

Drexel Physics 2005 Modern Qual Solutions

2014 entering class

September 15, 2015

1 Problem 1

We need expected value of the particle energy in the well to be less than $E = h\nu$. The energy levels of the finite square well are XXX. We then compute the partition function and take the derivative to find $\langle E \rangle$ as a function of temperature.

2 Problem 2

With $S = kT \ln \Omega$ and $\frac{\partial S}{\partial E} = \frac{1}{T}$, we need to add:

$$E = kT^2 \ln d\Omega = (\text{Joules}) = (\text{calories}) \quad (2.1)$$

to increase the number of accessible states by 1000.

3 Problem 3

Electron tunneling in transistors.

Angular momentum quantization as measured in the Stern-Gerlach experiment. When a beam of spin $\frac{1}{2}$ particles passes through a magnetic field they are split into two distinct beams instead of smeared over an angular region.

4 Problem 4

To compute the normalization constant we need to solve:

$$C^2 \int_V R^2 = 1 \quad (4.1)$$

$$\int (e^{-r/a})^2 r^2 dr = \frac{1}{C^2} \quad (4.2)$$

$$\int r^2 e^{-2r/a} dr = \frac{1}{C^2} \quad (4.3)$$

$$\frac{2}{(2/a)^3} = \frac{1}{C^2} \quad (4.4)$$

$$C = 2 \left(\frac{Z}{a_0} \right)^{3/2} \quad (4.5)$$

5 Problem 5

Free expansion of a gas is one irreversible process. The theory of microscopic reversability only holds for systems that remain at equilibrium.

6 Problem 6

The cloud of gas also has gravitational potential energy and electronic potential energy, so the total system energy may decrease even if the kinetic energy increases.

7 Problem 7

8 Problem A1

We see that $S_x = \frac{S_+ + S_-}{2}$. We find the eigenvalues in the usual manner:

$$S_x = \frac{\hbar}{2} \begin{bmatrix} 0 & \sqrt{2} & 0 \\ \sqrt{2} & 0 & \sqrt{2} \\ 0 & \sqrt{2} & 0 \end{bmatrix} \quad (8.1)$$

$$\begin{vmatrix} -\lambda & \sqrt{2}/2\hbar & 0 \\ \sqrt{2}/2\hbar & -\lambda & \sqrt{2}/2\hbar \\ 0 & \sqrt{2}/2\hbar & -\lambda \end{vmatrix} = 0 \quad (8.2)$$

$$-\lambda(\lambda^2 - \hbar^2/2) - \hbar\sqrt{2}/2(-\lambda\hbar\sqrt{2}/2) = 0 \quad (8.3)$$

$$-\lambda^3 + \lambda\hbar^2 = 0 \quad (8.4)$$

$$\lambda = \{-\hbar, \hbar\} \quad (8.5)$$

We now solve $S_x \mathbf{e} = \lambda \mathbf{e}$ for the eigenvectors.

$$e_1 = e_3 \quad (8.6)$$

$$e_2 = \frac{1}{\lambda} \sqrt{2}\hbar e_1 \quad (8.7)$$

$$\lambda = -\hbar : \mathbf{e}_{-\hbar} = \{1, -\sqrt{2}, 1\} \frac{1}{2} \quad (8.8)$$

$$\lambda = \hbar : \mathbf{e}_{\hbar} = \{1, \sqrt{2}, 1\} \frac{1}{2} \quad (8.9)$$

We now calculate the probability of measuring $S_x = \hbar$. (The question says $S + x = 1$, but 1 is not an eigenvalue of S_x .)

$$P(S_x = \hbar) = \langle e_{\hbar} | u \rangle^2 = \left(\frac{1}{2\sqrt{2}c} (1 * 1 + 4 * \sqrt{2} - 3 * 1) \right)^2 \quad (8.10)$$

$$P(S_x = \hbar) = \frac{1}{2c} (9 - 4\sqrt{2}) \quad (8.11)$$

After the measurement the system has a definite value of $S_x = \hbar$ and is in state e_{\hbar} . The eigenvector corresponding to $S_z = \hbar$ is $(1, 0, 0)$, so we calculate the probability the same way.

$$P(S_z = \hbar) = \langle \mathbf{e}_1 | \mathbf{e}_{\hbar} \rangle^2 = \left(1 * \frac{1}{2} + 0 + 0 \right)^2 \quad (8.12)$$

$$P(S_z = \hbar) = \frac{1}{4} \quad (8.13)$$

9 Problem A2

a) We open up the brackets and find $\Psi(x, 0) = 2x^3 - 10x^2 + 12x$.

$$\int_0^3 (C\Psi(x, 0))^2 dx = 1 \quad (9.1)$$

$$\frac{1}{C^2} = \left(\int_0^3 2x^3 - 10x^2 + 12x dx \right)^2 \quad (9.2)$$

$$\frac{1}{C^2} = \left(-\frac{1}{2}\right)^2 = \frac{1}{4} \quad (9.3)$$

$$C = 2 \quad (9.4)$$

$$\Psi(x, 0) = 2(2x^3 - 10x^2 + 12x) \quad (9.5)$$

b) The wavefunction has a single 0 in the range at $x = 2$, so it most closely resembles the standard wavefunction $\sin \frac{2\pi}{3}x$. c) The wavefunctions can be derived from the Schrodinger equation:

$$-\frac{\hbar^2}{2m} \frac{d^2\psi}{dx^2} = E\psi \quad (9.6)$$

$$\frac{d^2\psi}{dx^2} = -\frac{2mE}{\hbar^2}\psi \quad (9.7)$$

With $k \equiv \frac{\sqrt{2mE}}{\hbar}$, the solutions are of the form $A \cos kx + B \sin kx$. The infinite square well requires $\psi(0) = 0$, so $A = 0$. This square well also requires $\psi(3) = 0$, so k takes on discrete values determined by:

$$\sin \frac{n\pi x}{3} = 0 \quad (9.8)$$

$$k = \frac{n\pi}{3}, \quad n = 1, 2, 3, \dots \quad (9.9)$$

$$\frac{\sqrt{2mE}}{\hbar} = \frac{n\pi}{3} \quad (9.10)$$

$$E = \frac{n^2\pi^2\hbar^2}{18m}, \quad n = 1, 2, 3, \dots \quad (9.11)$$

So we can estimate the expectation value of the energy for $n = 2$ as $\frac{2\pi^2\hbar^2}{9m}$.

10 Problem A3

The energies of states are $E_0 = \frac{1}{2}\hbar\omega$ and $E_2 = \frac{5}{2}\hbar\omega$ so the time-evolving wavefunction is:

$$|\Psi(t)\rangle = C_0 e^{i\frac{1}{2}\omega t} |0\rangle + C_2 e^{i\frac{5}{2}\omega t} |2\rangle \quad (10.1)$$

b) We can write the momentum and position operators in terms of the creation and annihilation operators, which come from “factoring” the Hamiltonian.

$$H = \frac{p^2}{2m} + \frac{1}{2}kx^2 \quad (10.2)$$

$$a = \left(\frac{m\omega}{2\hbar}\right)^{1/2} \hat{X} + i \left(\frac{1}{2m\omega\hbar}\right)^{1/2} \hat{P} \quad (10.3)$$

$$a^\dagger = -\left(\frac{m\omega}{2\hbar}\right)^{1/2} \hat{X} + i \left(\frac{1}{2m\omega\hbar}\right)^{1/2} \hat{P} \quad (10.4)$$

$$\hat{P} = -\frac{i}{2}\sqrt{2m\omega\hbar}(a + a^\dagger) \quad (10.5)$$

$$\hat{X} = \sqrt{\frac{\hbar}{2m\omega}}(a + a^\dagger) \quad (10.6)$$

We now apply \hat{P} to the wavefunction.

$$\langle \hat{P} \rangle = \langle \Psi(t) | \hat{P} | \Psi(t) \rangle \quad (10.7)$$

$$\langle \hat{P} \rangle = \left(-\frac{1}{2}m\omega\hbar\right)\Psi(t) \left((a + a^\dagger)(C_0 e^{i\frac{1}{2}\omega t} |0\rangle + C_2 e^{i\frac{5}{2}\omega t} |2\rangle) \right) \quad (10.8)$$

$$\langle \hat{P} \rangle = \left(-\frac{1}{2}m\omega\hbar\right)\Psi^*(t) \left(0 + C_2 e^{i\frac{5}{2}\omega t} (-\hbar\omega) |1\rangle + C_0 e^{i\frac{1}{2}\omega t} (\hbar\omega) |1\rangle + C_2 e^{i\frac{5}{2}\omega t} (\hbar\omega) |3\rangle \right) \quad (10.9)$$

$$\langle \hat{P} \rangle = 0 \quad (10.10)$$

In the last step we used the orthogonality of the harmonic oscillator eigenfunctions, since $\Psi(t)$ only includes states with energy levels 0 and 2. c) The expectation value of \hat{X} is:

$$\langle \hat{X} \rangle = \langle \Psi(t) | \hat{X} | \Psi(t) \rangle \quad (10.11)$$

$$(10.12)$$

Since $\hat{X} = \sqrt{\frac{\hbar}{2m\omega}}(a + a^\dagger)$, we again see the combination of creation and annihilation operators. Once again the orthogonality of the harmonic oscillator eigenfunctions and the fact that the initial wavefunction contains only states 0 and 2 cause the expectation value to be 0. The expectation of both \hat{P} and \hat{X} will be 0 whenever the initial state contains only odd levels or only even levels, corresponding to combinations of only odd/even wavefunctions.

11 Problem B1

a)

i) Each distinguishable particle may be in one of 4 states, so there are 16

total states. 4 states have energy 0, 8 states have energy ϵ , and 4 states have energy 2ϵ .

$$Z = \sum_1^8 e^{-\beta\epsilon} + \sum_1^8 e^{-2\beta\epsilon} + \sum_1^8 e^{-4\beta\epsilon} \quad (11.1)$$

$$Z = 4(1 + 2e^{-\beta\epsilon} + e^{-2\beta\epsilon}) \quad (11.2)$$

ii) For Bosons there are three spin combinations and three energy combinations for a total of 9 states. Three of the states have energy 0, three have energy ϵ and three have energy 2ϵ .

$$Z = 3(1 + e^{-\beta\epsilon} + e^{-2\beta\epsilon}) \quad (11.3)$$

iii) For Fermions the Pauli exclusion principle will eliminate the states with both particles having the same spin. There are now 3 states, with one for each possible value of the energy.

$$Z = 1 + e^{-\beta\epsilon} + e^{-2\beta\epsilon} \quad (11.4)$$

b) As $T \rightarrow \infty$, $\beta = \frac{1}{kT} \rightarrow 0$ and $e^{-\beta\epsilon} \rightarrow 1$. For the classical particles, each state will have probability $\frac{1}{16}$, so a quarter of the systems will have $E = 0$. For the bosons, each state will have probability $\frac{1}{9}$, so one third of the systems will have $E = 0$.

For the fermions each state will have probability $\frac{1}{3}$, so one third of the systems will have $E = 0$.

c) We can't use $\langle E \rangle = \frac{\partial \ln Z}{\partial (-\beta)}$ since we have a small number of states (?? I'm not actually sure if this is correct, I may have just screwed up the calculation). Instead, we calculate $\langle E \rangle$ directly.

For the classical particles:

$$\langle E \rangle = \frac{\sum E_s e^{-\beta E_s}}{Z} \quad (11.5)$$

$$= \frac{4 \cdot 0 + 8\epsilon e^{-\beta\epsilon} + 4 \cdot 2\epsilon e^{-2\beta\epsilon}}{4(1 + 2e^{-\beta\epsilon} + e^{-2\beta\epsilon})} \quad (11.6)$$

$$= 2\epsilon \frac{e^{-\beta\epsilon} + e^{-2\beta\epsilon}}{1 + 2e^{-\beta\epsilon} + e^{-2\beta\epsilon}} \quad (11.7)$$

For the bosons:

$$\langle E \rangle = \frac{\sum E_s e^{-\beta E_s}}{Z} \quad (11.8)$$

$$= \frac{3 \cdot 0 + 3\epsilon e^{-\beta\epsilon} + 3 \cdot 2\epsilon e^{-2\beta\epsilon}}{3(1 + e^{-\beta\epsilon} + e^{-2\beta\epsilon})} \quad (11.9)$$

$$= \epsilon \frac{e^{-\beta\epsilon} + 2e^{-2\beta\epsilon}}{1 + e^{-\beta\epsilon} + e^{-2\beta\epsilon}} \quad (11.10)$$

For the fermions:

$$\langle E \rangle = \frac{\sum E_s e^{-\beta E_s}}{Z} \quad (11.11)$$

$$= \frac{0 + \epsilon e^{-\beta \epsilon} + 2\epsilon e^{-2\beta \epsilon}}{1 + e^{-\beta \epsilon} + e^{-2\beta \epsilon}} \quad (11.12)$$

$$= \epsilon \frac{e^{-\beta \epsilon} + 2e^{-2\beta \epsilon}}{1 + e^{-\beta \epsilon} + e^{-2\beta \epsilon}} \quad (11.13)$$

d) In the high temperature limit, $\langle E \rangle = \epsilon$ for all types of particles.

12 Problem B2

a) The heat capacity $C_v = (\frac{\partial U}{\partial T})_V$. If the heat capacity is constant then $U \sim T$. Then $A = U - TS = CT - TS = T(C - S)$ with C some constant.

b) For a monatomic ideal gas the Hamiltonian is $H = \frac{p^2}{2m}$. The partition function of one atom is:

$$Z_1 = \frac{1}{h^3} \int d^3x d^3p e^{-\beta \frac{p^2}{2m}} \quad (12.1)$$

$$Z_1 = \frac{V}{h^3} \int d^3p e^{-\beta \frac{p^2}{2m}} \quad (12.2)$$

Each momentum component integral $\int_0^\infty e^{-\beta \frac{p^2}{2m}} dp = \sqrt{2m\pi kT}$. With three independent components, the total partition function of one atom is:

$$Z_1 = \frac{V}{h^3} (2m\pi kT)^{3/2} \quad (12.3)$$

$$(12.4)$$

The system partition function (indistinguishable particles) is:

$$Z_N = \frac{1}{N!} \left(\frac{V}{h^3} \right)^N (2m\pi kT)^{3N/2} \quad (12.5)$$

The Helmholtz free energy A is $A = -kT \ln Z_N$.

$$A = -NkT \left(\ln \frac{V}{N!} + \frac{3}{2} \ln \frac{2m\pi kT}{h^2} \right) \quad (12.6)$$

c) The energy levels of the harmonic oscillator are $(n + \frac{1}{2})\hbar\omega$. The partition function for one atom is:

$$\sum_{s=0}^{\infty} e^{-\beta(s+\frac{1}{2})\hbar\omega} = e^{-\beta\frac{1}{2}\hbar\omega} (1 + e^{-\beta\hbar\omega} + e^{-2\beta\hbar\omega} + \dots) \quad (12.7)$$

$$= e^{-\beta\frac{1}{2}\hbar\omega} \left(\frac{1}{1 - e^{-\beta\hbar\omega}} \right) \quad (12.8)$$

The partition function for N atoms is $Z_1^N = \left(e^{-\beta \frac{1}{2} \hbar \omega} \left(\frac{1}{1 - e^{-\beta \hbar \omega}} \right) \right)^N$. The Helmholtz free energy is:

$$A = -kT \ln Z_N \quad (12.9)$$

$$A = -NkT \left(-\beta \frac{1}{2} \hbar \omega + \beta \hbar \omega \right) \quad (12.10)$$

$$A = -\frac{N}{2} \hbar \omega \quad (12.11)$$