_{j+\frac{1}{2}}^{k-1}}{4\Delta z_t^2} \end{split}$$

First law of thermodynamics

$$\partial_t \Theta_j^k = -\frac{1}{a} v_{j+\frac{1}{2}}^k \frac{\Theta_{j+1}^k - \Theta_{j-1}^k}{2\Delta \varphi_j} - w_{j+\frac{1}{2}}^k \frac{\Theta_j^{k+1} - \Theta_j^{k-1}}{2\Delta z_k} - (\Theta_j^k - \Theta_E) \tau^{-1} + \nu \frac{\Theta_j^{k+1} - 2\Theta_j^k + \Theta_j^{k-1}}{4\Delta z_k^2}$$

Vertical winds:

$$\partial_t w_{j+\frac{1}{2}}^k = -\frac{1}{a} v_j^k \frac{w_{j+\frac{3}{2}}^k - w_{j-\frac{1}{2}}^k}{2\Delta \varphi_i} - w_{j+\frac{1}{2}}^k \frac{w_{j+\frac{1}{2}}^{k+1} - w_{j+\frac{1}{2}}^{k-1}}{2\Delta z_k} - \frac{\Phi_j^{k+1} - \Phi_j^{k-1}}{2\Delta z_k} + \frac{g\Theta_j^k}{\Theta_0}$$

where

$$f_j = 2\Omega \sin(\varphi_j)$$

3.2 Time

[2] use a backward difference formula, same as [3], who do backward Euler steps with a Δt of 15 minutes.

4 Resolution

I see no reason why we should use a non-equidistant grid. [1] use 90 grid-points in the vertical direction and 50 from equator to pole. [3] use a way coarser grid, 7.826° latitude and 9 vertical levels.

5 Reformulation

$$\frac{u_{k,j+\frac{1}{2}}^{n+1}-u_{k,j+\frac{1}{2}}^{n}}{\Delta t} = -\frac{1}{a}v_{k,j+\frac{1}{2}}^{n}\frac{u_{k,j+\frac{3}{2}}^{n}-u_{k,j-\frac{1}{2}}^{n}}{2\Delta \varphi} + f_{j+\frac{1}{2}}v_{k,j+\frac{1}{2}}^{n} + \frac{u_{k,j+\frac{1}{2}}^{n}v_{k,j+\frac{1}{2}}^{n}\tan\varphi_{j+\frac{1}{2}}}{a} \\ -w_{k,j+\frac{1}{2}}^{n}\frac{u_{k+1,j+\frac{1}{2}}^{n}-u_{k-1,j+\frac{1}{2}}^{n}}{2\Delta z} + \nu\frac{u_{k+1,j+\frac{1}{2}}^{n}-2u_{k,j+\frac{1}{2}}^{n}+u_{k-1,j+\frac{1}{2}}^{n}}{4\Delta z^{2}}$$

$$\begin{split} &\frac{v_{k,j+\frac{1}{2}}^{n+1}-v_{k,j+\frac{1}{2}}^{n}}{\Delta t} = -\frac{1}{a}v_{k,j+\frac{1}{2}}^{n} \frac{v_{k,j+\frac{3}{2}}^{n}-v_{k,j-\frac{1}{2}}^{n}}{2\Delta\varphi} - f_{j+\frac{1}{2}}u_{k,j+\frac{1}{2}}^{n} - \frac{\left(u_{k,j+\frac{1}{2}}^{n}\right)^{2}\tan\varphi_{j+\frac{1}{2}}}{a} \\ &-w_{k,j+\frac{1}{2}}^{n} \frac{v_{k+1,j+\frac{3}{2}}^{n}-v_{k-1,j-\frac{1}{2}}^{n}}{2\Delta z} - \frac{1}{a} \frac{\Phi_{k,j+1}^{n}-\Phi_{k,j-1}^{n}}{2\Delta\varphi} + \nu \frac{v_{k+1,j+\frac{1}{2}}^{n}-2v_{k,j+\frac{1}{2}}^{n}+v_{k-1,j+\frac{1}{2}}^{n}}{4\Delta z^{2}} \end{split}$$

$$\frac{w_{k,j+\frac{1}{2}}^{n+1}-w_{k,j+\frac{1}{2}}^{n}}{\Delta t}=-\frac{1}{a}v_{k,j}^{n}\frac{w_{k,j+\frac{3}{2}}^{n}-w_{k,j-\frac{1}{2}}^{n}}{2\Delta \varphi}-w_{k,j+\frac{1}{2}}^{n}\frac{w_{k+1,j+\frac{1}{2}}^{n}-w_{k-1,j+\frac{1}{2}}^{n}}{2\Delta z}-\frac{\Phi_{k+1,j}-\Phi_{k-1,j}}{2\Delta z}+\frac{g\Theta_{k,j}^{n}}{\Theta_{0}}+\frac{g\Theta_{k,j}^{n}}{2\Delta z}+\frac{g\Theta_{k,j}^{n}}{2\Delta z$$

$$\frac{\Theta_{k,j+\frac{1}{2}}^{n+1} - \Theta_{k,j+\frac{1}{2}}^{n}}{\Delta t} = -\frac{1}{a} v_{k,j+\frac{1}{2}}^{n} \frac{\Theta_{k,j+1}^{n} - \Theta_{k,j-1}^{n}}{2\Delta \varphi} - w_{k,j+\frac{1}{2}}^{n} \frac{\Theta_{k+1,j}^{n} - \Theta_{k-1,j}^{n}}{2\Delta z} - (\Theta_{k,j}^{n} - \Theta_{E})\tau^{-1} + \nu \frac{\Theta_{k+1,j}^{n} - 2\Theta_{k,j}^{n} + \Theta_{k-1,j}^{n}}{4\Delta z^{2}}$$

References

[1] Isaac M. Held and Arthur Y. Hou. Nonlinear Axially Symmetric Circulations in a Nearly Inviscid Atmosphere. *Journal of the Atmospheric Sciences*, 37(3):515–533, March 1980. Publisher: American Meteorological Society Section: Journal of the Atmospheric Sciences.

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