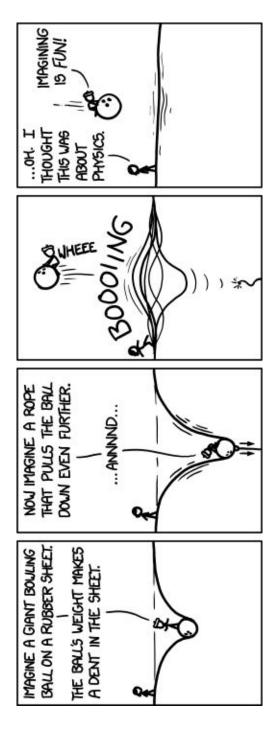
#### Maxwell's Equations in General Relativity

Joseph P. Glaser



# Background on Tensors

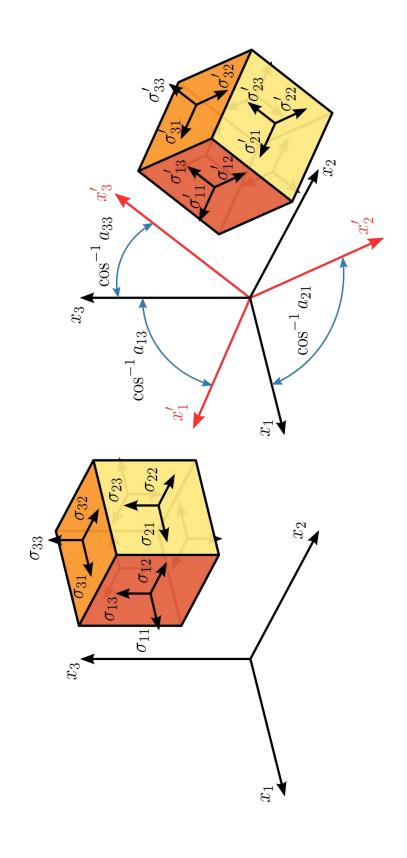






#### Tensor Definition

Tensors are simply arrays of numbers, or functions, that transform according to certain rules under a change of coordinates.





### **Examples of Tensors**

Tensors are classified as: (n, m) Tensor of Rank (n + m), where:

- n is the number of contravariant indices
- m is the number of covariant indices

m = 2	Metric Tensor		Elasticity Tensor
<i>m</i> = 1	Covector	Linear Transform	
m=0	Scalar	Vector	Inverse Metric Tensor
n, m	0 = u	<i>n</i> = 1	n = 2



## Useful Operations on Tensors

Einstein Summation Notation:

$$\geq$$

$$y = \sum_{i=0}^N v_i w^i = v_i w^i$$

Raising & Lowering a Tensor's Index:

$$g^{ij}A_j = A^i \quad g_{ij}A^j = A_i$$

# The Electromagnetic Field Tensor





### Maxwell's Equations

We want the following to hold in General Relativity:

 $abla \cdot ec{E} = rac{
ho}{\epsilon_0}$ 

Gauss's Law [Magnetism] 
$$\rightarrow \nabla \cdot \vec{B} =$$



## Introducing the Four-Current

Assume a charge density,  $\rho$ , is moving with some velocity,  $\mathbf{v}$ .

The four-current density is thus:

$$J^i = \gamma[\rho, \vec{J}]$$

where the electrical current density as seen in the lab-frame is:

$$ec{J}=ec{eta}
ho$$





$$\frac{\nabla \cdot \vec{E} = 4\pi k\rho}{\partial x} + \frac{\partial E_y}{\partial y} + \frac{\partial E_z}{\partial z} = 4\pi k\rho$$

$$\frac{\partial F^{00}}{\partial x} + \frac{\partial F^{01}}{\partial y} + \frac{\partial F^{02}}{\partial z} = 4\pi kJ^0$$

$$\frac{\partial F^{00}}{\partial t} + \frac{\partial F^{01}}{\partial x} + \frac{\partial F^{02}}{\partial y} + \frac{\partial F^{03}}{\partial z} = 4\pi kJ^0$$



## Satisfying the Coulomb Force

$$\frac{\mathrm{d}\vec{p}}{\mathrm{d}t} = q\vec{E}$$

$$rac{\mathrm{d} p^i}{\mathrm{d} au} = q F^{ij} u_j$$

where:

$$p^i=mu^i$$

$$u^i = \gamma[1, \vec{v}]$$

Noting the relationship:

$$2p_i \frac{\mathrm{d}p^i}{\mathrm{d}\tau} = \frac{\mathrm{d}\vec{p} \cdot \vec{p}}{\mathrm{d}\tau} = 0$$



## Satisfying the Coulomb Force

$$2p_i \frac{\mathrm{d}p^i}{\mathrm{d}\tau} = 2(mu_i)qu_j F^{ij} = 0$$

If we suppose that F is anti-symmetric and zero along the diagonal, the above equation holds for all four-velocities.

$$\begin{bmatrix}
0 & E_x & E_y & E_z \\
-E_x & 0 & a & b \\
-E_y & -a & 0 & c \\
-E_y & -b & -c & 0
\end{bmatrix}$$



### Satisfying Lorentz Force

$$\frac{\mathrm{d}p^{x}}{\mathrm{d}\tau} = qF^{ij}u_{j}$$

$$\frac{\mathrm{d}p^{x}}{\mathrm{d}\tau} = q(-E_{x}u_{0} + 0 + au_{2} + bu_{3})$$

$$= q\gamma(E_{x} + av_{y} + bv_{z})$$

$$= q\gamma(E_{x} + B_{z}v_{y} + (-B_{y})v_{z})$$

$$= q\gamma(E_{x} + B_{z}v_{y} + (-B_{y})v_{z})$$

$$= q\gamma(E_{x} + [\vec{v} \times \vec{B}]_{x})$$



# The Electromagnetic Field Tensor

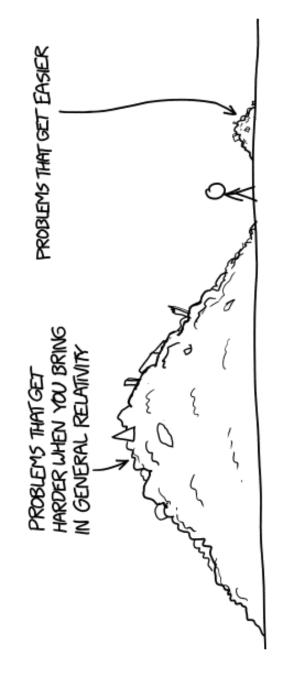


### Confirming Ampere's Law

Try out the x-component of the four-current density:

Try out the x-component of the four-current density: 
$$4\pi kJ^1=\partial_jF^{1j}$$
 
$$=\partial_jF^{10}+\partial_jF^{12}+\partial_3F^{13}$$
 
$$=\partial_0E_x+\partial_2B_z-\partial_3B_y$$
 
$$=-\frac{\partial E_x}{\partial t}+[\nabla\times\vec{B}]_x$$

#### The Reissner-Nordström Charged Mass Solution







#### The R-N Metric

$$= \begin{bmatrix} -f(r) & 0 & 0 & 0 \\ 0 & 1/f(r) & 0 & 0 \\ 0 & 0 & r^2 & 0 \\ 0 & 0 & r^2 sin(\theta) \end{bmatrix}$$

$$f(r) = \left(1 - \frac{2M}{r} + \frac{Q^2}{r^2}\right)$$



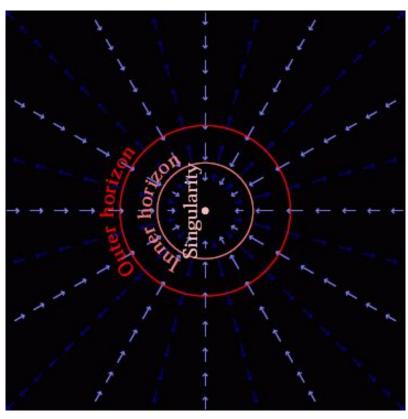
# Application to Charged Black Holes

$$r_{outer} = M + \sqrt{M^2 - Q^2}$$

$$r_{inner} = M - \sqrt{M^2 - Q^2}$$

$$v = \sqrt{\frac{2M(r)}{r}} = \frac{1}{\sqrt{1 - f(r)^2}}$$

$$r_{stop} = rac{Q^2}{2M}$$



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#### **Thanks!**

