

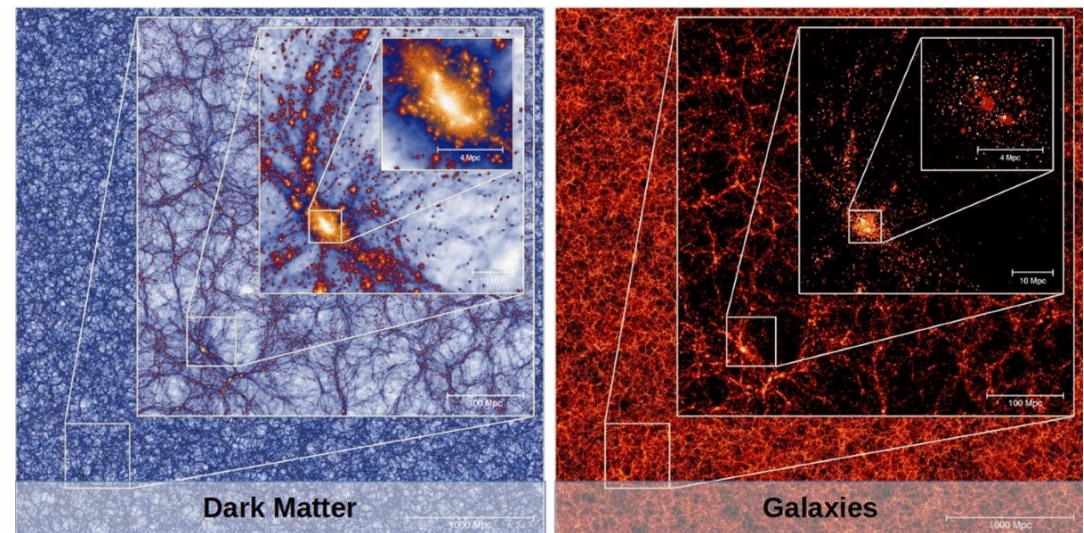


Australian
National
University

ASTRONOMICAL SIMULATIONS

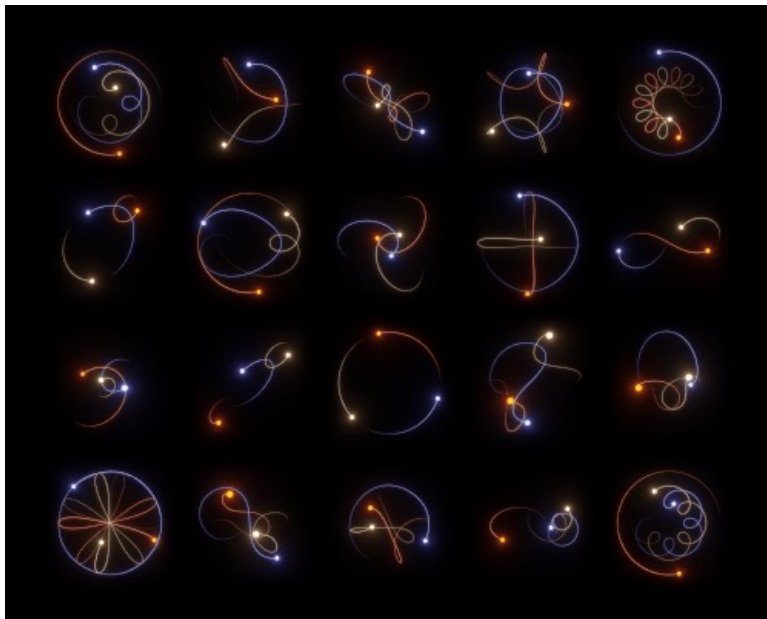
Yuxiang.Qin@anu.edu.au

Duffield Building D.115



N-body simulation

Isaac Newton gained renown for solving the two-body problem, showing how gravitational attraction binds the Earth and the Sun into elliptical orbits.



Moving on to the chaotic 3-body problem:

Leonhard Euler studied solutions when the three masses are in the same line.

Joseph-Louis Lagrange later discovered special solutions when the three masses form an equilateral triangle.

Then we have computers...



N-body simulation

Pseudocode

```
def n_body_simulation(N, dt, total_time):  
    positions = np.random.rand(N, 3) * volume_size # initialize positions  
    velocities = np.random.rand(N, 3) * velocity_scale # initialize velocities  
    masses = np.ones(N) * mass_scale # initialize masses  
  
    for step in range(int(total_time / dt)): # loop through all steps  
        forces = calculate_forces(positions, N) # gravity!  
  
        # Update velocities and position based on forces  
        for i in range(N):  
            velocities[i] += forces[i] / masses[i] * dt #  $v_{\text{new}} = v_{\text{old}} + (F/m) * dt$   
            positions[i] += velocities[i] * dt #  $r_{\text{new}} = r_{\text{old}} + v * dt$   
            positions %= box_size # periodic boundary conditions  
  
    return positions, velocities
```



N-body simulation

Pseudocode

$$\ddot{\vec{r}}_j = -G \sum_{i \neq j}^N \frac{m_i}{|\vec{r}_i - \vec{r}_j|^3} (\vec{r}_i - \vec{r}_j)$$

```
def calculate_forces(positions, N):
    forces = np.zeros((N, 3)) # Initialize force array

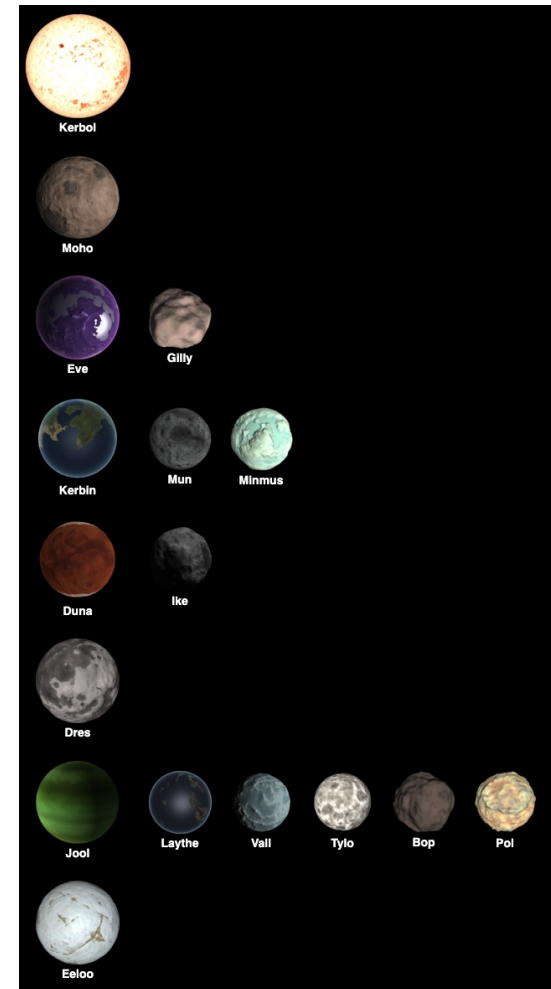
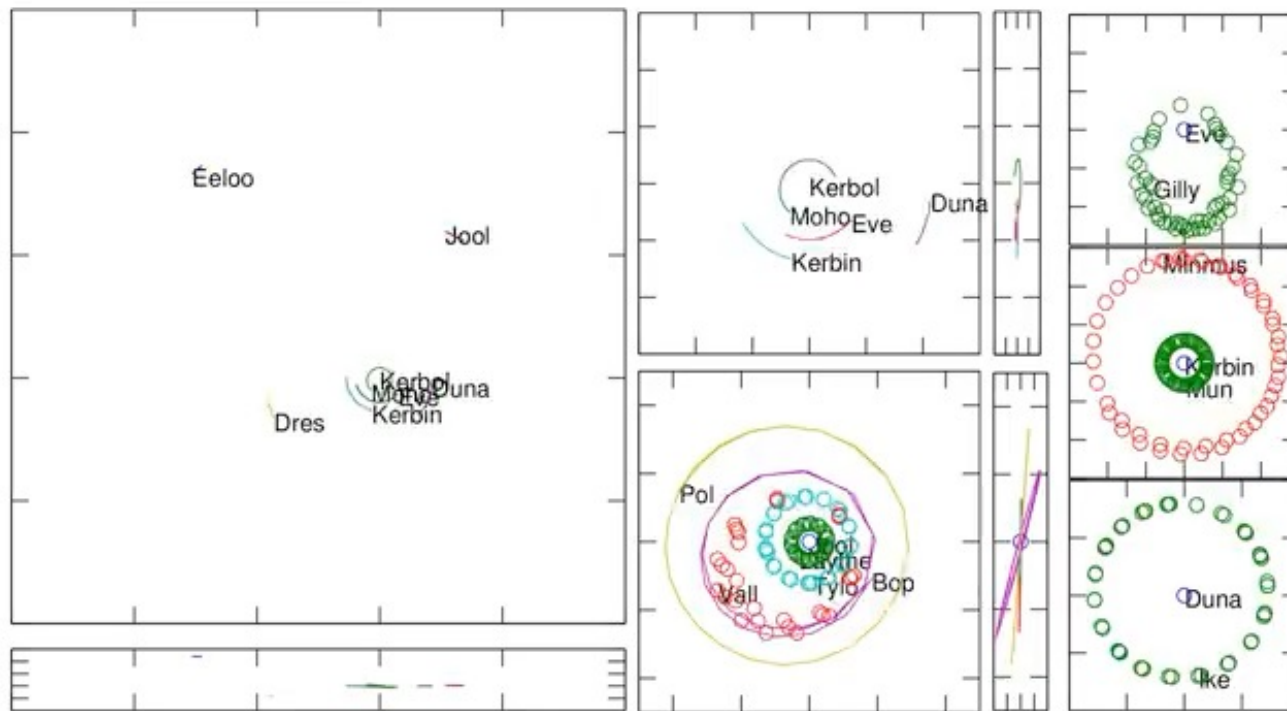
    for i in range(N):
        for j in range(i + 1, N):
            # Calculate gravitational force between particles i and j
            distance_vector = positions[j] - positions[i]
            distance = np.linalg.norm(distance_vector)
            force_magnitude = G * (masses[i] * masses[j]) / (distance ** 2 + soften**2)
            force_vector = force_magnitude * distance_vector / distance
            forces[i] += force_vector # Force on particle i
            forces[j] -= force_vector # Equal and opposite force on particle j

    return forces
```



Planetary system

Kerbal Space Program



Star cluster

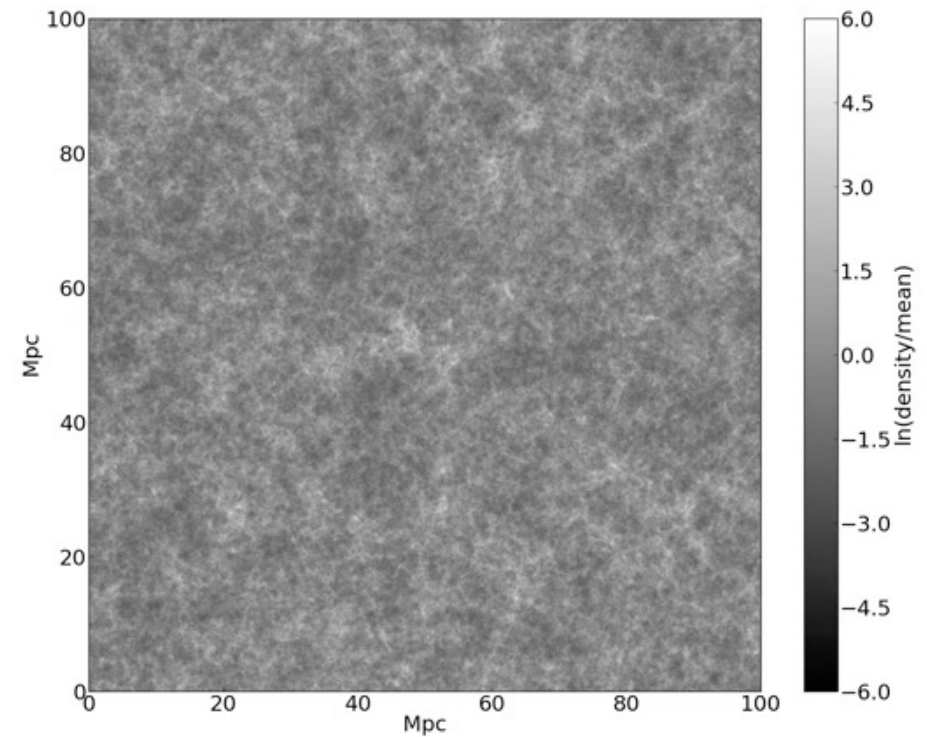
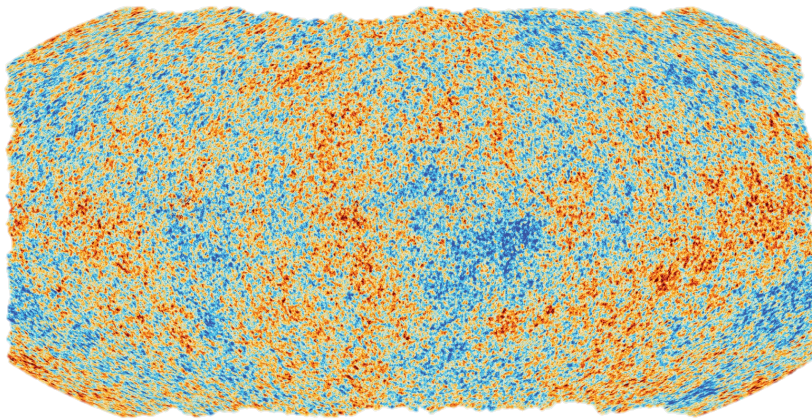
Stellar group R136 in the 30 Doradus nebula, in LMC



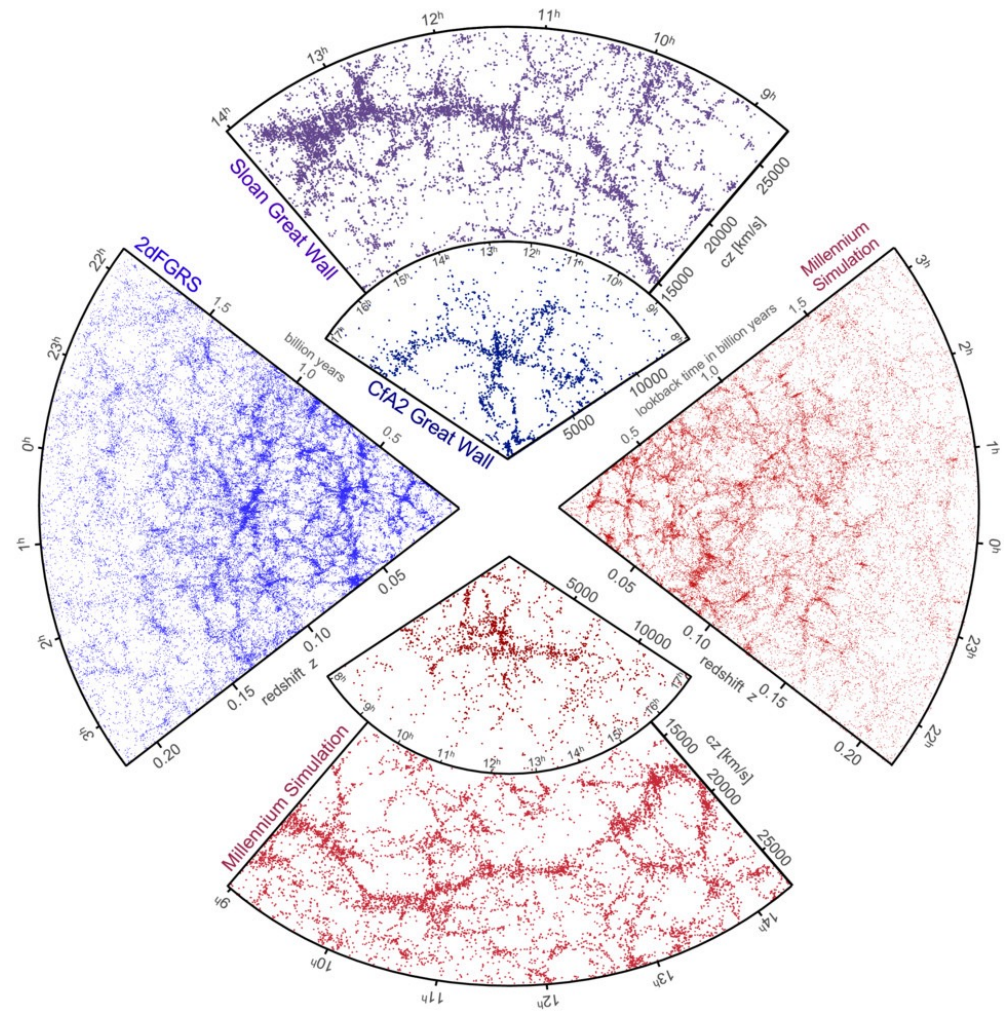
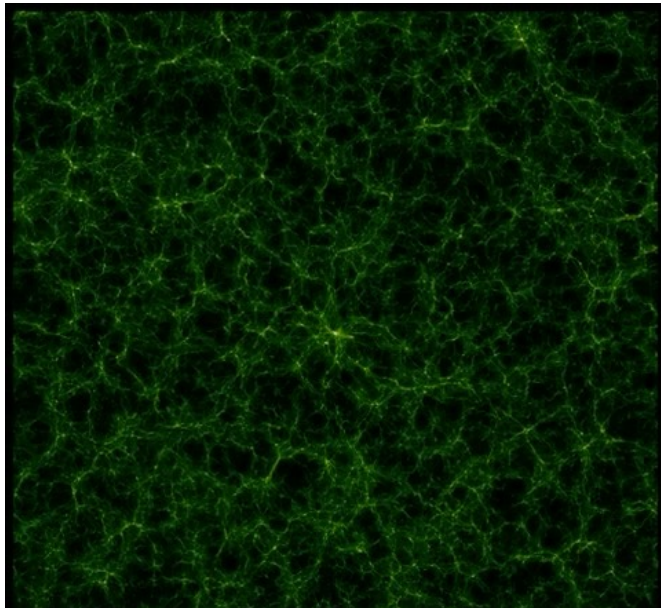
Andromeda colliding with the Milky Way



Cosmological N-body simulation



Cosmological N-body simulations



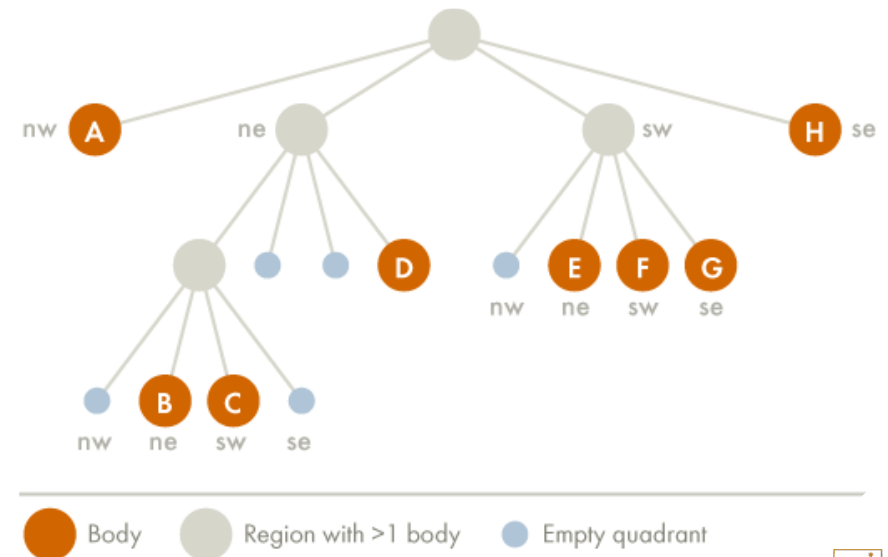
open simulations-part1.ipynb



The Barnes-Hut Algorithm

When particles are sufficiently far away, we can approximate their gravitational force as a group using their centre of *mass*.

Octree structure: the computational domain is hierarchically partitioned into a sequence of cubes, where each cube contains 8 siblings & each with half the side-length of the parent cube.



The Barnes-Hut Algorithm

-- Domain Decomposition

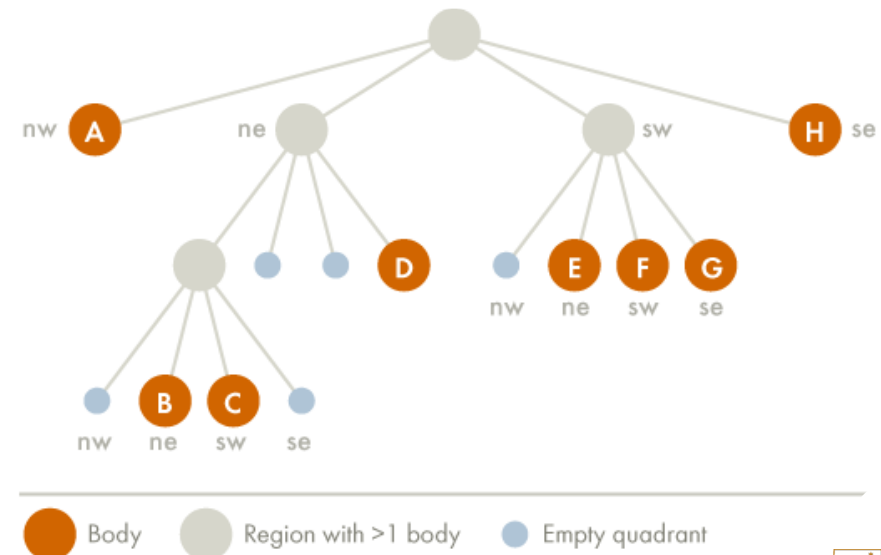
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    return (x>=cx) + 2*(y>=cy) + 4*(z>=cz)

def build_tree(positions, masses):
    root = Node(bounding_box_of_all)
    for each particle p:
        insert(root, p)
    compute_mass(root)
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def insert(node, p):
    if node.is_leaf and few_particles:
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The Barnes-Hut Algorithm

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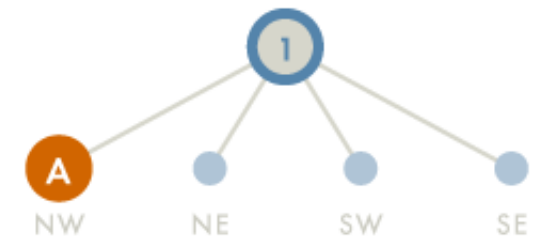
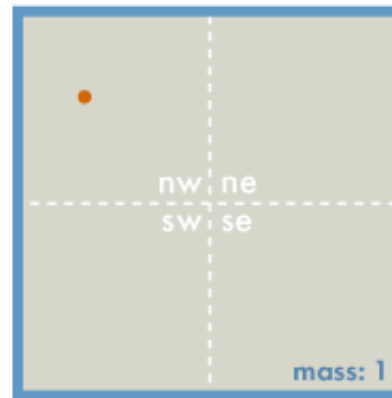
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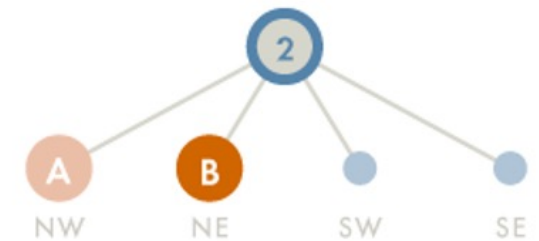
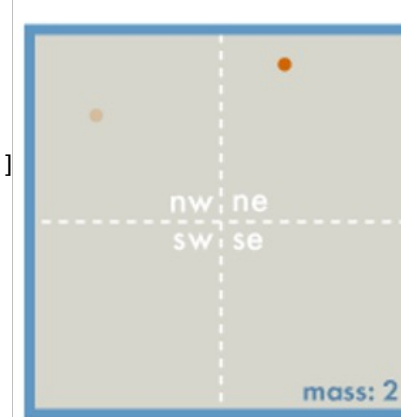
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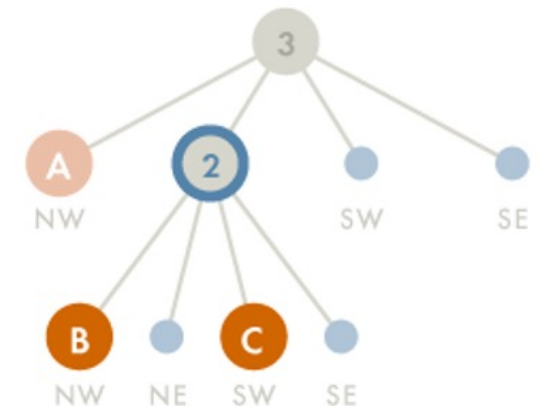
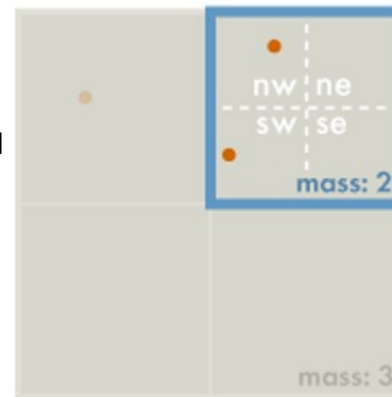
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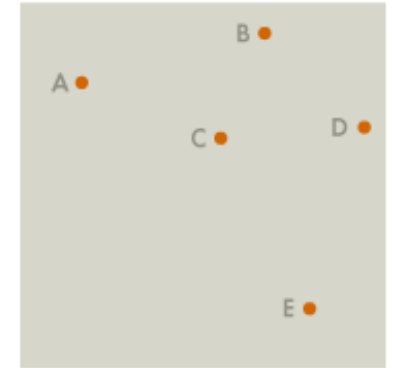
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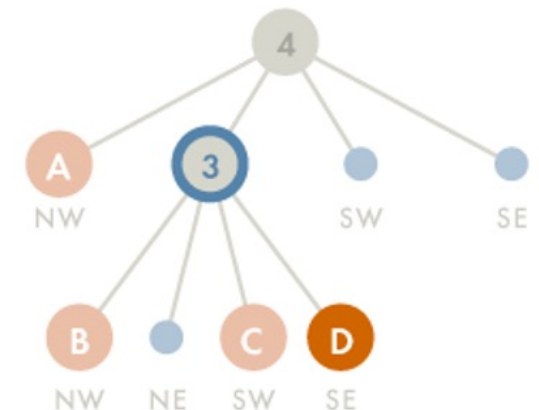
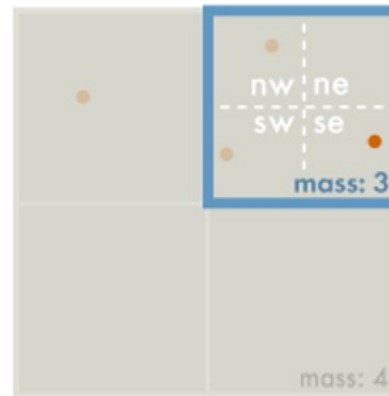
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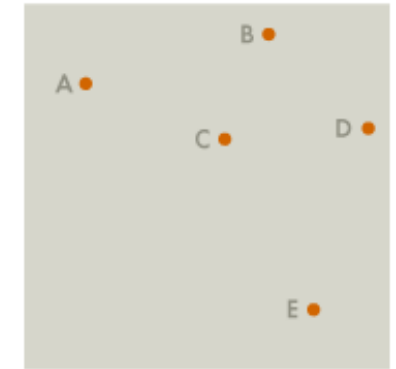
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-- Domain Decomposition



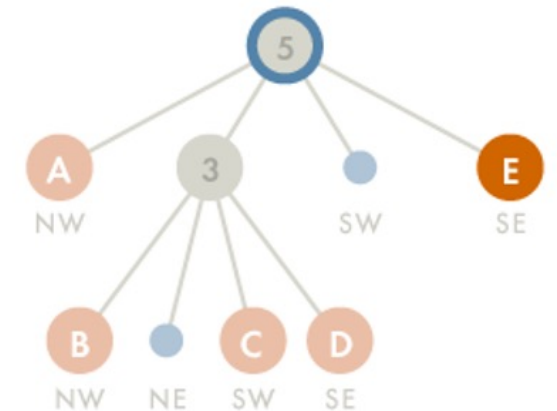
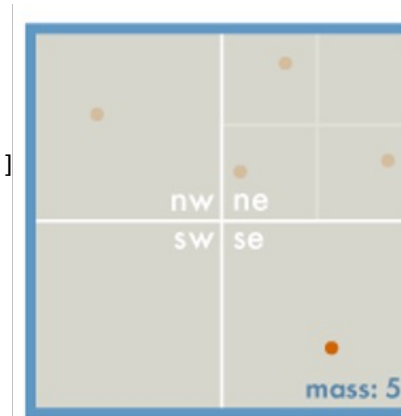
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The Barnes-Hut Algorithm

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def calculate_forces(pos, mass, G, softening):
    acc = zeros_like(pos)
    root = build_octree(pos, mass)
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    def add_from_node(i, node):
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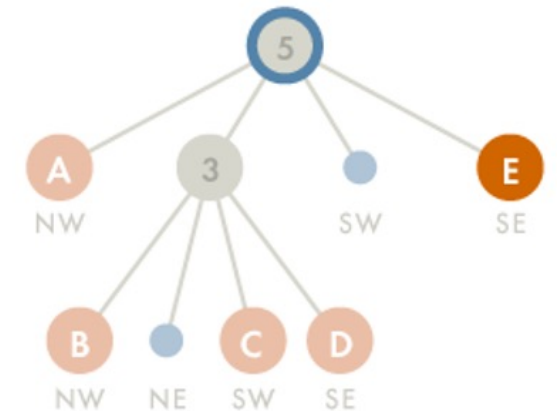
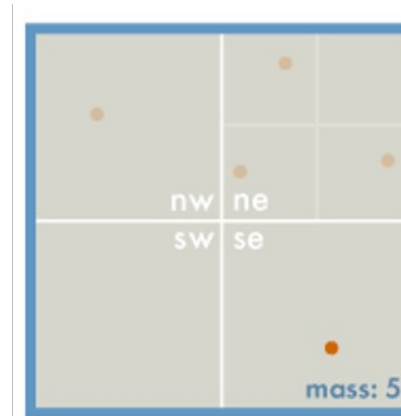
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            for j in node.idx:
                if j == i: continue
                dr = pos[j] - pos[i]
                dist = norm(r)
                inv = 1 / (dist**2 + soft2)**1.5
                acc[i] += G * mass[j] * dr * inv
        else:
            dr = node.com - pos[i]
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            if node.size()/dist < theta:
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    for i in range(len(pos)):
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-- Calculate Forces



The Barnes-Hut Algorithm

-- Calculate Forces

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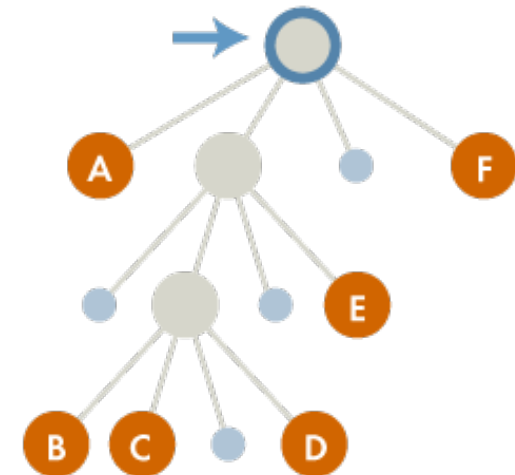
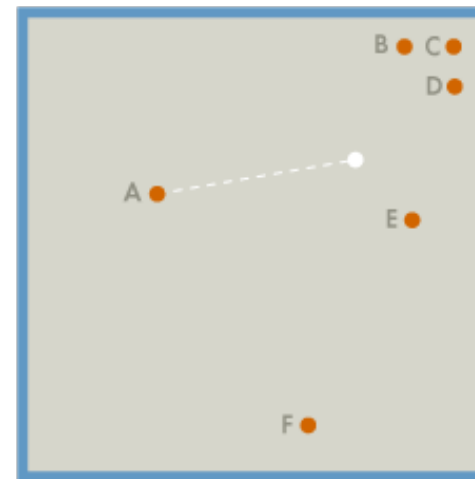
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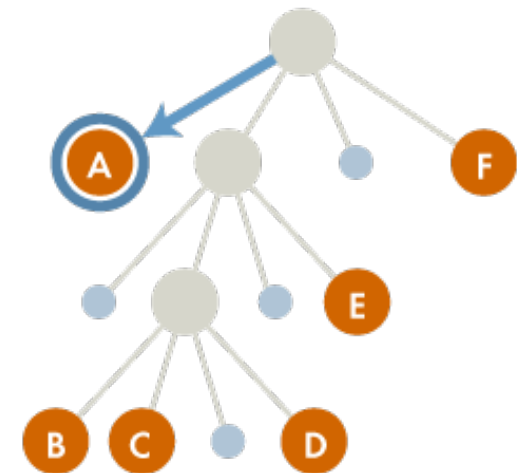
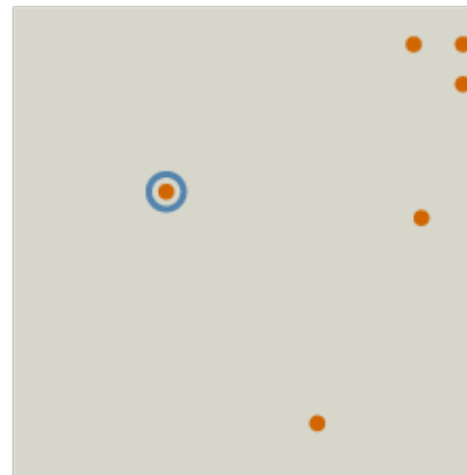
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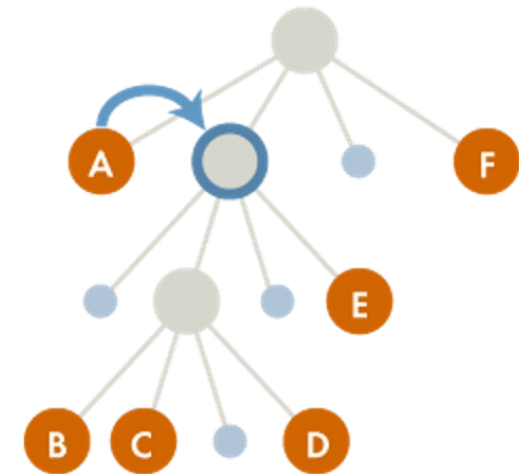
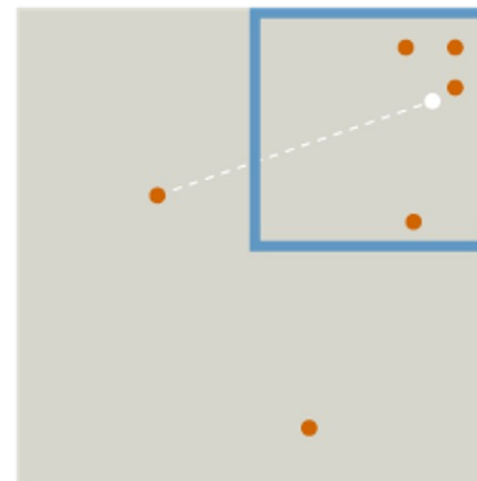
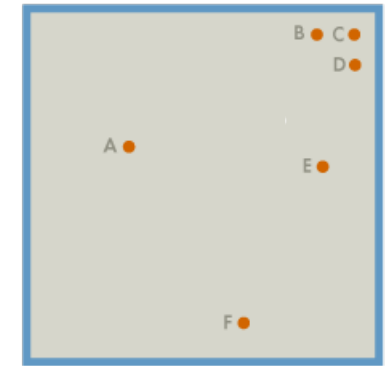
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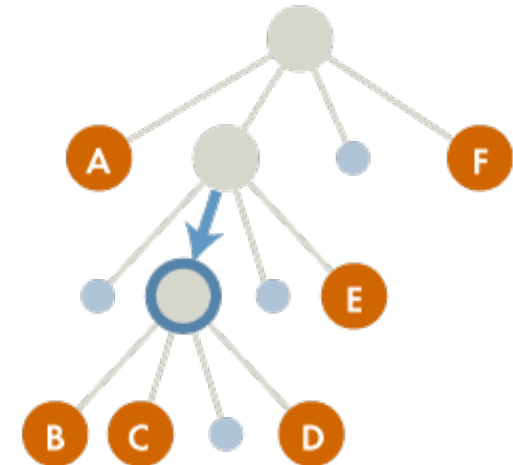
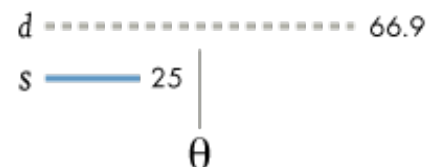
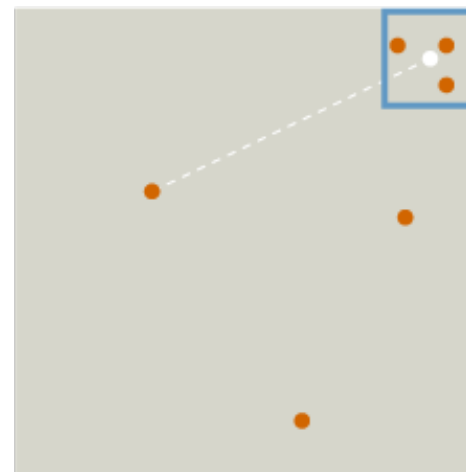
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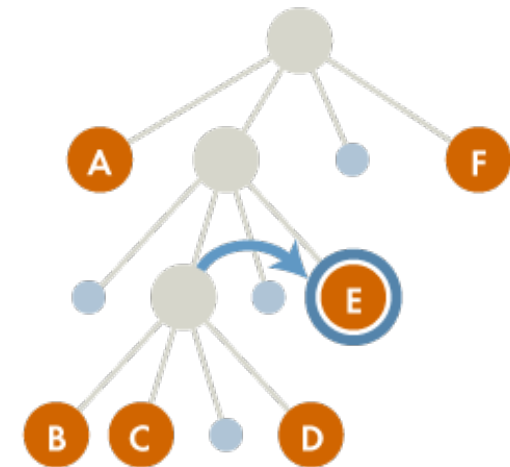
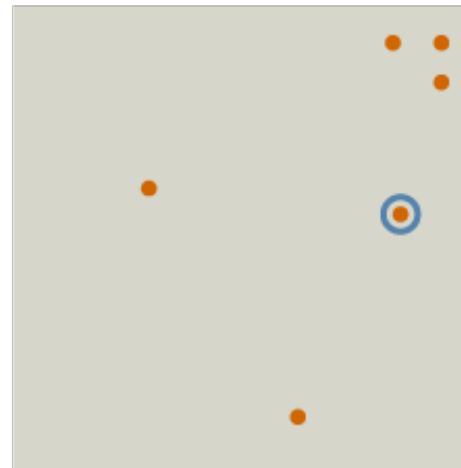
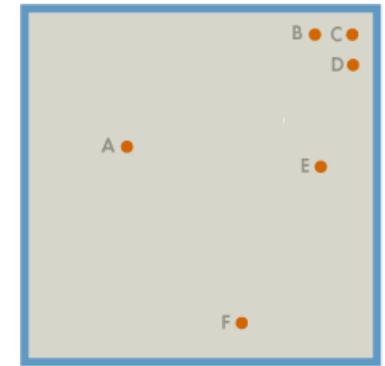
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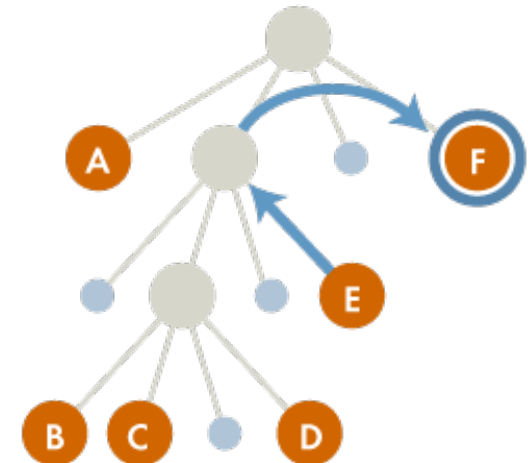
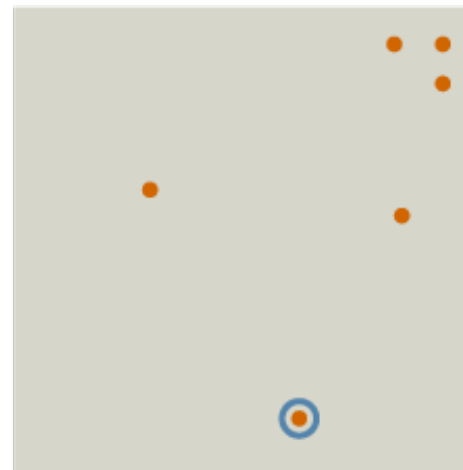
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            if node.size()/dist < theta:
                inv = 1 / (dist**3 + soft2**1.5)
                acc[i] += G * node.mass * dr * inv
            else:
                for ch in node.children:
                    add_from_node(i, ch)

    for i in range(len(pos)):
        add_from_node(i, root)

    return acc
```

-- Calculate Forces



open simulations-part2.ipynb

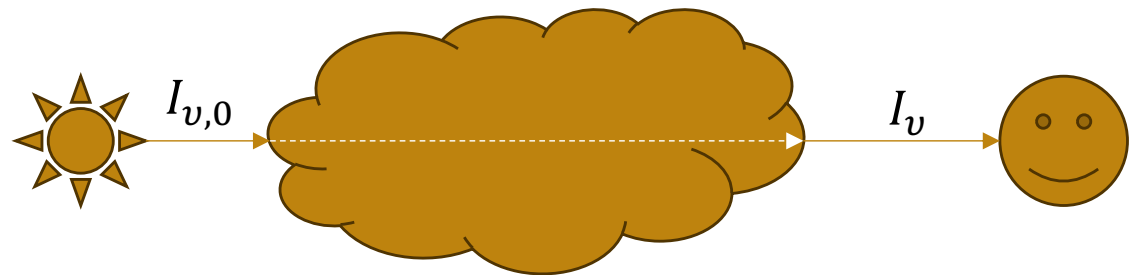


Radiative Transfer

$$\frac{dI_v}{ds} = j_v - \alpha_v I_v$$

Optical depth: $\tau_v = \int_0^s ds' \alpha_v$

$$I_v = I_{v,0} e^{-\tau_v} + \int_0^s ds' j_v e^{-\tau_v(s')}$$



Radiative Transfer Pseudocode

$$\frac{dI_\nu}{ds} = j_\nu - \alpha_\nu I_\nu$$

$$\text{Optical depth: } \tau_\nu = \int_0^s ds' \alpha_\nu$$

$$I_\nu = I_{\nu,0} e^{-\tau_\nu} + \int_0^s ds' j_\nu e^{-\tau_\nu(s')}$$

```
grid = initialize_grid(GRID_SIZE)
```

```
# Main ray tracing loop
for i in range(NUM_RAYS):
    ray = initialize_ray()
    trace_ray(ray, grid)
```

```
# Apply recombination in the grid after processing all photons
for cell in grid:
    cell["ionization_fraction"] *= (1 - RECOMBINATION_RATE)
```

```
def initialize_ray():
    # Initialize ray properties (starting position, direction)
    ray = {
        "position": SOURCE_POSITION,
        "direction": random_direction(),
        "distance_traveled": 0.0,
        "absorbed": False
    }
    return ray

def random_direction():
    # Generate random spherical coordinates
    theta = np.random.uniform(0, 2 * np.pi) # azimuthal angle
    phi = np.random.uniform(0, np.pi) # polar angle

    # Convert spherical coordinates to Cartesian coordinates
    x = np.sin(phi) * np.cos(theta)
    y = np.sin(phi) * np.sin(theta)
    z = np.cos(phi)

    # Normalize the vector to ensure it has a length of 1
    direction = np.array([x, y, z])
    return direction
```



Radiative Transfer Pseudocode

```
def trace_ray(ray, grid):
    while ray.distance_traveled < MAX_DISTANCE:
        # Update ray position based on direction
        ray.position += ray.direction * TIME_STEP # Move ray forward
        ray.distance_traveled += TIME_STEP

        # Check for grid boundaries
        cell = get_cell_at_position(grid, ray.position)
        if cell:
            # Check for absorption based on grid properties
            if check_for_absorption(ray, cell):
                ray.absorbed = True
                cell["ionization_fraction"] += (CELL_VOLUME * IONIZATION_EFFICIENCY) # update ionization fraction
                break # Stop tracing this ray if absorbed

def get_cell_at_position(grid, position):
    for cell in grid:
        if (cell["x_min"] <= position[0] <= cell["x_max"] and
            cell["y_min"] <= position[1] <= cell["y_max"] and
            cell["z_min"] <= position[2] <= cell["z_max"]):
            return cell
    return None

def check_for_absorption(ray, cell):
    # Determine if the ray is absorbed based on the cell properties
    tau = cell["optical_depth"] * cell["cross_section"] * distance
    absorption_probability = 1 - np.exp(-tau)
    return random.uniform(0, 1) < absorption_probability(cell)
```



open simulations-part3.ipynb



prerequisite

We will start at 1:05pm. Before that, please

git clone/pull from <https://github.com/qyx268/astr4004-8004-2025.git>

Install ipympl

