

# Red-Sequence Evolution of Galaxy Clusters

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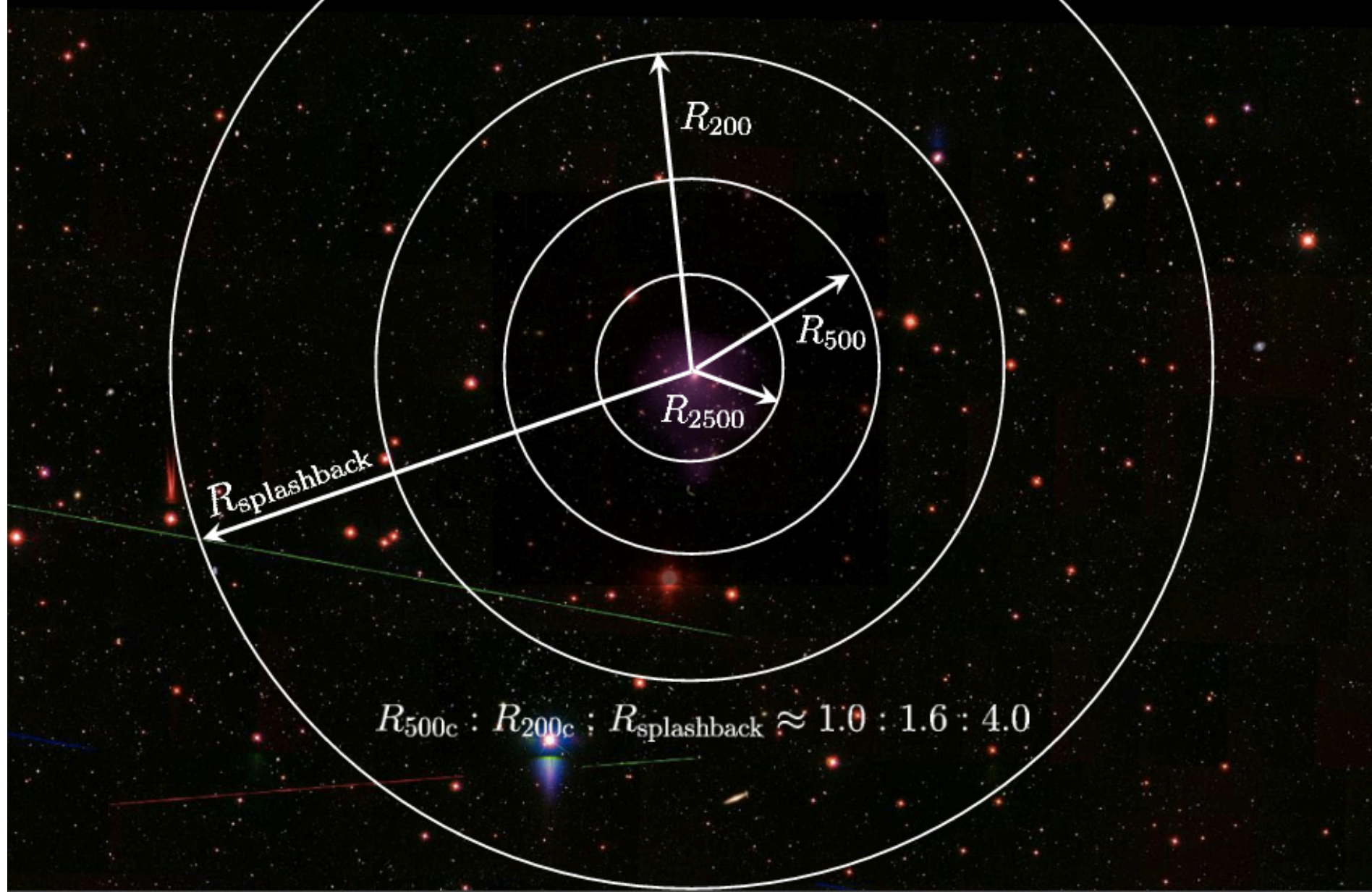
# Outline : Astrophysics and Cosmology

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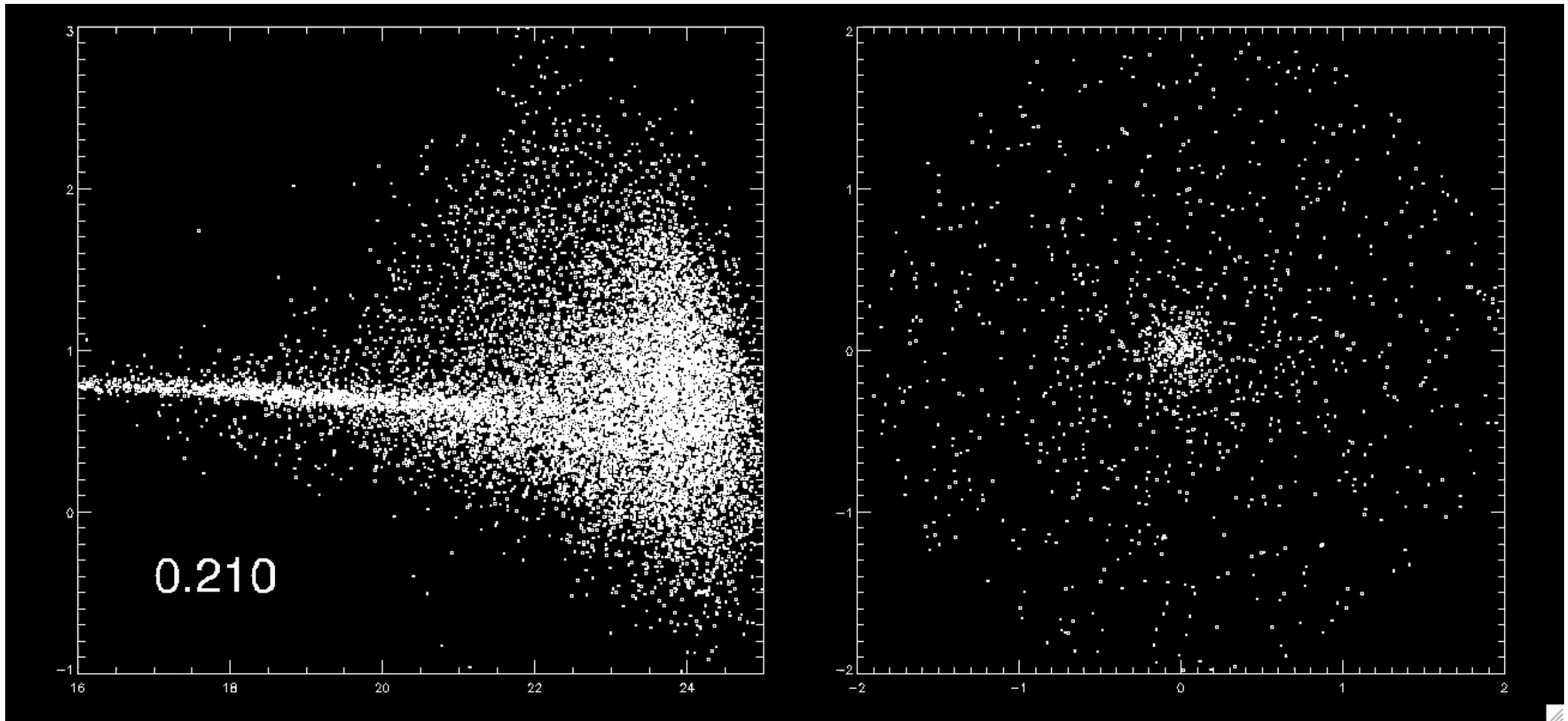
- Motivation
  - SPT,DES Projects
  - A DES Class Project Illustration, Autumn 2016, University of Chicago
- Cosmology with Galaxy Clusters – Membership, Richness, Mass
- This work
  - Datasets
  - Datacuts – selection functions
  - Analysis
- Red Sequence Evolution in motion

$$M_{\Delta} = \frac{4\pi}{3} \Delta \rho_{\text{crit}}(z) R_{\Delta}^3$$

$$\rho_{\text{crit}}(z) = \frac{3H^2(z)}{8\pi G}$$



$$R_{500c} : R_{200c} : R_{\text{splashback}} \approx 1.0 : 1.6 : 4.0$$



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Gladders et al. – Red Sequence Cluster Survey

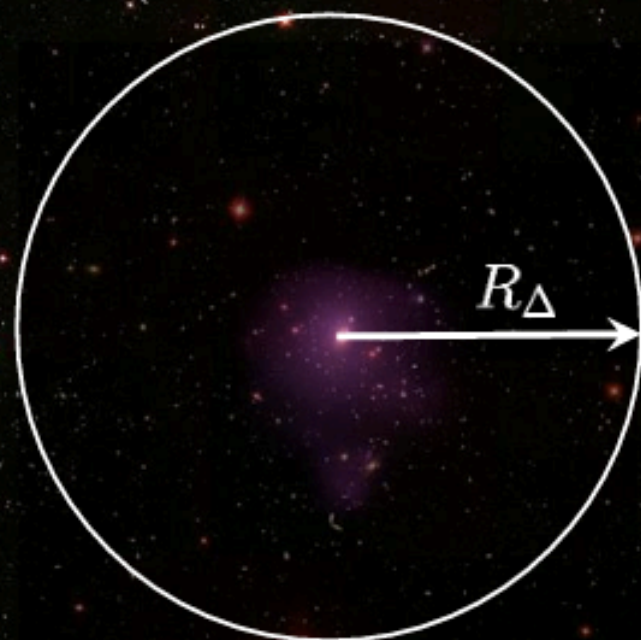


# Concepts: An Overview

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$$M_{\Delta} = \frac{4\pi}{3} \Delta \rho_{\text{ref}}(z) R_{\Delta}^3$$

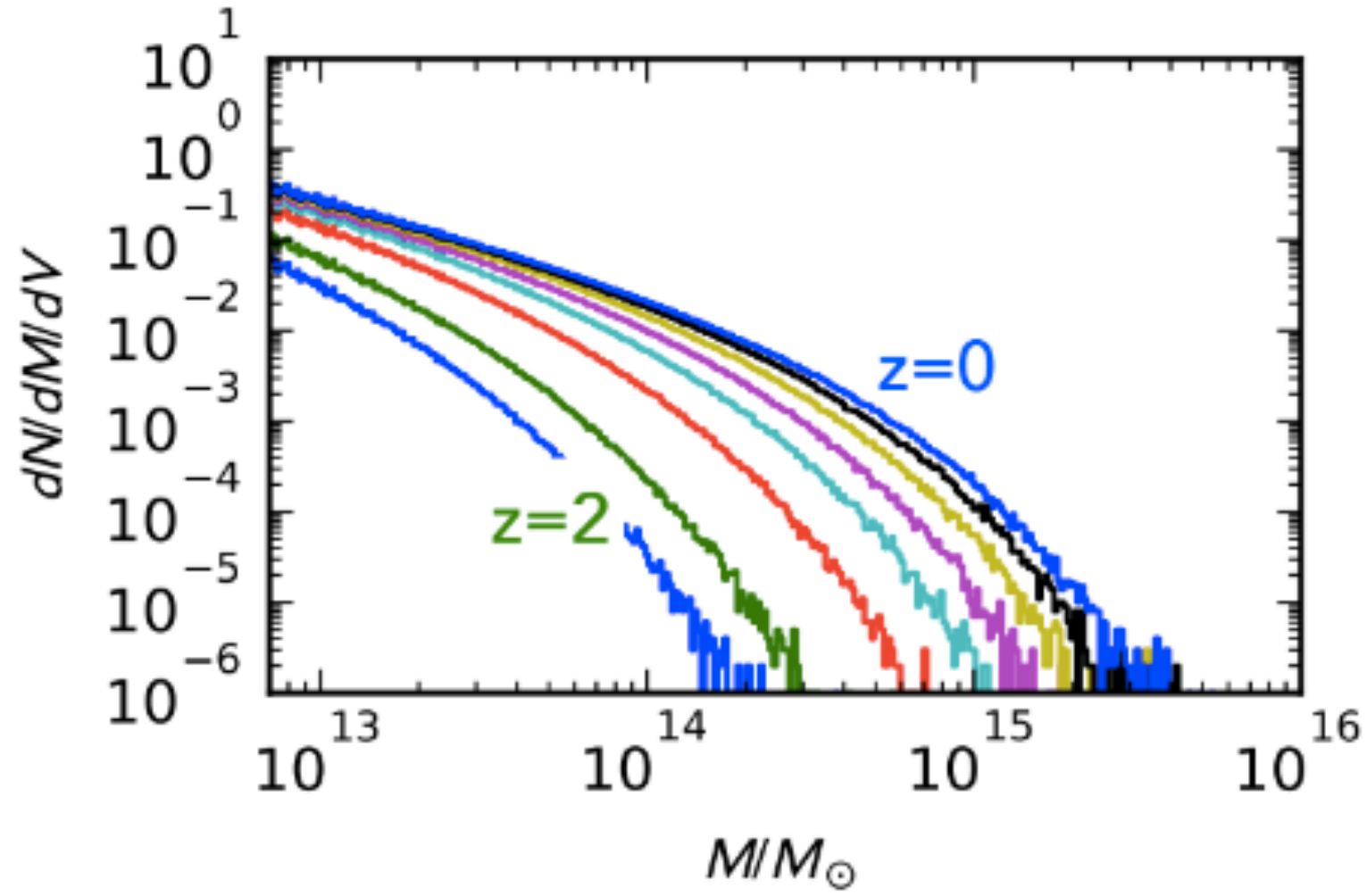
$$T \propto \frac{M_{\Delta}}{R_{\Delta}} \propto \Delta^{1/3} \rho_{\text{ref}}^{1/3}(z) M_{\Delta}^{2/3} \left\{ \begin{array}{l} \propto \rho_{\text{crit}}^{1/3}(z) M_{\Delta\text{c}}^{2/3} \propto E^{2/3}(z) M_{\Delta\text{c}}^{2/3} \\ \propto \rho_{\text{mean}}^{1/3}(z) M_{\Delta\text{m}}^{2/3} \propto (1+z) M_{\Delta\text{m}}^{2/3} \end{array} \right.$$



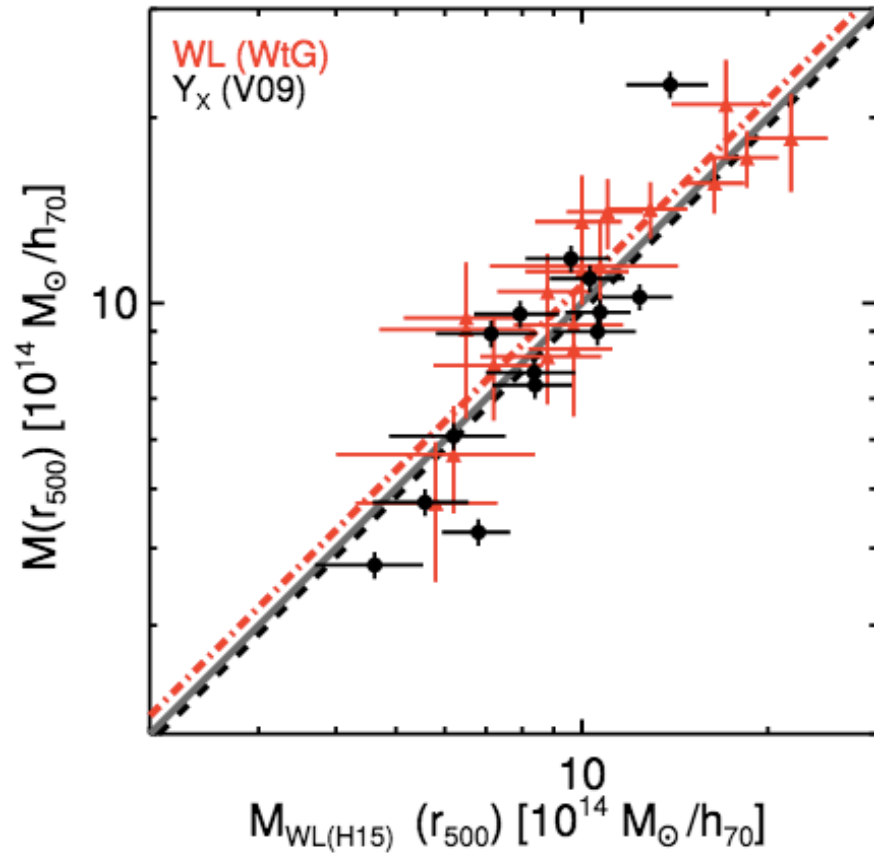
$$\mu m_p \frac{GM_{\text{HE}}(< R)}{R} = kT(R) \left[ -\frac{d \ln \rho_g}{d \ln R} - \frac{d \ln T}{d \ln R} \right]$$

$$\frac{GM_{\text{J}}(< R)}{R} = \sigma_r^2(R) \left[ -\frac{d \ln n_{\text{gal}}}{d \ln R} - \frac{d \ln \sigma_r^2}{d \ln R} - 2\beta(R) \right]$$

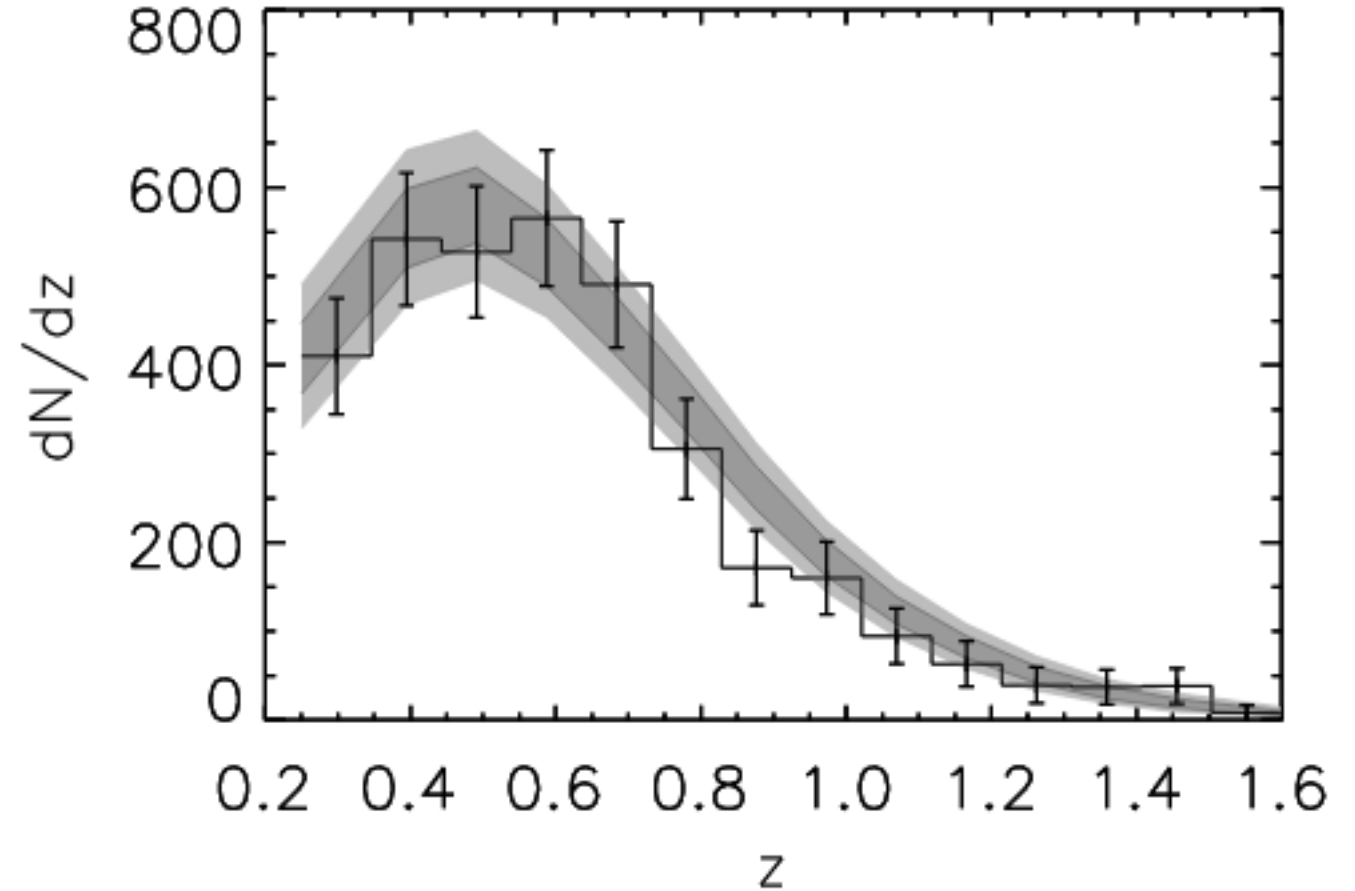
# Halo Mass Function



# Current Constraints



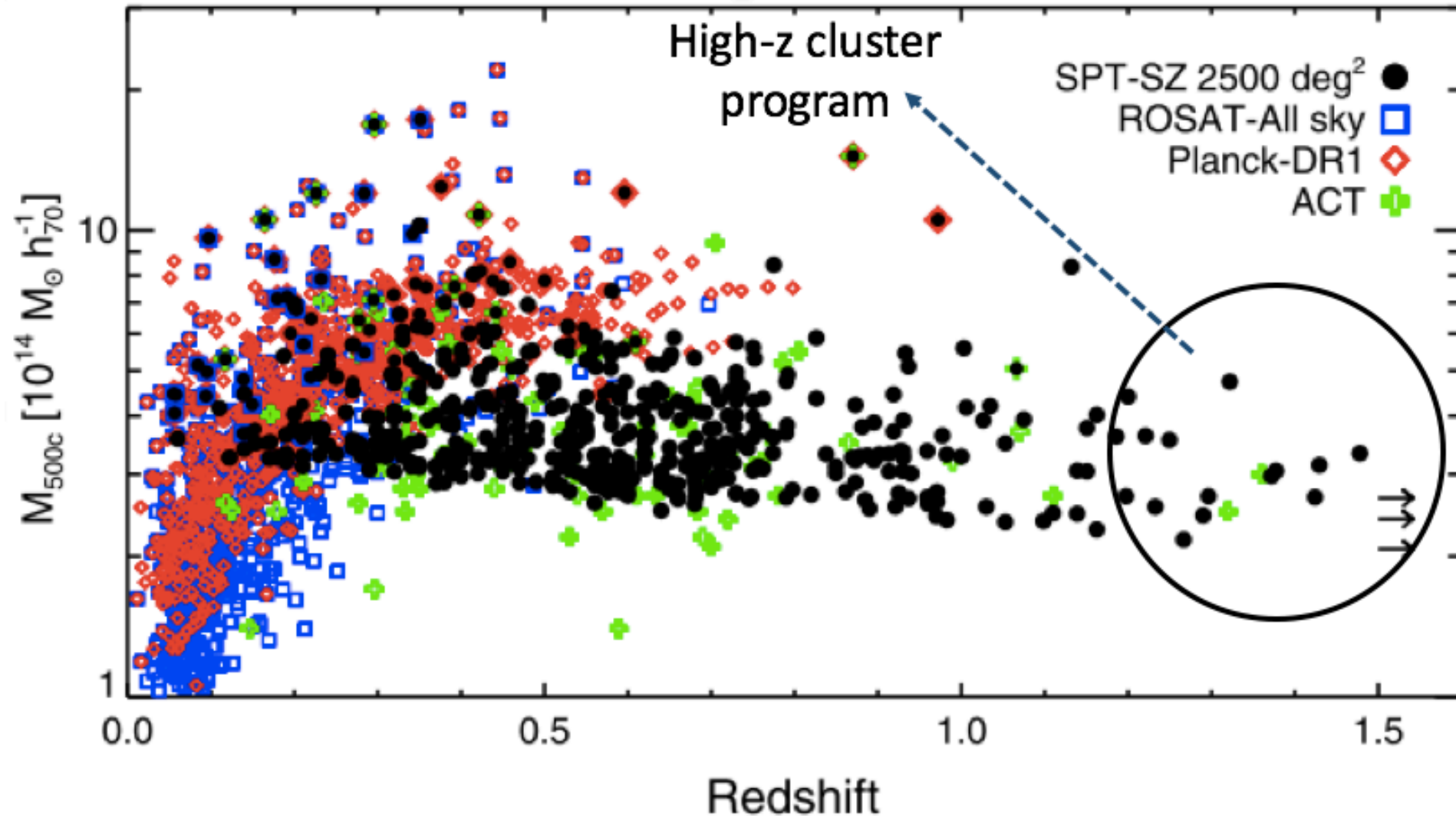
WtG, Allen, Mantz et al



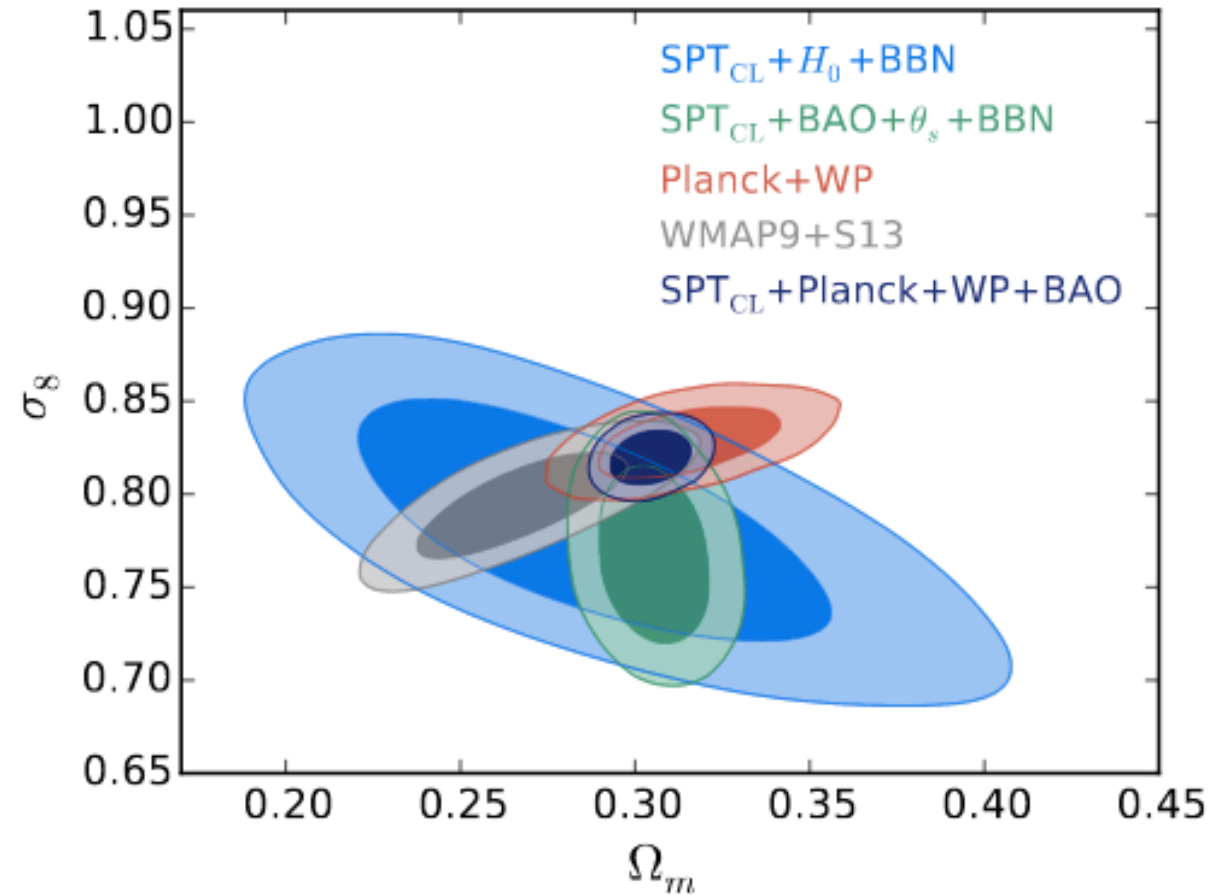
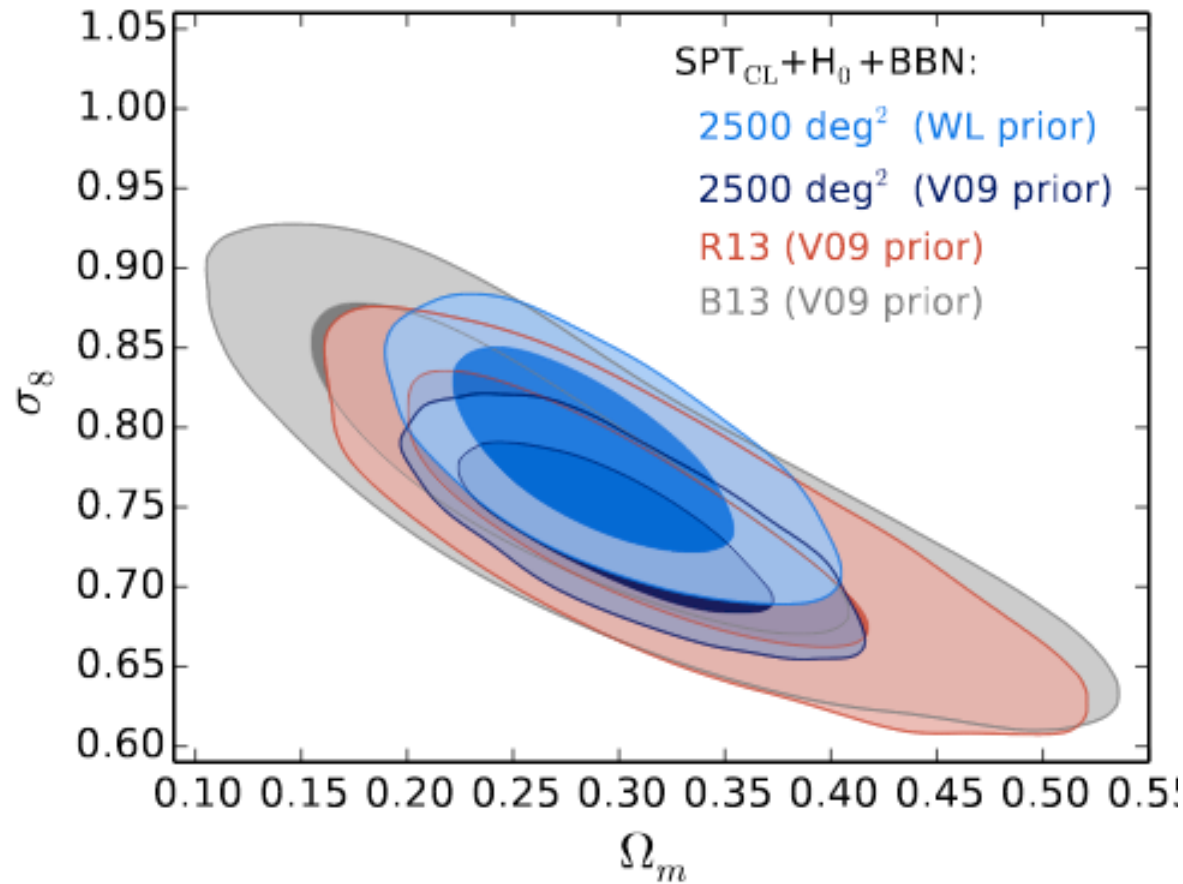
SPT, deHaan et al. 2017

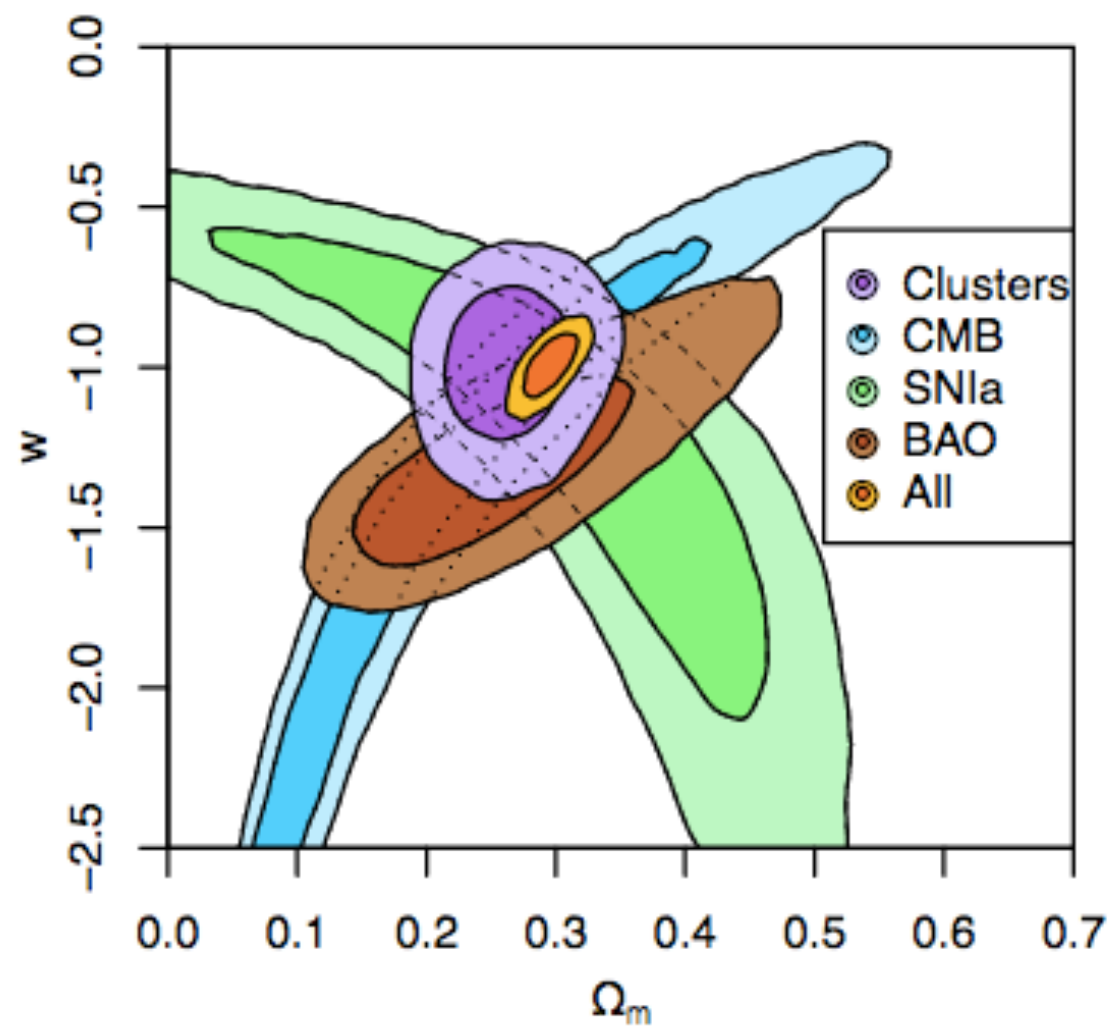
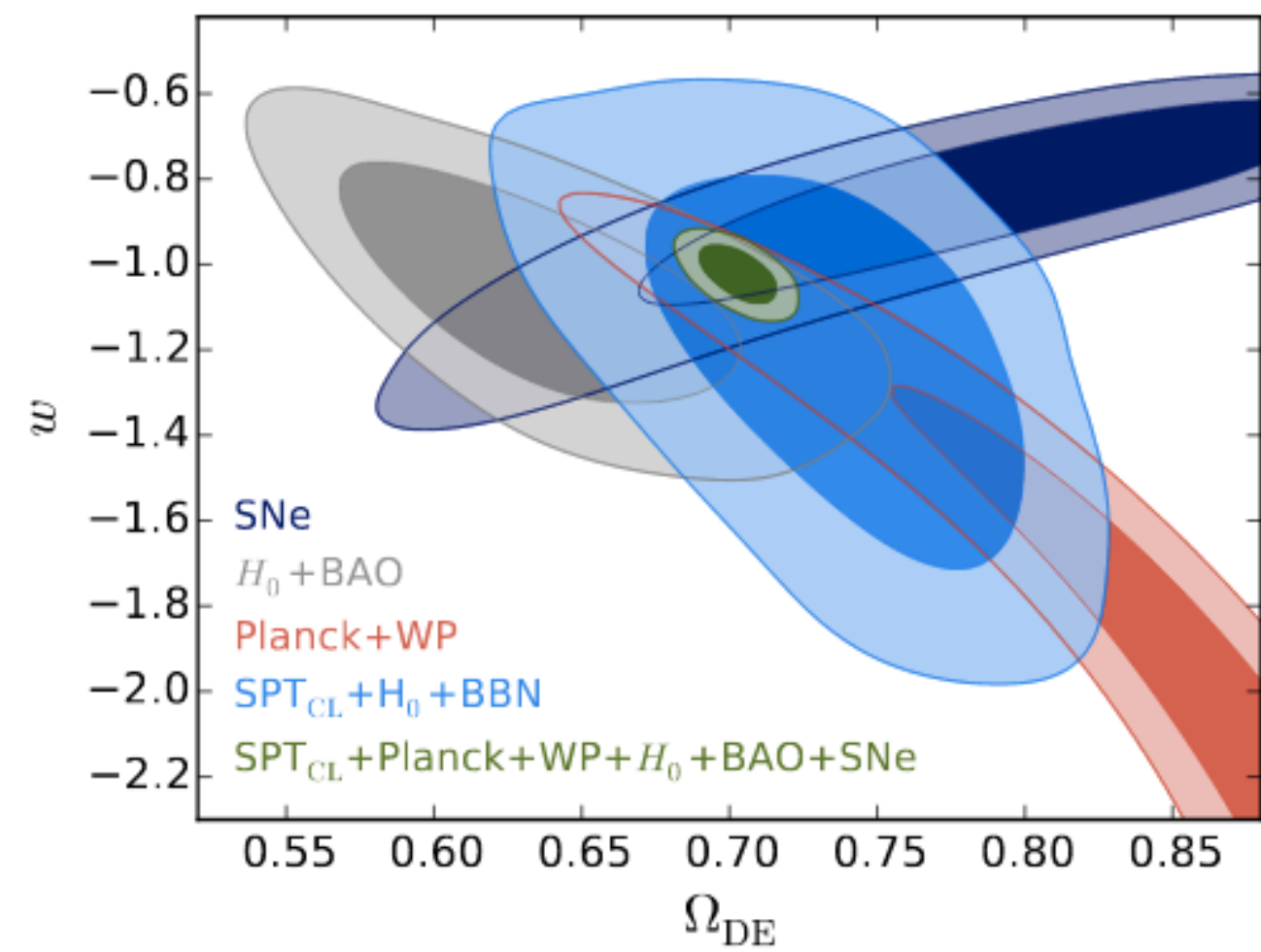


# SPT-SZ Cosmology



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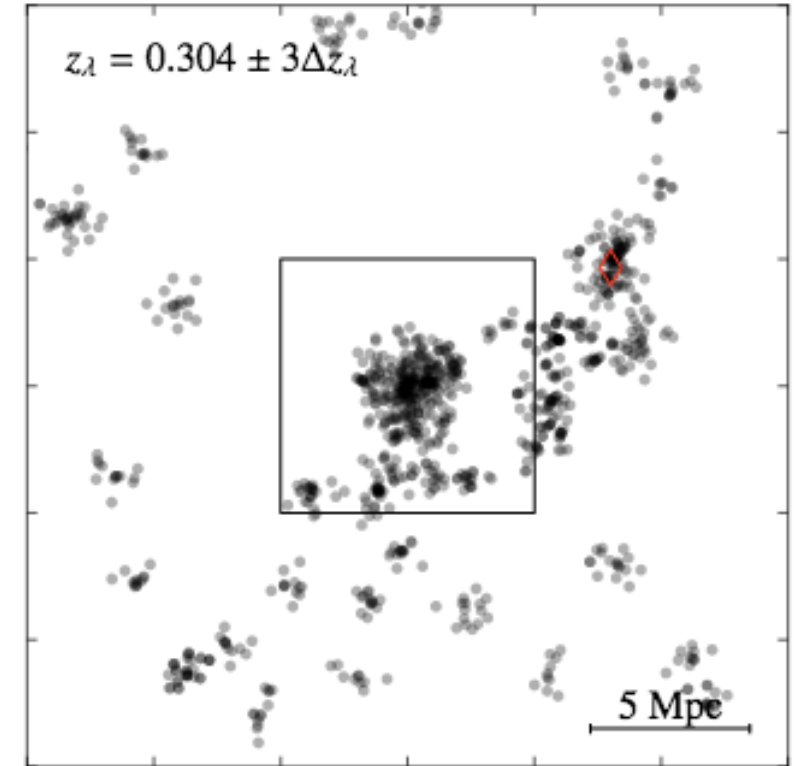
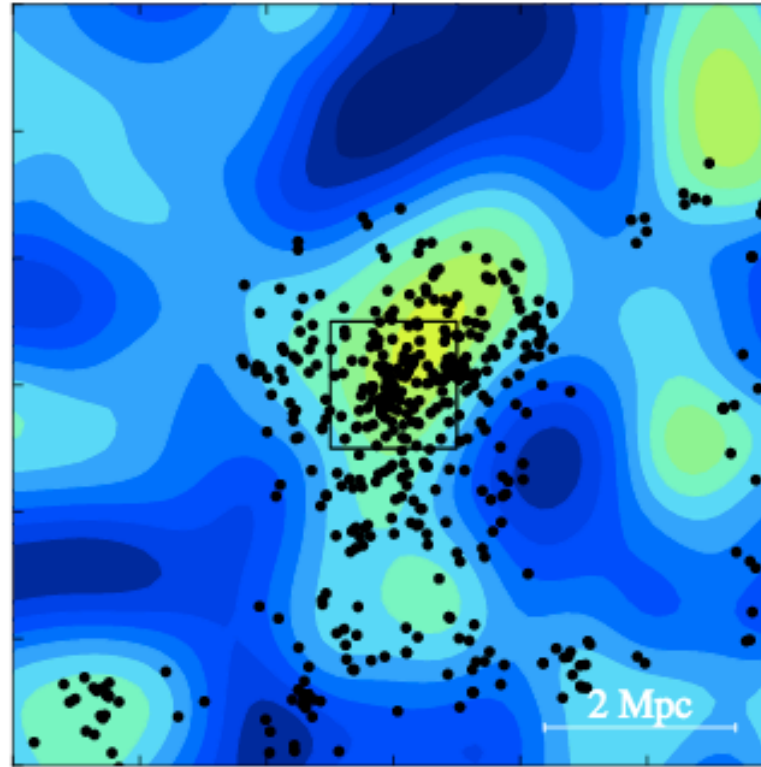
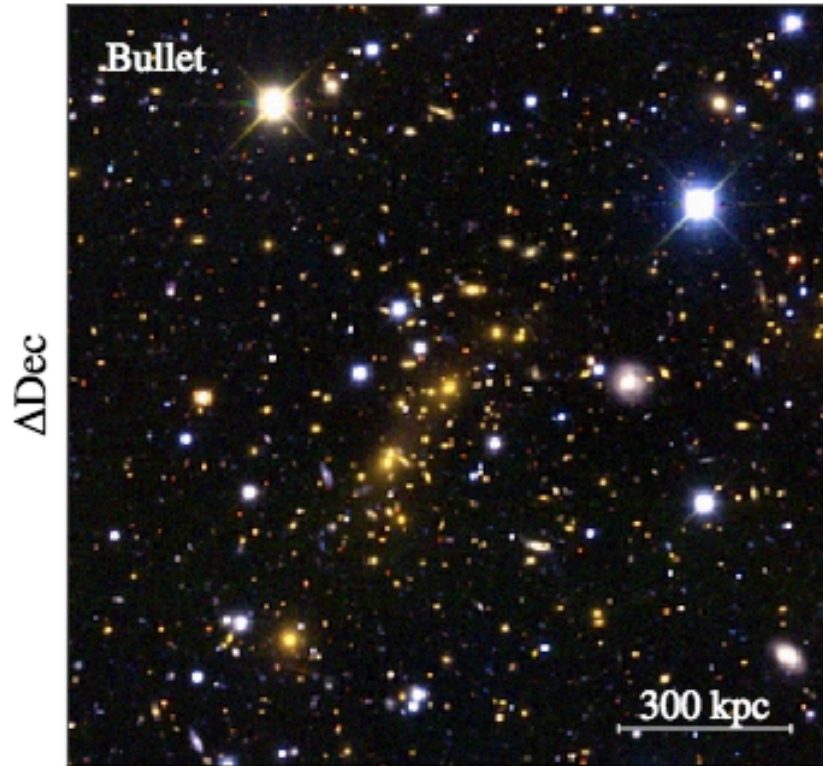


This Work

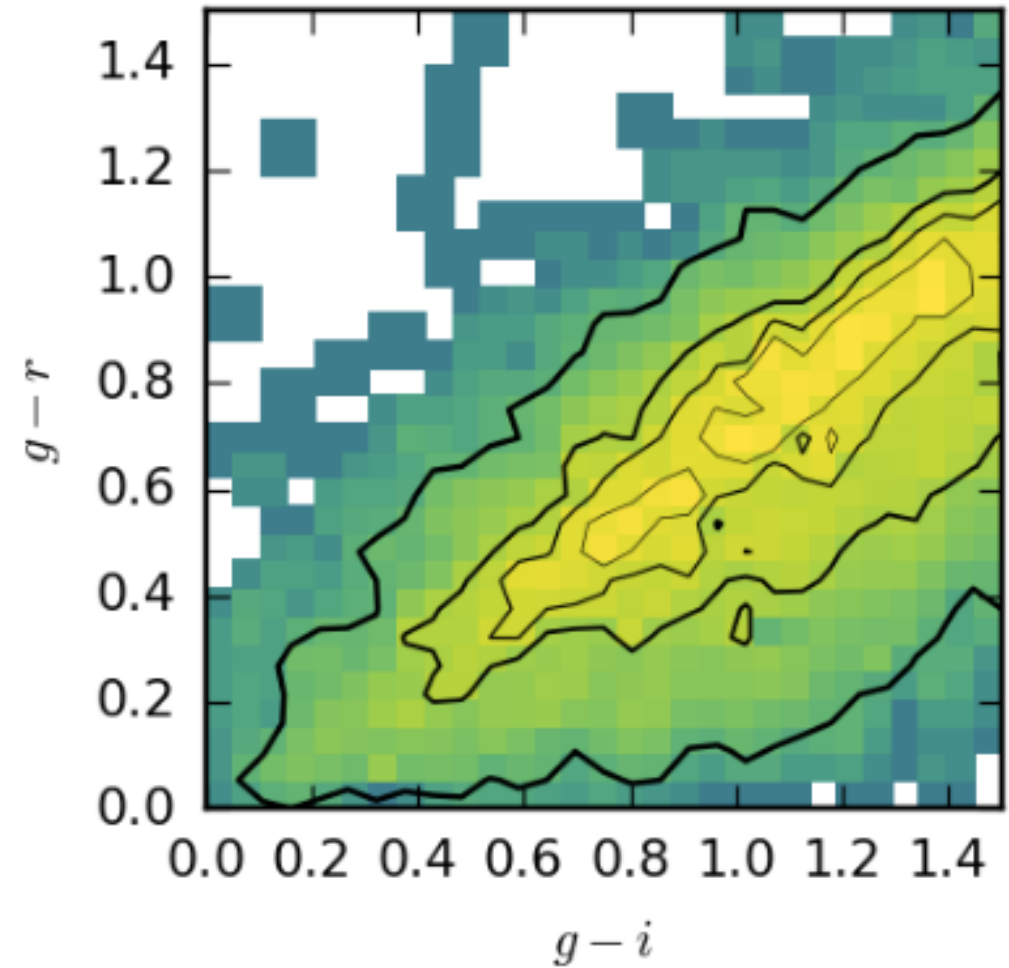
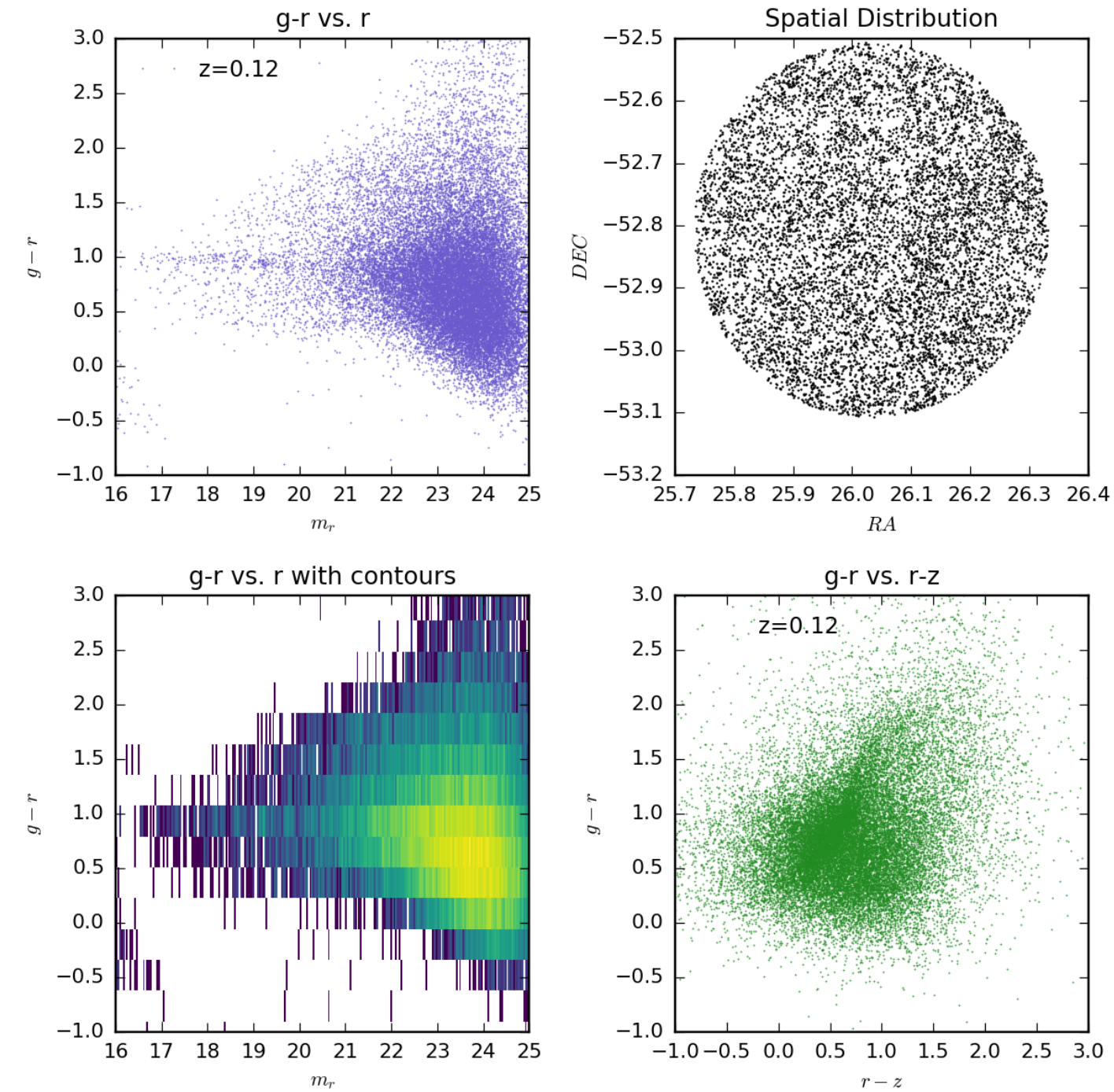
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# DES RedMapper Galaxy Clusters

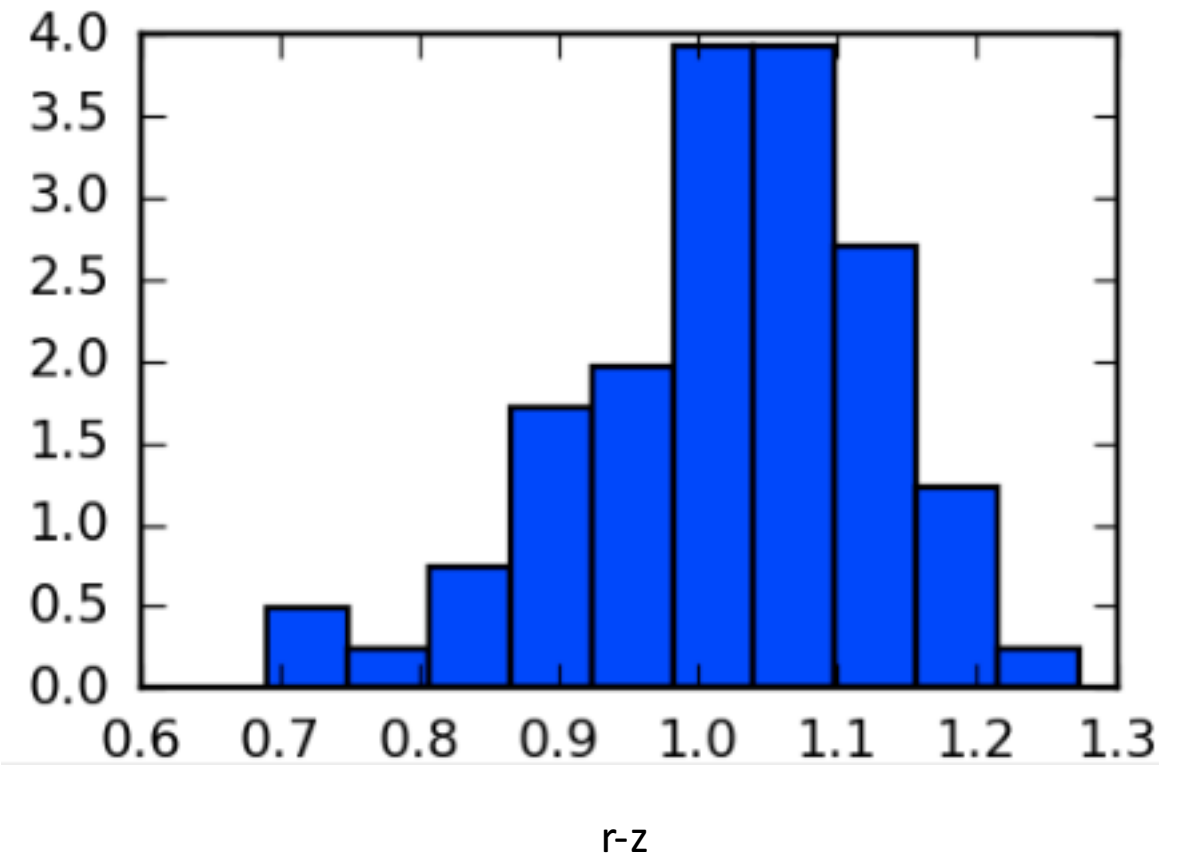
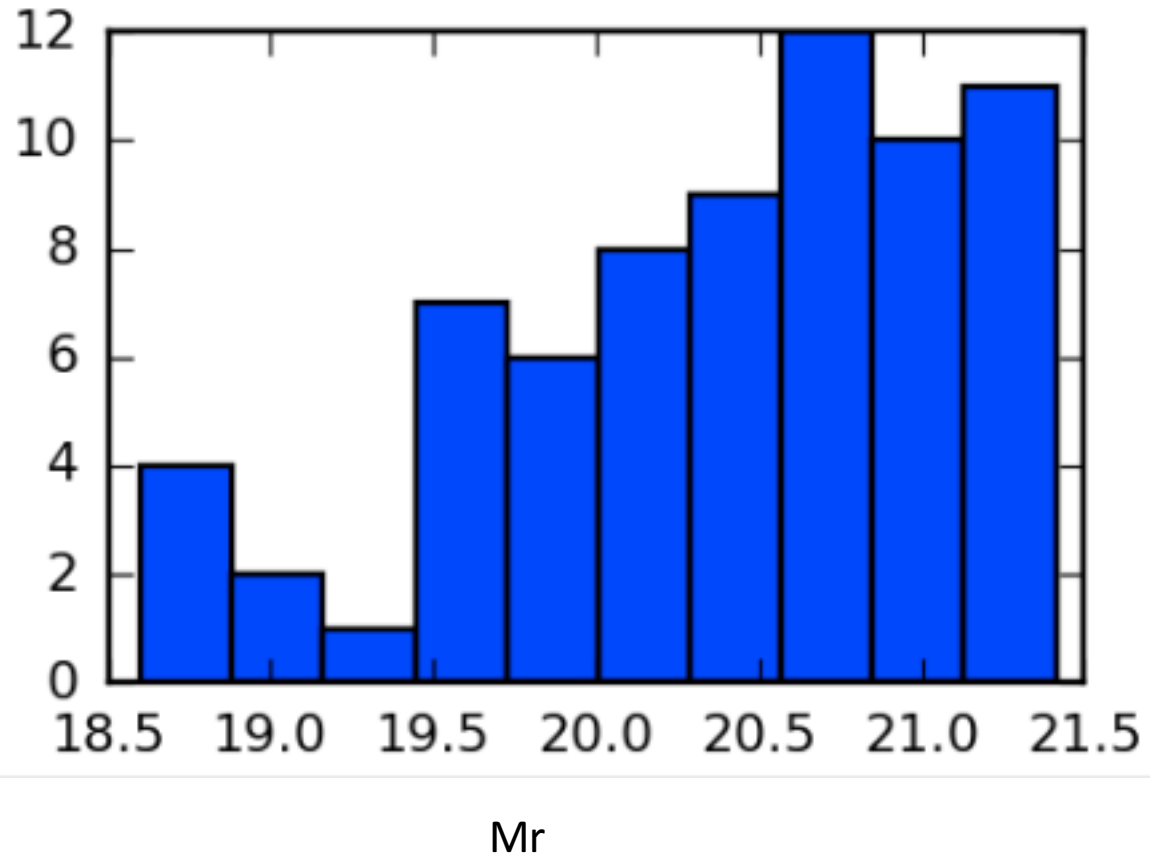


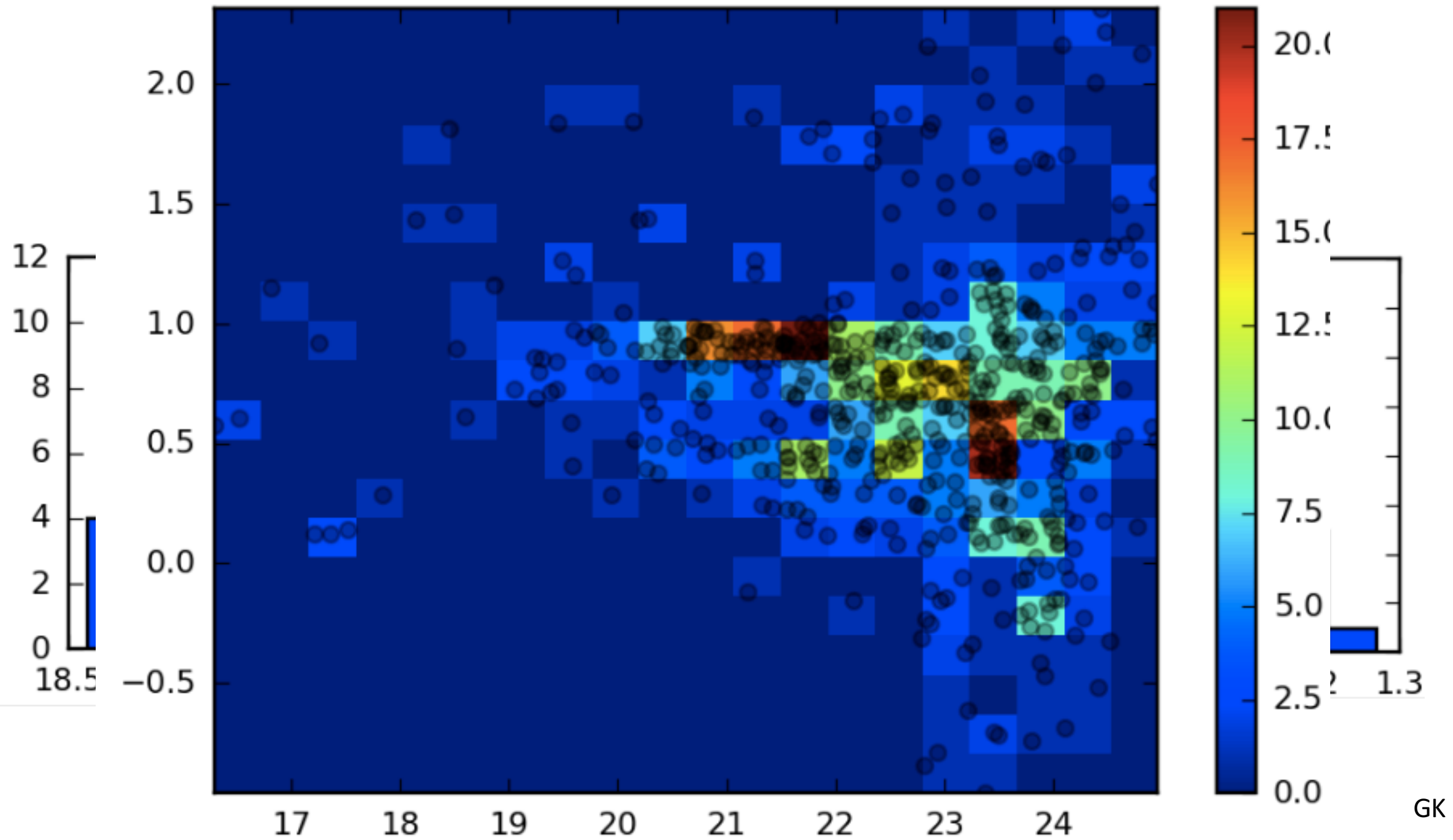
# Example: SPT Cluster in DES



# Distribution in color-magnitude space : 1 Cluster Tile

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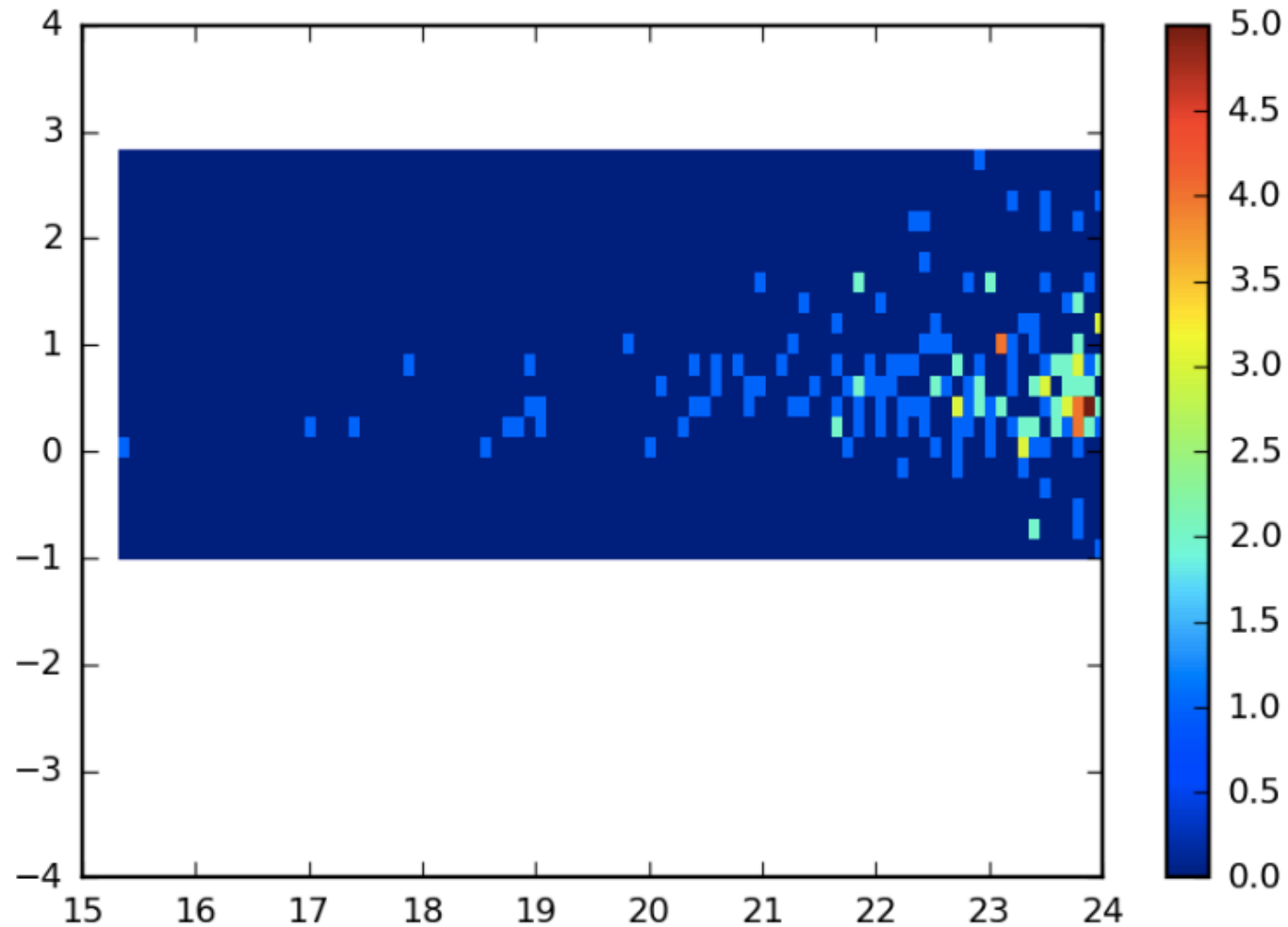






# Background contamination

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Final product: Red Sequence Evolution with  $z$

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