

$$\text{Q1W } f(k) = \frac{g}{(2\pi)^3} \frac{1}{\left(\exp\left(\frac{\epsilon - \mu}{k_B T}\right) \pm 1\right)}$$

$$\epsilon = \sqrt{k^2 + m^2}$$

$$2dk = kdk$$

$$P = \int \epsilon f(k) d^3k = \frac{g}{2\pi^2} \int \frac{k^3 \epsilon dk}{\left(\exp\left(\frac{\epsilon - \mu}{k_B T}\right) \pm 1\right)}$$

$$n = \int f(k) d^3k = \frac{g}{2\pi^2} \int \frac{k^3 dk}{\exp\left(\frac{\epsilon - \mu}{k_B T}\right) \pm 1}$$

$$P = \int \frac{k^2}{3\epsilon} f(k) d^3k = \frac{g}{6\pi^2} \int \frac{k^4 dk}{\epsilon \left(\exp\left(\frac{\epsilon - \mu}{k_B T}\right) \pm 1\right)}$$

$$\frac{dP}{dT} = \frac{g}{6\pi^2} \int \frac{k^4 dk}{\epsilon \left(\exp\left(\frac{\epsilon - \mu}{k_B T}\right) \pm 1\right)^2} \times e^{\left(\frac{\epsilon - \mu}{k_B T}\right)} \times \frac{(\epsilon - \mu)}{k_B T^2} = \frac{g}{6\pi^2} \int \frac{k^4 dk}{\epsilon \left(\exp\left(\frac{\epsilon - \mu}{k_B T}\right) \pm 1\right)^2} \times \frac{e^{\frac{\epsilon - \mu}{k_B T}} (\epsilon - \mu)}{k_B T^2}$$

$$\frac{d}{dT} \left( \frac{1}{\left(\exp\left(\frac{\epsilon - \mu}{k_B T}\right) \pm 1\right)} \right) = - \frac{e^{\frac{\epsilon - \mu}{k_B T}}}{\left(\exp\left(\frac{\epsilon - \mu}{k_B T}\right) \pm 1\right)^2} \frac{1}{k_B T} \frac{d\epsilon}{d\mu} dk$$

$$\text{Let } \left( \frac{1}{\left(\exp\left(\frac{\epsilon - \mu}{k_B T}\right) \pm 1\right)} \right) = A$$

Then

$$\frac{dP}{dT} = - \frac{g}{6\pi^2} \int k^3 dA (\epsilon - \mu)$$

$$= - \frac{g}{6\pi^2} \left[ \int \epsilon k^3 dA - \mu \int k^3 dA \right] = \frac{g}{6\pi^2} \left[ \int k^3 \epsilon dA - \mu \int k^3 dA \right]$$

$$\epsilon dA = d(\epsilon A) - A d\epsilon$$

$$= - \frac{g}{6\pi^2} \left[ \int k^3 d(\epsilon A) - A d\epsilon - \mu \int k^3 dA \right]$$

$$= - \frac{g}{6\pi^2} \left[ \int k^3 d\epsilon A - \int k^3 A d\epsilon - \mu \int k^3 dA \right]$$

$$= -\frac{g}{6\pi^2 T} \left[ \cancel{\int_0^\infty k^3 \epsilon dk} - \int_0^\infty 3k^2 \epsilon dk - \mu \cancel{\int_0^\infty k^3 dk} + \mu \int_0^\infty 3k^2 dk - \int_0^\infty k^4 dk \frac{1}{\epsilon} \right]$$

$$= \frac{g}{6\pi^2 T} \left[ \int_0^\infty \frac{3k^2 \epsilon dk}{(\exp \frac{\epsilon - \mu}{k_B T} \pm 1)} + \int_0^\infty \frac{k^4 dk}{\epsilon (\exp \frac{\epsilon - \mu}{k_B T} \pm 1)} - \mu \int_0^\infty \frac{3k^2 dk}{(\exp \frac{\epsilon - \mu}{k_B T} \pm 1)} \right]$$

$$= \frac{1}{T} \left[ \frac{g}{2\pi^2} \int_0^\infty \frac{k^3 \epsilon dk}{(\exp \frac{\epsilon - \mu}{k_B T} \pm 1)} + \frac{g}{6\pi^2} \int_0^\infty \frac{k^4 dk}{\epsilon (\exp \frac{\epsilon - \mu}{k_B T} \pm 1)} - \mu \frac{g}{2\pi^2} \int_0^\infty \frac{k^3 dk}{(\exp \frac{\epsilon - \mu}{k_B T} \pm 1)} \right]$$

$$= \frac{1}{T} (\mathcal{P} + P - \mu n)$$

$$\text{So } \frac{dp}{dT} = \frac{(\mathcal{P} + P - \mu n)}{T}$$

$$\sigma = \frac{dp}{dT} = \left( \frac{\mathcal{P} + P - \mu n}{T} \right)$$

1. Given cross section times velocity of the equation

$$\langle \sigma v \rangle \sim 7.4 \times 10^{-26} \text{ cm}^3 \text{ s}^{-1}$$

The rate of interaction is  $\Gamma = n \langle \sigma v \rangle$

The rate per neutron of  $np \rightarrow {}^2\text{H} \gamma$  is:

$$\Gamma = n_p \langle \sigma v \rangle = \eta n_\gamma \langle \sigma v \rangle$$

[ we can take  $\eta \sim \frac{n_p}{n_\gamma}$  ]

$$= \eta n_\gamma \times 7.4 \times 10^{-26} \text{ m}^3 \text{ sec}^{-1}$$

Substituting for  $n_\gamma = 2.85 \times 10^{37} \left( \frac{T}{1 \text{ MeV}} \right)^3$  we get.

$$\Gamma = \eta \times 2 \times 10^{12} \text{ sec}^{-1} \left( \frac{T}{1 \text{ MeV}} \right)^3$$

We know that the expansion rate is given by

$$H \sim 0.6 \text{ sec}^{-1} \left( \frac{T}{1 \text{ MeV}} \right)^2$$

Substituting for  $T = 60 \text{ keV}$

$$\Gamma = \eta \times 2 \times 10^{12} \text{ sec}^{-1} \left( \frac{60 \text{ keV}}{1 \text{ MeV}} \right)^3$$

$$H = 0.6 \text{ sec}^{-1} \left( \frac{60 \text{ keV}}{1 \text{ MeV}} \right)^2$$

To find the criteria for  $\eta$  where nucleosynthesis is possible, we need  $\Gamma > H$  i.e.

$$\text{we get } \eta > 4 \times 10^{-12}$$

⑦ Given  $n > 4 \times 10^{-12}$  and  $\mu_2 = \mu_p + \mu_n$ .

Given the Saha equation:

$$\frac{n_2}{n_p n_n} = \left( \frac{2\pi}{m_n T} \frac{2\pi}{m_p T} \frac{m_2 T}{2\pi} \right)^{3/2} e^{B/T}$$

$$\frac{n_2}{n_n} = \left( \frac{T}{m_p} \right)^{3/2} e^{B/T} \cdot \left( \frac{4\pi^2}{m_n T^2} \right)^{3/2} n_p$$

Substituting for  $n_p = g \left( \frac{1.2}{\pi^2} \right) T^3 \times \frac{3}{4}$  & using the formula for chemical potential we get.

$$\frac{n_2}{n_n} = \left( \frac{T}{m_p} \right)^{3/2} e^{B/T}$$

=

2. For a collisionless motion, particle freely propagate in phase space according to the Liouville equation i.e. if  $f$  is the distribution function, then,

$$\frac{df}{d\lambda} = 0.$$

Let  $f$  be a function of  $x$  and  $p$  i.e.  $f = f(x^\mu, p^\mu)$ .

Then  $\frac{df}{d\lambda} = \frac{\partial f}{\partial x^\mu} \frac{dx^\mu}{d\lambda} + \frac{\partial f}{\partial p^\mu} \frac{dp^\mu}{d\lambda}$ , where  $\lambda$  is the affine parameter.

Now let us consider the geodesic equation on momentum, i.e.,

$$\frac{dp^\mu}{d\lambda} + \Gamma_{\alpha\beta}^\mu p^\alpha p^\beta = 0$$

$$\Rightarrow \frac{dp^\mu}{d\lambda} = -\Gamma_{\alpha\beta}^\mu p^\alpha p^\beta$$

$$\Rightarrow \frac{df}{d\lambda} = \frac{\partial f}{\partial x^\mu} \frac{dx^\mu}{d\lambda} + \frac{\partial f}{\partial p^\mu} (-\Gamma_{\alpha\beta}^\mu p^\alpha p^\beta)$$

~~also we can write~~  $\frac{dx^\mu}{d\lambda} = p^\mu$ , then,

$$\frac{df}{d\lambda} = 0 = \frac{\partial f}{\partial x^\mu} p^\mu - \Gamma_{\alpha\beta}^\mu p^\alpha p^\beta \frac{\partial f}{\partial p^\mu} = 0$$

$$\Rightarrow \frac{\partial f}{\partial x^\mu} p^\mu - \Gamma_{\alpha\beta}^\mu p^\alpha p^\beta \frac{\partial f}{\partial p^\mu} = 0 \quad \text{--- (a)}$$

We know that  $g_{\mu\nu} p^\mu p^\nu = -m^2$  for massive particle and

Let us find out  $g_{\mu\nu} p^\mu p^\nu = 0$  for massless particle.

Then we are considering photons, so  $g_{\mu\nu} p^\mu p^\nu = 0$

Let us find what  $\Gamma_{ij}^0 p^i p^j$  is.

$$\Gamma_{ij}^0 p^i p^j = -g^{00} g_{ij} \delta_{ij} \frac{a(t)}{a(t)} p^i p^j \\ = -g^{00} \delta_{ij} \frac{a(t)}{a(t)} p^i p^j = 0 \quad (\text{for photons})$$

So, eqn (a) becomes,

$$p^4 \frac{\partial f}{\partial x^4} - \Gamma_{\alpha\beta}^i p^\alpha p^\beta \frac{\partial f}{\partial p^i} = 0 \quad - (b)$$

For homogeneous universe, the distribution function is independent of position. Then eqn (b) becomes,

$$p^0 \frac{\partial f}{\partial x^0} - \Gamma_{\alpha\beta}^i p^\alpha p^\beta \frac{\partial f}{\partial p^i} = 0 \quad - (c)$$

For a flat universe in FRW metric, ( $k=0$ ), non zero  $\Gamma$  are

$$\text{for } \Gamma_{ij}^i = \frac{1}{2} g^{ii} \partial_0 g_{ij} \delta_{ij} \\ = \frac{1}{2} g^{ii} g_{ij}^2 \frac{\dot{a}(t)}{a(t)} \delta_{ij} = \frac{\dot{a}(t)}{a(t)} = H.$$

rest all terms are 0.  $[ds^2 = c^2 dt^2 - a^2(t)(dx^2 + dy^2 + dz^2)]$

Then eqn (c) becomes,

$$p^0 \frac{\partial f}{\partial x^0} - \Gamma_{ij}^i p^i p^j \frac{\partial f}{\partial p^i} = 0$$

$$\Rightarrow \frac{\partial f}{\partial t} - \Gamma_{0i}^i p^i \frac{\partial f}{\partial p^i} = 0 \quad [i=j, \text{ then only } \Gamma_{0j}^j \text{ exist}]$$

$$\Rightarrow \frac{\partial f}{\partial t} - \frac{\dot{a}}{a} p^i \frac{\partial f}{\partial p^i} = 0$$



$$= a \frac{\partial f}{\partial t} \frac{\partial f}{\partial t} = -\frac{a}{a} p_i \frac{\partial f}{\partial p_i} \quad \text{---(d)}$$

The number density  $n = \int \frac{d^3p}{(2\pi)^3} f = \frac{1}{2\pi^2} \int p^2 f dp$

$$= \frac{1}{2\pi^2} \left[ \frac{p^3}{3} f \right]_0^\infty - \frac{1}{2\pi^2} \int p^3 \frac{\partial f}{\partial p} dp$$

$$p^3 f \rightarrow 0 \text{ as } p \rightarrow 0$$

$$p^3 f \rightarrow 0 \text{ as } p \rightarrow \infty \text{ as } f \rightarrow 0$$

$$\text{So } n = -\frac{1}{3} \frac{1}{2\pi^2} \int p^3 \frac{\partial f}{\partial p} dp = -\frac{1}{3} \frac{1}{2\pi^2} \int p \frac{\partial f}{\partial p} d^3p$$

$$\Rightarrow -3n = \frac{1}{2\pi^2} \int p \frac{\partial f}{\partial p} d^3p \quad \text{---(e)}$$

So in equation d, if we integrate by  $d^3p$  on both sides, we get

$$\int \frac{1}{(2\pi)^3} \frac{\partial f}{\partial t} d^3p = -\frac{a}{a} \int p_i \frac{\partial f}{\partial p_i} \frac{d^3p}{(2\pi)^3} \quad \text{---(f)}$$

$$n = \int \frac{d^3p}{(2\pi)^3} f$$

$$\Rightarrow \frac{\partial n}{\partial t} = \int \frac{d^3p}{(2\pi)^3} \frac{\partial f}{\partial t} \quad \text{---(g)}$$

Using eqn (g) and (e) on eqn (f) we get

$$\frac{\partial n}{\partial t} = -\frac{a}{a} - 3n$$

$$= -3Hn$$

$$\Rightarrow \dot{n} + 3Hn = 0$$

$$\frac{dn}{n} = -3 \frac{da}{a} \Rightarrow \ln n = \ln a^{-3} + k$$

$$\Rightarrow n = K a^{-3} \quad n \propto a^{-3}$$