

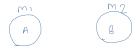
Gravitation

The phenomenon of attraction 

Between any two objects in the Universe is called gravitation.

Gravity — attractive force which is exerted by earth on any body.

Universal Law of Gravitation.



$$F \propto m_1 m_2$$

$$F \propto \frac{1}{r^2}$$

$$F = G \times \frac{m_1 m_2}{r^2}$$

Everybody in the universe attracts every other body with an force which is directly proportional to the product of their masses & inversely proportional to the square of the distance between them.

Universal gravitational constant

$$G = \frac{(F \times r^2)}{m_1 m_2}$$

$$G = \frac{N \cdot m^2}{kg^2}$$

$$G = 6.67 \times 10^{-11} N \cdot m^2 \cdot kg^{-2}$$

SI unit of G

Universal gravitational constant is numerically equal to the gravitational force of attraction between two bodies each of mass 1 kg kept at unit distance from each other.

$$G = F$$

$$G = \frac{F \cdot r^2}{m_1 m_2}$$

$$G = F \times (m)^2$$

$$1 kg \times 1 kg$$

$$G = F$$

Why gravitational constant (G) is known as Universal gravitational constant?

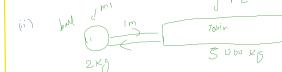
\rightarrow The value of G doesn't depend on mass of two bodies, distance b/w two bodies, nature, medium, shape/size.

Conditions

It is within two objects each of 1 kg and 1 m apart.

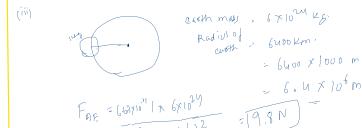


$$F = G \times \frac{m_1 m_2}{r^2} = 6.67 \times 10^{-11} \times \frac{1 \times 1}{1} = 6.67 \times 10^{-11} N$$



$$\text{Force of Ball to Train} = G \times \frac{m_1 m_2}{r^2} = 6.67 \times 10^{-11} \times \frac{1 \times 5000}{1^2} = 1000.6 N$$

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$$F_{AE} = G \times \frac{m_1 m_2}{r^2} = \frac{6.67 \times 10^{-11} \times (6 \times 10^{24})}{(6 \times 10^6)^2} = 19.3 N$$

$$F_{AE} = 9.8 N$$

Newton's 2nd law of motion

$$F = ma$$

$$\therefore \frac{F}{m} = a = 9.8 = 9.8 m s^{-2}$$

$$\therefore a_{AE} = \frac{F}{m} = \frac{9.8}{6 \times 10^{24}} = \frac{1.63}{\times 10^{-24}} m s^{-2}$$



Before the person is released, the earth moves in a circular path with certain speed and changes direction at every certain point due to centripetal force or acceleration. The force that causes this acceleration is called centripetal force. The centripetal force is acting towards the centre. This is the reason why we feel like our heart is seeking here. In the absence of it,

The radius of the circle traced by the earth is the radius of the elliptical orbit of the earth. If there were no such force, the earth would move in a straight line.

i) Two particles A and B of mass m_1 and m_2 respectively are placed on some distance r . If the mass of each of the two particles is doubled, keeping the distance b/w them unchanged, the gravitational force b/w them will be:

$$F_1 = G \times \frac{m_1 m_2}{r^2}$$

$$F_2 = G \times \frac{2m_1 2m_2}{r^2}$$

$$F_2 = G \times \frac{4m_1 m_2}{r^2}$$

$$F_2 = 4 \times G \times \frac{m_1 m_2}{r^2}$$

$$F_2 = 4 F_1$$

$$F_2 = \frac{G \times m_1 m_2}{r^2}$$

$$F_2 = \frac{G \times 2m_1 2m_2}{r^2}$$

$$F_2 = \frac{4G \times m_1 m_2}{r^2}$$

$$F_2 = 4 F_1$$

2) mass of sun is $2 \times 10^{30} \text{ kg}$ & mass of earth is $6 \times 10^{24} \text{ kg}$.
At avg distance b/w the sun and the earth is $1.5 \times 10^8 \text{ km}$.
Calculate the gravitational force b/w them.



3) The gravitational force b/w two objects is F . If masses of both objects are halved w/o changing distance b/w them, gravitational force would become.

$$F_1 = G M_1 M_2 / R^2$$

$$F_2 = G (M_1/2)(M_2/2) / R^2 = \frac{G M_1 M_2}{R^2} / 4 = F_1/4$$

4) Gravitational force b/w two objects weighing 20g and 15g is $1.8 \times 10^{-10} \text{ N}$. Find distance b/w two objects.

Scientists → Arrangement of planets

Ptolemy → Geocentric theory
Two AD

Copernicus → 1543
Heliocentric theory
Sun at the center

Tycho Brahe → 1548
Tycho's system

Galileo → 1609
Sun is at the center
Heliocentric theory
Telescope

Earth is at center & all planets revolve around

Moon is at the center.

Kepler's law of planetary motion

He stated three laws which govern the motion of planets around the sun. They are known as Kepler's law of planetary motion.



Circle → 1 center
1 focus (radius are equal)

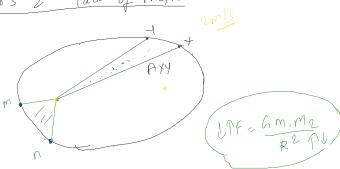


Ellipse → 2 centers
foci, foci

Kepler's 1st law of motion

→ Planets move in elliptical orbits around the sun, with the sun at one of two foci of the elliptical orbit.

Kepler's 2nd law of motion



$$\downarrow F = G M_1 M_2 / R^2 \uparrow$$

→ Each planet revolves around the sun in such a way that the line joining the planet to the sun sweeps over equal areas in equal intervals of time.

→ A planet moves faster when it is closer to the sun & moves slowly when it is farther away from the sun.

Kepler's 3rd law of motion

→ The cube of the mean distance of a planet from the sun is directly proportional to the square of time it takes to move around the sun.

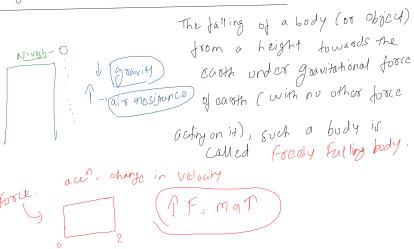
$$T^2 \propto R^3$$

$$T^2 = \text{constant } R^3$$

$$T^2 / R^3 = \text{constant}$$

mean distance of planet from sun.
T = time period of the planet around the sun.

Free fall and acceleration due to gravity (g)

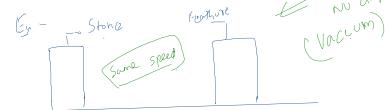


Acc due to gravity (g)

→ The uniform acc produced in a freely falling body due to gravitational force of the earth is known as acc due to gravity.

→ SI unit = m s^{-2}

→ g doesn't depend on the mass of the falling object.



Calculation of acc due to gravity

→ Force exerted by earth on

the body

$$F = G M m / R^2 = m a$$

→ $F = m a$ → $a = ?$

$$G M m / R^2 = m a$$

$$a = G M / R^2$$

Now, $a = G M / R^2$.

$$g = G M / R^2$$

$$g = (6.67 \times 10^{-11} \text{ N m kg}^{-2}) \times (6 \times 10^{24} \text{ kg}) / (6.4 \times 10^6)^2$$

$$g = 9.8 \text{ m s}^{-2}$$

This force by earth produces acc in the stone due to which stone moves downwards.