

1 Expanding the conservation laws

The conservation laws are given below in terms of $\mu = 1$. For numerical convenience, place a dimensionless μ in front of each \vec{B} appearance and let $\mu = 0$ be plain compressible hydrodynamics.

$$\begin{aligned}\frac{\partial \rho}{\partial t} &= -\vec{\nabla} \cdot (\rho \vec{v}) \\ \frac{\partial(\rho \vec{v})}{\partial t} &= -\vec{\nabla} \cdot (\rho \vec{v} \vec{v} - \vec{B} \vec{B}) - \vec{\nabla} P^* \\ \frac{\partial \vec{B}}{\partial t} &= \vec{\nabla} \times (\vec{v} \times \vec{B}) \\ \frac{\partial E}{\partial t} &= -\vec{\nabla} \cdot \left((E + P^*) \vec{v} - \vec{B} (\vec{B} \cdot \vec{v}) \right)\end{aligned}\tag{1}$$

where $P^* = (\gamma - 1) \left(E - \rho \frac{v^2}{2} - \frac{B^2}{2} \right) + \frac{B^2}{2}$. Note that $\vec{B} \vec{B}$ is a dyadic tensor $\mathbf{B} = \vec{B} \vec{B}$ such that $\mathbf{B}_{ij} = B_i B_j$. Then the divergence $\vec{\nabla} \cdot \mathbf{B} = \partial^i B_i B_j \hat{e}_j = B_i \partial_i B_j \hat{e}_j = \left(\vec{B} \cdot \vec{\nabla} \right) \vec{B}$. On the other hand, $\vec{\nabla} \cdot \rho \vec{v} \vec{v} = \left(\vec{\nabla} \cdot \rho \vec{v} \right) \vec{v} + (\vec{v} \cdot \vec{\nabla}) \rho \vec{v}$ where in general we allow $\vec{\nabla} \cdot \rho \vec{v} \neq 0$.

Some intermediate steps: $\vec{\nabla} \cdot (\rho \vec{v} \vec{v}) = \frac{\rho \vec{v}}{\rho} \cdot \vec{\nabla} (\rho \vec{v})$. Additionally, $\vec{\nabla} \times (\vec{v} \times \vec{B}) = (\vec{B} \cdot \vec{\nabla}) \vec{v} - (\vec{v} \cdot \vec{\nabla}) \vec{B} - \vec{B} (\vec{\nabla} \cdot \vec{v})$ since only $\vec{\nabla} \cdot \vec{B} = 0$.

Construct the vector field $\vec{u}(x, y, z)$ where $\vec{u} = (\rho, \rho v_x, \rho v_y, \rho v_z, B_x, B_y, B_z, E)$ and assume we have point-by-point cached values of $-(E + P^*) \vec{v} + \vec{B} (\vec{B} \cdot \vec{v}) = EPmDV_i$ (along with P^* , four values), then we can represent these evolution equations as (two lines b/c too damn long)

$$\begin{aligned}\frac{\partial \vec{u}}{\partial t} = f(\vec{u}) = & \begin{pmatrix} -\frac{\partial u_2}{\partial x} - \frac{\partial u_3}{\partial y} - \frac{\partial u_4}{\partial z} \\ -\frac{u_2}{u_1} \frac{\partial u_2}{\partial x} - \frac{u_3}{u_1} \frac{\partial u_2}{\partial y} - \frac{u_4}{u_1} \frac{\partial u_2}{\partial z} + u_5 \frac{\partial u_2}{\partial x} + u_6 \frac{\partial u_2}{\partial y} + u_7 \frac{\partial u_2}{\partial z} - \frac{\partial P^*}{\partial x} \\ -\frac{u_2}{u_1} \frac{\partial u_3}{\partial x} - \frac{u_3}{u_1} \frac{\partial u_3}{\partial y} - \frac{u_4}{u_1} \frac{\partial u_3}{\partial z} + u_5 \frac{\partial u_3}{\partial x} + u_6 \frac{\partial u_3}{\partial y} + u_7 \frac{\partial u_3}{\partial z} - \frac{\partial P^*}{\partial y} \\ -\frac{u_2}{u_1} \frac{\partial u_4}{\partial x} - \frac{u_4}{u_1} \frac{\partial u_2}{\partial y} - \frac{u_4}{u_1} \frac{\partial u_4}{\partial z} + u_5 \frac{\partial u_4}{\partial x} + u_6 \frac{\partial u_4}{\partial y} + u_7 \frac{\partial u_4}{\partial z} - \frac{\partial P^*}{\partial z} \\ u_5 \frac{\partial u_2}{\partial x} + u_6 \frac{\partial u_2}{\partial y} + u_7 \frac{\partial u_2}{\partial z} - u_2 \frac{\partial u_5}{\partial x} - u_3 \frac{\partial u_5}{\partial y} - u_4 \frac{\partial u_5}{\partial z} - u_5 \left(\frac{\partial u_2}{\partial x} + \frac{\partial u_3}{\partial y} + \frac{\partial u_4}{\partial z} \right) \\ u_1 \\ u_5 \frac{\partial u_3}{\partial x} + u_6 \frac{\partial u_3}{\partial y} + u_7 \frac{\partial u_3}{\partial z} - u_2 \frac{\partial u_6}{\partial x} - u_3 \frac{\partial u_6}{\partial y} - u_4 \frac{\partial u_6}{\partial z} - u_6 \left(\frac{\partial u_2}{\partial x} + \frac{\partial u_3}{\partial y} + \frac{\partial u_4}{\partial z} \right) \\ u_1 \\ u_5 \frac{\partial u_4}{\partial x} + u_6 \frac{\partial u_4}{\partial y} + u_7 \frac{\partial u_4}{\partial z} - u_2 \frac{\partial u_7}{\partial x} - u_3 \frac{\partial u_7}{\partial y} - u_4 \frac{\partial u_7}{\partial z} - u_7 \left(\frac{\partial u_2}{\partial x} + \frac{\partial u_3}{\partial y} + \frac{\partial u_4}{\partial z} \right) \\ u_1 \\ \frac{\partial EPmDV_1}{\partial x} + \frac{\partial EPmDV_2}{\partial y} + \frac{\partial EPmDV_3}{\partial z} \end{pmatrix} \\ & + \begin{pmatrix} 0 \\ -\frac{u_2}{u_1} \left(\frac{\partial u_2}{\partial x} + \frac{\partial u_3}{\partial y} + \frac{\partial u_4}{\partial z} \right) \\ -\frac{u_3}{u_1} \left(\frac{\partial u_2}{\partial x} + \frac{\partial u_3}{\partial y} + \frac{\partial u_4}{\partial z} \right) \\ -\frac{u_4}{u_1} \left(\frac{\partial u_2}{\partial x} + \frac{\partial u_3}{\partial y} + \frac{\partial u_4}{\partial z} \right) \end{pmatrix}\end{aligned}\tag{2}$$

Note that all derivatives are given by the expression $\frac{\partial f_{x,y,z}}{\partial x} = \frac{f_{x+1,y,z} - f_{x-1,y,z}}{2 \, dx}$, which is the slope of the extrapolated quadratic fit between points $(x+1, y, z), (x, y, z), (x-1, y, z)$.

2 Numerics Considerations

- Runge-Kutta takes the entire $12L^3$ size \vec{u} (including the $EPmDV_i, P^*$ variables) and updates it at once by computing a $d\vec{u}$. This requires three temporary arrays before obtaining $d\vec{u}$. Since this is evolving the entire array forward at each timestep, we require $4N$ operations to move forward a timestep, the same as if we run R-K at each point individually (fixing all other points on the grid constant) just requiring more space. However, the upshot is that we require storing updates to the whole grid to compute $d\vec{u}$ accurately using R-K, while using something like Euler we don't need temporary values and *can* ignore updates to the grid with no accuracy penalty.

GPU optimizations must be considered. Since GPU-side storage is limited and computing all these temporary arrays seems slow, it would seem that some local approximation to the R-K at each pixel would be preferable. While this loses some accuracy and actually requires at least as many operations, it may be preferred for its decreased memory footprint.