

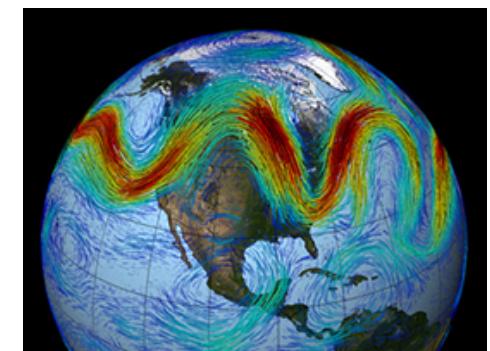
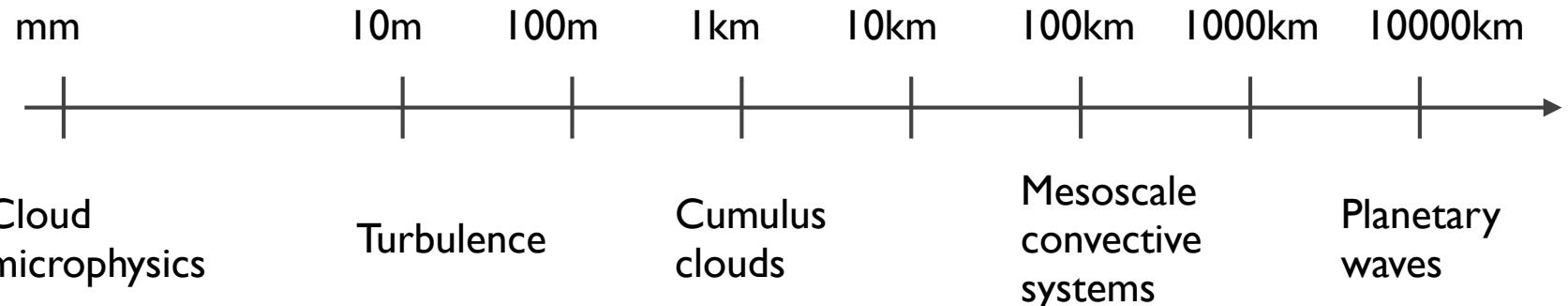


Estimating Stochastic Closures Using Sparsity-Promoting Ensemble Kalman Inversion

Jinlong Wu, Tapio Schneider, Andrew Stuart

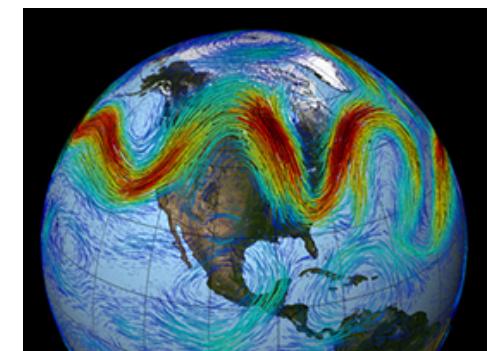
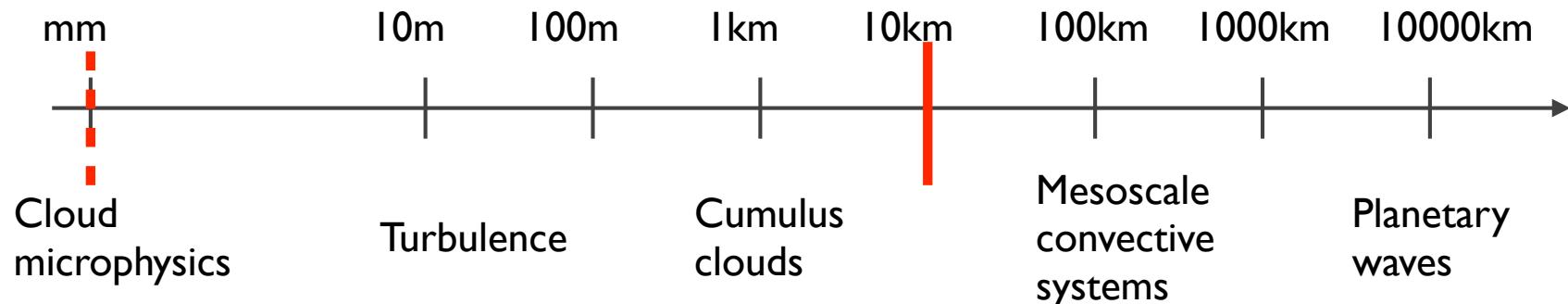
Joint IS-ENES3/ESIWACE2 Virtual Workshop, March 17th, 2021

Motivation: Climate Change



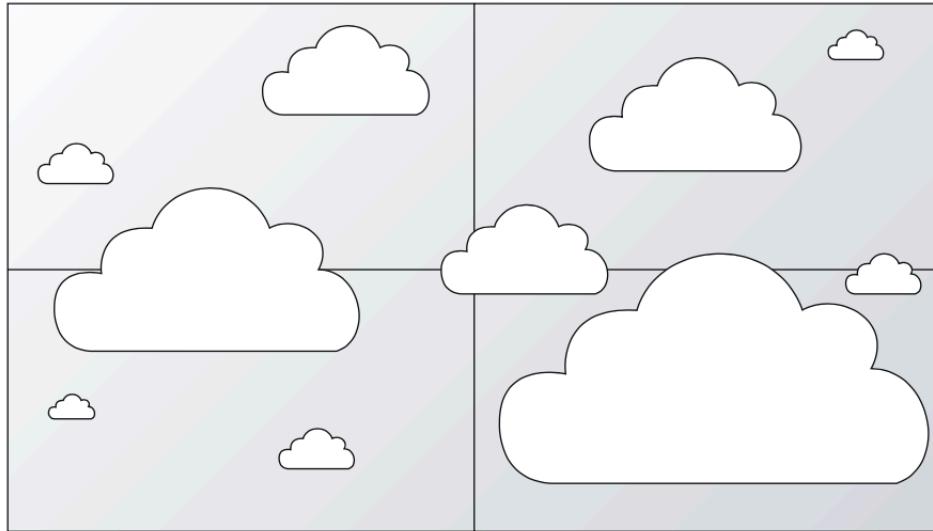
Motivation: Climate Change

Closure models are needed
for unresolved processes



Closure Problem

Figure source: Palmer, Nature Reviews Physics (2019)

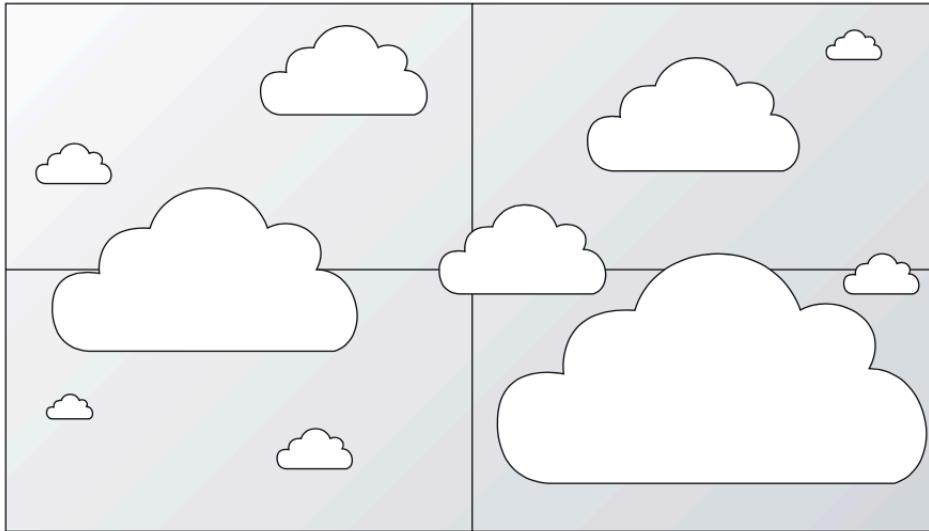


$$\frac{\partial u}{\partial t} = f(u)$$

- ❖ In practical simulations, it's usually impossible to resolve all the scales.

Closure Problem

Figure source: Palmer, Nature Reviews Physics (2019)



$$\frac{\partial u}{\partial t} = f(u) \quad U = \mathcal{F}(u)$$

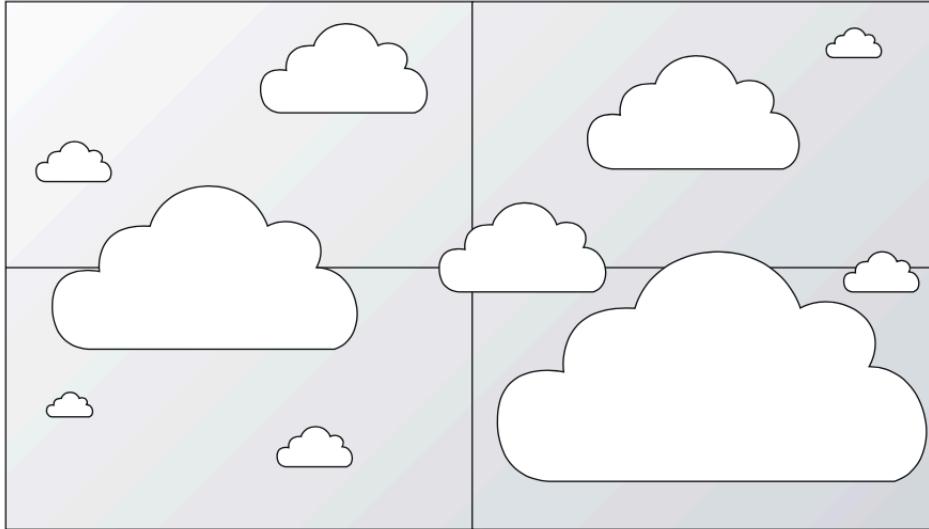
Filter

A mathematical model of atmospheric dynamics. The first equation shows the time derivative of a variable u is equal to a function $f(u)$. The second equation shows the total variable U is the result of a filter \mathcal{F} applied to u . An orange arrow labeled "Filter" points from the word "Filter" to the symbol \mathcal{F} .

- ❖ In practical simulations, it's usually impossible to resolve all the scales.

Closure Problem

Figure source: Palmer, Nature Reviews Physics (2019)



Filter
↑

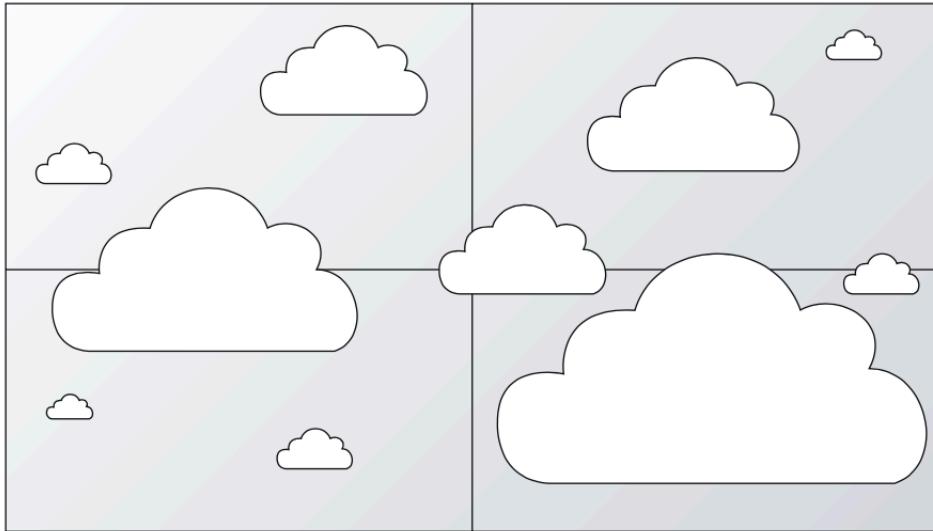
$$\frac{\partial u}{\partial t} = f(u) \quad U = \mathcal{F}(u)$$

$$\frac{\partial U}{\partial t} = \bar{f}(U; u)$$

- ❖ In practical simulations, it's usually impossible to resolve all the scales.
- ❖ To describe the impact of unresolved physics, closure models are needed.

Closure Problem

Figure source: Palmer, Nature Reviews Physics (2019)



Filter
↑

$$\frac{\partial u}{\partial t} = f(u) \quad U = \mathcal{F}(u)$$

$$\frac{\partial U}{\partial t} = \bar{f}(U; u)$$

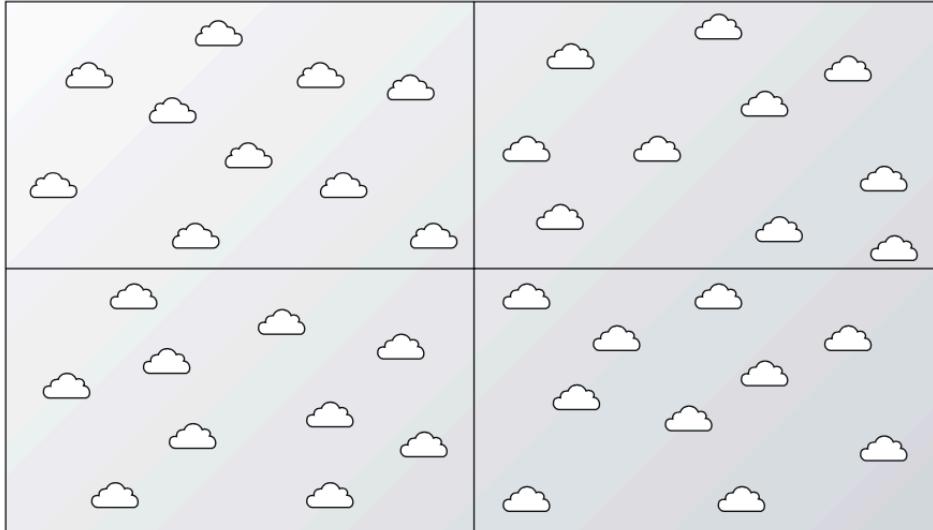
$$\frac{\partial U}{\partial t} = \tilde{f}(U)$$

$$\bar{f}(U; u) \approx \tilde{f}(U)$$

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Closure Problem

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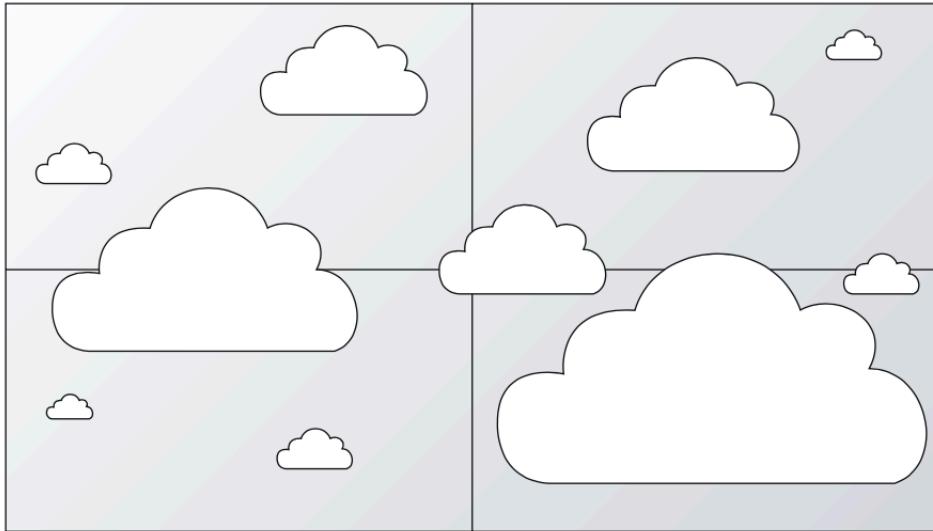
$$\frac{\partial u}{\partial t} = f(u) \quad U = \mathcal{F}(u)$$
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$$\bar{f}(U; u) \approx \tilde{f}(U)$$

- ❖ In practical simulations, it's usually impossible to resolve all the scales.
- ❖ To describe the impact of unresolved physics, closure models are needed.

Closure Problem

Figure source: Palmer, Nature Reviews Physics (2019)



Filter
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$$\frac{\partial u}{\partial t} = f(u) \quad U = \mathcal{F}(u)$$

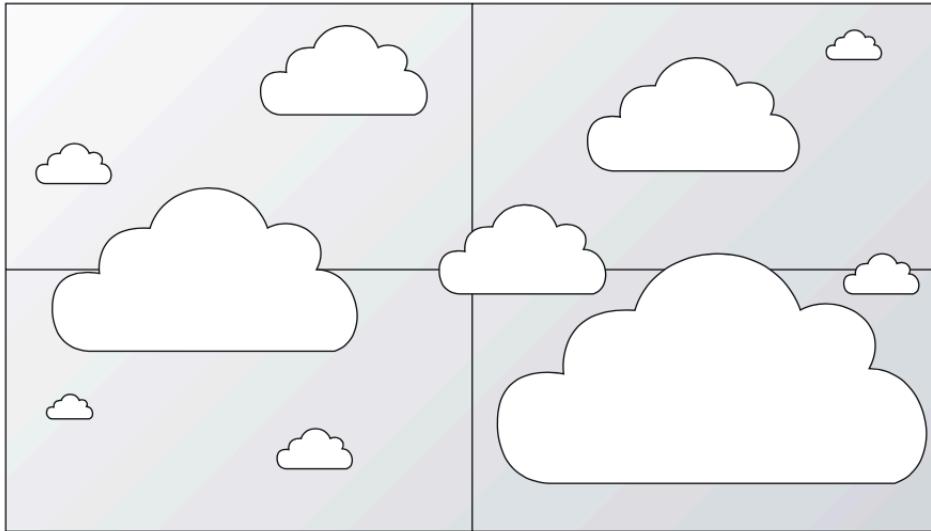
$$\frac{\partial U}{\partial t} = \bar{f}(U; u)$$

$$\frac{\partial U}{\partial t} = \tilde{f}(U)$$

$$\bar{f}(U; u) \approx \tilde{f}(U) + M(U; \alpha) + \sqrt{\sigma(U; \gamma)} \dot{W}$$

Closure Problem

Figure source: Palmer, Nature Reviews Physics (2019)



Filter
↑

$$\frac{\partial u}{\partial t} = f(u) \quad U = \mathcal{F}(u)$$

$$\frac{\partial U}{\partial t} = \bar{f}(U; u)$$

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$$\bar{f}(U; u) \approx \tilde{f}(U) + M(U; \alpha) + \sqrt{\sigma(U; \gamma)} \dot{W}$$

Dictionary learning

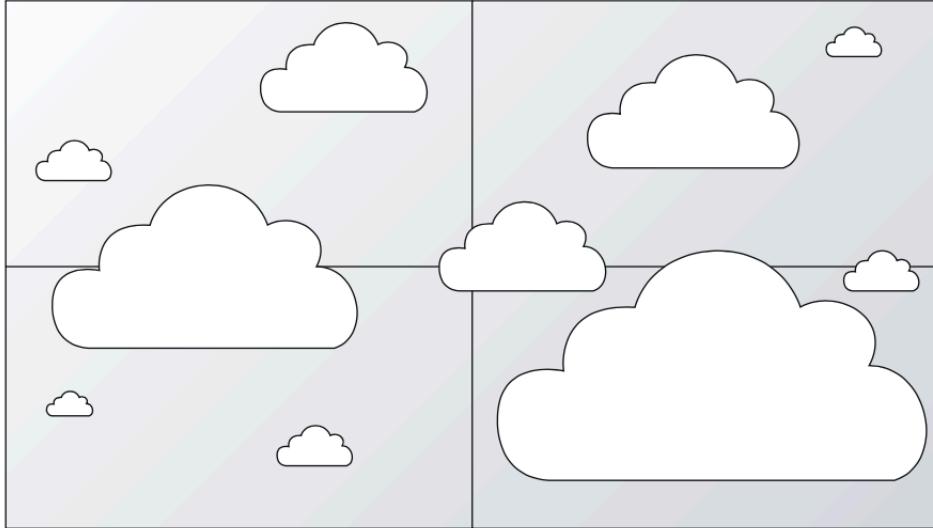
RKHS

Neural network

...

Closure Problem

Figure source: Palmer, Nature Reviews Physics (2019)



Filter
↑

$$\frac{\partial u}{\partial t} = f(u) \quad U = \mathcal{F}(u)$$

$$\frac{\partial U}{\partial t} = \bar{f}(U; u)$$

$$\frac{\partial U}{\partial t} = \tilde{f}(U)$$

$$\bar{f}(U; u) \approx \tilde{f}(U) + M(U; \alpha) + \sqrt{\sigma(U; \gamma)} \dot{W}$$

$$\theta = \{\alpha, \gamma\}$$

- ❖ We estimate the unknown parameters θ from some data y using ensemble Kalman inversion.

Ensemble Kalman Inversion (EKI)

- ❖ Inverse problem to be solved:

Data $y = G(\theta) + \eta, \quad \eta \sim N(0, \Gamma)$

- ❖ EKI formula (j: index of ensemble, m: index of EKI step):

$$\theta_{m+1}^{(j)} = \theta_m^{(j)} + C_m^{\theta G} \left(C_m^{GG} + \Gamma \right)^{-1} \left(y_{m+1}^{(j)} - G(\theta_m^{(j)}) \right)$$

- ❖ Advantage of EKI:

- ❖ Derivative-free optimization;
- ❖ Robust to noisy evaluations of the forward map G ;
- ❖ The core task is equivalent to a convex optimization problem.

Learning Stochastic Closures

Case I: Noisy Lorenz 63 System

True System: $\dot{x} = \tilde{f}(x) + \sqrt{\sigma}W$

Where $\tilde{f}(x)$ is the standard Lorenz 63 system :

$$\frac{dx_1}{dt} = \alpha(x_2 - x_1)$$

$$\frac{dx_2}{dt} = x_1(\rho - x_3) - x_2$$

$$\frac{dx_3}{dt} = x_1x_2 - \beta x_3$$

Case I: Noisy Lorenz 63 System

Modeled System: $\dot{x} = \tilde{f}(x) + \boxed{\sqrt{\sigma}W}$

Where $\tilde{f}(x)$ is the standard Lorenz 63 system :

$$\frac{dx_1}{dt} = \alpha(x_2 - x_1)$$

$$\frac{dx_2}{dt} = x_1(\rho - x_3) - \boxed{M(x_2)}$$

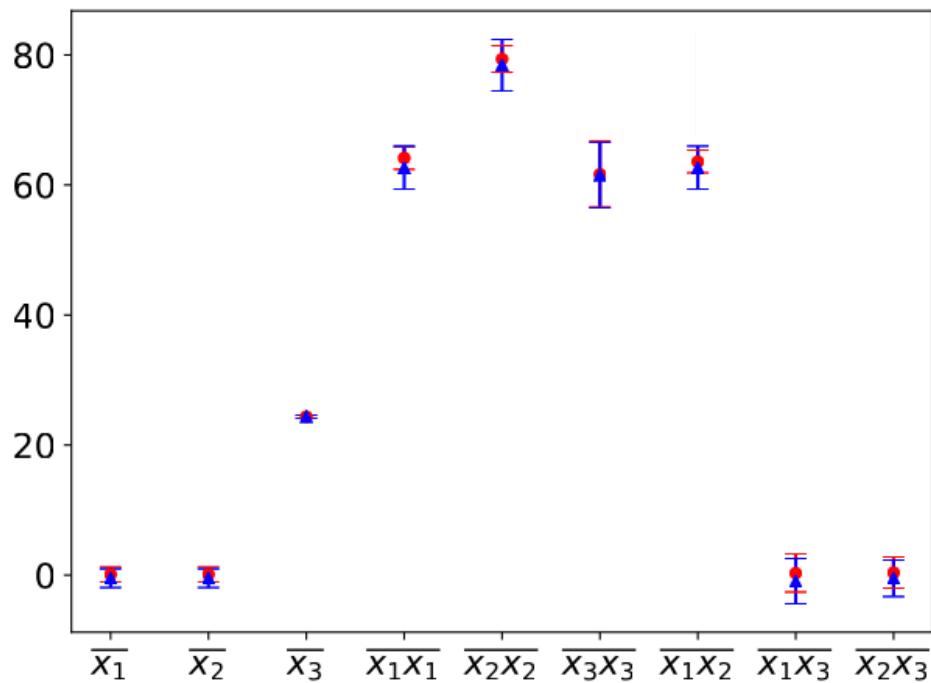
$$\frac{dx_3}{dt} = x_1x_2 - \beta x_3$$

Data: $\{\overline{x_1}, \overline{x_2}, \overline{x_3}, \overline{x_1x_1}, \overline{x_2x_2}, \overline{x_3x_3}, \overline{x_1x_2}, \overline{x_1x_3}, \overline{x_2x_3}\}$

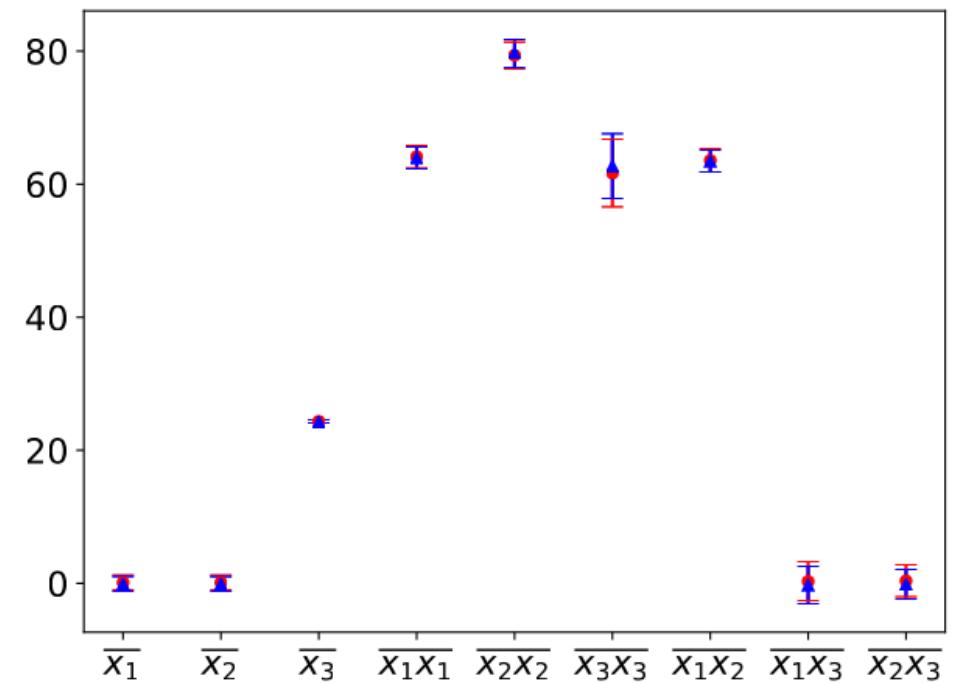
Case I: Moments

Red:True system

Blue: Modeled system



Without stochastic term

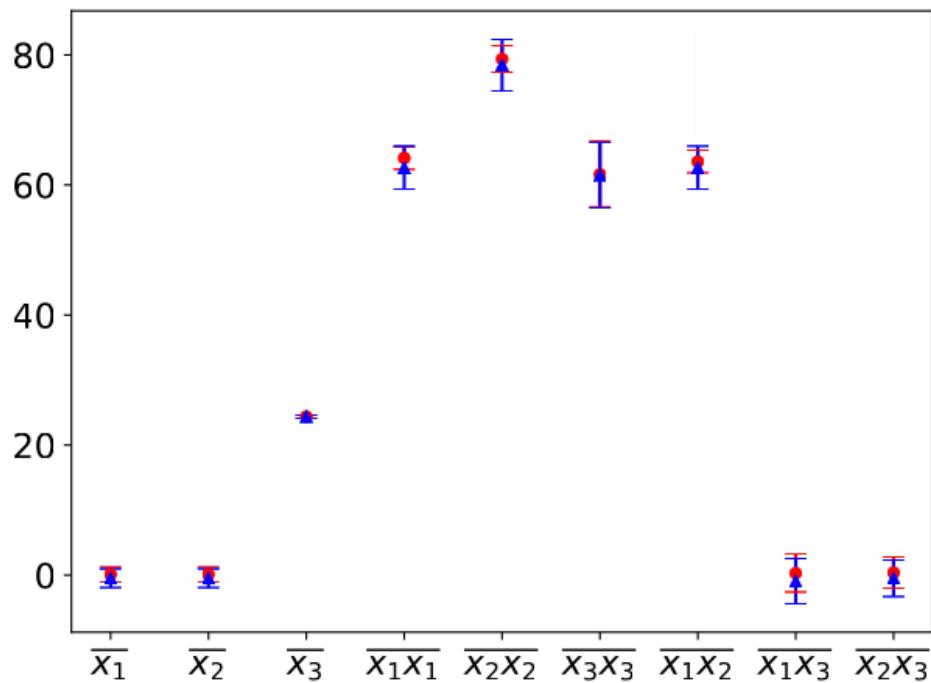


With stochastic term

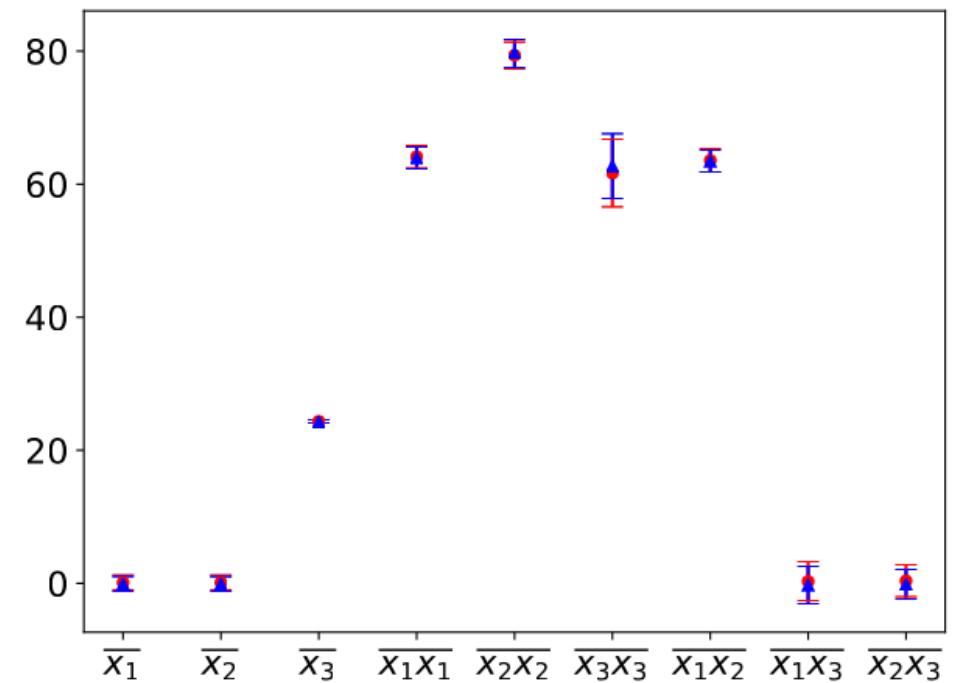
Case I: Moments

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Without stochastic term



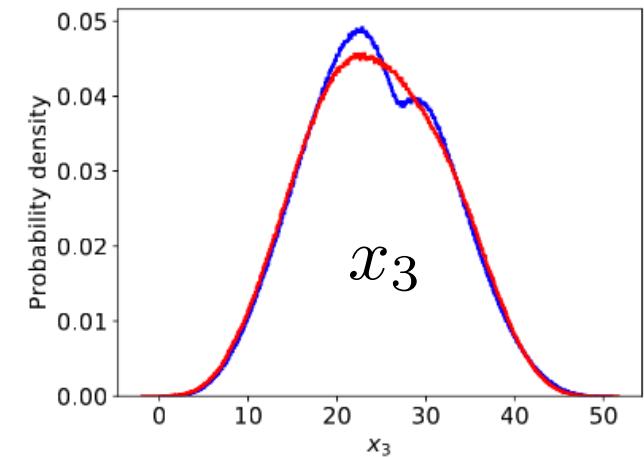
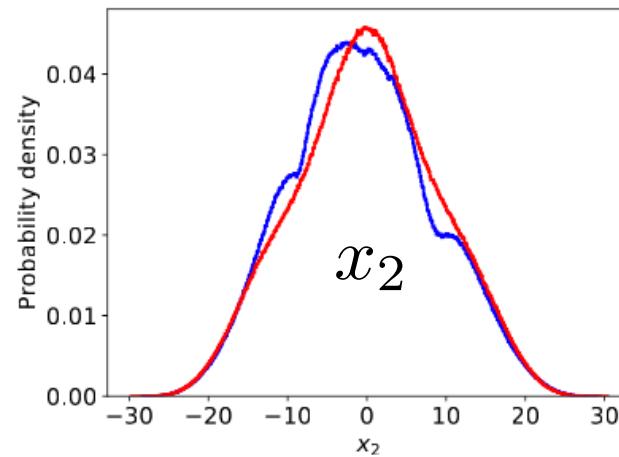
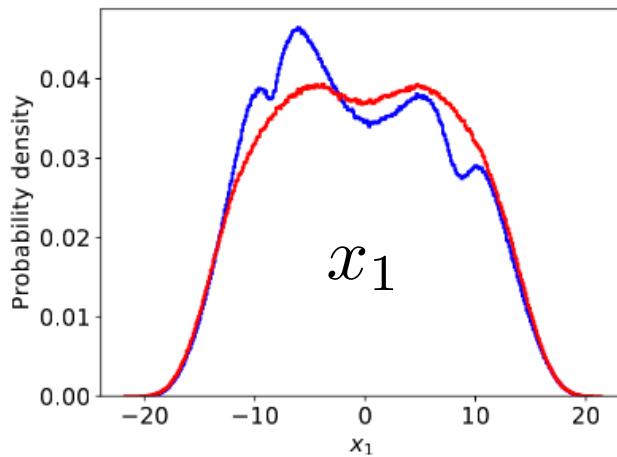
With stochastic term

- ❖ Does it mean the two models are equally good?

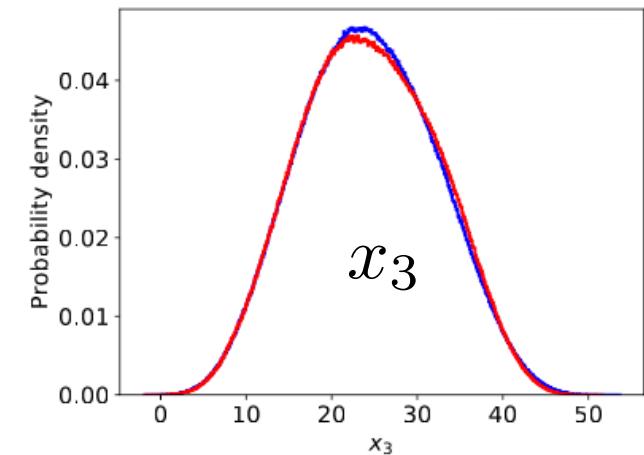
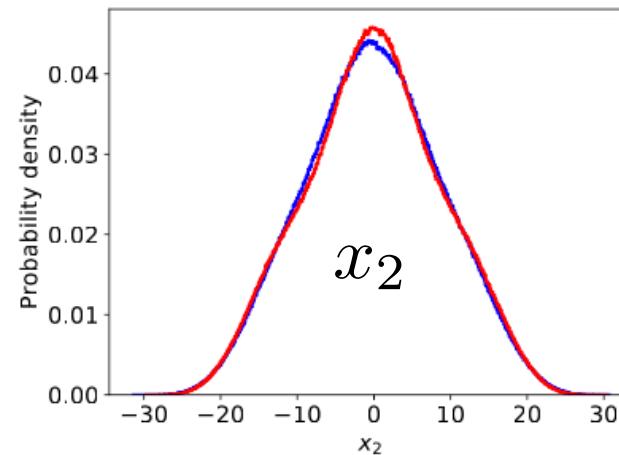
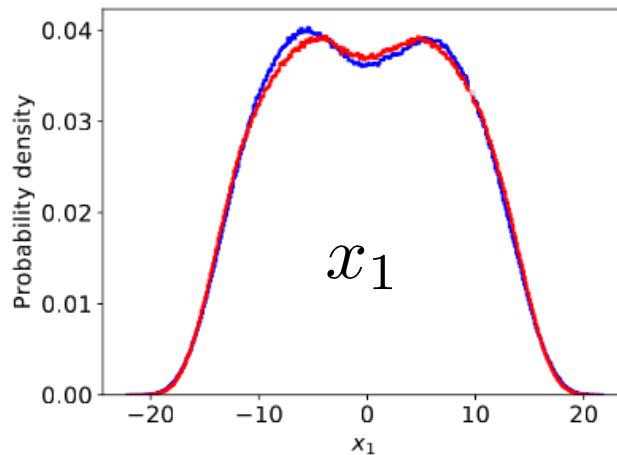
Case I: Invariant Measures

Red: True system

Blue: Modeled system



Without stochastic term



With stochastic term

Case II: Reduced Lorenz 63 System

- ❖ Lorenz 63 system in transformed coordinates $\dot{a} = g(a)$:

$$g_1(a) = 2.3a_1 - 6.2a_3 - 0.49a_1a_2 - 0.57a_2a_3,$$

$$g_2(a) = -62 - 2.7a_2 + 0.49a_1^2 - 0.49a_3^2 + 0.14a_1a_3,$$

$$g_3(a) = -0.63a_1 - 13a_3 + 0.43a_1a_2 + 0.49a_2a_3.$$

- ❖ Reduced-order model:

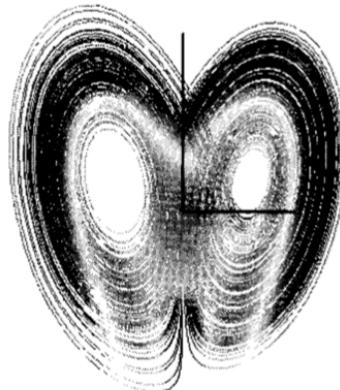
$$\dot{a}_1 = 2.3a_1 - 0.49a_1a_2$$

$$\dot{a}_2 = -62 - 2.7a_2 + 0.49a_1^2$$

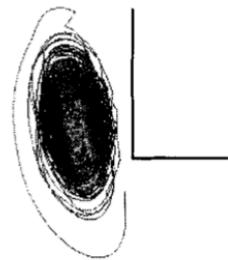
Case II: Reduced Lorenz 63 System

- ❖ Comparison of trajectories:

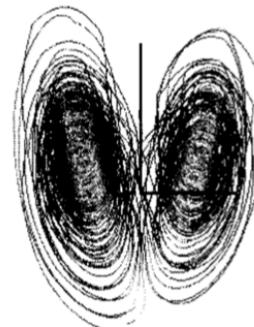
True system



Reduced-order system



Reduced-order system with simple stochastic term (white noises)



Case II: Reduced Lorenz 63 System

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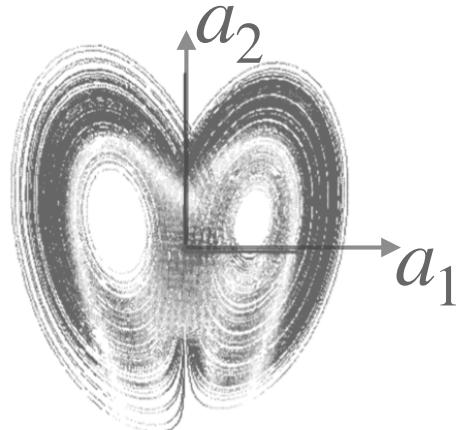
- ❖ Reduced-order model:

$$\dot{a}_1 = 2.3a_1 - 0.49a_1a_2 + M_1(a_1, a_2) + \sqrt{\sigma_1(a_1, a_2)}\dot{W}$$

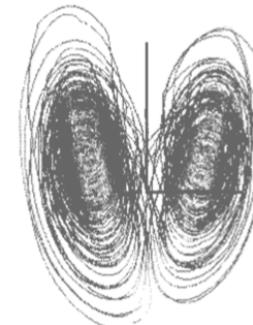
$$\dot{a}_2 = -62 - 2.7a_2 + 0.49a_1^2 + M_2(a_1, a_2) + \sqrt{\sigma_2(a_1, a_2)}\dot{W}$$

Case II: Reduced Lorenz 63 System

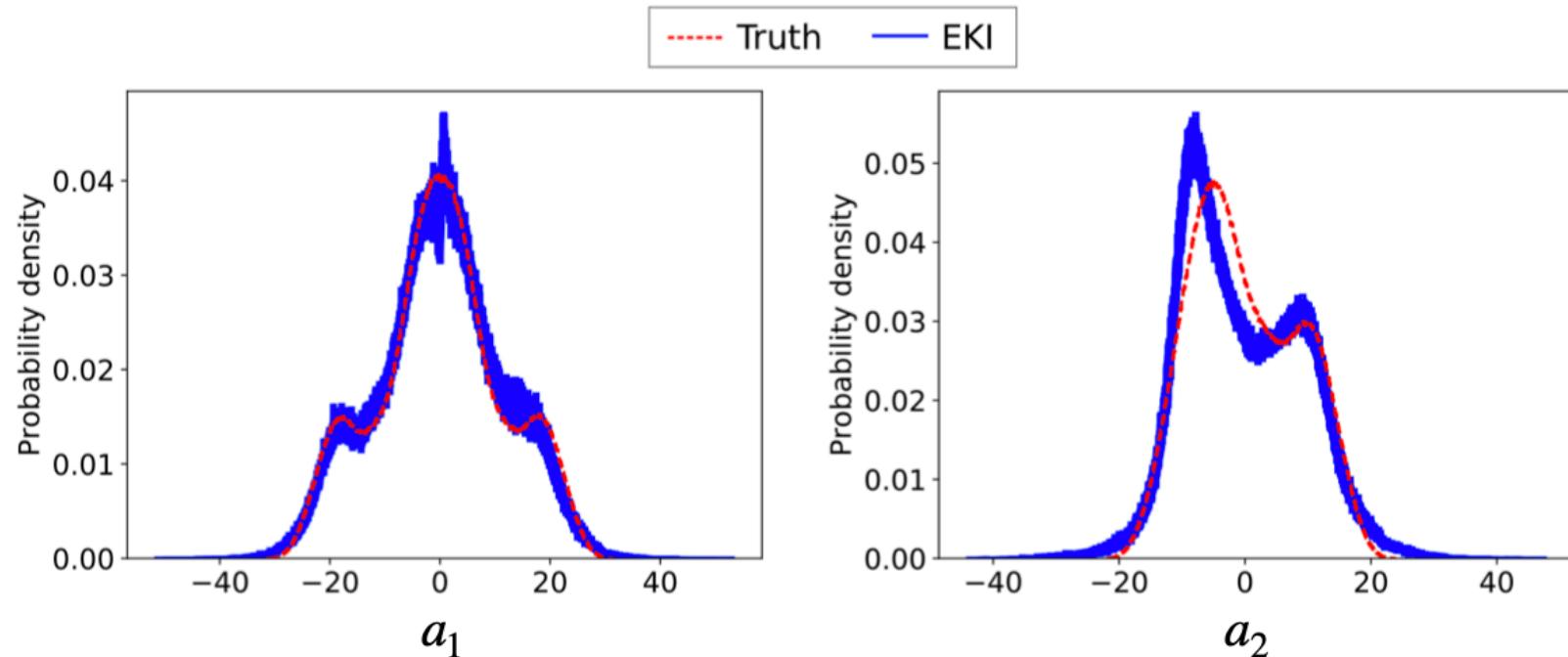
True system



Reduced-order system with simple stochastic term (white noises)



Palmer, Quarterly Journal of the Royal Meteorological Society, 127(572), 279-304 (2001)



Sparsity-Promoting Ensemble Kalman Inversion

Ensemble Kalman Inversion (EKI)

- ❖ Inverse problem to be solved:

$$y = G(\theta) + \eta, \quad \eta \sim N(0, \Gamma)$$

- ❖ EKI formula (j: index of ensemble, m: index of EKI step):

$$\theta_{m+1}^{(j)} = \theta_m^{(j)} + C_m^{\theta G} \left(C_m^{GG} + \Gamma \right)^{-1} \left(y_{m+1}^{(j)} - G(\theta_m^{(j)}) \right)$$

- ❖ Advantage of EKI:

- ❖ Derivative-free optimization;
- ❖ Robust to noisy evaluations of the forward map G ;
- ❖ The core task is equivalent to a convex optimization problem.

Ensemble Kalman Inversion (EKI)

- ❖ Introduce a new variable $w = G(\theta)$:

$$\begin{aligned}v &= (\theta, w)^\top, \\ \Psi(v) &= (\theta, G(\theta))^\top.\end{aligned}$$

- ❖ Formulating a noisy observed dynamical system ($Hv = w$, $H^\perp v = \theta$):

$$\begin{aligned}v_{m+1} &= \Psi(v_m) \\ y_{m+1} &= Hv_{m+1} + \eta_{m+1}.\end{aligned}$$

Ensemble Kalman Inversion (EKI)

- ❖ Introduce a new variable $v = G(\theta)$:

$$v = (\theta, w)^\top,$$
$$\Psi(v) = (\theta, G(\theta))^\top.$$

- ❖ Formulating a noisy observed dynamical system ($Hv = w, H^\perp v = \theta$):

$$v_{m+1} = \Psi(v_m)$$
$$y_{m+1} = Hv_{m+1} + \eta_{m+1}.$$

- ❖ Solving the optimization problem:

$$J_m^{(j)}(v) := \frac{1}{2} |y_{m+1}^{(j)} - Hv|_\Gamma^2 + \frac{1}{2} |v - \Psi(v_m^{(j)})|_{C_m^{\Psi\Psi}}^2,$$
$$v_{m+1}^{(j)} = \arg \min_v J_m^{(j)}(v).$$

Sparse EKI

- ❖ Objective function of standard EKI:

$$\frac{1}{2}v^\top \left(H^\top \Gamma^{-1} H + C_m^{-1} \right) v - \left(C_m^{-1} \Psi(v_m^{(j)}) + H^\top \Gamma^{-1} y^{(j)} \right)^\top v.$$

- ❖ Imposing sparsity within EKI (recall that $H^\perp v = \theta$):

$$\begin{aligned} \min_v \quad & \frac{1}{2}v^\top Qv + q^\top v \\ \text{s.t.} \quad & |H^\perp v|_{\ell_1} \leq \gamma. \end{aligned}$$

Sparse EKI

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- ❖ Introduce variables $v_i = v_i^+ - v_i^-$ and $|v_i| = v_i^+ + v_i^-$:

$$\begin{aligned} \min_{v^+, v^-} \quad & \frac{1}{2} (v^+ - v^-)^\top Q (v^+ - v^-) + q^\top (v^+ - v^-) \\ \text{s.t.} \quad & H^\perp (v^+ + v^-) \leq \gamma, \quad v^+ \geq 0, \quad v^- \geq 0. \end{aligned}$$

Case III: Noisy Lorenz 63 System

True System: $\dot{x} = \tilde{f}(x) + \sqrt{\sigma}W$

Where $\tilde{f}(x)$ is the standard Lorenz 63 system :

$$\frac{dx_1}{dt} = \alpha(x_2 - x_1)$$

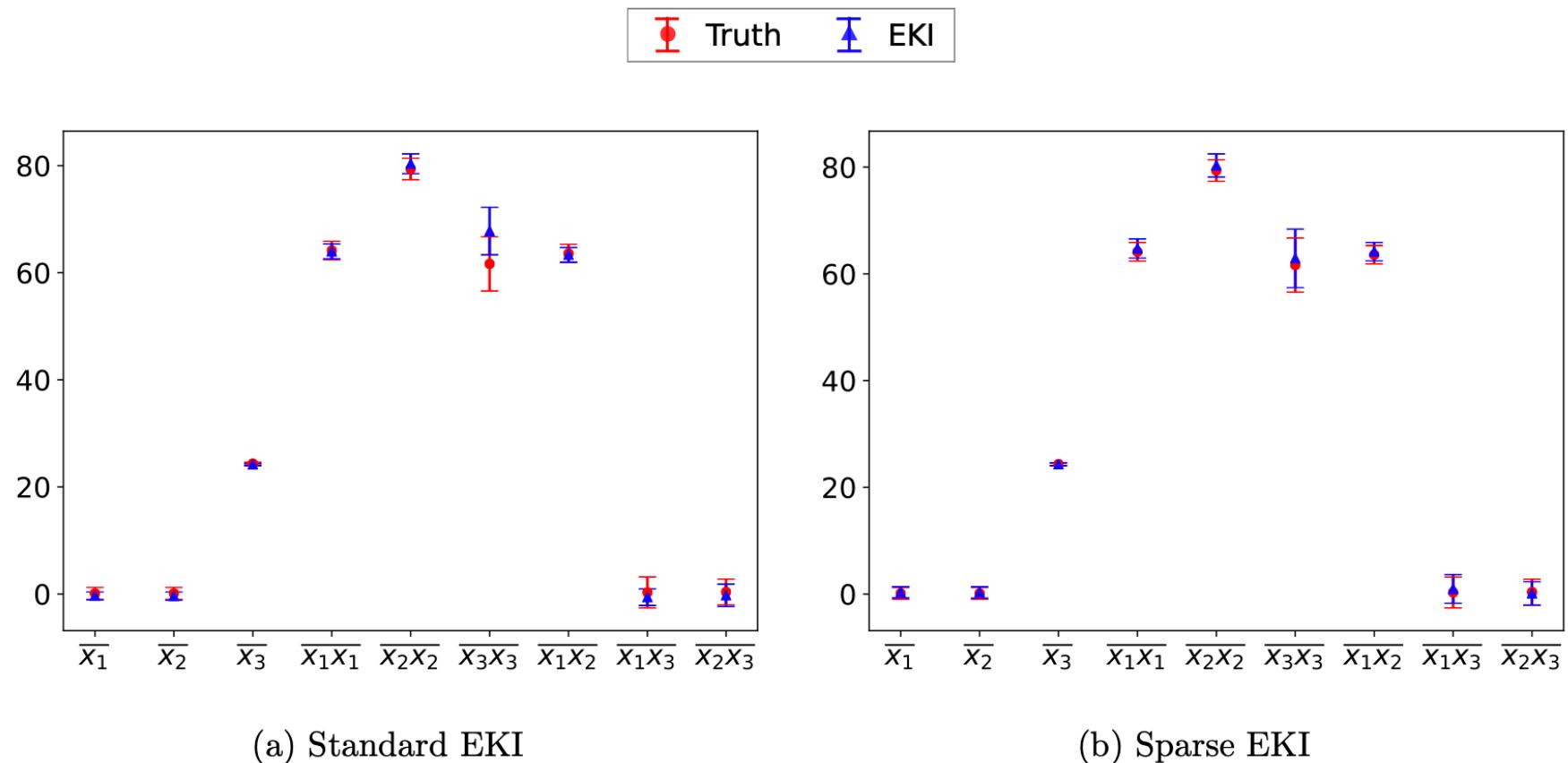
$$\frac{dx_2}{dt} = x_1(\rho - x_3) - x_2$$

$$\frac{dx_3}{dt} = x_1x_2 - \beta x_3$$

Modeled system: $\dot{X}_k = \sum_{i=1}^9 \theta_{ki} \phi_i(X) + \sqrt{\sigma}W_k, \quad k = 1, 2, 3$

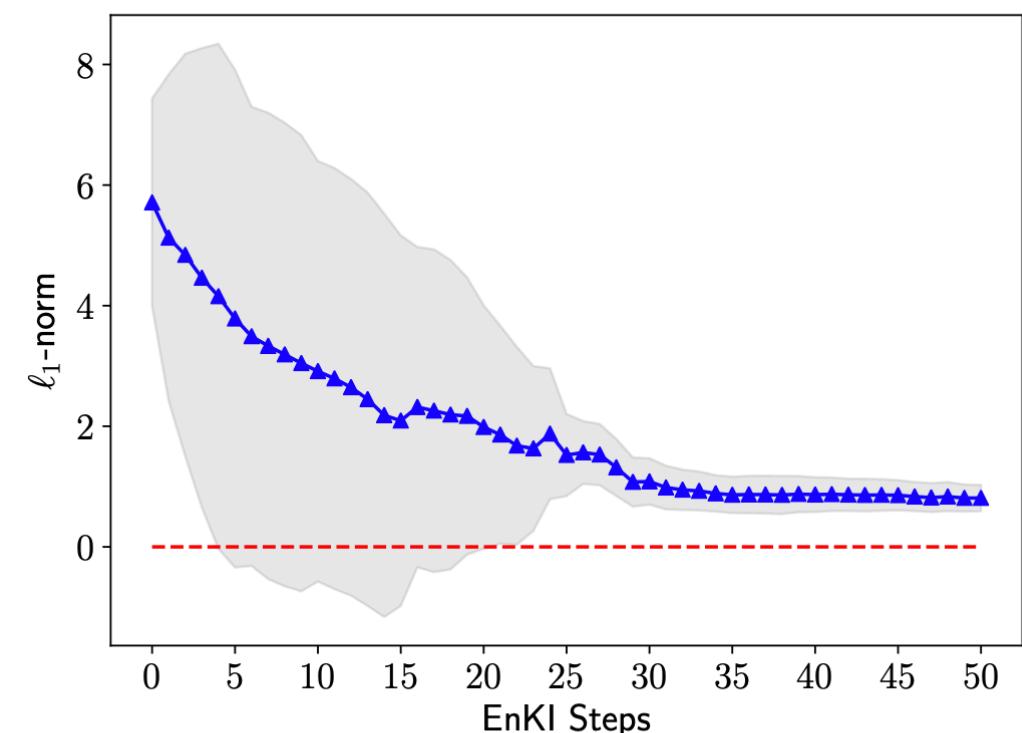
$\{\phi_i\}$ contains all the first and second-order polynomial basis functions

Case III: Moments

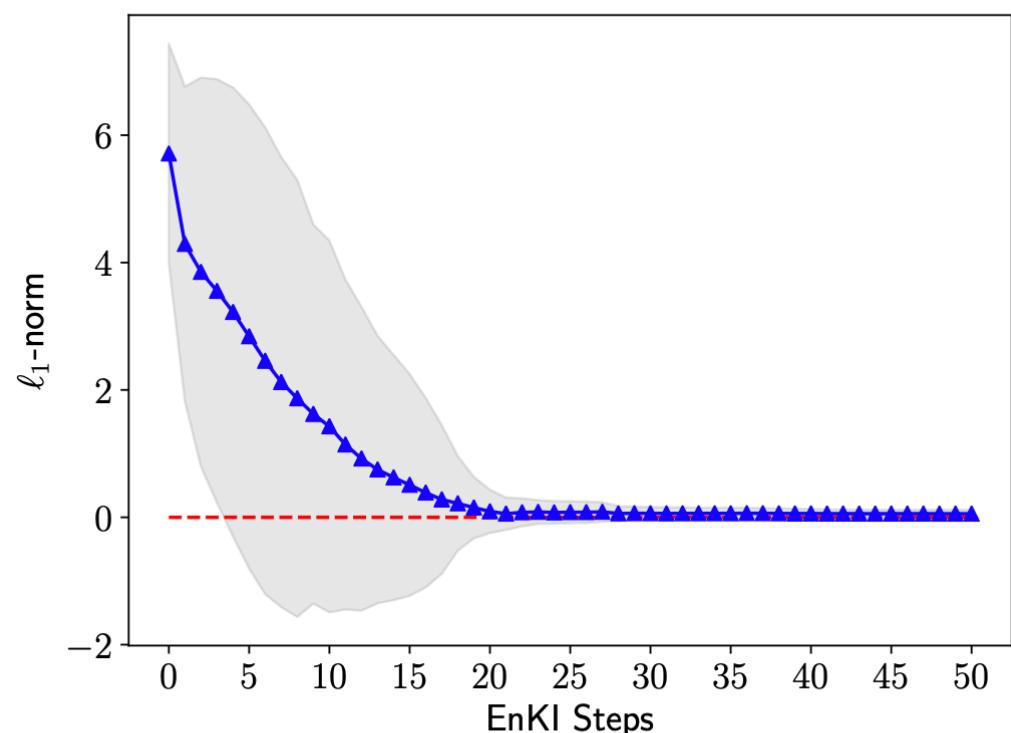


Case III: Redundant Coefficients

Ensemble mean $\pm 2\sigma_{\text{std}}$ Truth



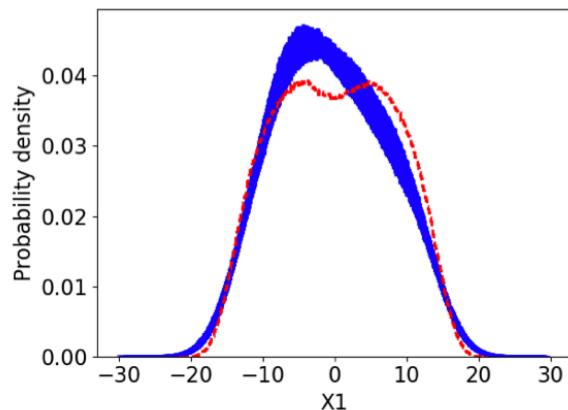
(a) Standard EKI



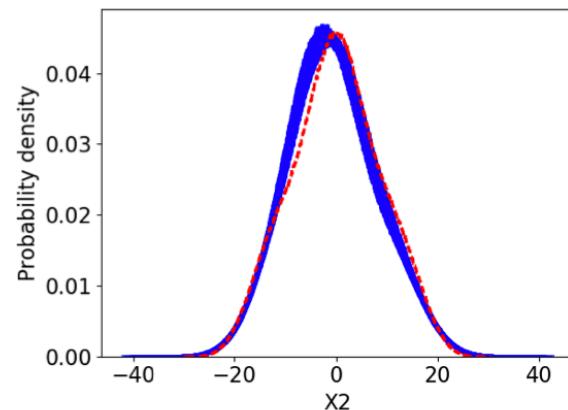
(b) Sparse EKI

Case III: Invariant Measures

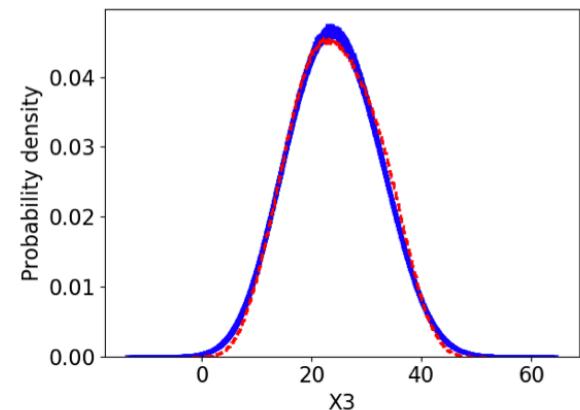
--- Truth — EKI



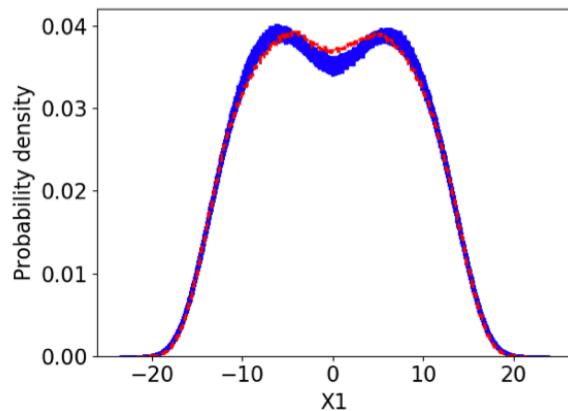
(a) X_1 (Standard EKI)



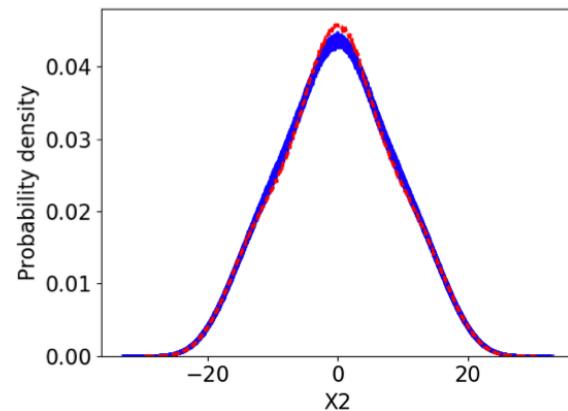
(b) X_2 (Standard EKI)



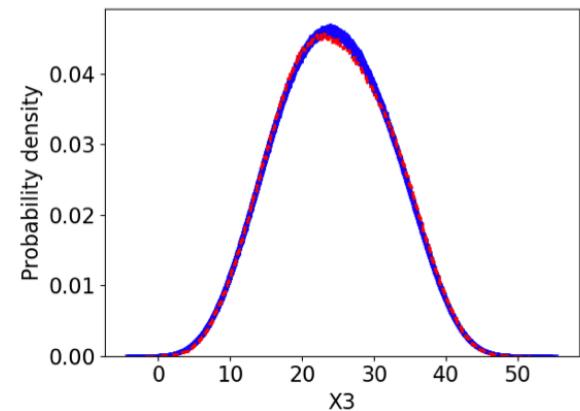
(c) X_3 (Standard EKI)



(d) X_1 (Sparse EKI)



(e) X_2 (Sparse EKI)



(f) X_3 (Sparse EKI)

Summary

- ❖ Many complex systems (e.g., earth system) involve **multi-scale dynamics**, and there is usually **no clear scale separation** between resolved and unresolved scales, which justifies the use of **stochastic closures**.
- ❖ We demonstrated the learning of stochastic closures based on **time-averaged statistics** using ensemble Kalman inversion (EKI).
- ❖ With the assumption that simpler explanation is often better (Occam's razor), we imposed **sparsity constraint** into EKI and showed improved performance of learning stochastic closures.

More details can be found at:

Schneider, Stuart, Wu. Learning Stochastic Closures Using Ensemble Kalman Inversion. arXiv:2004.08376 (2020)

Schneider, Stuart, Wu. Ensemble Kalman Inversion for Sparse Learning of Dynamical Systems from Time-Averaged Data. arXiv:2007.06175 (2020)

Thanks!