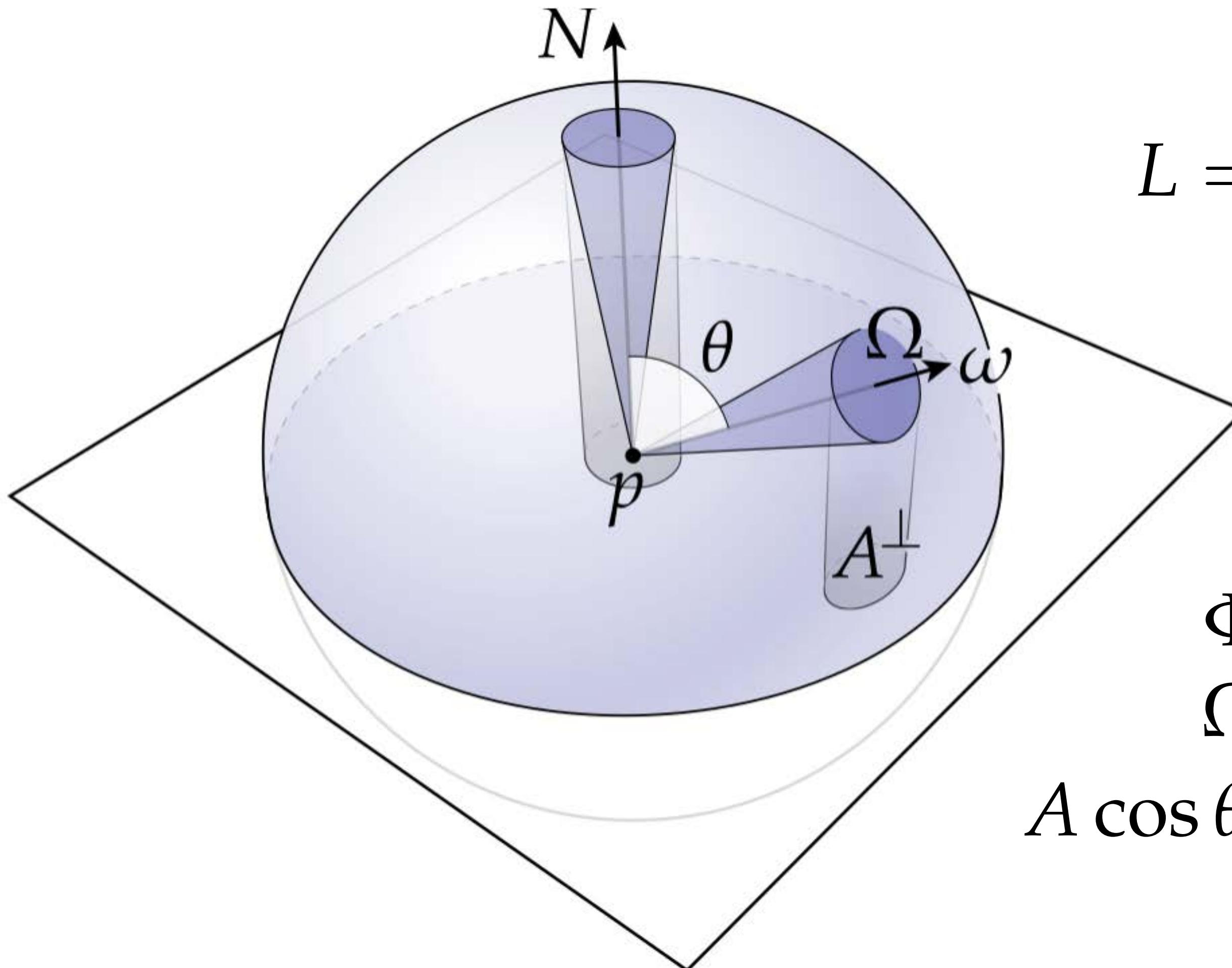


The Rendering Equation

Computer Graphics
CMU 15-462/15-662, Fall 2016

Review: What is *radiance*?

- Radiance at point p in direction N is radiant energy (“#hits”) per unit time, per solid angle, per unit area perpendicular to N .



$$L = \frac{\partial^2 \Phi}{\partial \Omega \partial A \cos \theta}$$

Φ — radiant energy

Ω — solid angle

$A \cos \theta$ — projected area

Intuition: Radiance and Irradiance



$$E = \int_{H^2} L(\omega) \cos \theta \, d\omega$$

irradiance
radiance in direction ω
angle between ω and normal

The equation $E = \int_{H^2} L(\omega) \cos \theta \, d\omega$ is shown with three red arrows pointing to its components: 'irradiance' points to the variable E , 'radiance in direction ω ' points to the function $L(\omega)$, and 'angle between ω and normal' points to the term $\cos \theta$.



Incident vs. Exitant Radiance

INCIDENT



EXITANT



In both cases: intensity of illumination is highly dependent on *direction* (not just location in space or moment in time).

The Rendering Equation

- Core functionality of photorealistic renderer is to estimate radiance at a given point p , in a given direction ω_o
- Summed up by the *rendering equation* (Kajiya):

$$L_o(p, \omega_o) = L_e(p, \omega_o) + \int_{\mathcal{H}^2} f_r(p, \omega_i) \rightarrow \omega_o L_i(p, \omega_i) \cos \theta d\omega_i$$

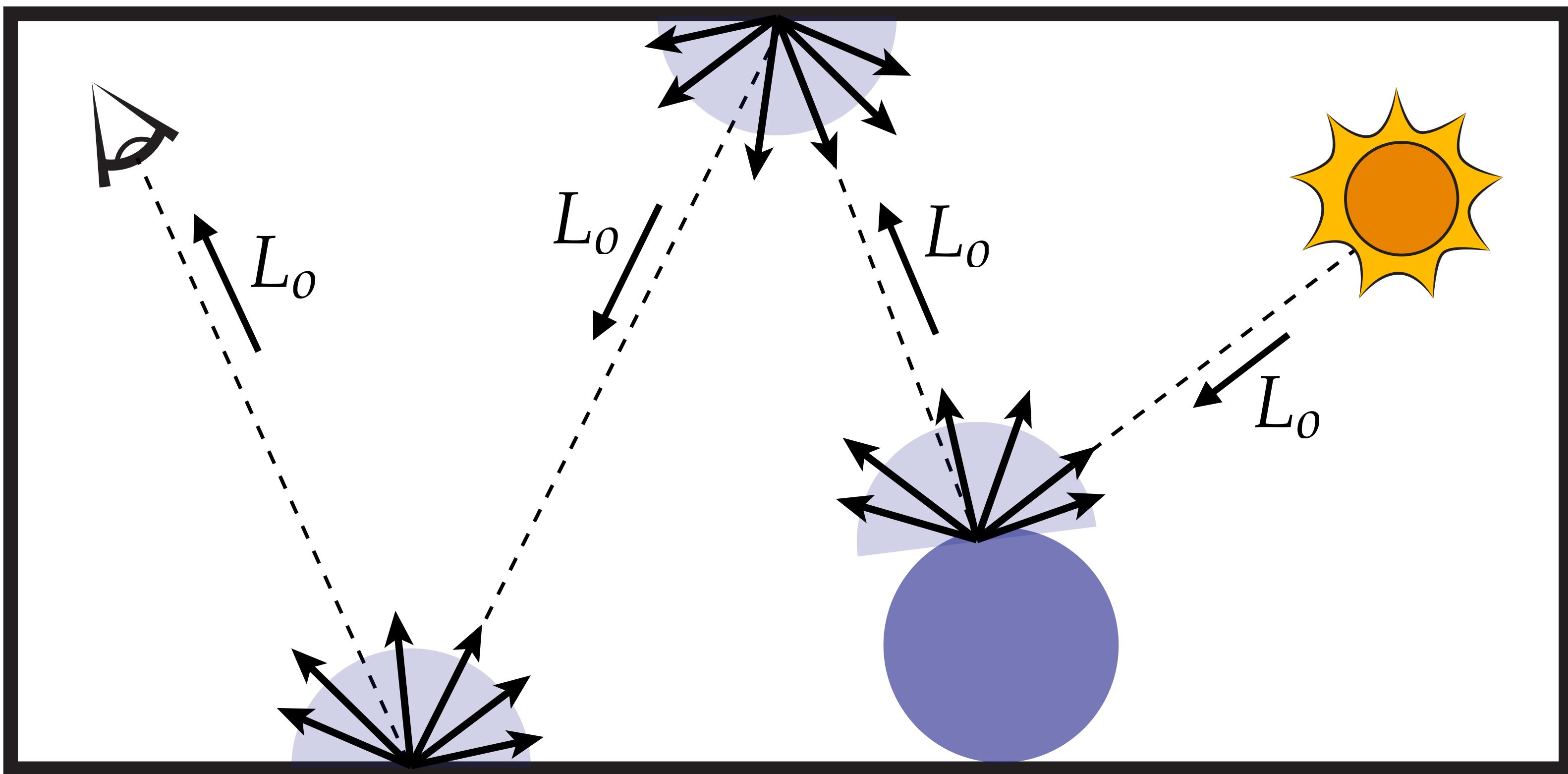
Diagram illustrating the components of the rendering equation:

- outgoing/observed radiance**: $L_o(p, \omega_o)$
- emitted radiance (e.g., light source)**: $L_e(p, \omega_o)$
- point of interest**: p
- direction of interest**: ω_o
- all directions in hemisphere**: \mathcal{H}^2
- scattering function**: $f_r(p, \omega_i)$
- incoming radiance**: $L_i(p, \omega_i)$
- angle between incoming direction and normal**: $\cos \theta$
- incoming direction**: ω_i

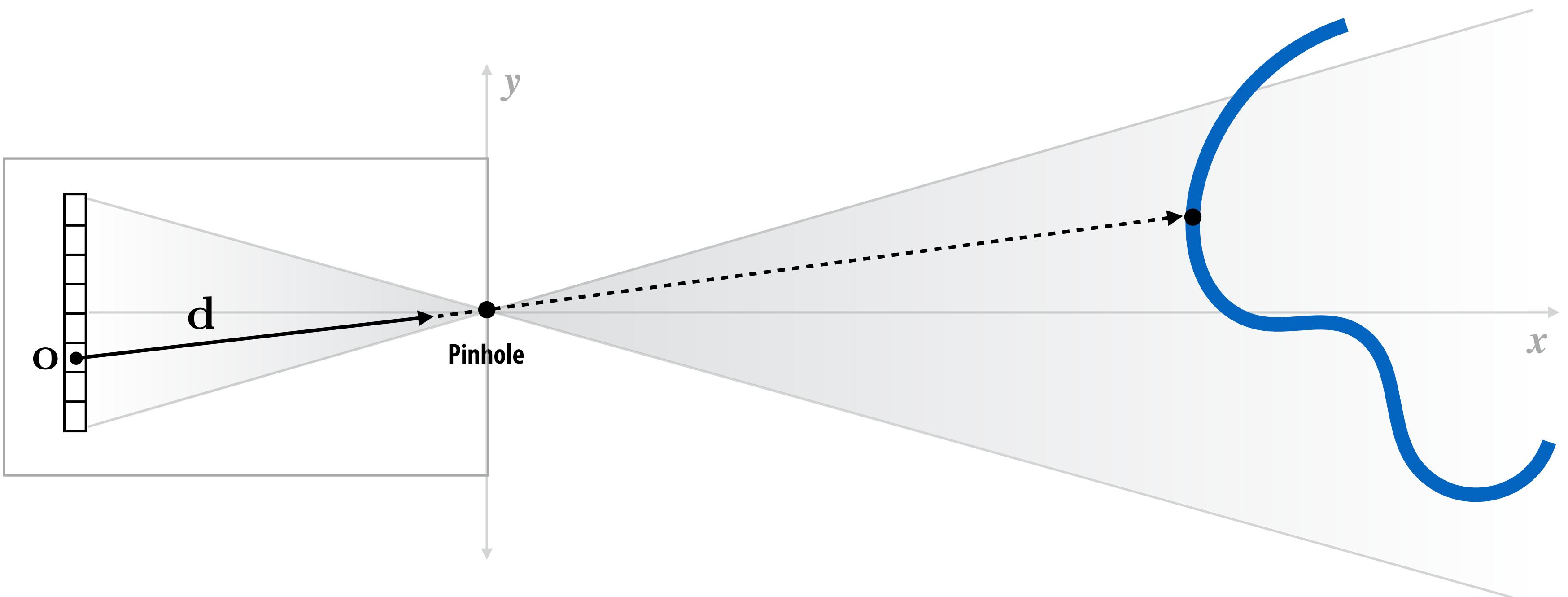
Key challenge: to evaluate incoming radiance, we have to compute yet another integral. I.e., rendering equation is *recursive*.

Recursive Raytracing

- Basic strategy: recursively evaluate rendering equation!

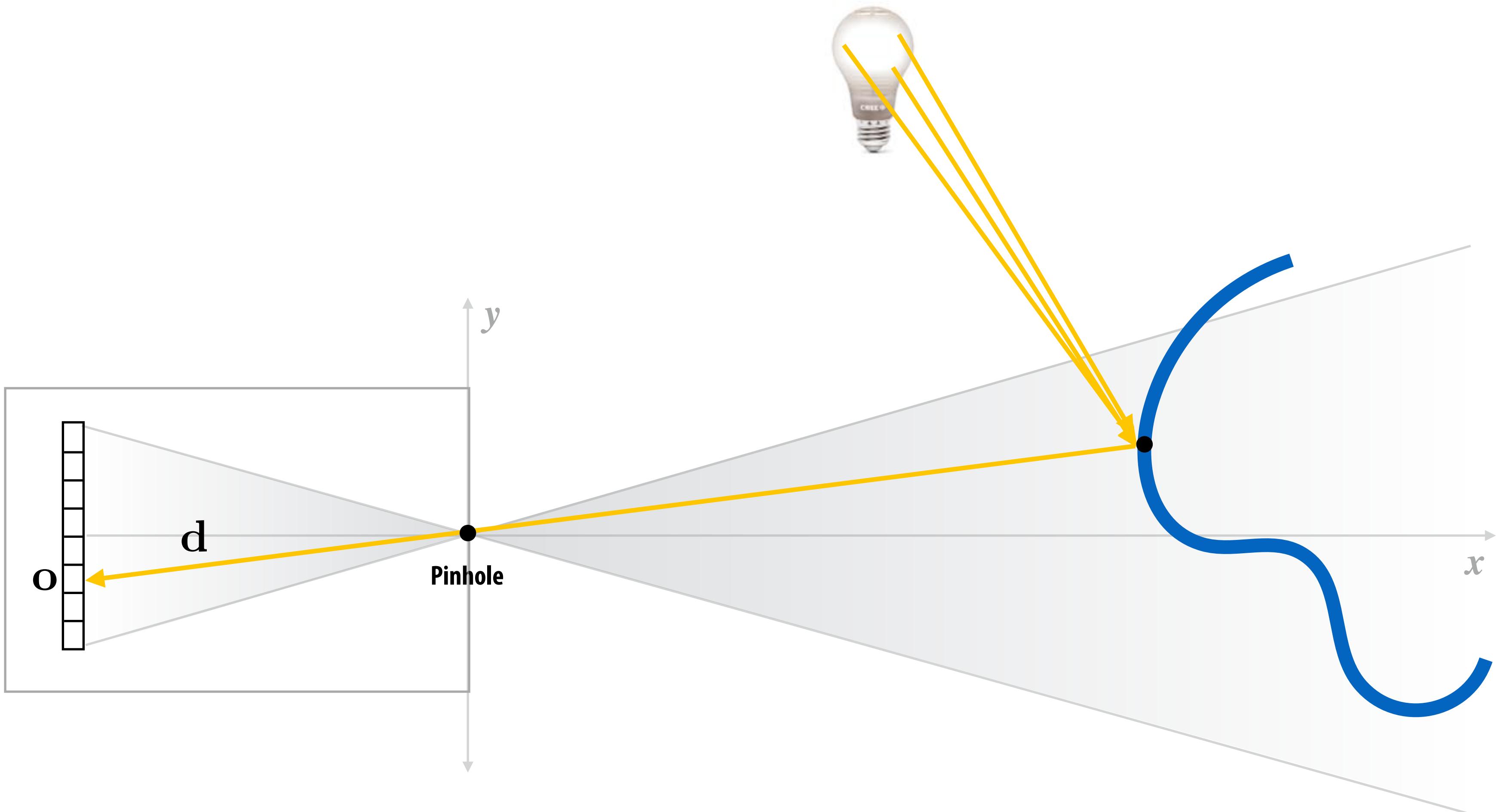


Renderer measures radiance along a ray

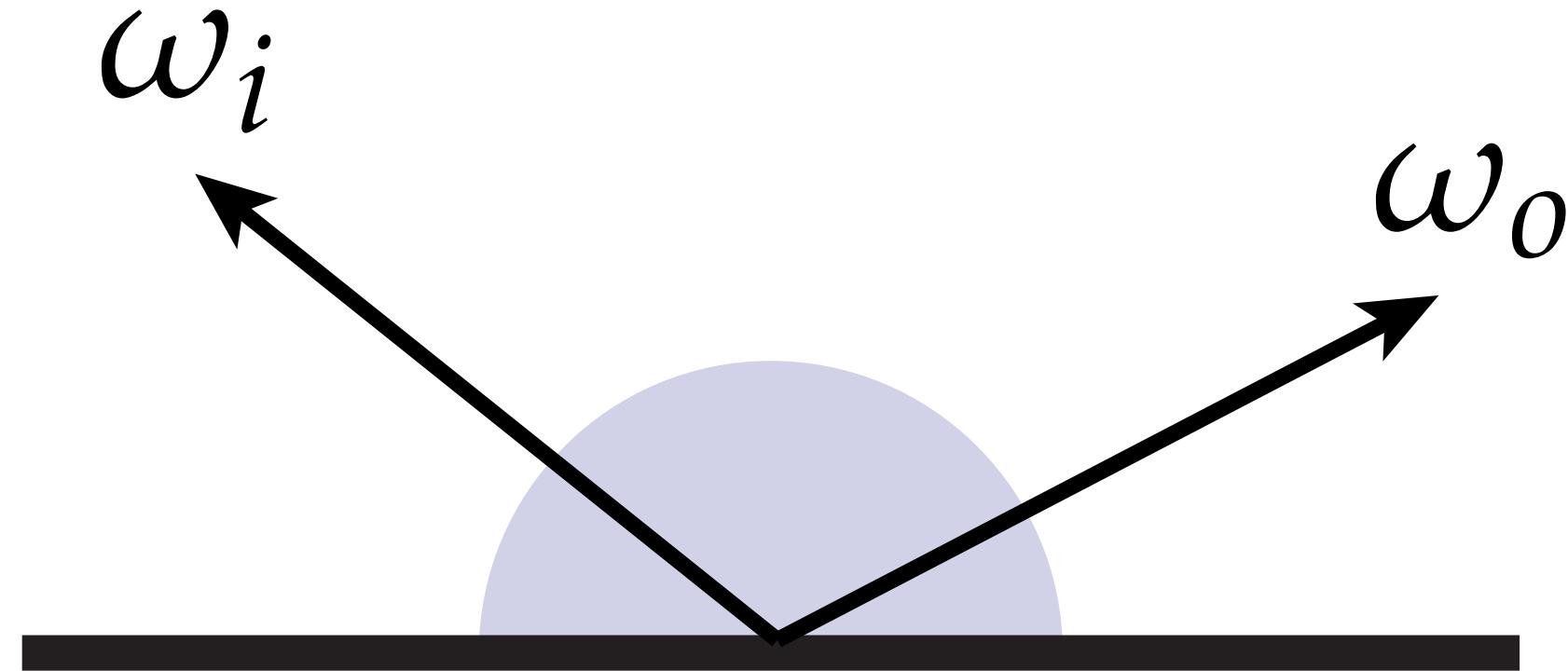


At each “bounce,” want to measure radiance traveling in the direction opposite the ray direction.

Renderer measures radiance along a ray



Radiance entering camera in direction d = light from scene light sources that is reflected off surface in direction d .



How does *reflection* of light affect
the outgoing radiance?

$$L_o(\mathbf{p}, \omega_o) = \int_{\mathcal{H}^2} f_r(\mathbf{p}, \omega_i, \omega_o) L_i(\mathbf{p}, \omega_i) \cos \theta d\omega_i$$

A red oval highlights the term $f_r(\mathbf{p}, \omega_i, \omega_o)$.

Reflection models

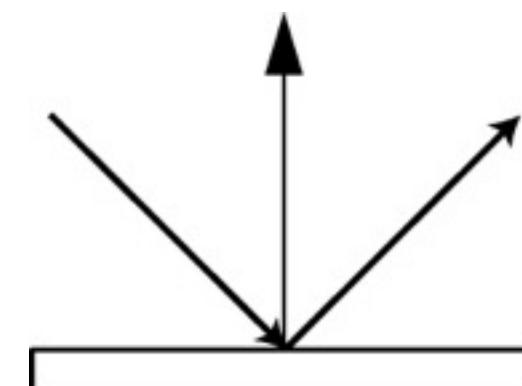
- **Reflection is the process by which light incident on a surface interacts with the surface such that it leaves on the incident (same) side without change in frequency**
- **Choice of reflection function determines surface appearance**



Categories of reflection functions

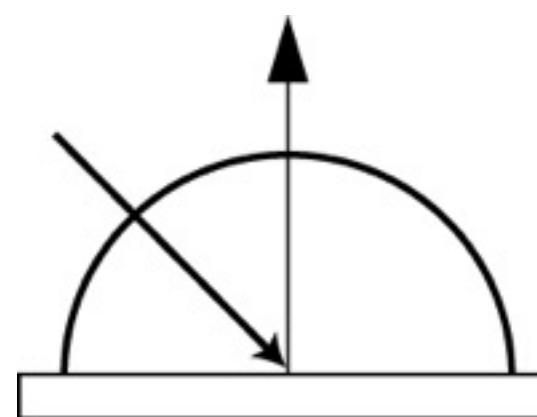
■ Ideal specular

Perfect mirror



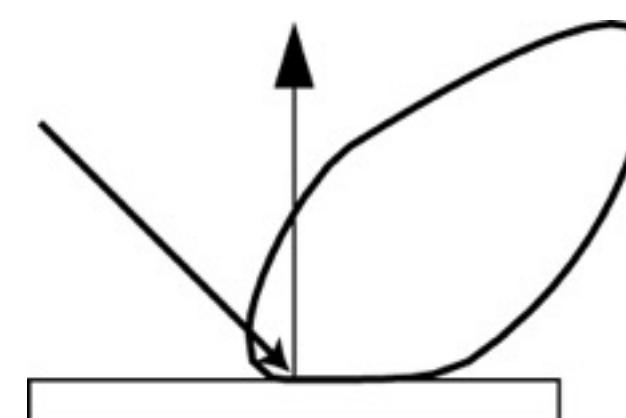
■ Ideal diffuse

Uniform reflection in all directions



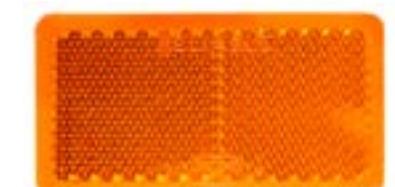
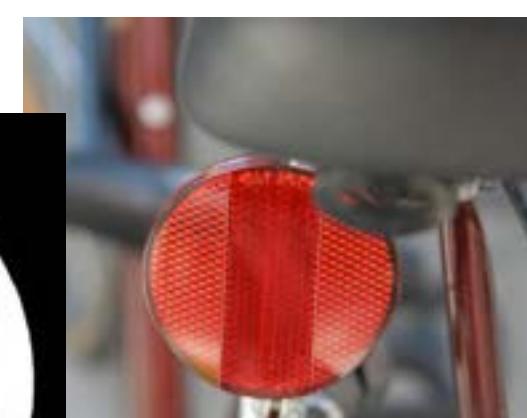
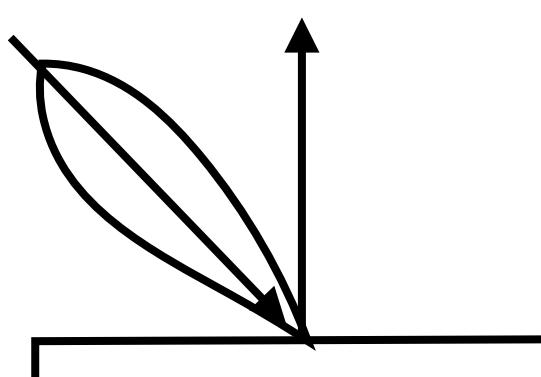
■ Glossy specular

Majority of light distributed in reflection direction



■ Retro-reflective

Reflects light back toward source



Diagrams illustrate how incoming light energy from given direction is reflected in various directions.

Materials: diffuse



Materials: plastic



Materials: red semi-gloss paint



Materials: Ford mystic lacquer paint



Materials: mirror



Materials: gold



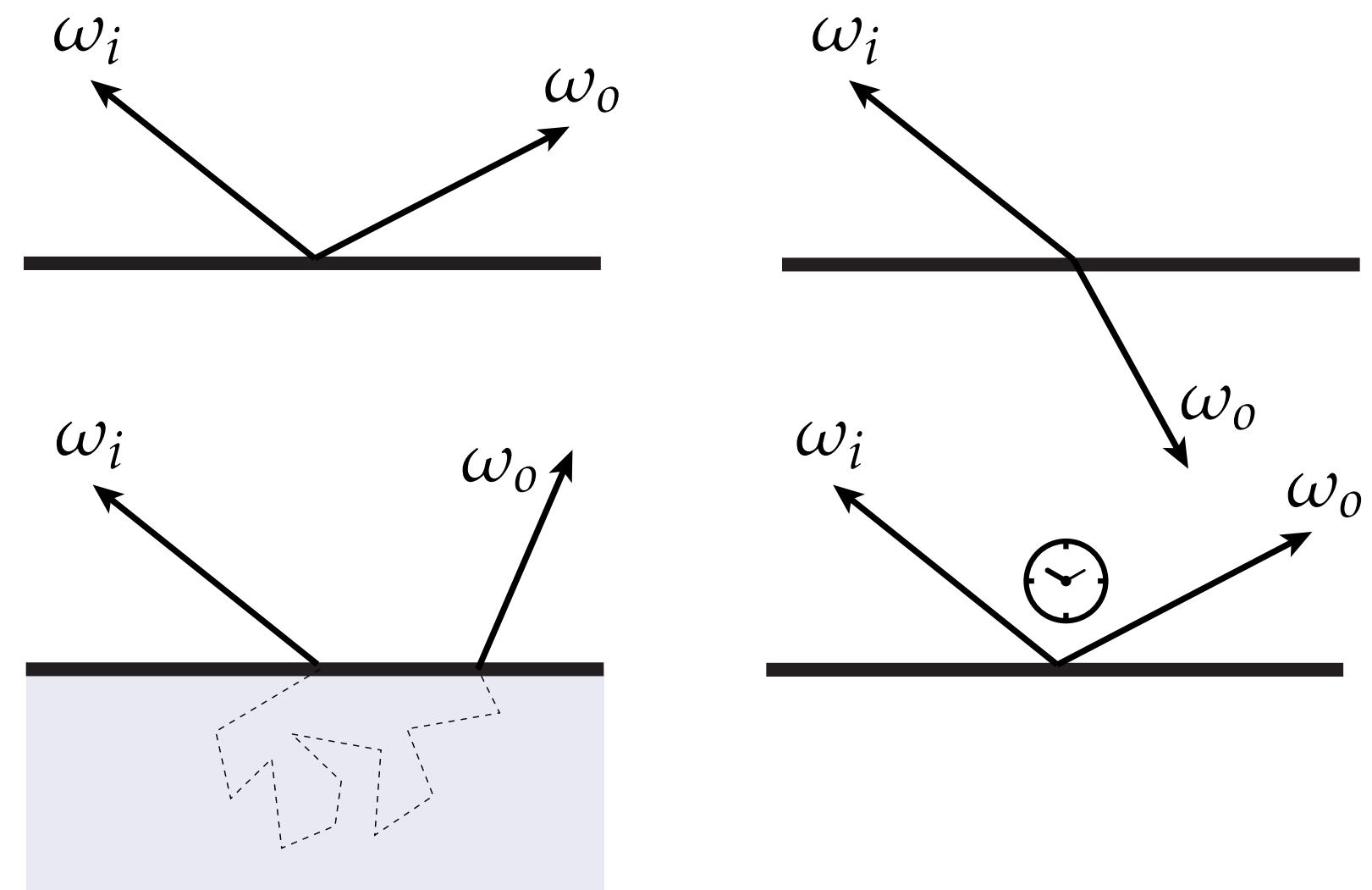
Materials



Models of Scattering

- How can we model “scattering” of light?
- Many different things that could happen to a photon:

- bounces off surface
- transmitted through surface
- bounces around inside surface
- absorbed & re-emitted
- ...

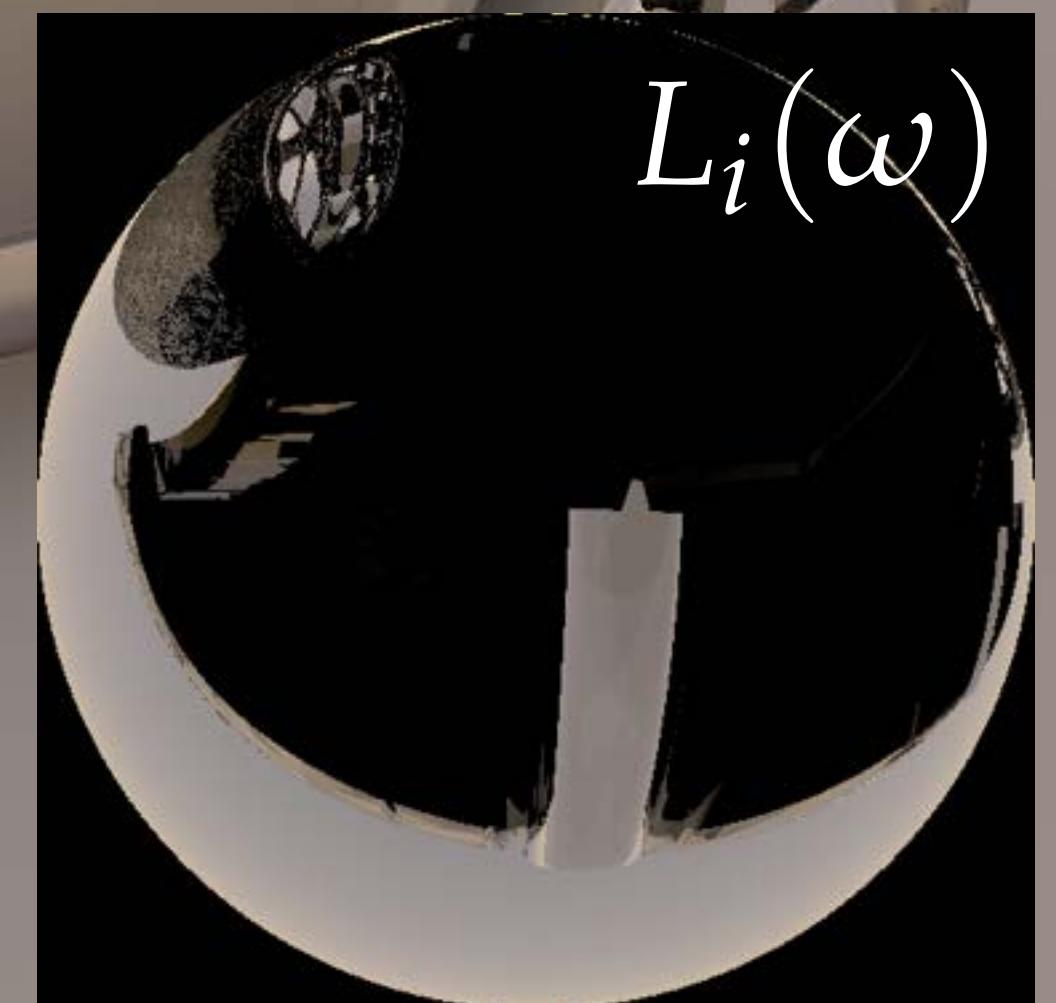
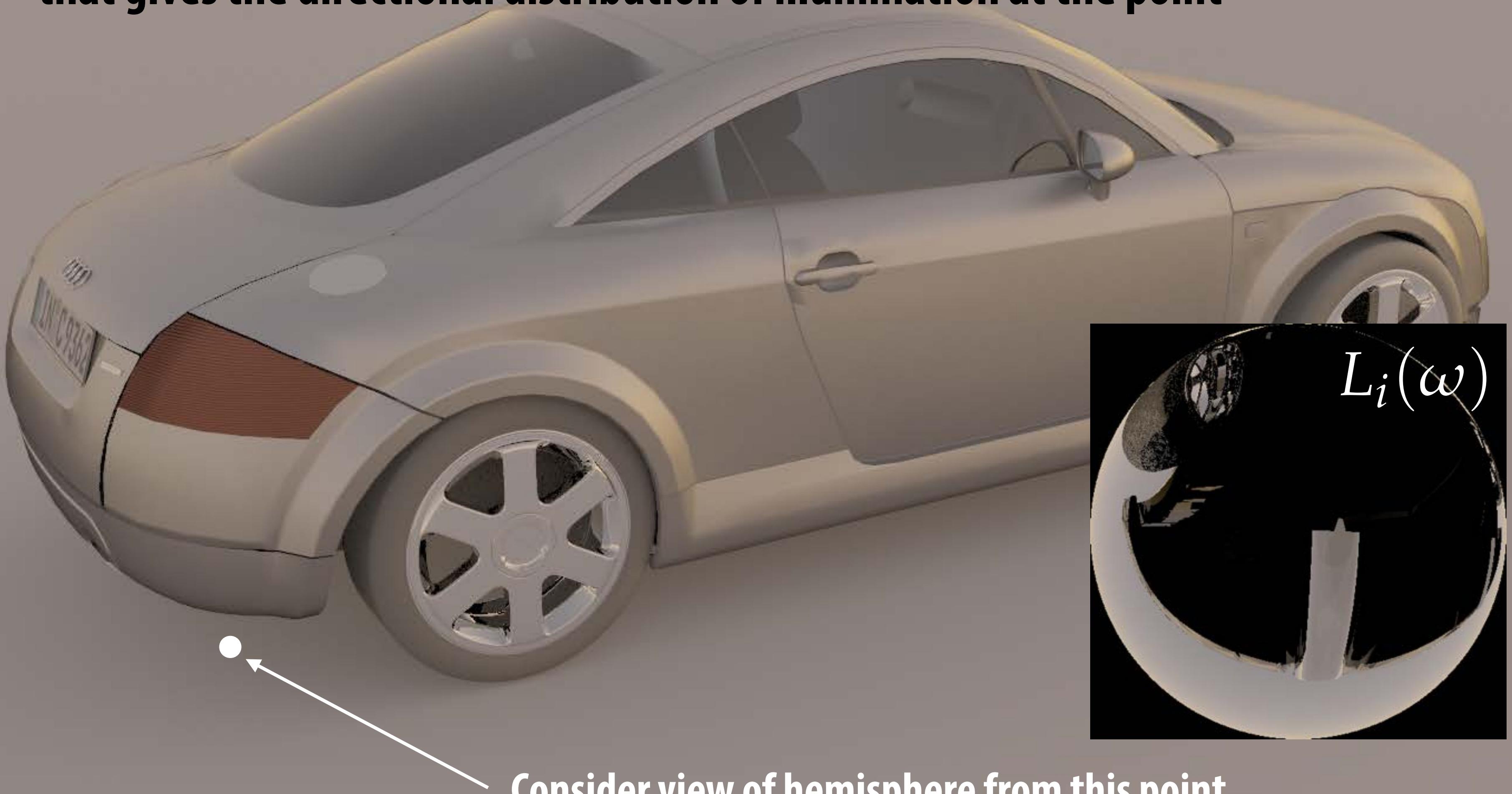


- What goes in must come out! (Total energy must be conserved)
- In general, can talk about “*probability**” a particle arriving from a given direction is scattered in another direction

*Somewhat more complicated than this, because some light is absorbed!

Hemispherical incident radiance

At any point on any surface in the scene, there's an incident radiance field that gives the directional distribution of illumination at the point



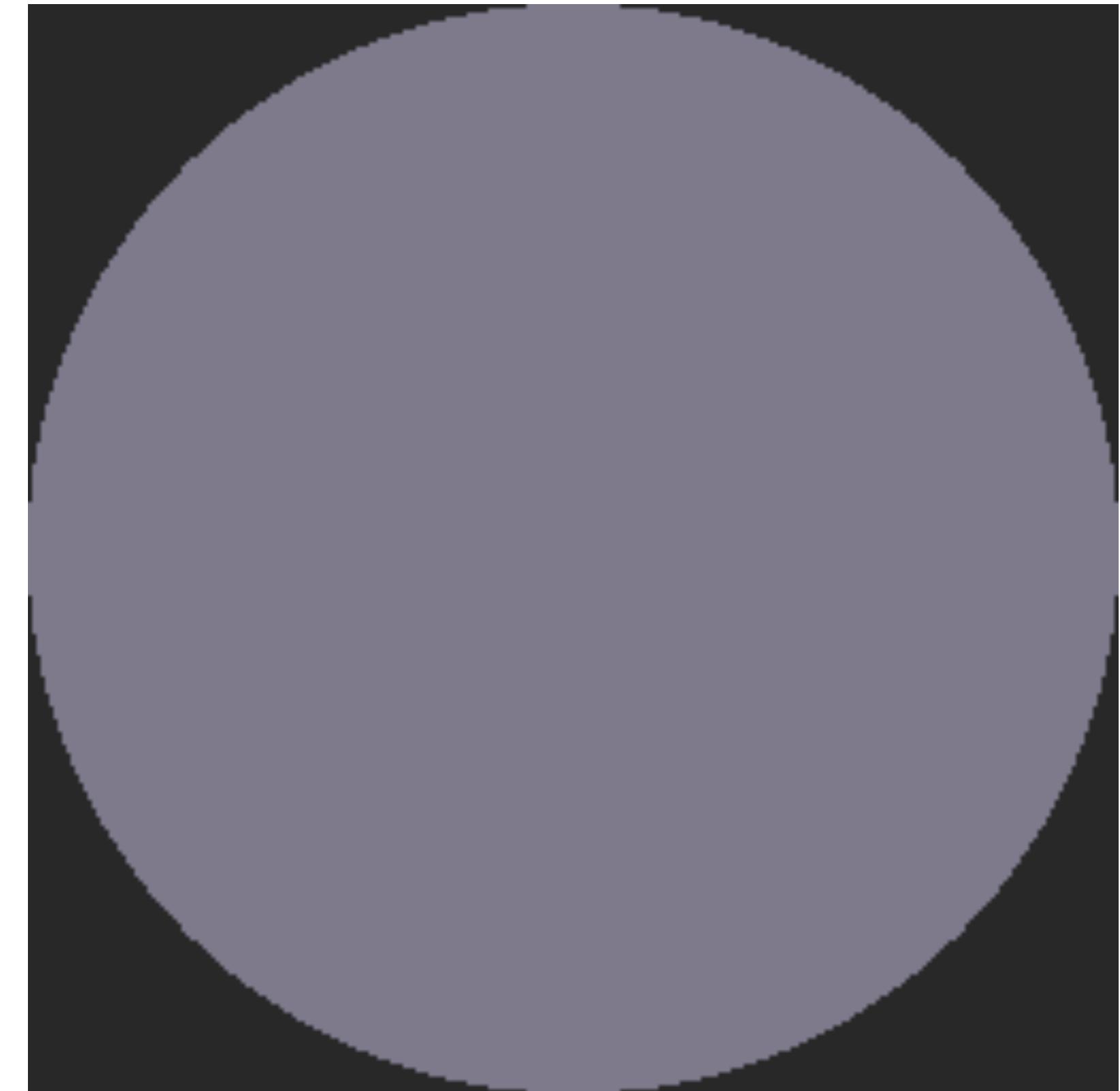
Consider view of hemisphere from this point

Diffuse reflection

Exitant radiance is the same in all directions



Incident radiance



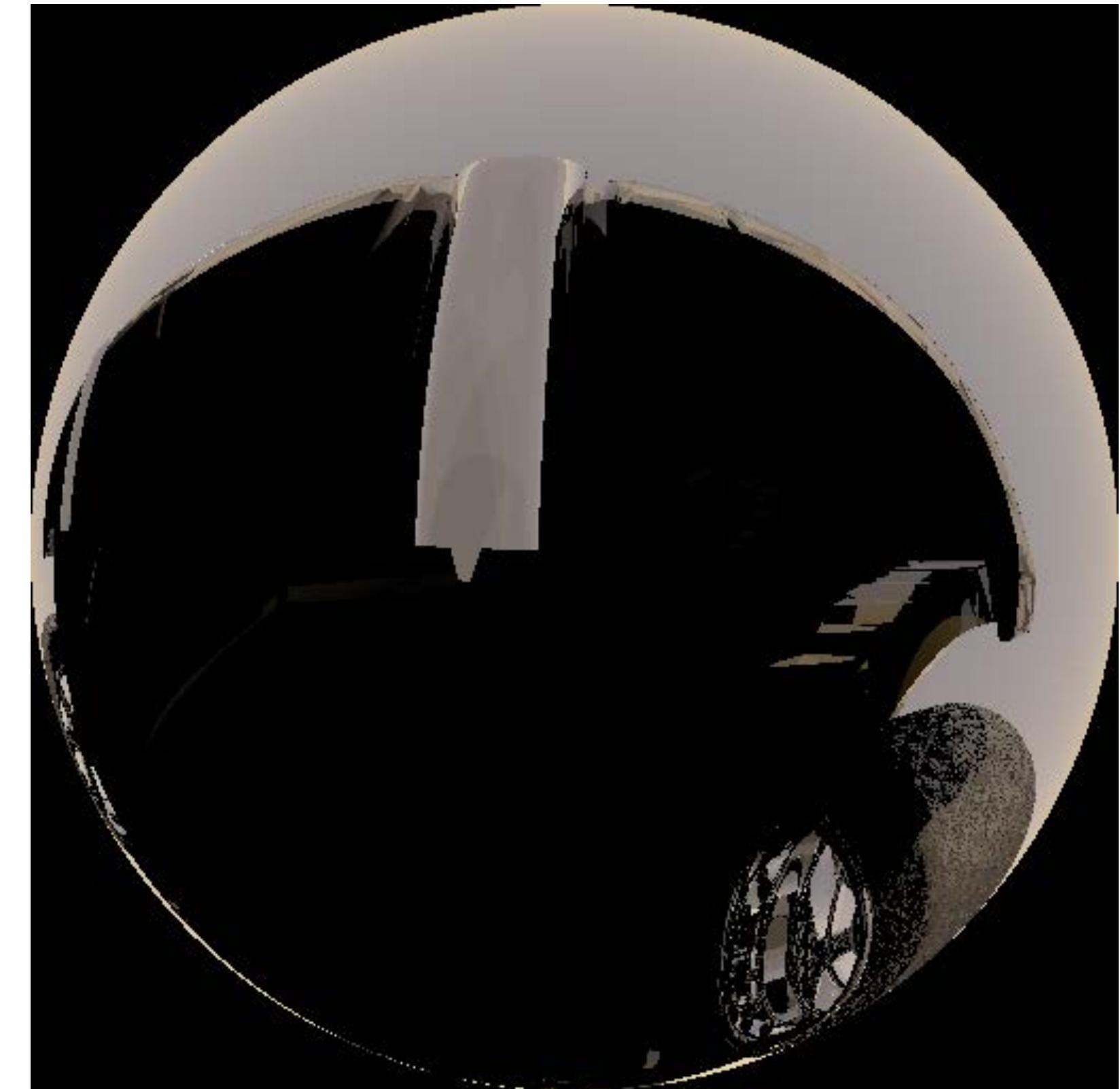
Exitant radiance

Ideal specular reflection

Incident radiance is “flipped around normal” to get exitant radiance



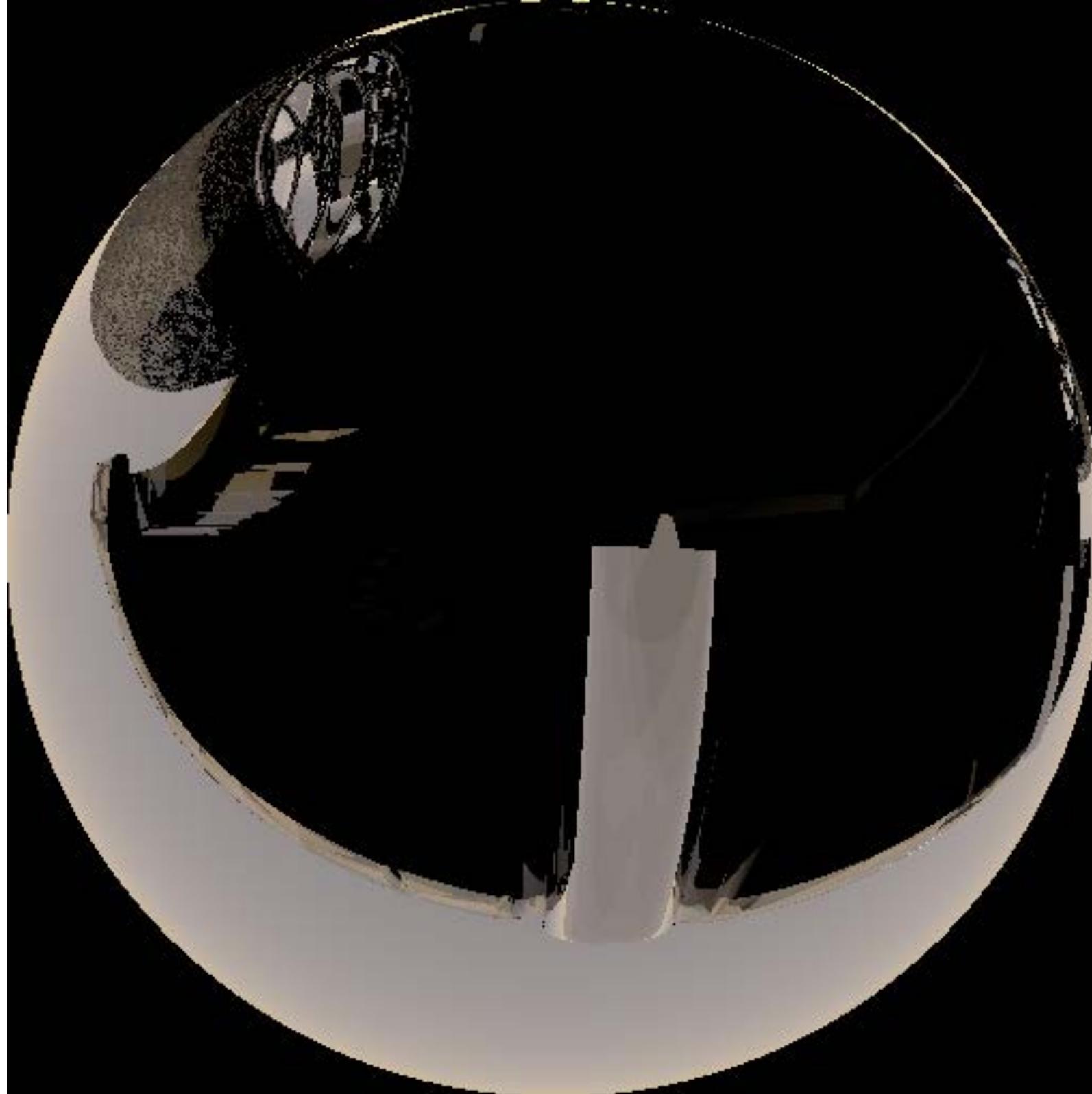
Incident radiance



Exitant radiance

Plastic

Incident radiance gets “flipped and blurred”



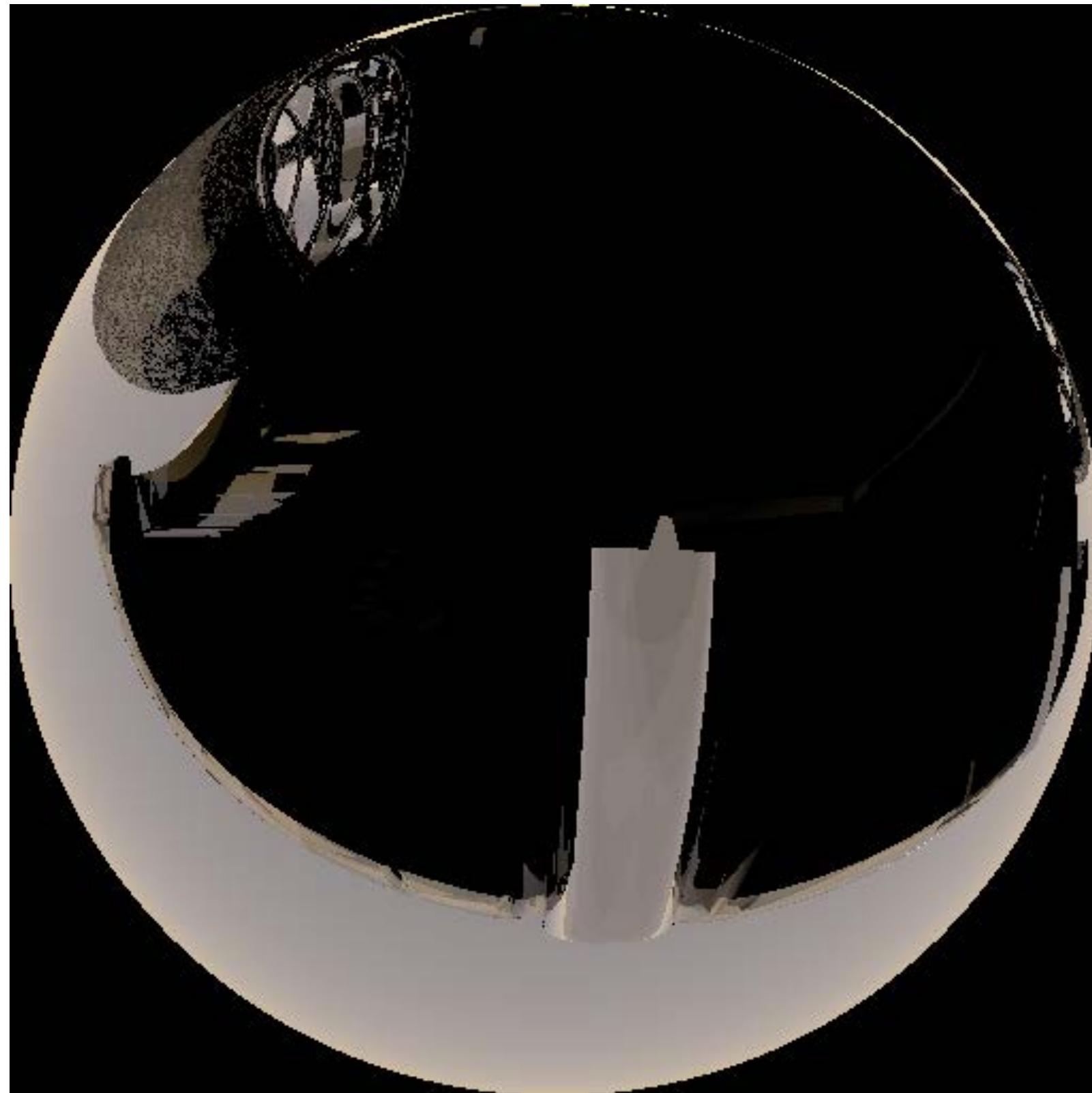
Incident radiance



Exitant radiance

Copper

More blurring, plus coloration (nonuniform absorption across frequencies)



Incident radiance



Exitant radiance

Scattering off a surface: the BRDF

- “Bidirectional reflectance distribution function”
- Encodes behavior of light that “bounces off” surface
- Given incoming direction ω_i , how much light gets scattered in any given outgoing direction ω_o ?
- Describe as distribution $f_r(\omega_i \rightarrow \omega_o)$

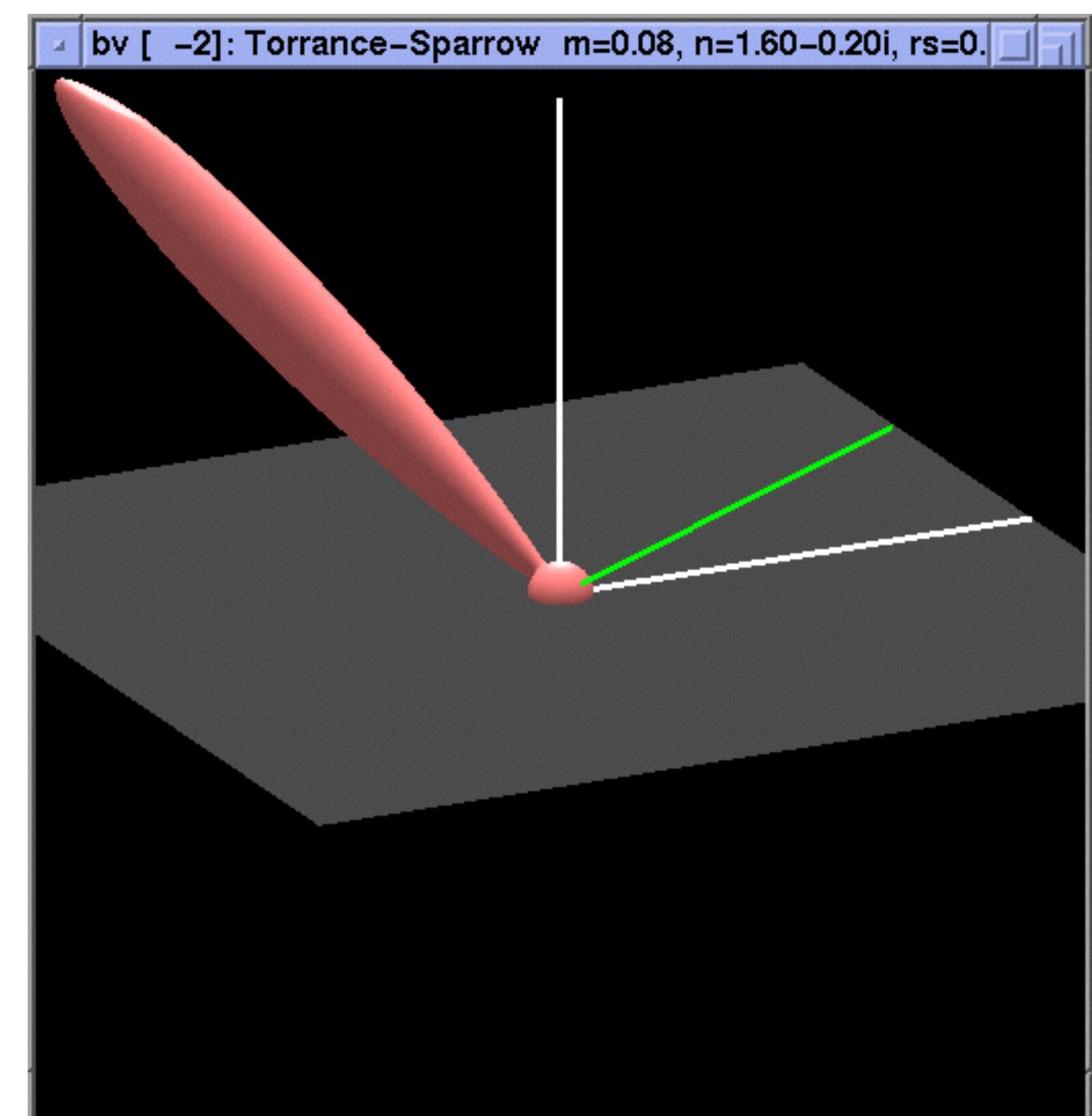
$$f_r(\omega_i \rightarrow \omega_o) \geq 0$$

$$\int_{\mathcal{H}^2} f_r(\omega_i \rightarrow \omega_o) \cos \theta d\omega_i \leq 1$$

why less than or equal?
where did the rest of the energy go?!

$$f_r(\omega_i \rightarrow \omega_o) = f_r(\omega_o \rightarrow \omega_i)$$

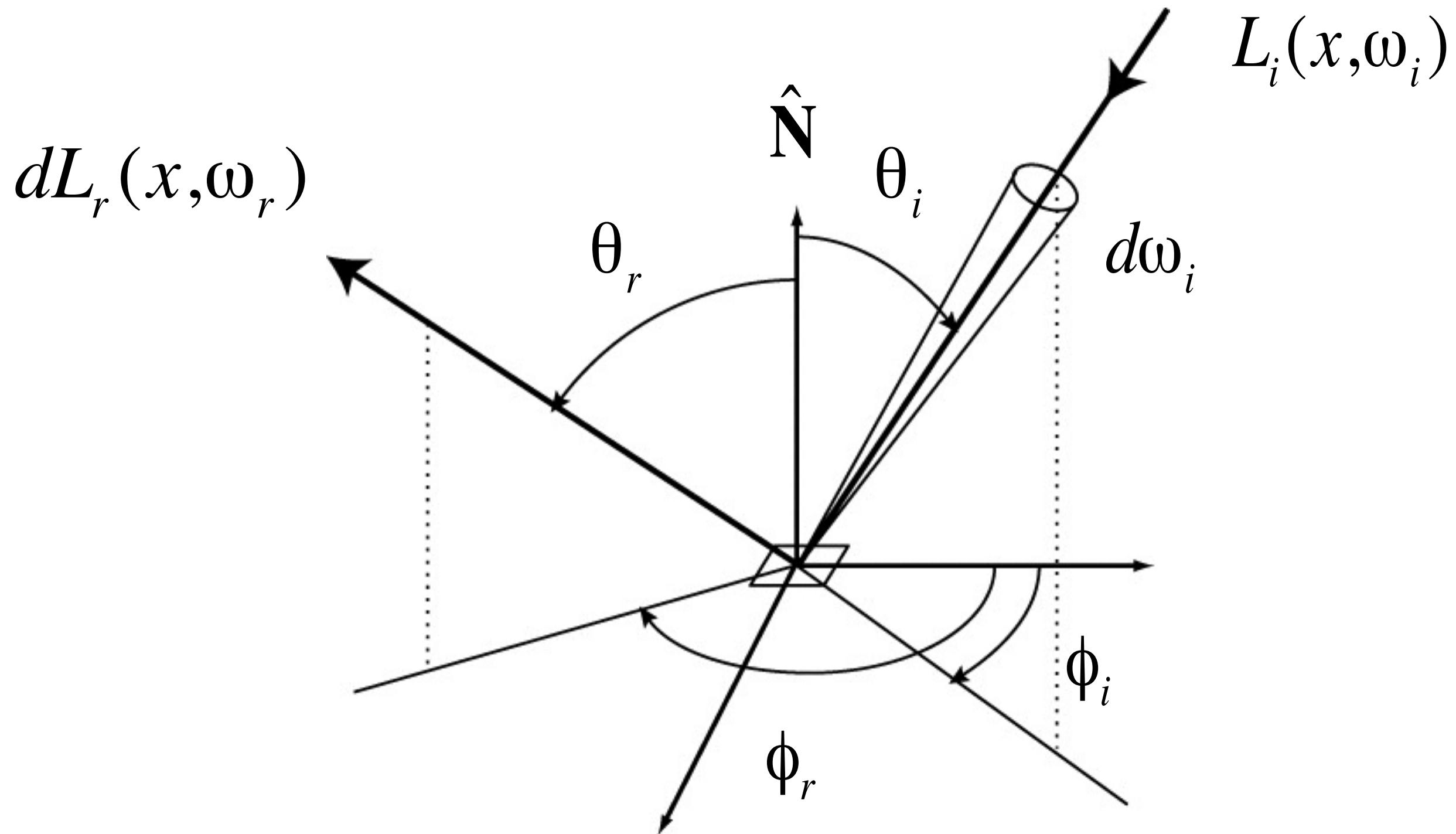
“Helmholtz reciprocity”



bv (Szymon Rusinkiewicz)

Q: Why should Helmholtz reciprocity hold? Think about little mirrors...

Radiometric description of BRDF

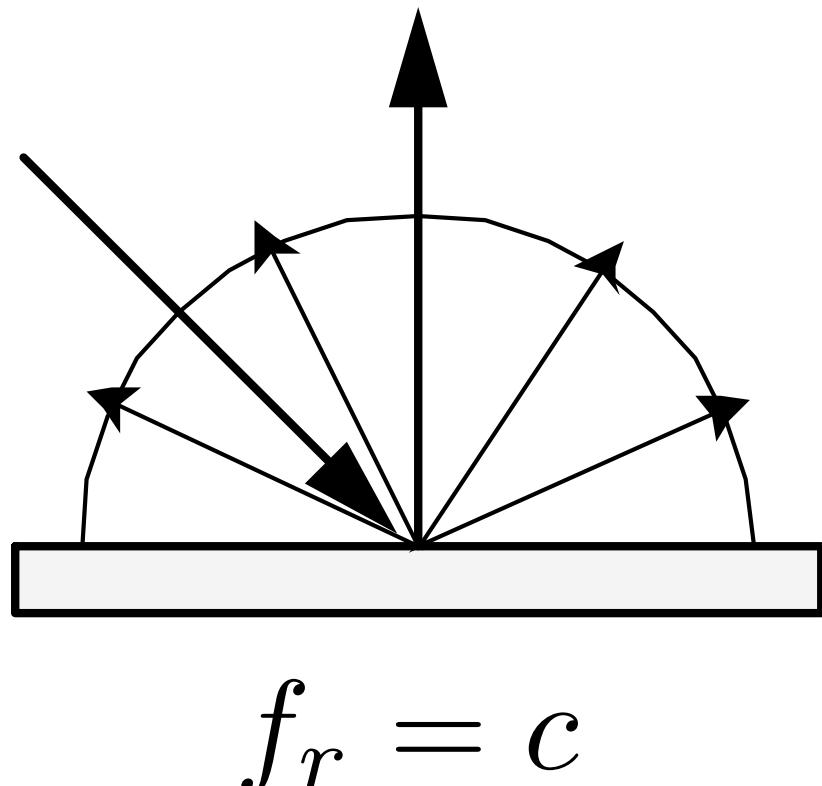


$$f_r(\omega_i \rightarrow \omega_o) = \frac{dL_o(\omega_o)}{dE_i(\omega_i)} = \frac{dL_o(\omega_o)}{dL_i(\omega_i) \cos \theta_i} \left[\frac{1}{sr} \right]$$

“For a given change in the incident irradiance, how much does the exitant radiance change?”

Example: Lambertian reflection

Assume light is equally likely to be reflected in each output direction

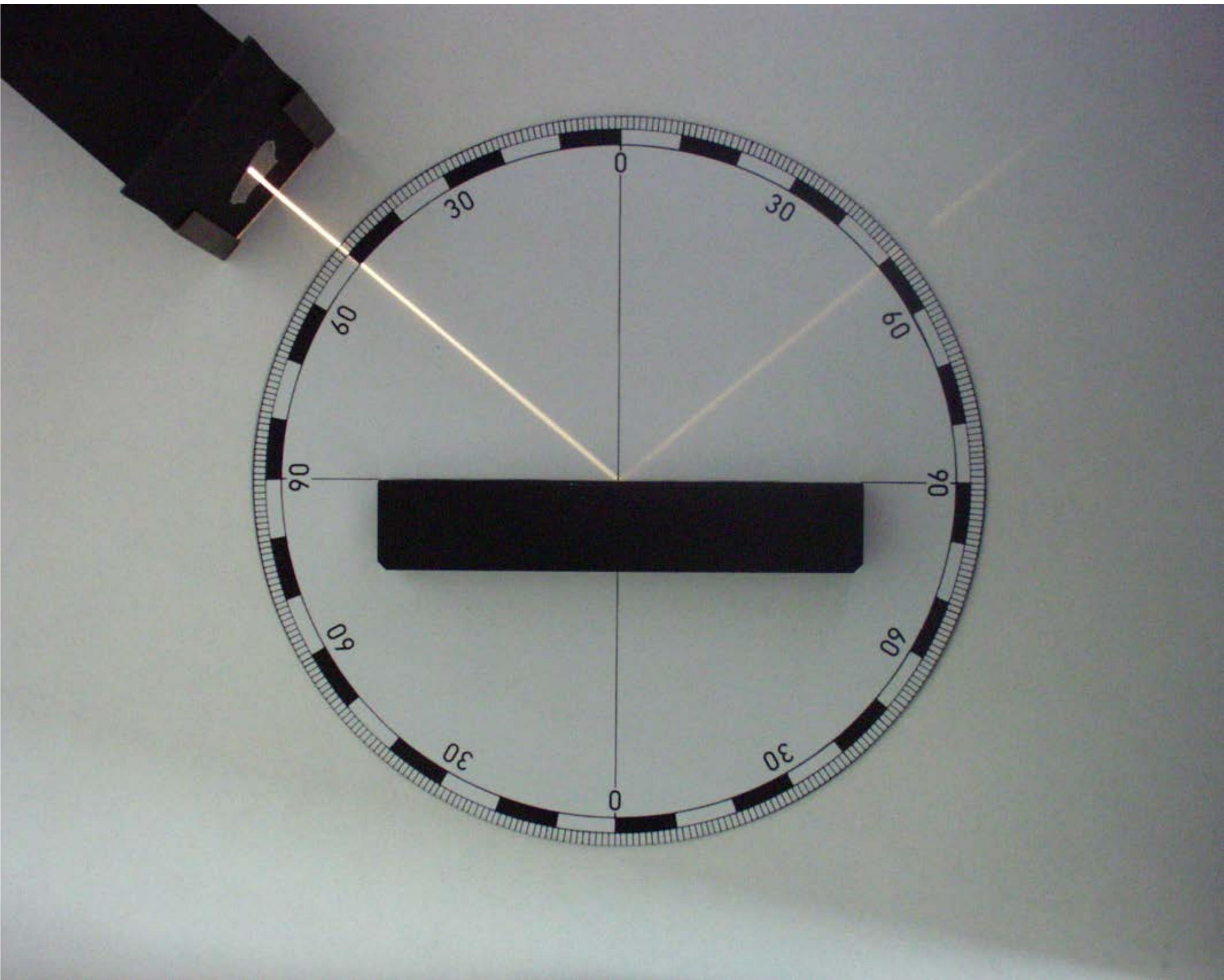


$$\begin{aligned} L_o(\omega_o) &= \int_{H^2} f_r L_i(\omega_i) \cos \theta_i d\omega_i \\ &= f_r \int_{H^2} L_i(\omega_i) \cos \theta_i d\omega_i \\ &= f_r E \end{aligned}$$

$$f_r = \frac{\rho}{\pi}$$



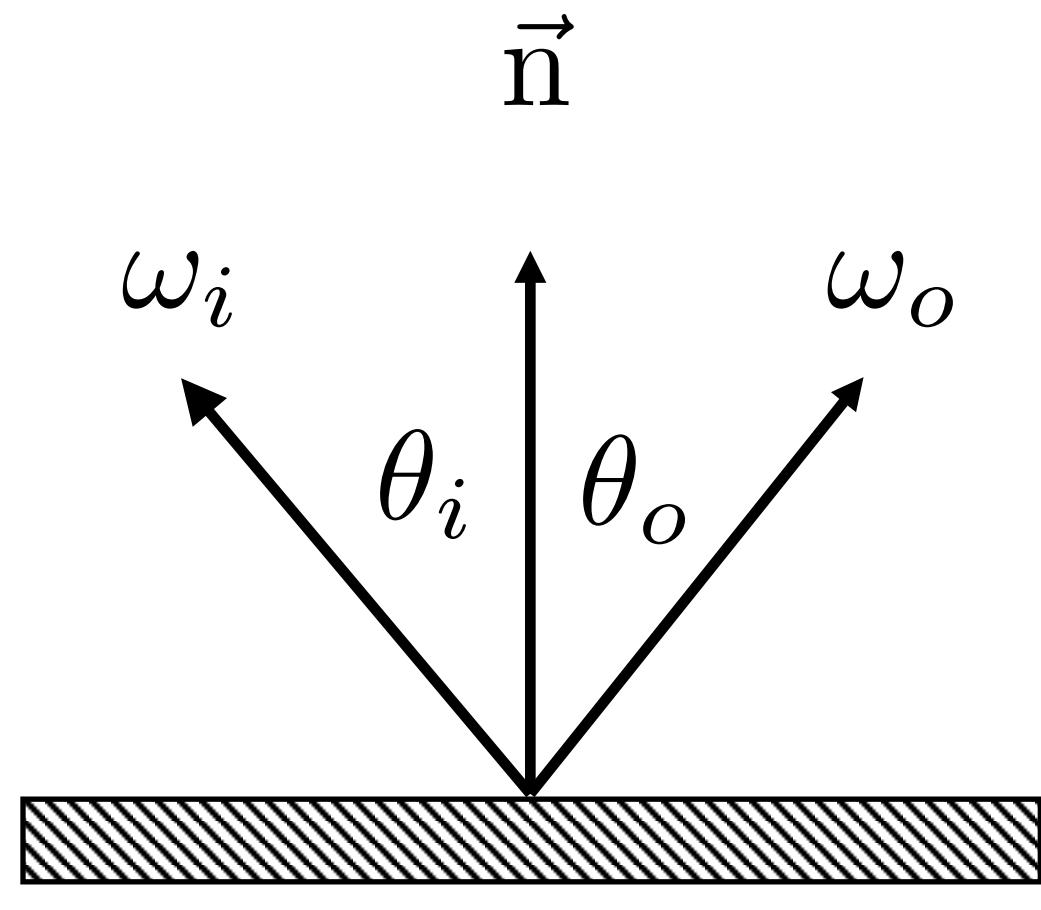
Example: perfect specular reflection



[Zátónyi Sándor]

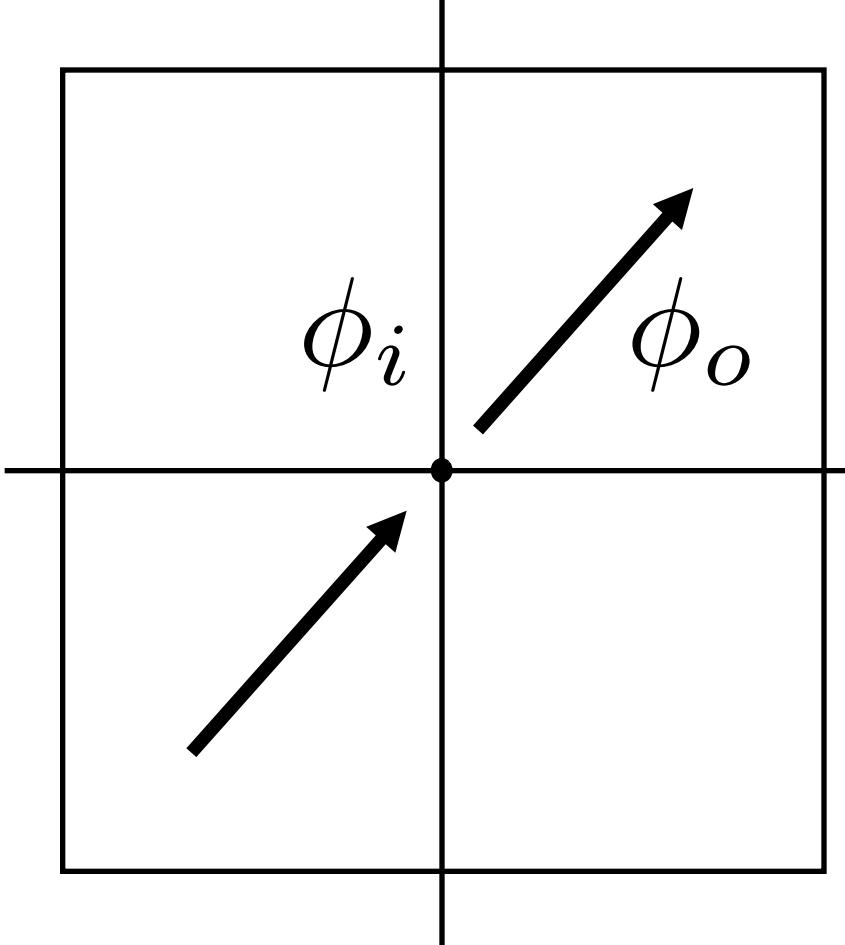
CMU 15-462/662, Fall 2015

Geometry of specular reflection



$$\theta = \theta_o = \theta_i$$

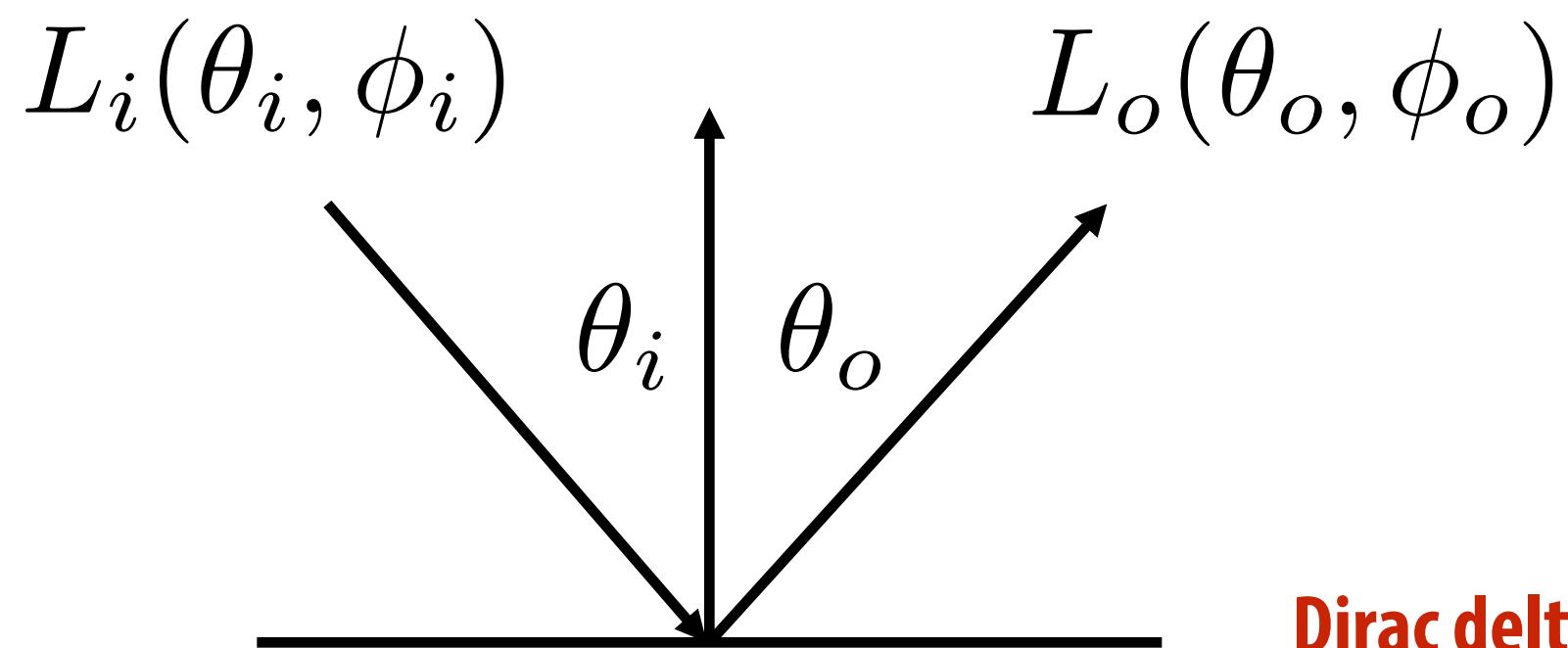
**Top-down view
(looking down on surface)**



$$\omega_o + \omega_i = 2 \cos \theta \vec{n} = 2(\omega_i \cdot \vec{n})\vec{n}$$

$$\omega_o = -\omega_i + 2(\omega_i \cdot \vec{n})\vec{n}$$

Specular reflection BRDF



$$L_o(\theta_o, \phi_o) = L_i(\theta_o, \phi_o \pm \pi)$$

Dirac delta

$$f_r(\theta_i, \phi_i; \theta_o, \phi_o) = \frac{\delta(\cos \theta_i - \cos \theta_o)}{\cos \theta_i} \delta(\phi_i - \phi_o \pm \pi)$$

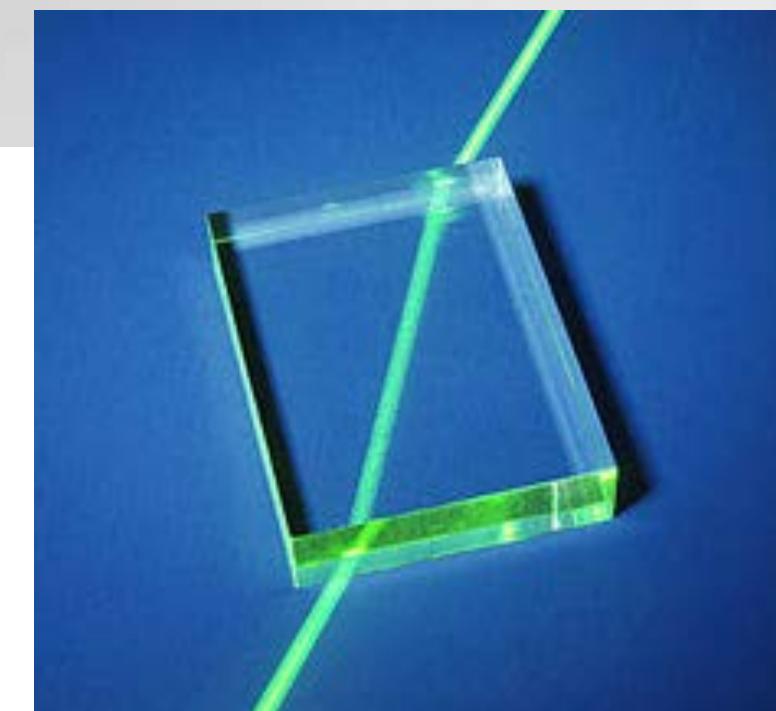
- Strictly speaking, f_r is a *distribution*, not a function
- In practice, no hope of finding reflected direction via random sampling; simply pick the reflected direction!



Transmission

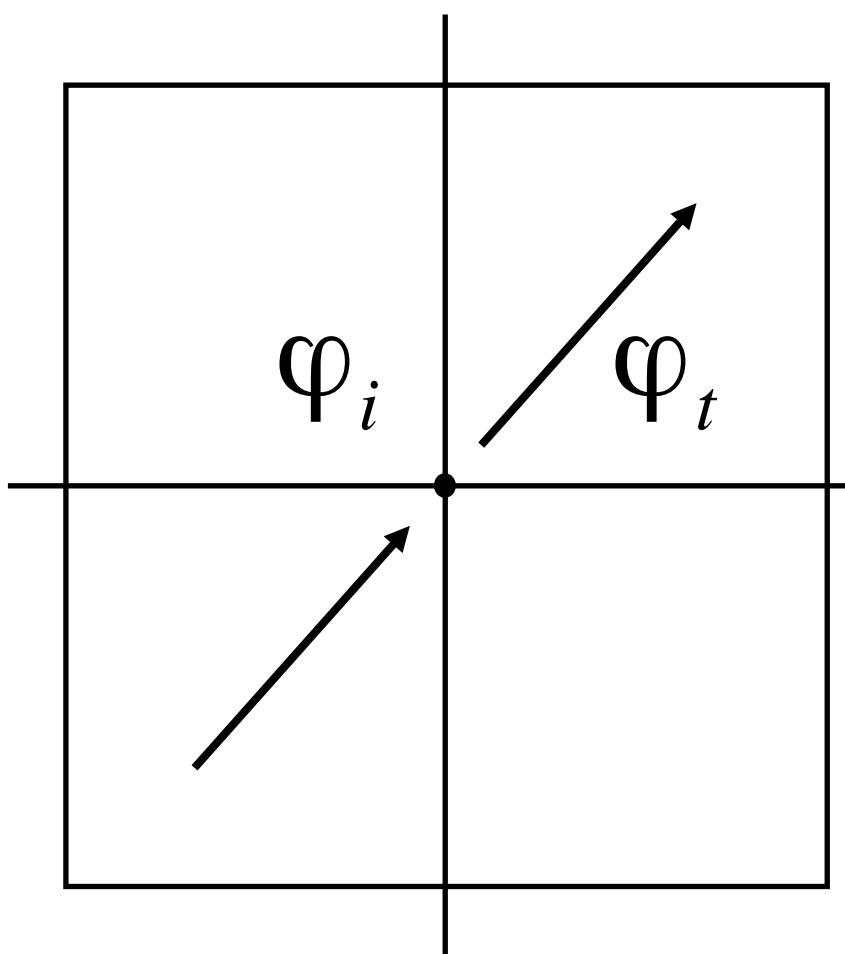
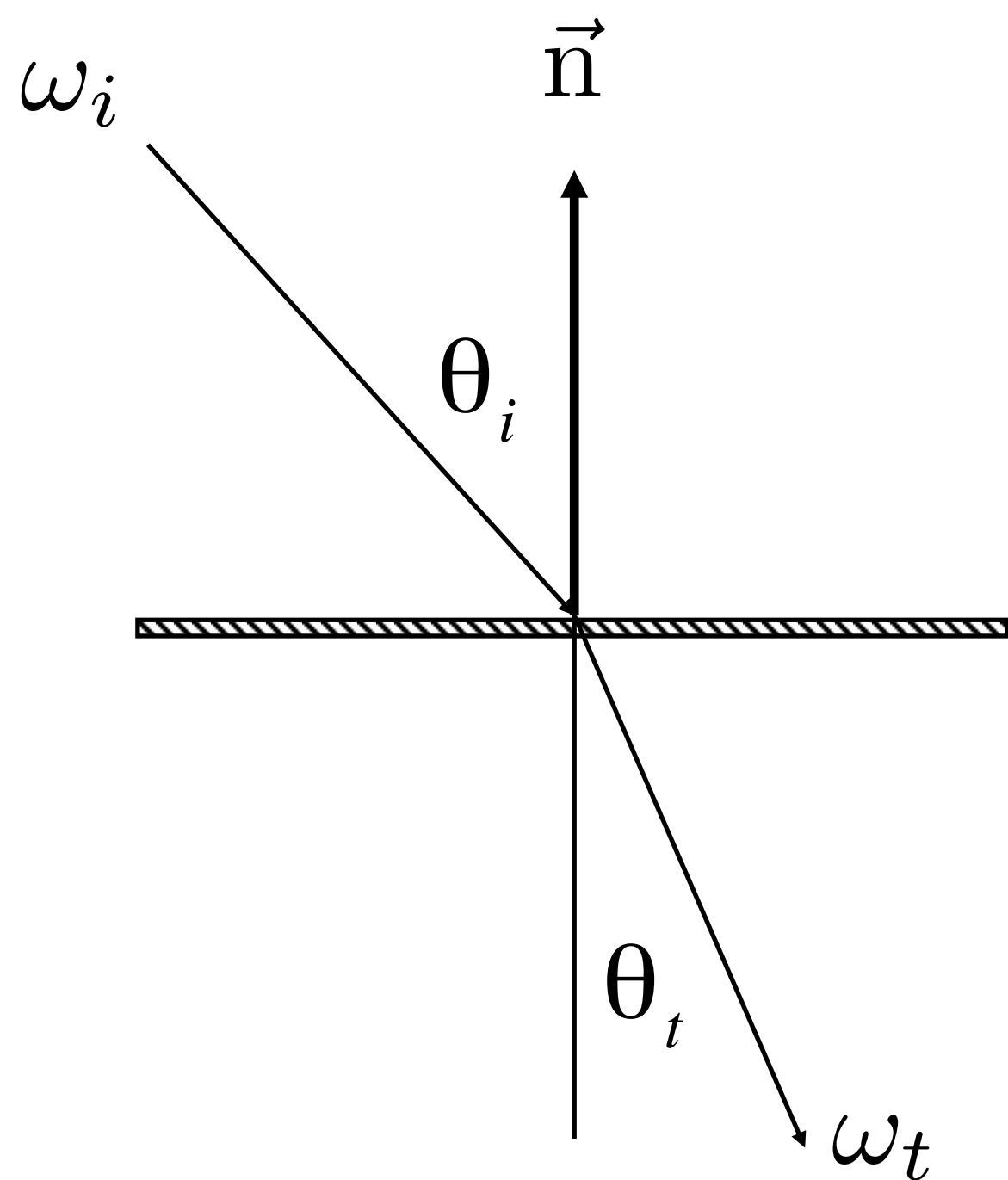
In addition to reflecting off surface, light may be transmitted through surface.

Light refracts when it enters a new medium.



Snell's Law

Transmitted angle depends on index of refraction of medium incident ray is in and index of refraction of medium light is entering.



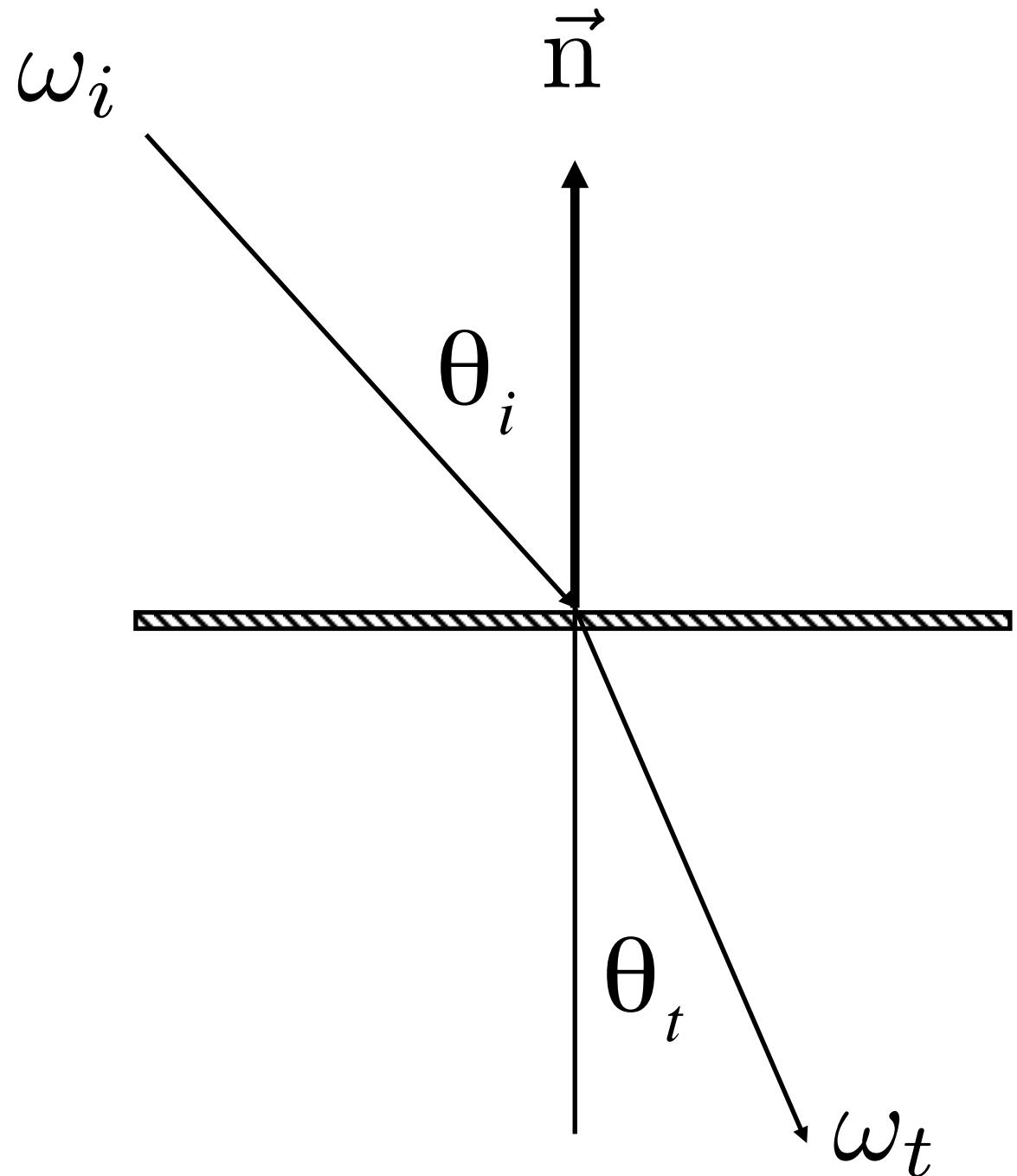
Medium	η^*
Vacuum	1.0
Air (sea level)	1.00029
Water (20°C)	1.333
Glass	1.5-1.6
Diamond	2.42

* index of refraction is wavelength dependent (these are averages)

$$\eta_i \sin \theta_i = \eta_t \sin \theta_t$$

Law of refraction

$$\eta_i \sin \theta_i = \eta_t \sin \theta_t$$



$$\cos \theta_t = \sqrt{1 - \sin^2 \theta_t}$$

$$= \sqrt{1 - \left(\frac{\eta_i}{\eta_t}\right)^2 \sin^2 \theta_i}$$

$$= \sqrt{1 - \left(\frac{\eta_i}{\eta_t}\right)^2 (1 - \cos^2 \theta_i)}$$

$$1 - \left(\frac{\eta_i}{\eta_t}\right)^2 (1 - \cos^2 \theta_i) < 0$$

Total internal reflection:

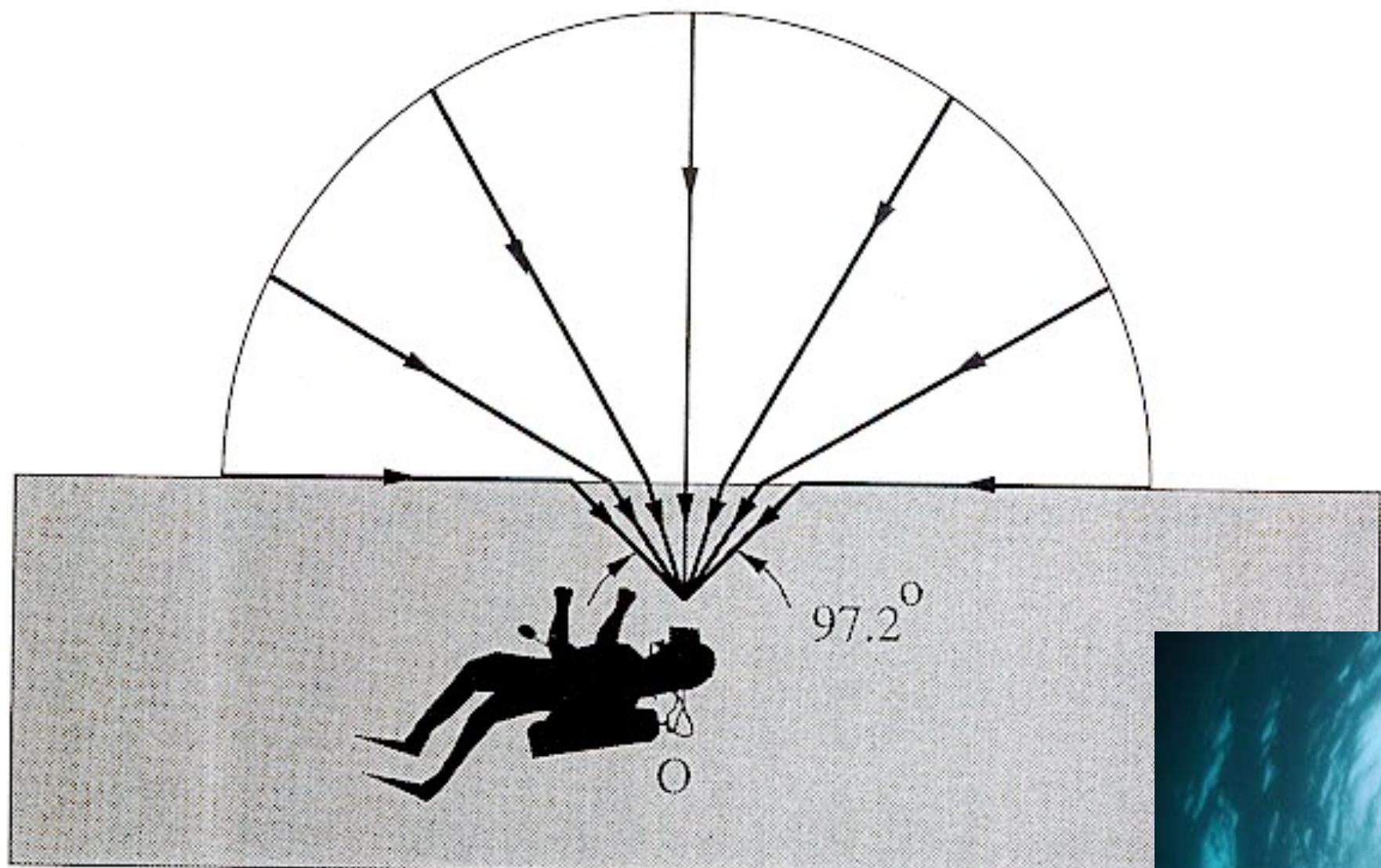
When light is moving from a more optically dense medium to a less optically dense medium:

$$\frac{\eta_i}{\eta_t} > 1$$

Light incident on boundary from large enough angle will not exit medium.

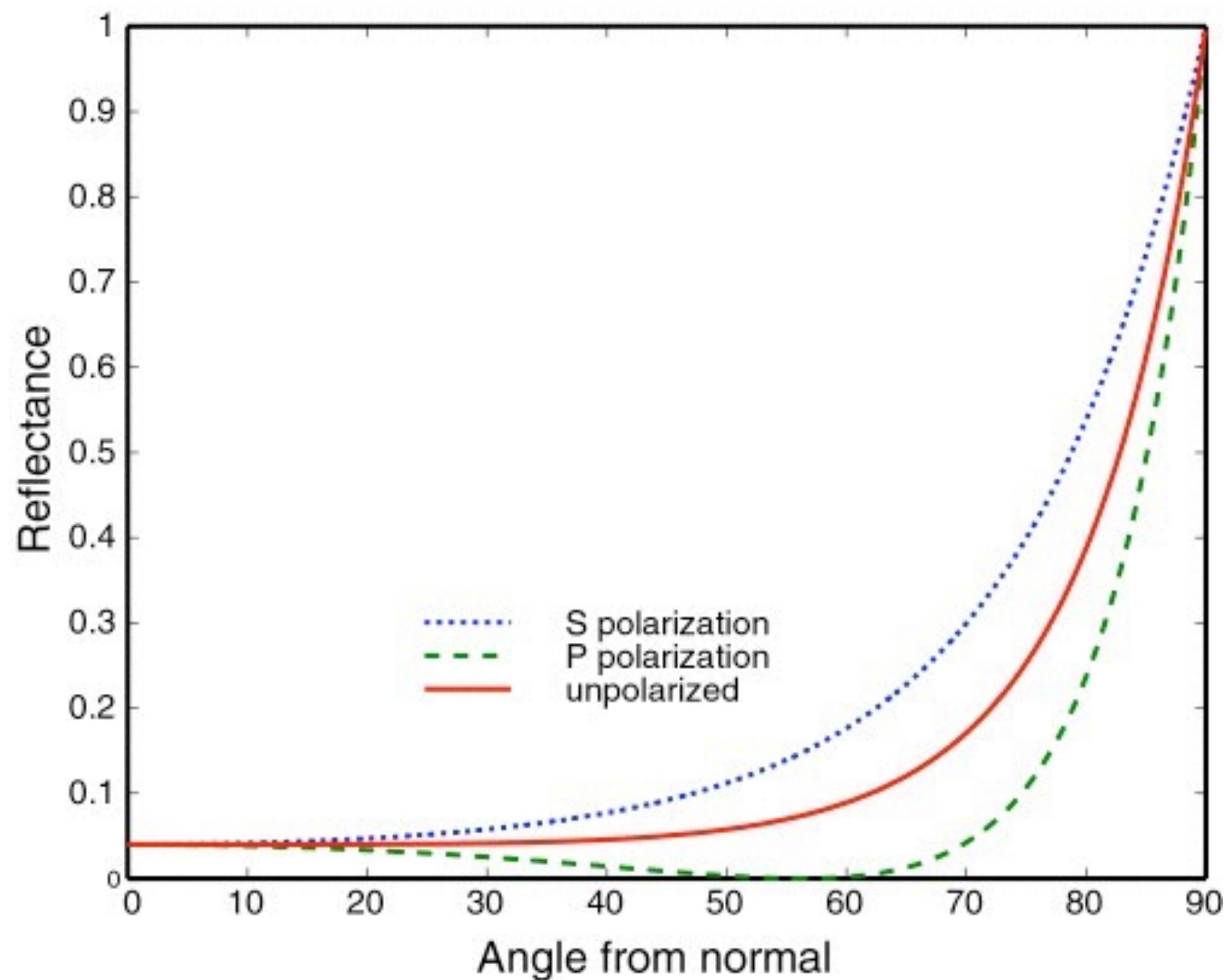
Optical manhole

Only small “cone” visible, due to total internal reflection (TIR)



Fresnel reflection

Many real materials:
reflectance increases w/
viewing angle



[Lafontaine et al. 1997]

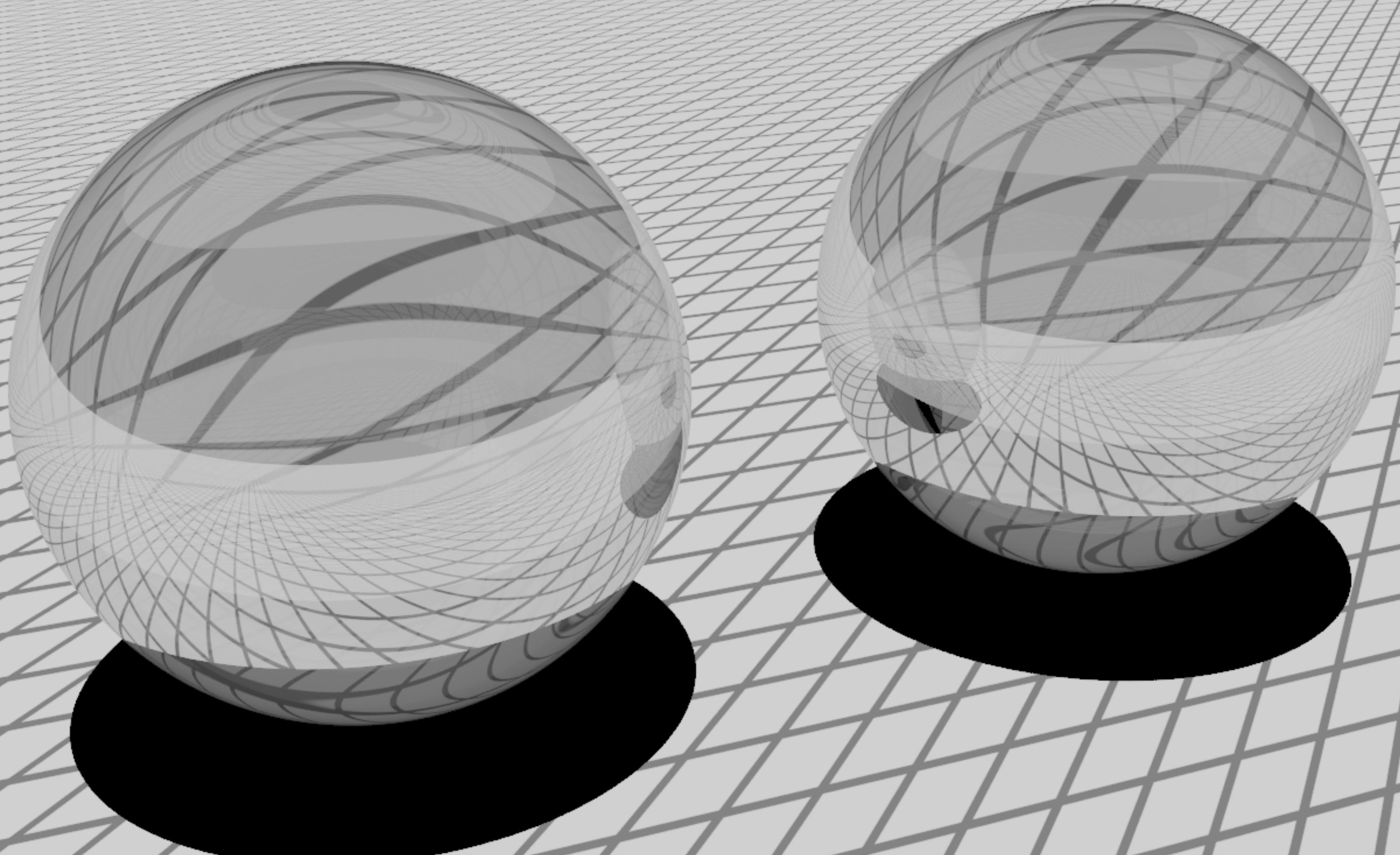
Snell + Fresnel: Example



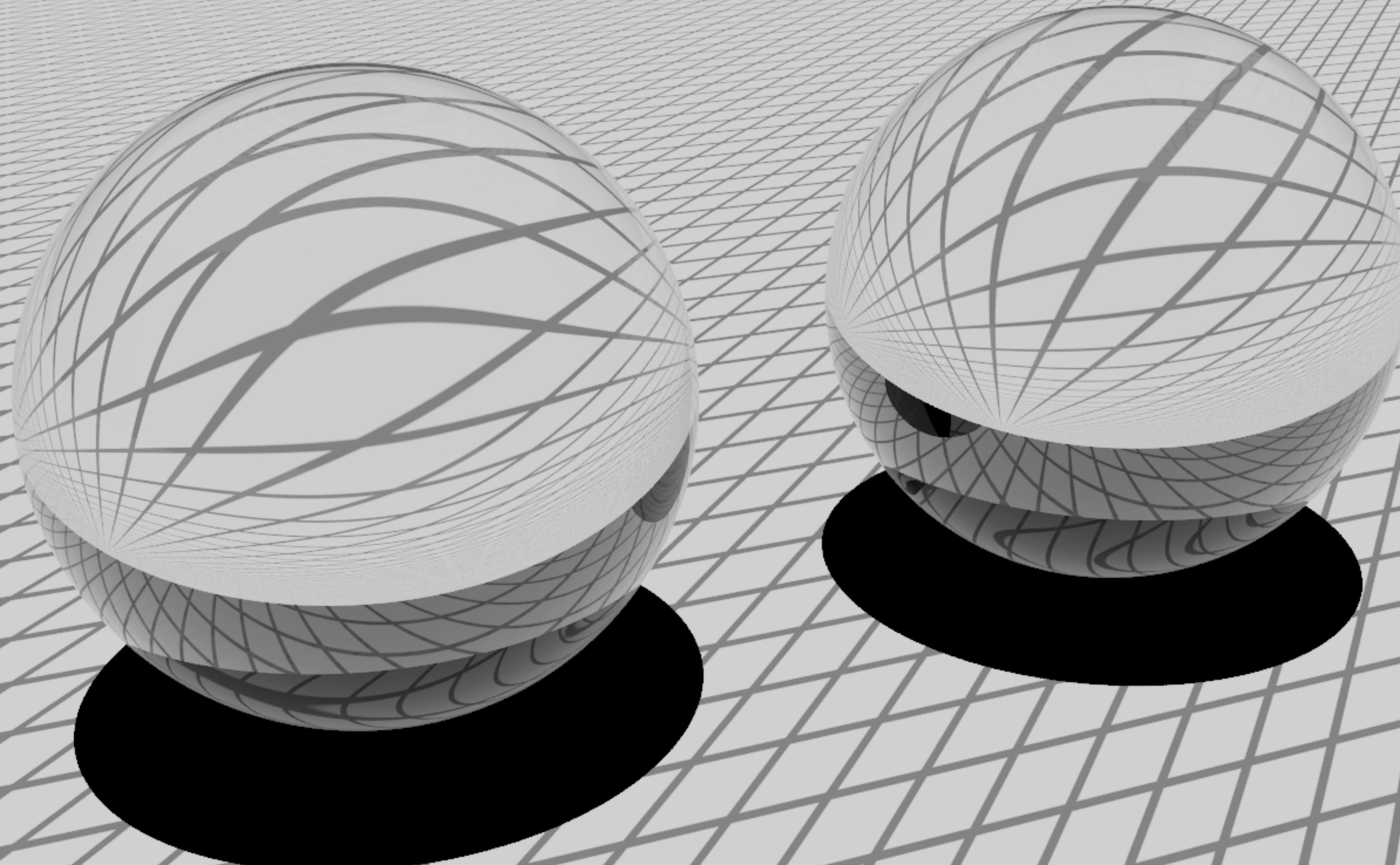
Refraction (Snell)

Reflection (Fresnel)

Without Fresnel (fixed reflectance/transmission)



Glass with Fresnel reflection/transmission

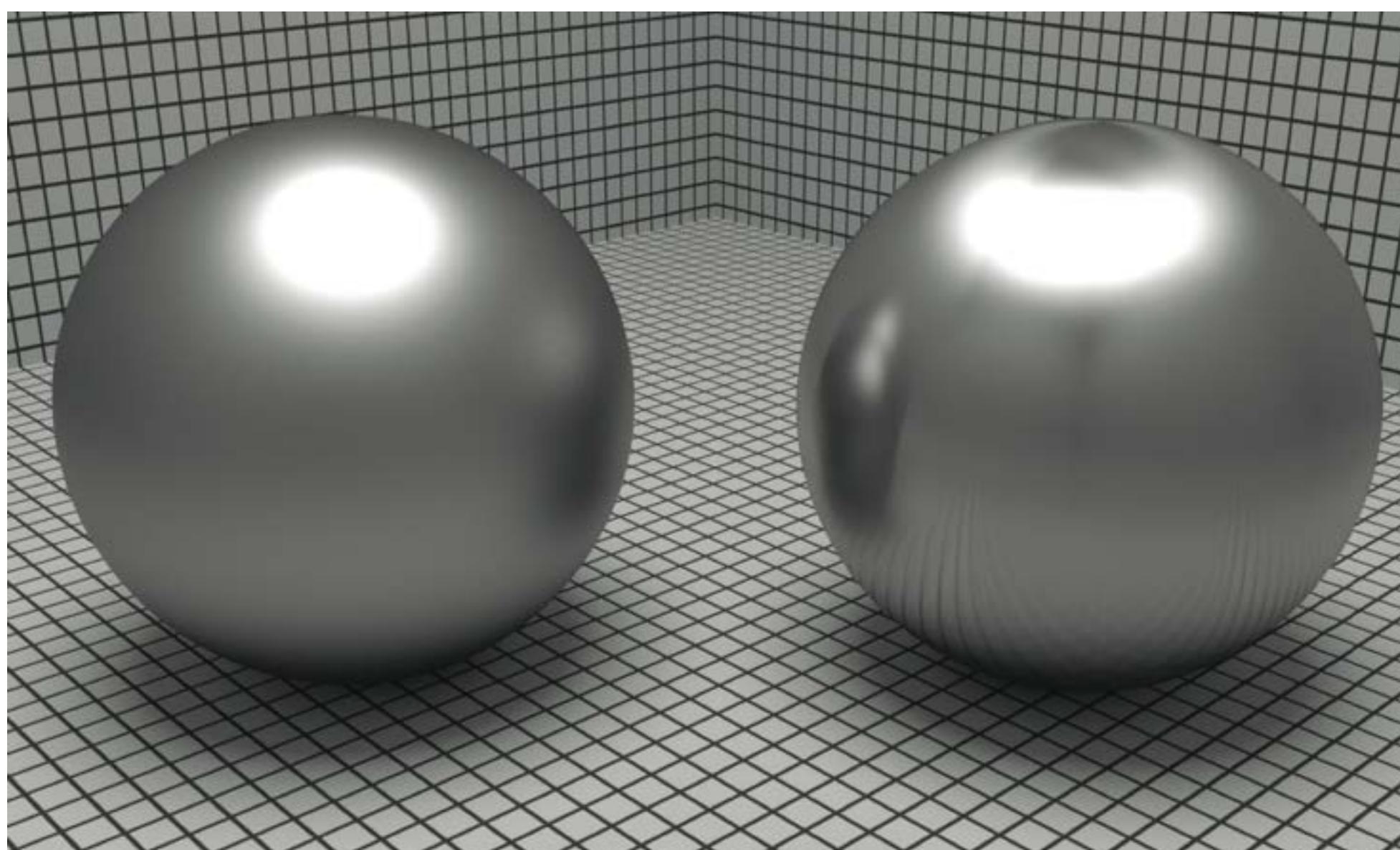


Anisotropic reflection

Reflection depends on azimuthal angle ϕ

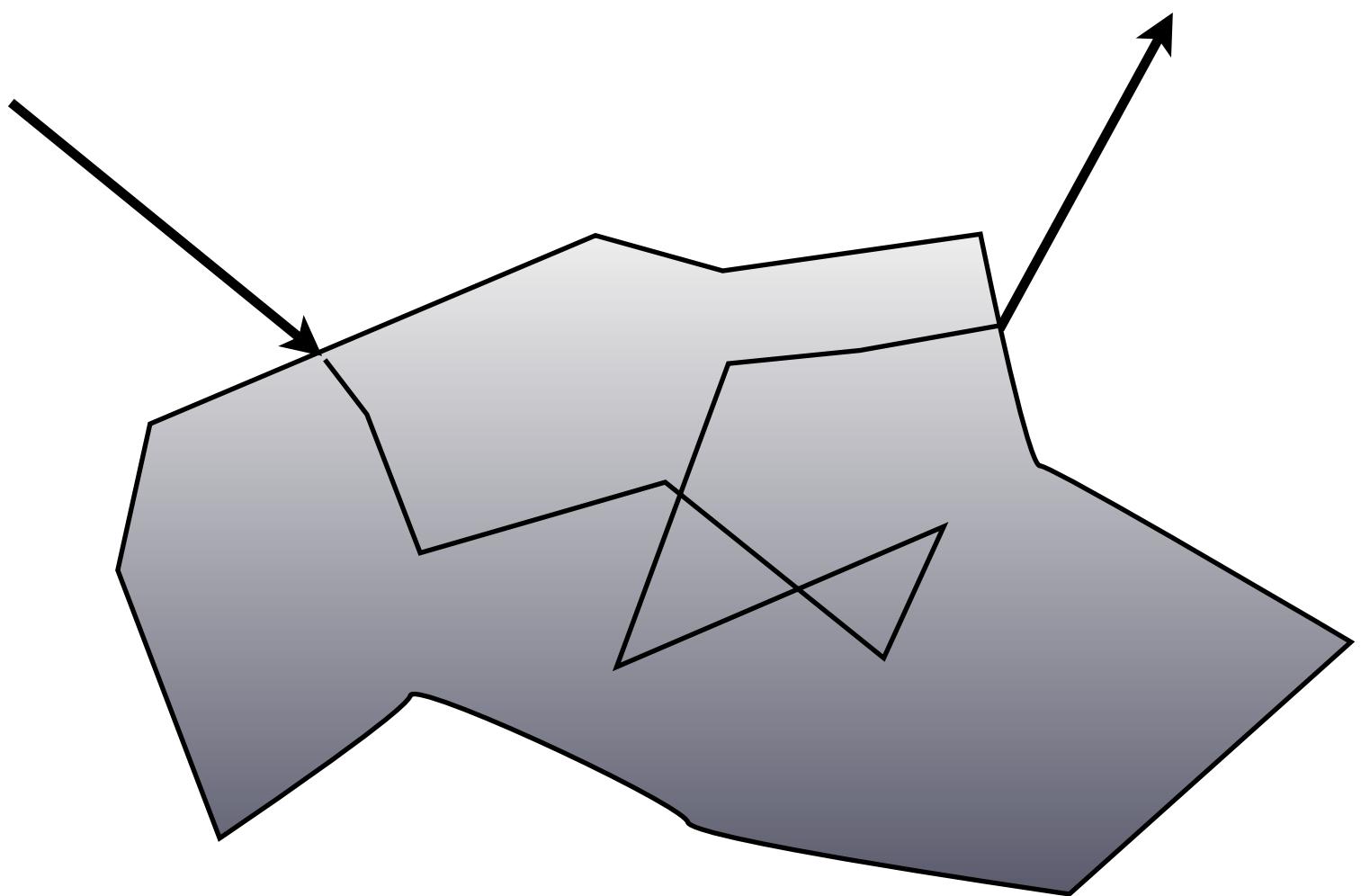


Results from oriented microstructure of surface
e.g., brushed metal



Subsurface scattering

- Visual characteristics of many surfaces caused by light entering at different points than it exits
 - Violates a fundamental assumption of the BRDF



[Jensen et al 2001]



[Donner et al 2008]

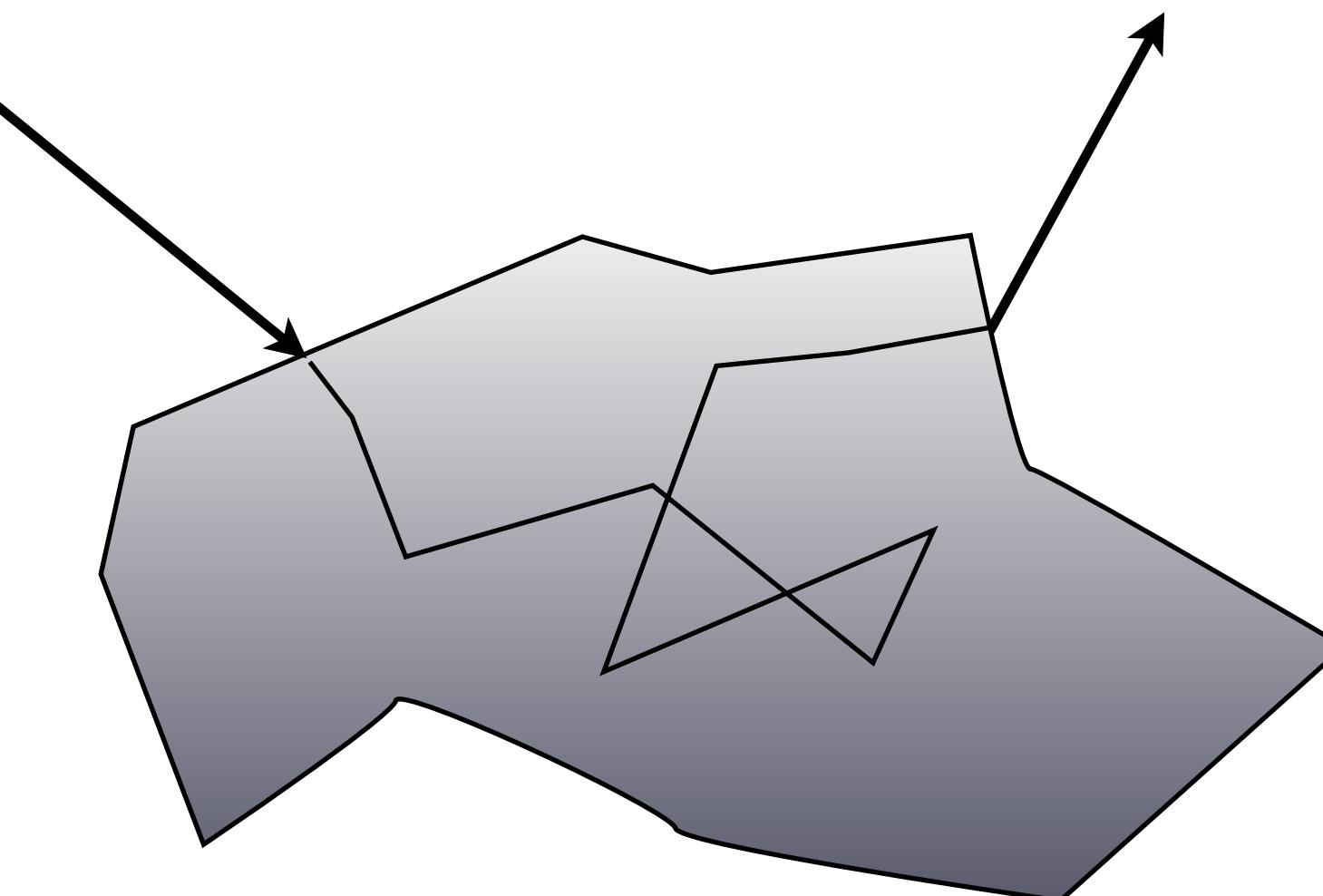
Scattering functions

- Generalization of BRDF; describes exitant radiance at one point due to incident differential irradiance at another point:

$$S(x_i, \omega_i, x_o, \omega_o)$$

- Generalization of reflection equation integrates over all points on the surface and all directions(!)

$$L(x_o, \omega_o) = \int_A \int_{H^2} S(x_i, \omega_i, x_o, \omega_o) L_i(x_i, \omega_i) \cos \theta_i d\omega_i dA$$



Translucent materials: Jade



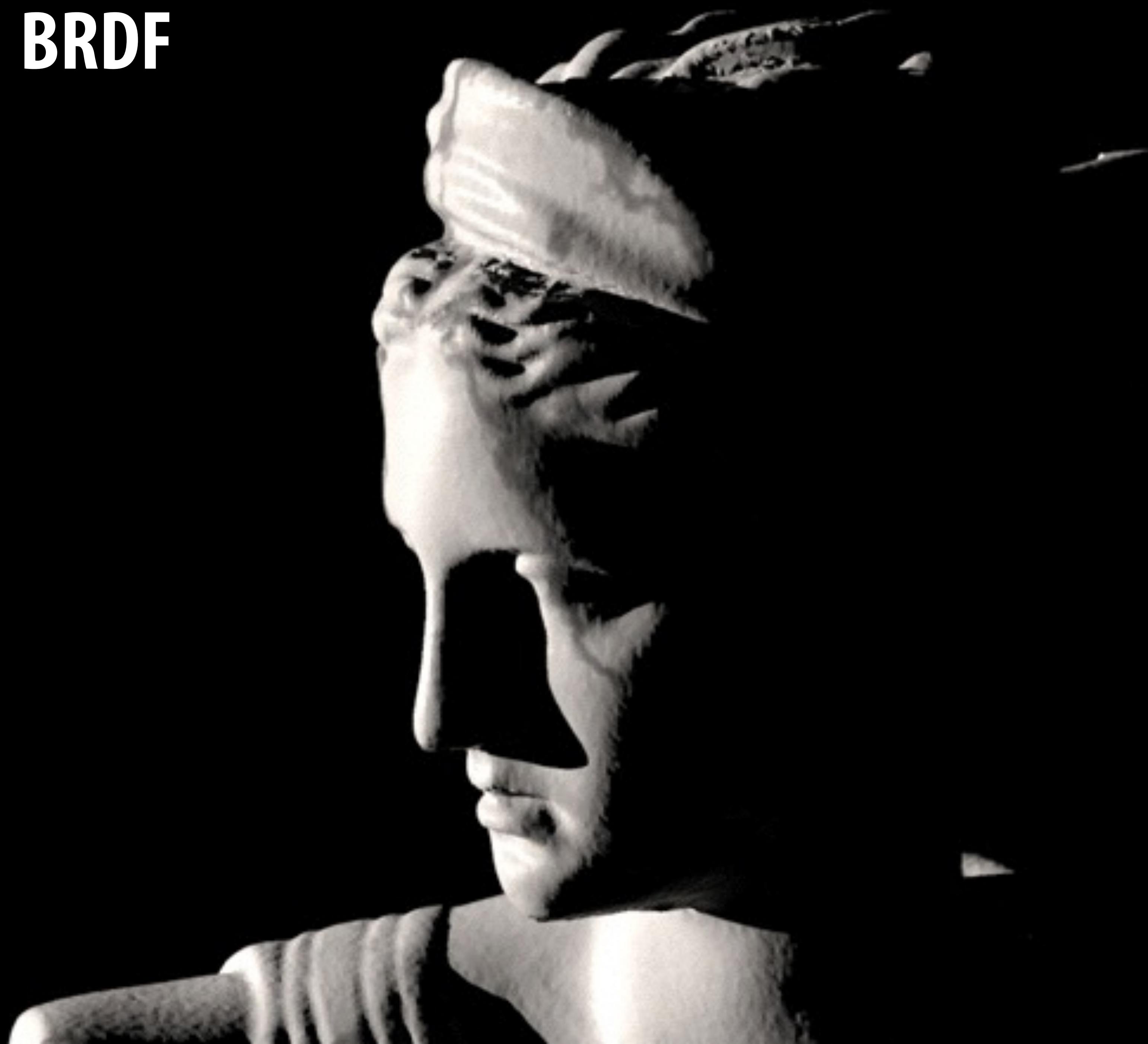
Translucent materials: skin



Translucent materials: leaves



BRDF



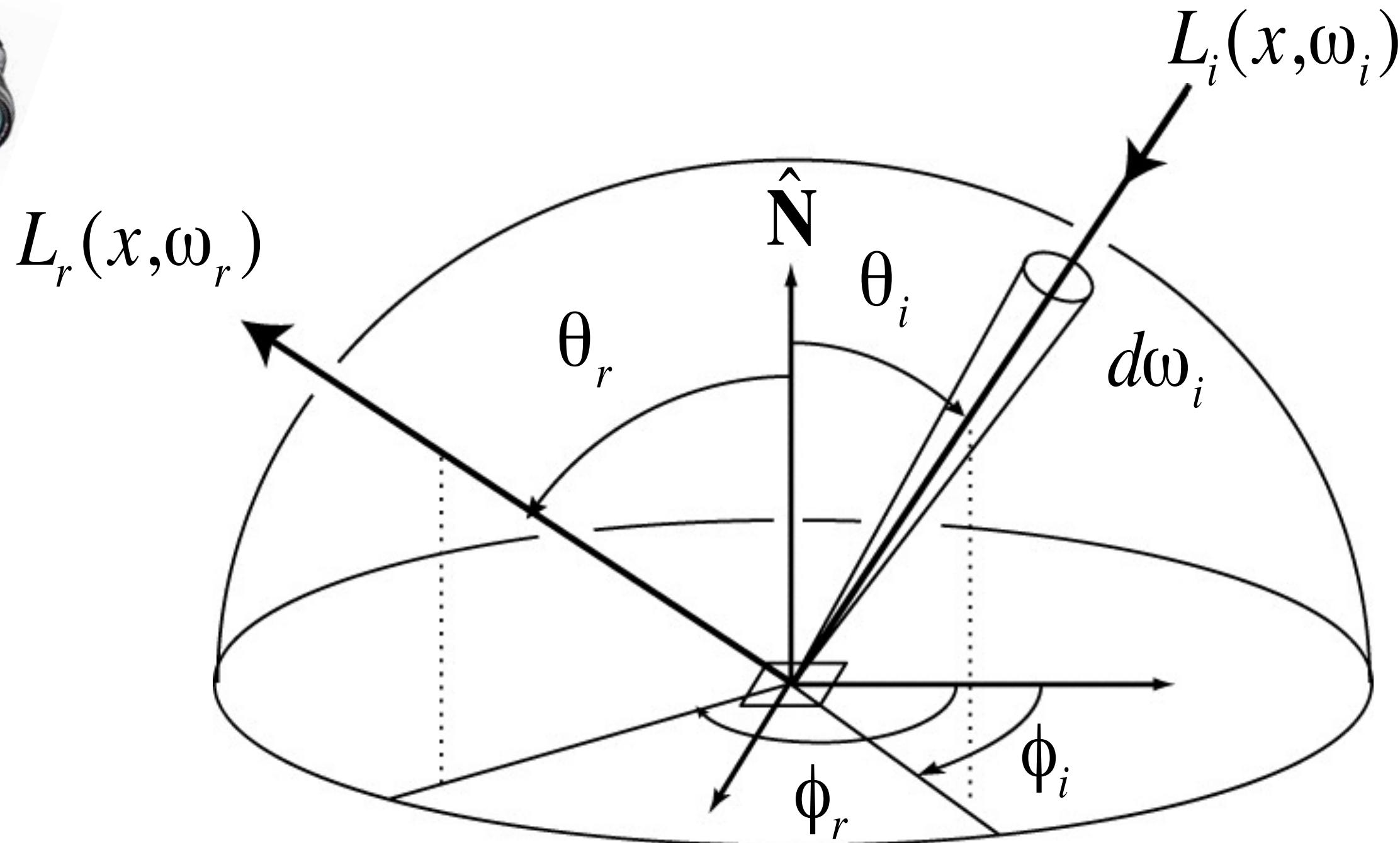
BSSRDF



Ok, so scattering is *complicated!*

**What's a (relatively simple) algorithm
that can capture all this behavior?**

The reflection equation



$$dL_r(\omega_r) = f_r(\omega_i \rightarrow \omega_r) dL_i(\omega_i) \cos \theta_i$$

$$L_r(p, \omega_r) = \int_{H^2} f_r(p, \omega_i \rightarrow \omega_r) L_i(p, \omega_i) \cos \theta_i d\omega_i$$

The reflection equation

- Key piece of overall rendering equation:

$$L_r(p, \omega_r) = \int_{H^2} f_r(p, \omega_i \rightarrow \omega_r) L_i(p, \omega_i) \cos \theta_i d\omega_i$$

- Approximate integral via Monte Carlo integration
- Generate directions ω_j sampled from some distribution $p(\omega)$
- Compute the estimator

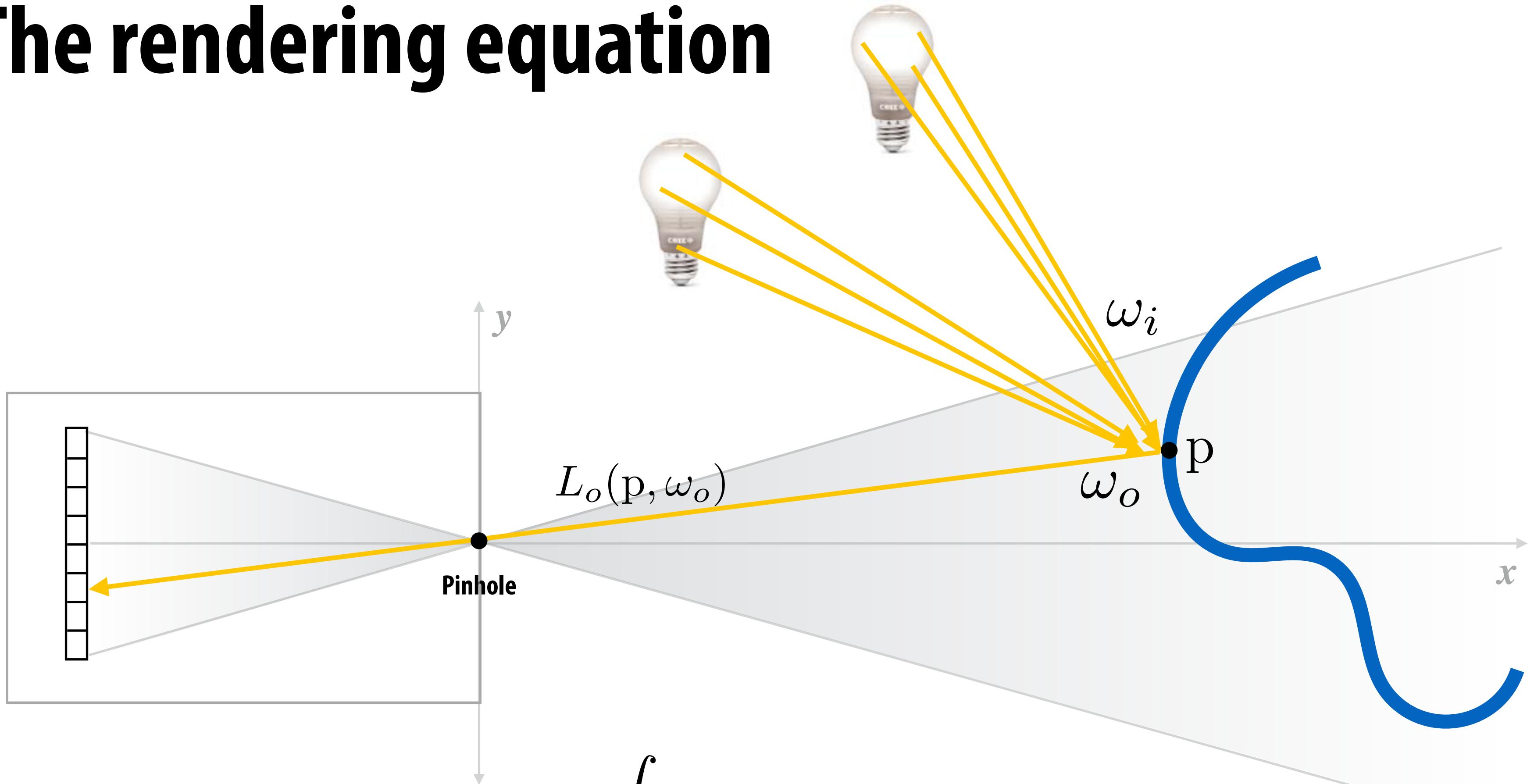
$$\frac{1}{N} \sum_{j=1}^N \frac{f_r(p, \omega_j \rightarrow \omega_r) L_i(p, \omega_j) \cos \theta_j}{p(\omega_j)}$$

- To reduce variance $p(\omega)$ should match BRDF or incident radiance function

Estimating reflected light

```
// Assume:  
// Ray ray hits surface at point hit_p  
// Normal of surface at hit point is hit_n  
  
Vector3D wr = -ray.d;    // outgoing direction  
Spectrum Lr = 0.;  
for (int i = 0; i < N; ++i) {  
    Vector3D wi;          // sample incident light from this direction  
    float pdf;             // p(wi)  
  
    generate_sample(brdf, &wi, &pdf);    // generate sample according to brdf  
  
    Spectrum f = brdf->f(wr, wi);  
    Spectrum Li = trace_ray(Ray(hit_p, wi)); // compute incoming Li  
    Lr += f * Li * fabs(dot(wi, hit_n)) / pdf;  
}  
return Lr / N;
```

The rendering equation



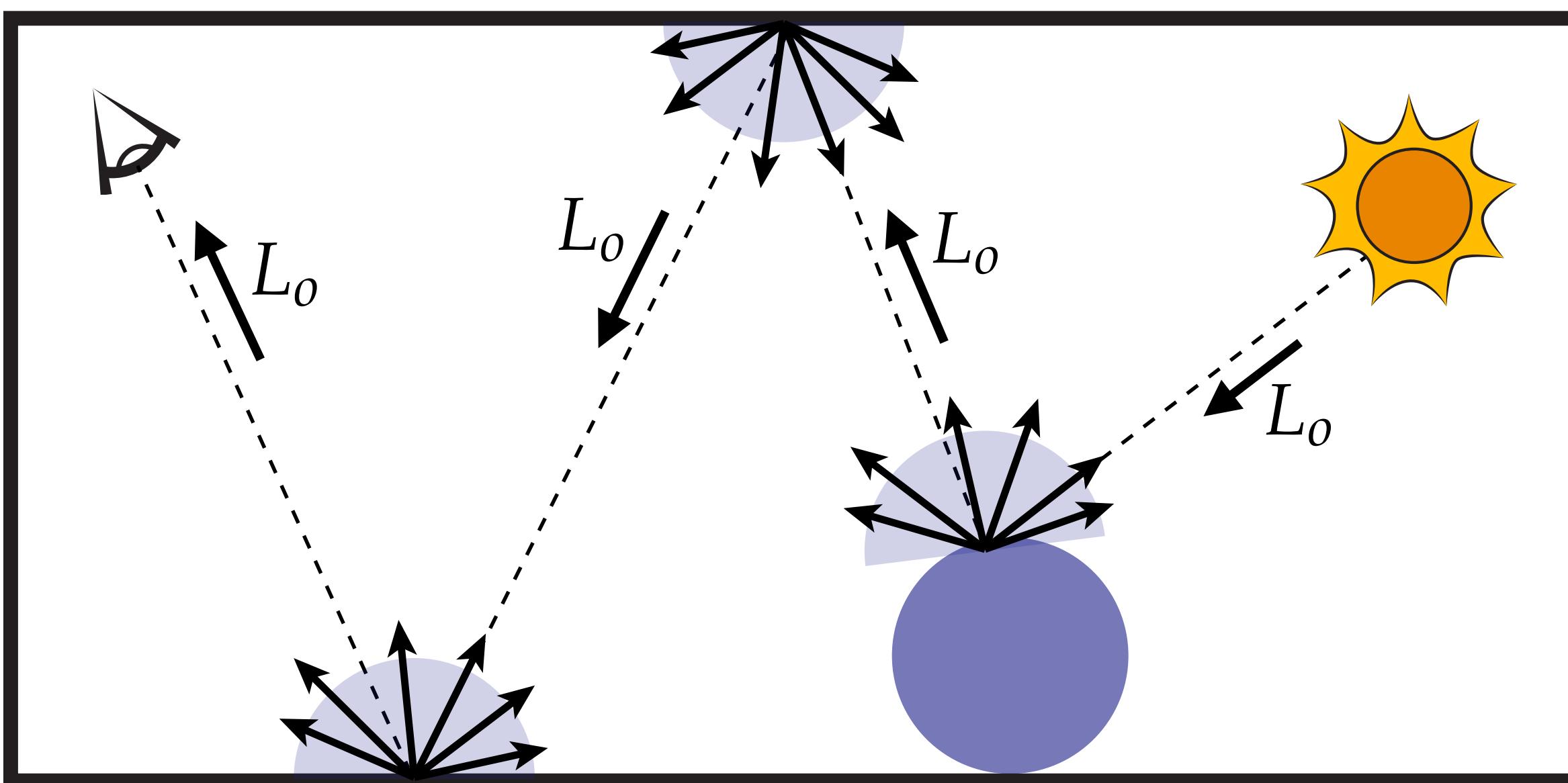
$$L_o(p, \omega_o) = L_e(p, \omega_o) + \int_{H^2} f_r(p, \omega_i \rightarrow \omega_o) L_i(p, \omega_i) \cos \theta_i d\omega_i$$

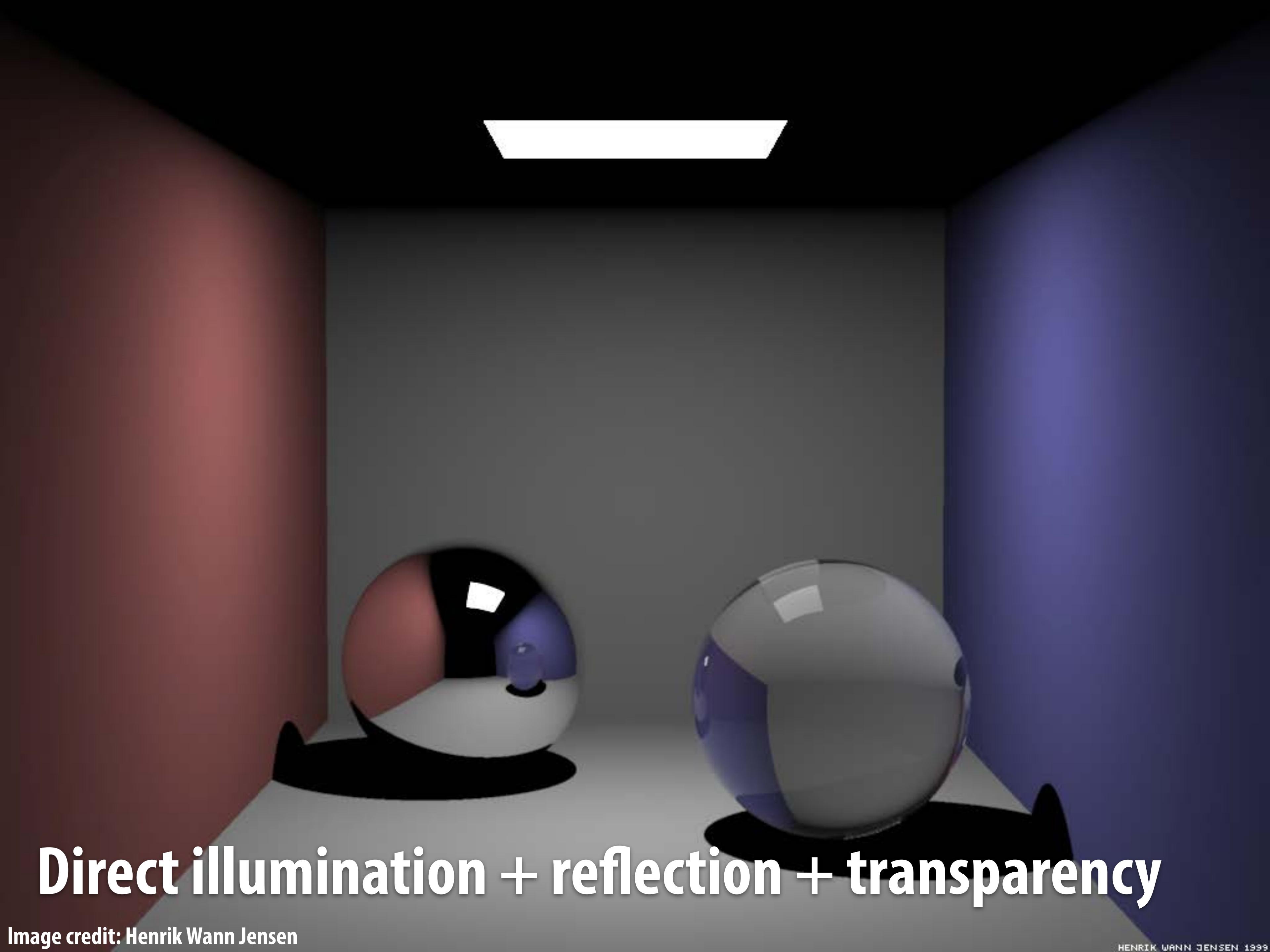
Now that we know how to handle reflection, how do we solve the full rendering equation? Have to determine incident radiance...

Key idea in (efficient) rendering: take advantage of special knowledge to break up integration into “easier” components.

Path tracing: overview

- Partition the rendering equation into direct and indirect illumination
- Use Monte Carlo to estimate each partition separately
 - One sample for each
 - Assumption: 100s of samples per pixel
- Terminate paths with *Russian roulette*





Direct illumination + reflection + transparency



Global illumination solution

Image credit: Henrik Wann Jensen

HENRIK WANN JENSEN 2000

Review: Monte Carlo integration

■ Definite integral

What we seek to estimate

$$\int_a^b f(x) dx$$

■ Random variables

X_i is the value of a random sample drawn from the distribution $p(x)$

Y_i is also a random variable.

$$X_i \sim p(x)$$

$$Y_i = f(X_i)$$

■ Expectation of f

$$E[Y_i] = E[f(X_i)] = \int_a^b f(x) p(x) dx$$

■ Estimator

Monte Carlo estimate of $\int_a^b f(x) dx$

$$F_N = \frac{b-a}{N} \sum_{i=1}^N Y_i$$

Assuming samples X_i drawn from uniform pdf. I will provide estimator for arbitrary PDFs later in lecture.

Basic Monte Carlo estimator

$$\lim_{N \rightarrow \infty} (b - a) \frac{1}{N} \sum_{i=1}^N f(X_i) = \int_a^b f(x) dx$$



**Assume uniform
probability density (for now)**

$$X_i \sim U(a, b)$$

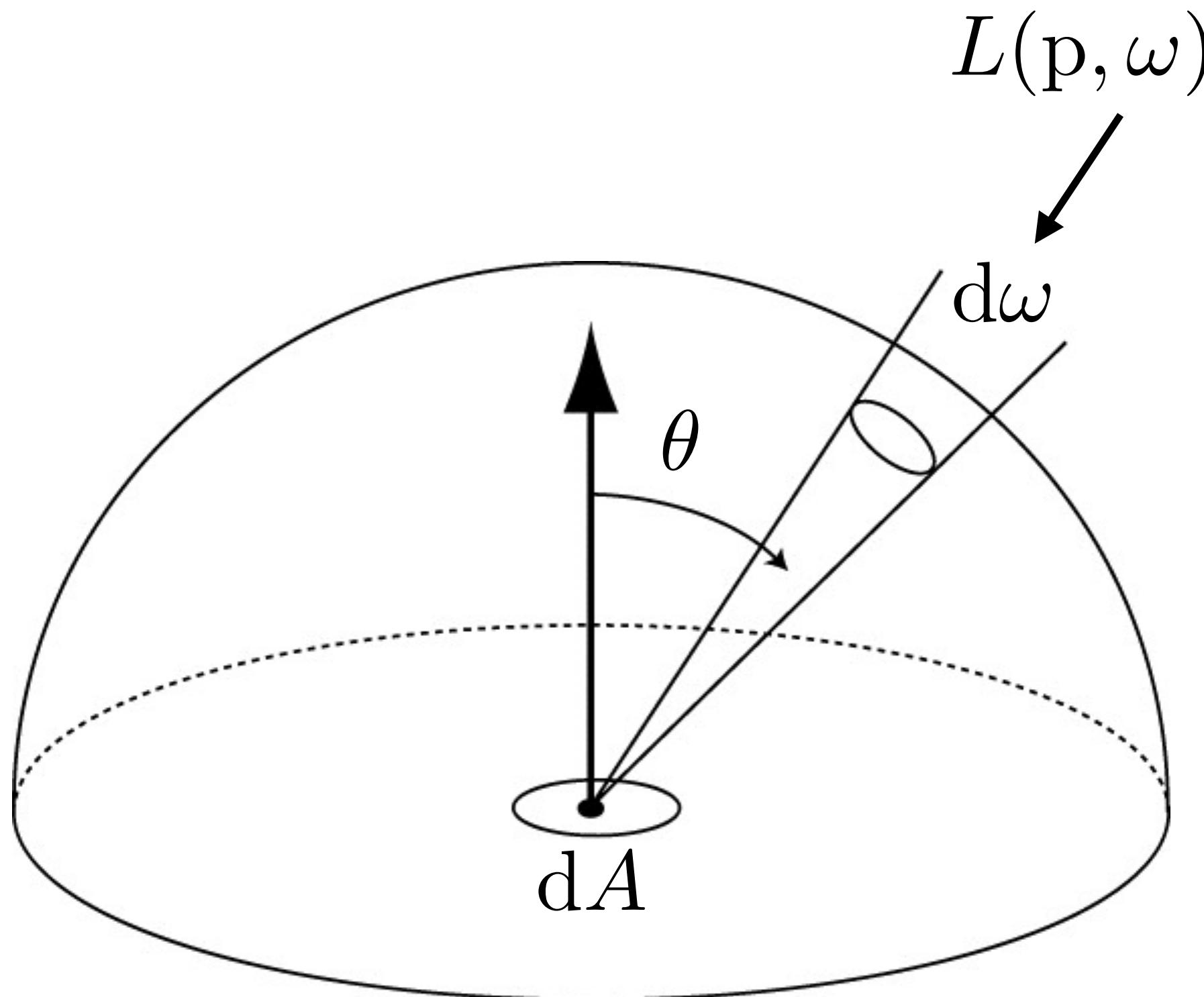
$$p(x) = \frac{1}{b - a}$$

Näive direct lighting estimate

Uniformly-sample hemisphere of directions with respect to solid angle

$$p(\omega) = \frac{1}{2\pi}$$

$$E(p) = \int L(p, \omega) \cos \theta d\omega$$



Estimator:

$$X_i \sim p(\omega)$$

$$Y_i = f(X_i)$$

$$Y_i = L(p, \omega_i) \cos \theta_i$$

$$F_N = \frac{2\pi}{N} \sum_{i=1}^N Y_i$$

Näive direct lighting estimate

Uniformly-sample hemisphere of directions with respect to solid angle

$$E(p) = \int L(p, \omega) \cos \theta d\omega$$

Given surface point p

For each of N samples:

Generate random direction: ω_i

Compute incoming radiance arriving L_i at p from direction: ω_i

Compute incident irradiance due to ray: $dE_i = L_i \cos \theta_i$

Accumulate $\frac{2\pi}{N} dE_i$ into estimator

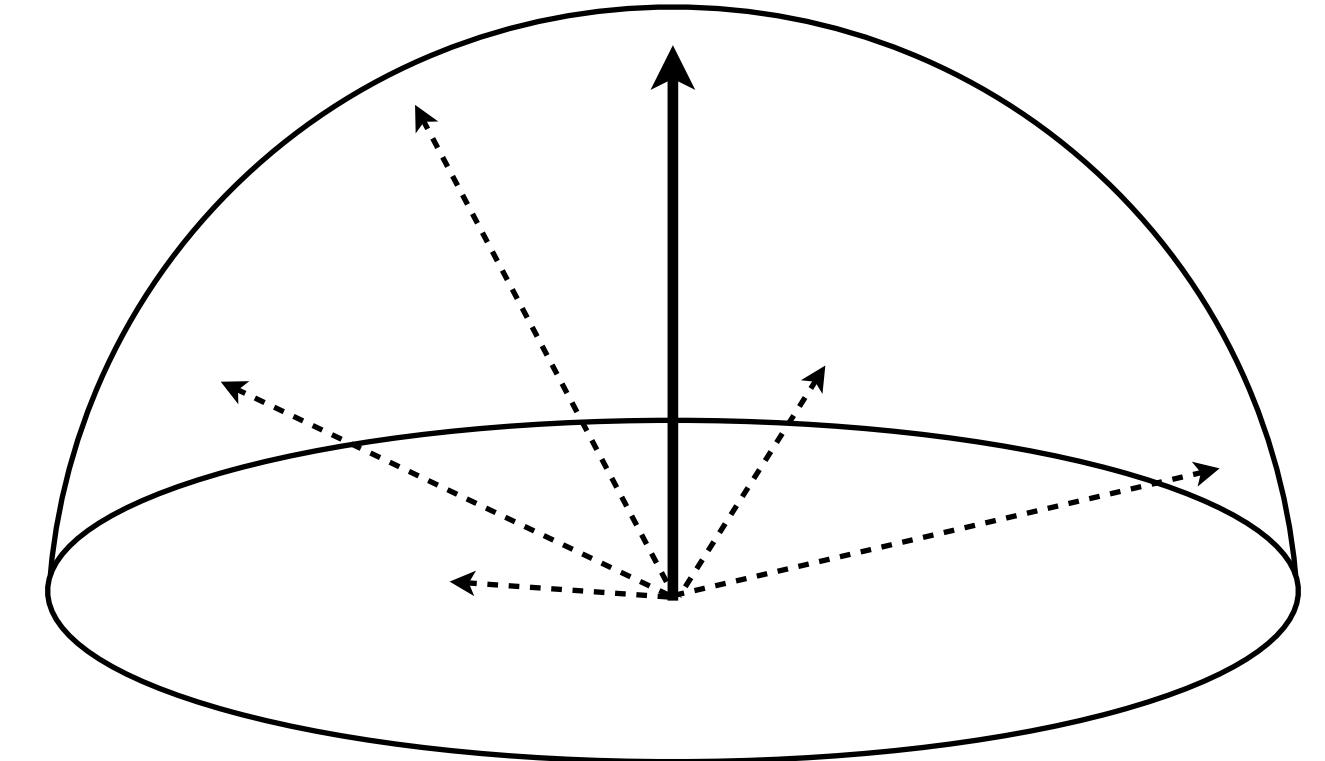
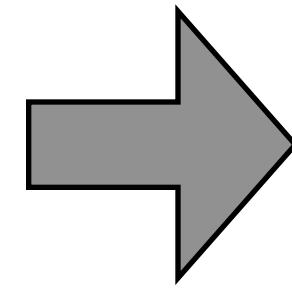
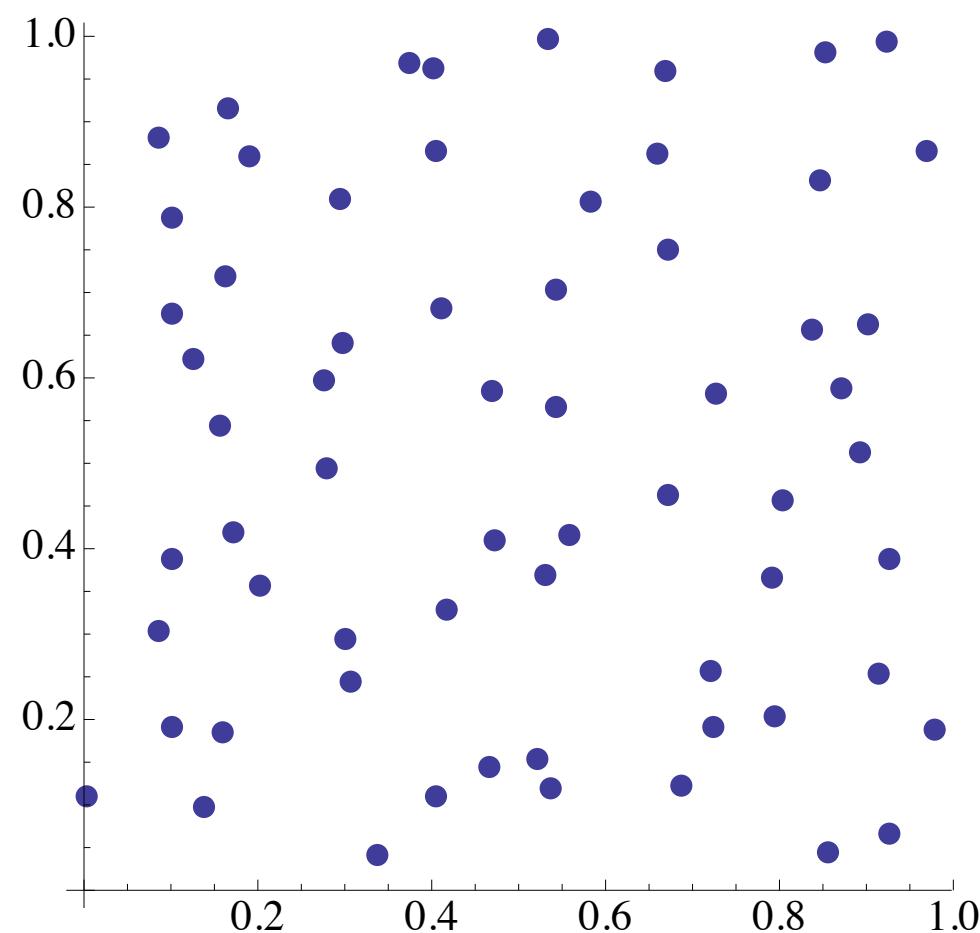
Uniform hemisphere sampling

Generate random direction on hemisphere (all directions equally likely)

$$p(\omega) = \frac{1}{2\pi}$$

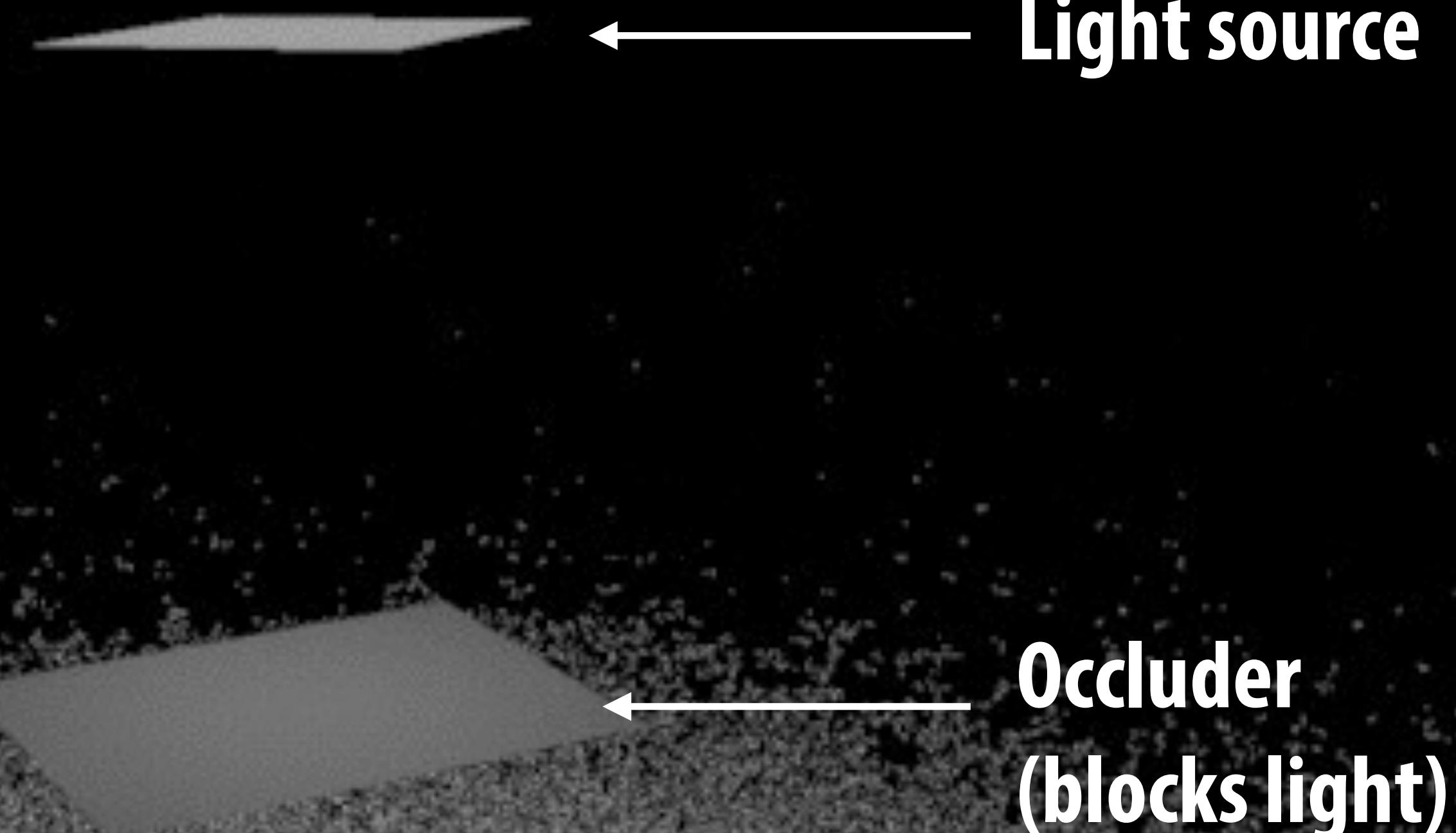
Direction computed from uniformly distributed point on 2D plane:

$$(\xi_1, \xi_2) = (\sqrt{1 - \xi_1^2} \cos(2\pi\xi_2), \sqrt{1 - \xi_1^2} \sin(2\pi\xi_2), \xi_1)$$

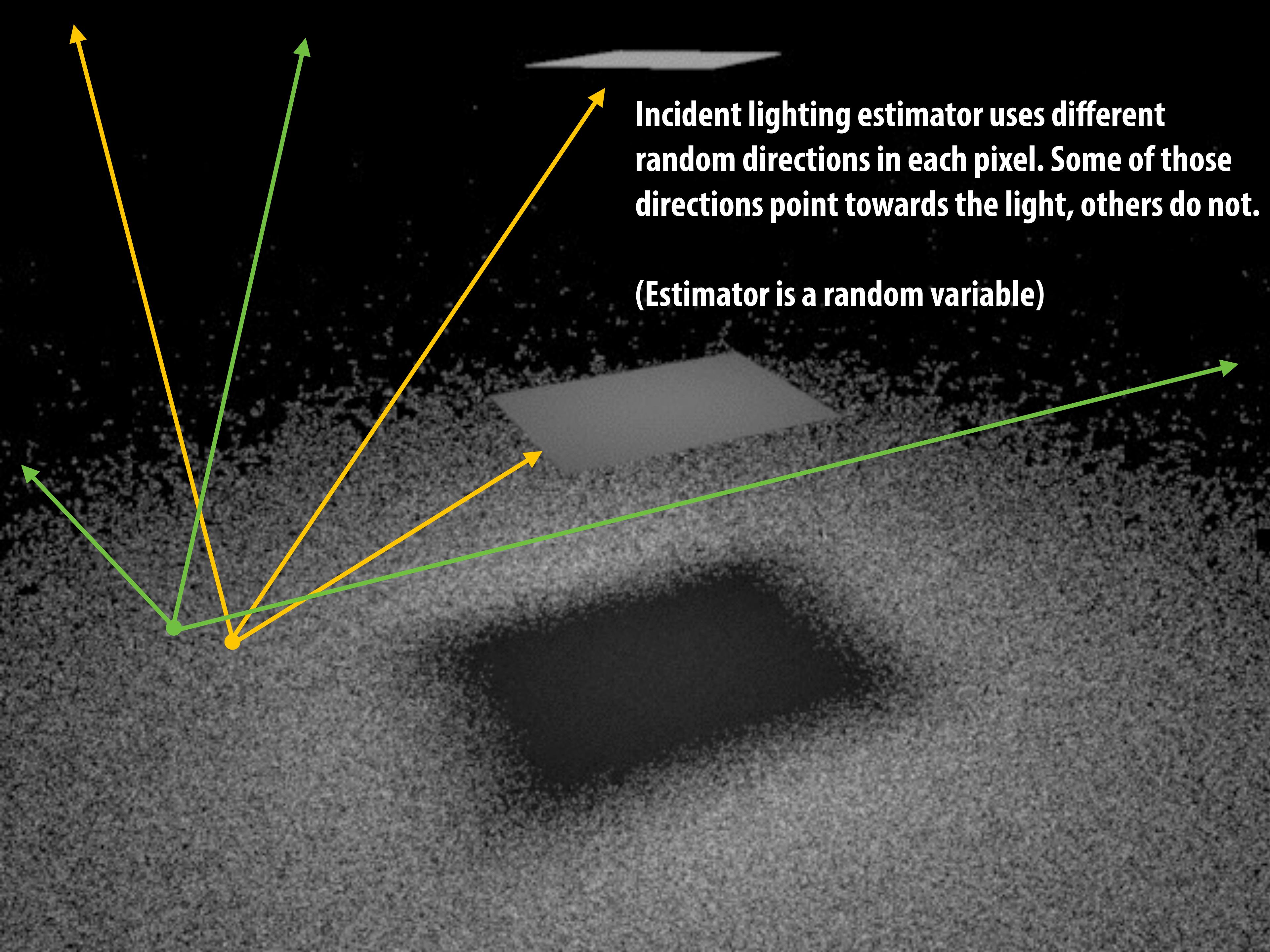


Try at home: derive from the inversion method [Arvo]

**Hemispherical solid angle
sampling, 100 sample rays
(random directions drawn
uniformly from hemisphere)**



Why is the image in the previous slide “noisy”?



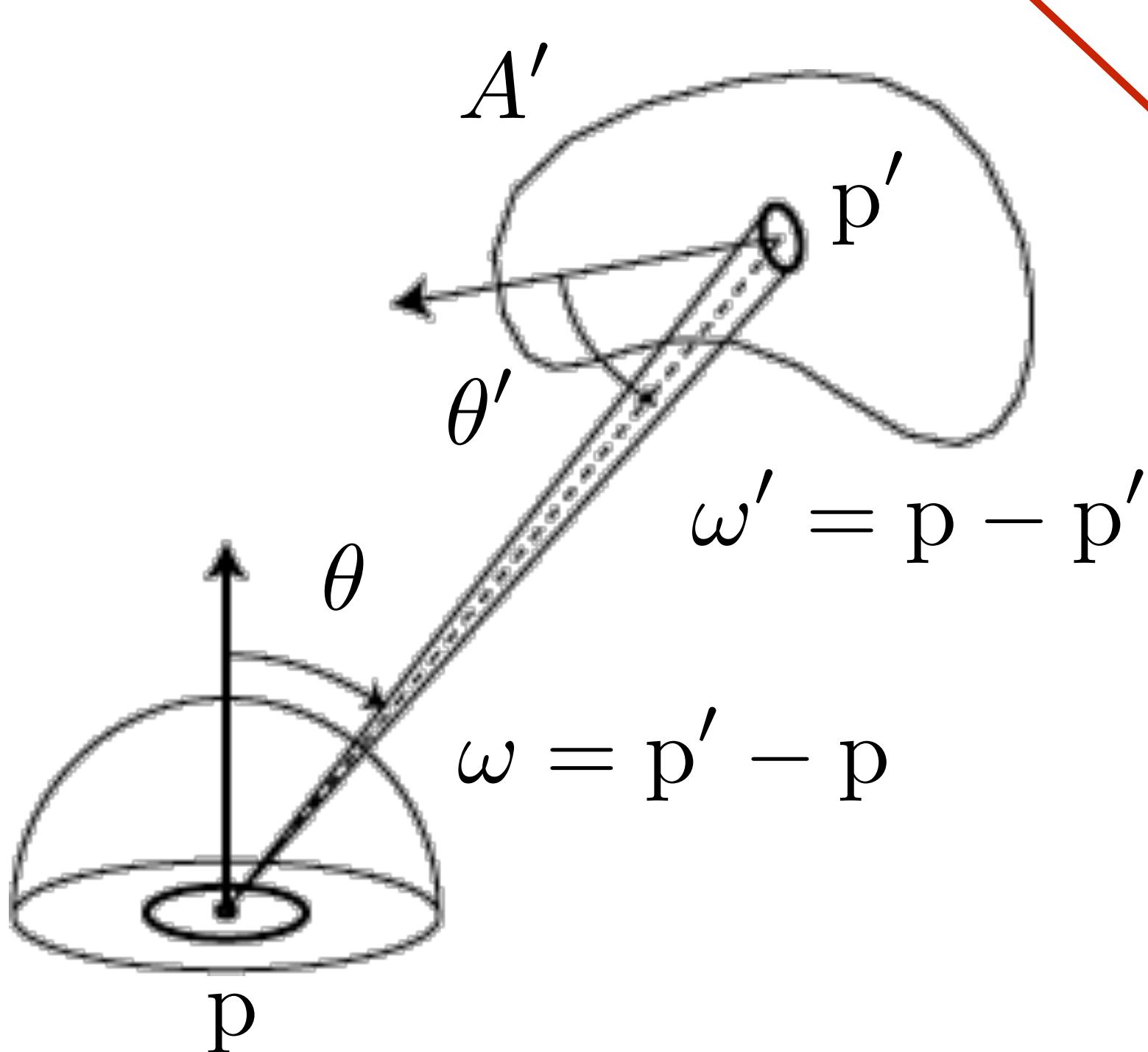
Idea: don't need to integrate over entire hemisphere of directions (incoming radiance is 0 from most directions)

**Only integrate over the area of the light
(directions where incoming radiance is non-zero)**

Direct lighting: area integral

$$E(p) = \int L(p, \omega) \cos \theta d\omega \quad \text{Integral over directions}$$

$$E(p) = \int_{A'} L_o(p', \omega') V(p, p') \frac{\cos \theta \cos \theta'}{|p - p'|^2} dA' \quad \text{Change of variables to integral over area of light}$$

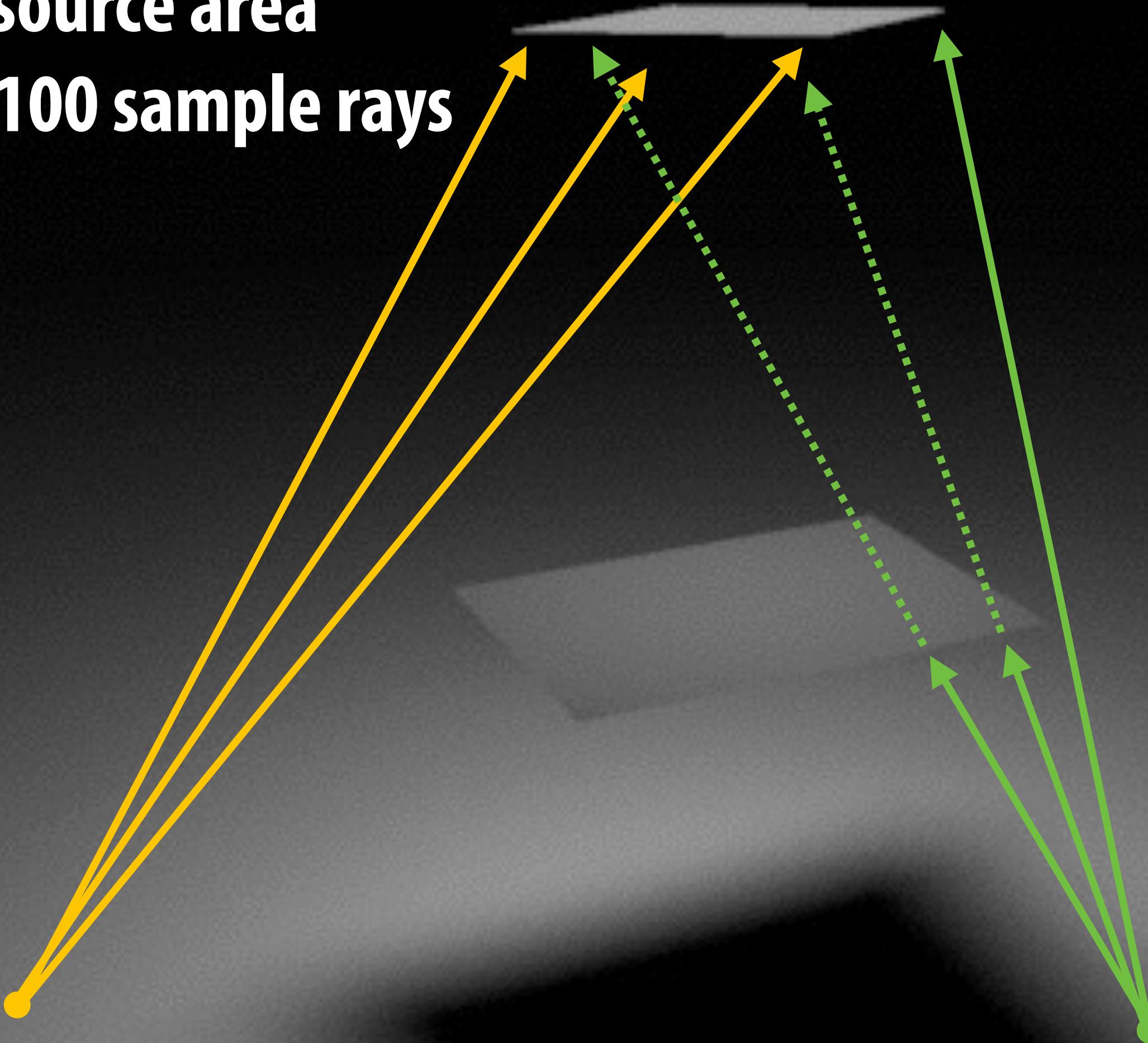


$$d\omega = \frac{dA}{|p' - p|^2} = \frac{dA' \cos \theta}{|p' - p|^2}$$

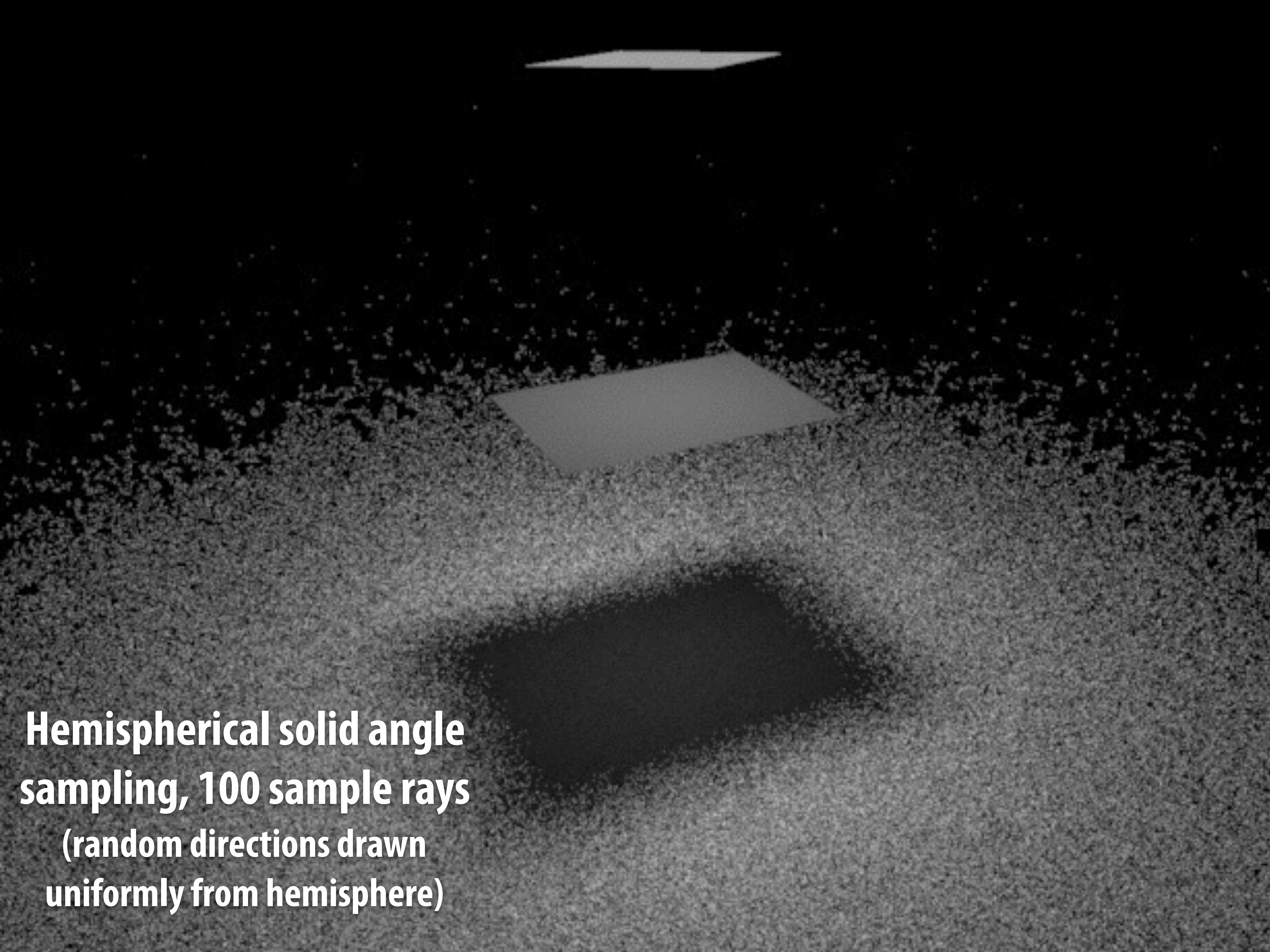
Binary visibility function:
1 if p' is visible from p , 0 otherwise
(accounts for light occlusion)

Outgoing radiance from light point p , in direction w' towards p

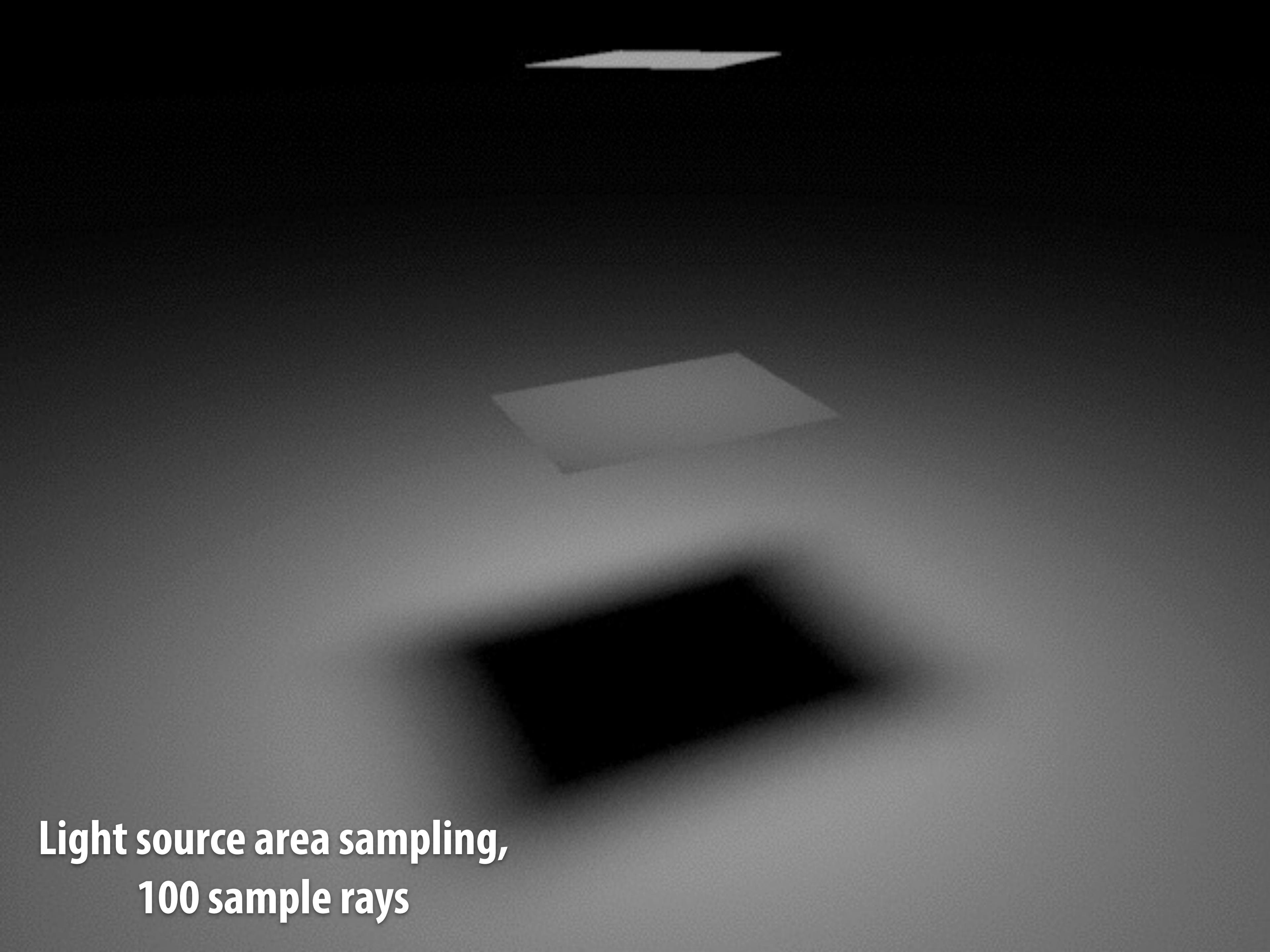
Light source area sampling, 100 sample rays



If no occlusion is present, all directions chosen in computing estimate “hit” the light source.
(Choice of direction only matters if portion of light is occluded from surface point p .)



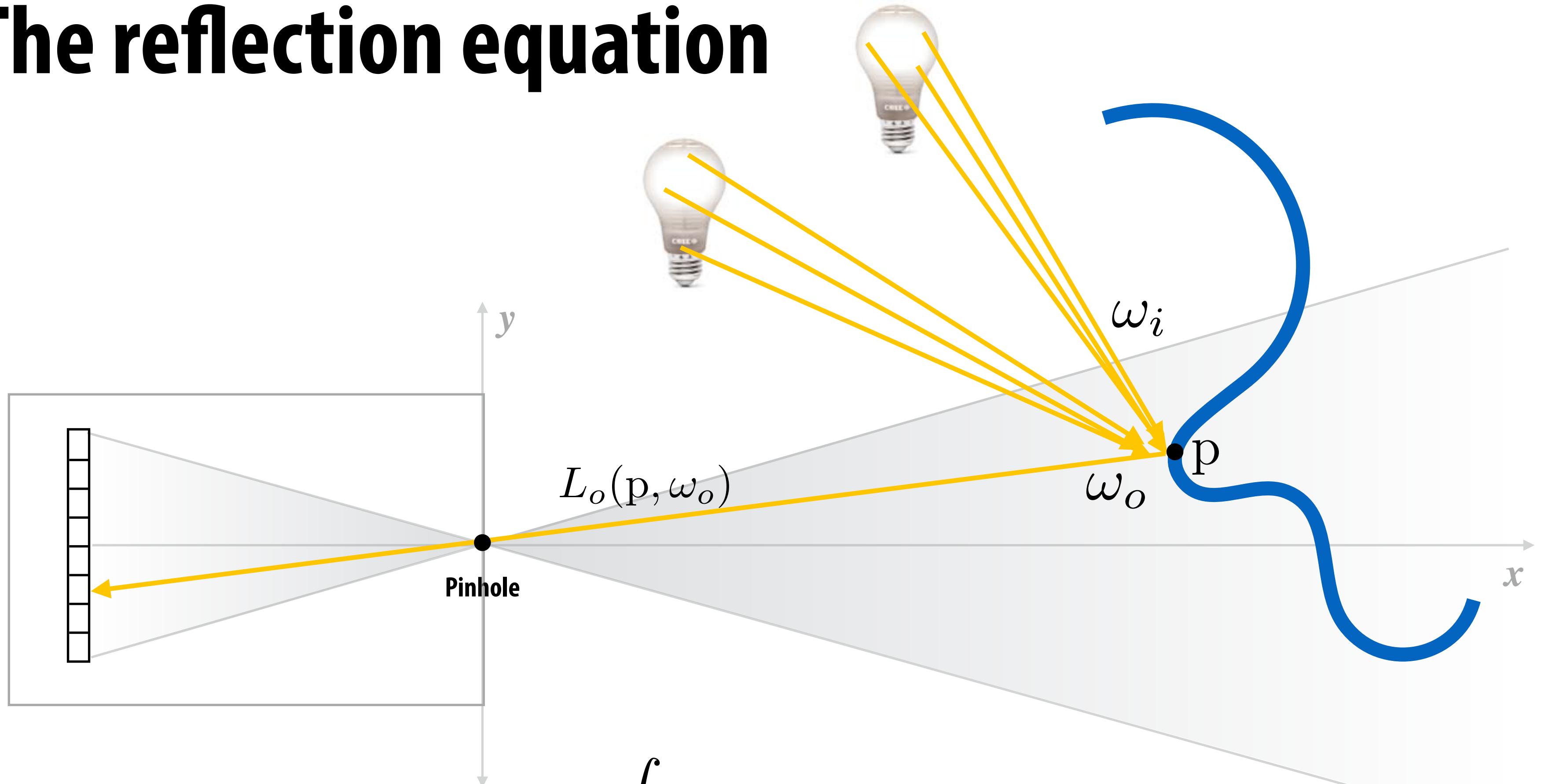
**Hemispherical solid angle
sampling, 100 sample rays
(random directions drawn
uniformly from hemisphere)**



**Light source area sampling,
100 sample rays**

Estimating indirect lighting

The reflection equation

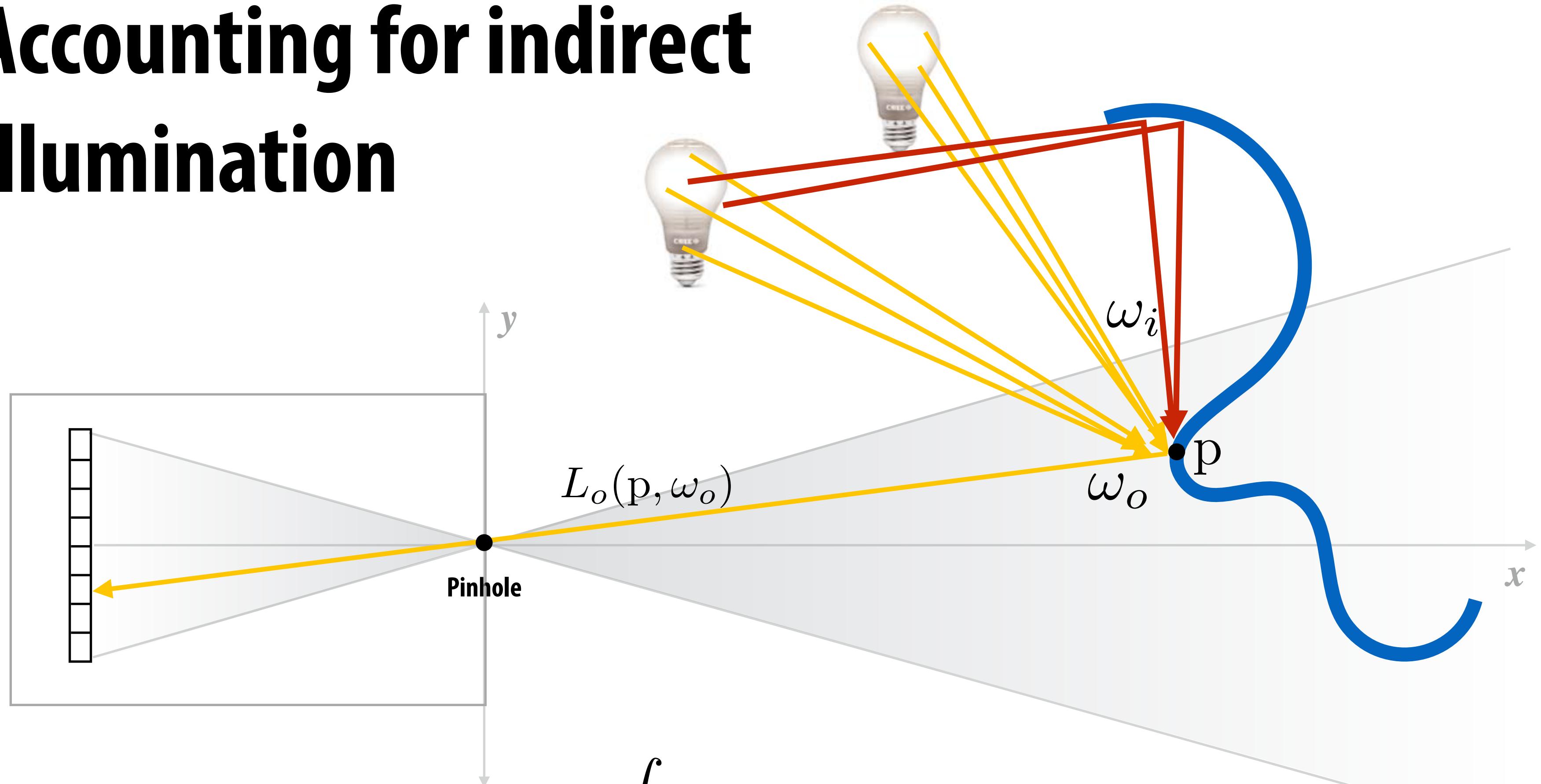


$$L_o(p, \omega_o) = L_e(p, \omega_o) + \int_{H^2} f_r(p, \omega_i \rightarrow \omega_o) L_i(p, \omega_i) \cos \theta_i d\omega_i$$

Need to know incident radiance.

So far, have only computed incoming radiance from scene light sources.

Accounting for indirect illumination



$$L_o(p, \omega_o) = L_e(p, \omega_o) + \int_{H^2} f_r(p, \omega_i \rightarrow \omega_o) L_i(p, \omega_i) \cos \theta_i d\omega_i$$

Incoming light energy from direction ω_i may be due to light reflected off another surface in the scene (not an emitter)

Path tracing: indirect illumination

$$\int_{H^2} f_r(\omega_i \rightarrow \omega_o) L_{o,i}(tr(p, \omega_i), -\omega_i) \cos \theta_i d\omega_i$$

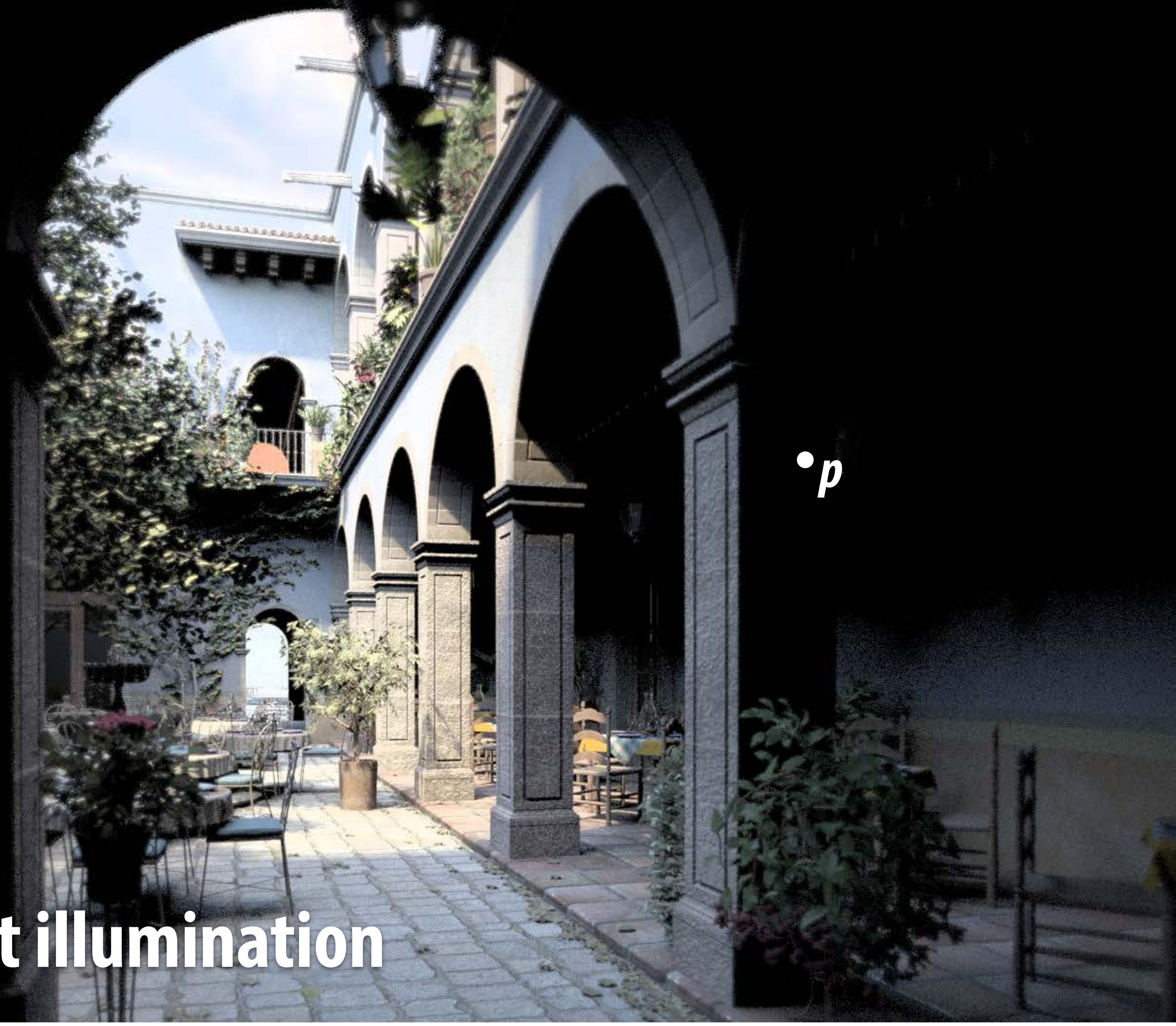
- Sample incoming direction from some distribution (e.g. proportional to BRDF):

$$\omega_i \sim p(\omega)$$

- Recursively call path tracing function to compute incident indirect radiance
- Monte Carlo estimator:

$$\frac{f_r(\omega_i \rightarrow \omega_o) L_{o,i}(tr(p, \omega_i), -\omega_i) \cos \theta_i}{p(\omega_i)}$$

Direct illumination





One-bounce global illumination



Two-bounce global illumination

Four-bounce global illumination



Eight-bounce global illumination

\bullet^p





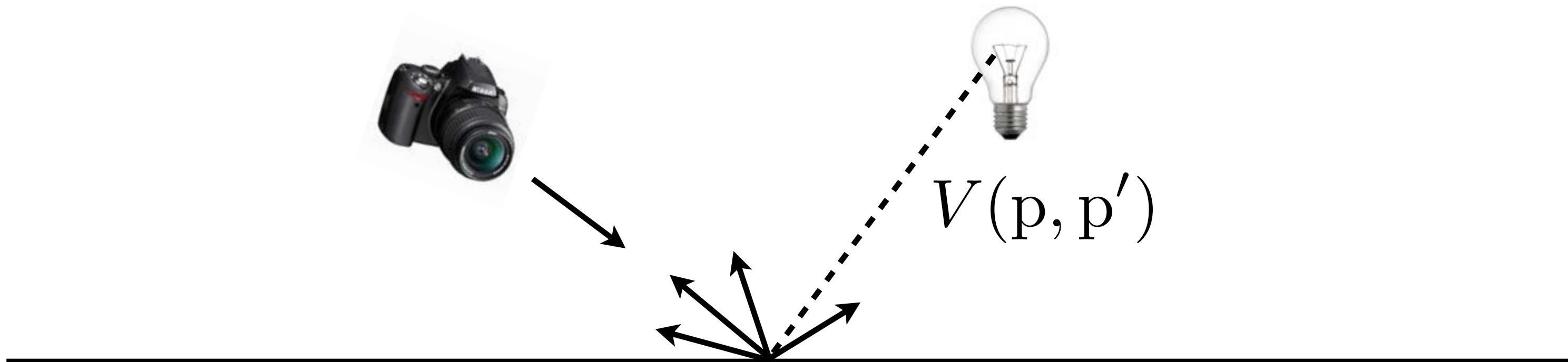
Sixteen-bounce global illumination

**Wait a minute...
When do we *stop*!?**

Russian roulette

- Idea: want to avoid spending time evaluating function for samples that make a **small contribution** to the final result
- Consider a low-contribution sample of the form:

$$L = \frac{f_r(\omega_i \rightarrow \omega_o) L_i(\omega_i) V(p, p') \cos \theta_i}{p(\omega_i)}$$



Russian roulette

$$L = \frac{f_r(\omega_i \rightarrow \omega_o) L_i(\omega_i) V(p, p') \cos \theta_i}{p(\omega_i)}$$



$$L = \left[\frac{f_r(\omega_i \rightarrow \omega_o) L_i(\omega_i) \cos \theta_i}{p(\omega_i)} \right] V(p, p')$$

- If tentative contribution (in brackets) is small, total contribution to the image will be small regardless of $V(p, p')$
- Ignoring low-contribution samples introduces systemic error
 - No longer an unbiased estimator
- Instead, randomly discard low-contribution samples in a way that leaves estimator unbiased

Russian roulette

- New estimator: evaluate original estimator with probability p_{rr} , reweight. Otherwise ignore.
- Same expected value as original estimator:

$$p_{\text{rr}}E\left[\frac{X}{p_{\text{rr}}}\right] + E[(1 - p_{\text{rr}})0] = E[X]$$



No Russian roulette: 6.4 seconds



**Russian roulette: terminate 50% of all contributions with
luminance less than 0.25: 5.1 seconds**



**Russian roulette: terminate 50% of all contributions with
luminance less than 0.5: 4.9 seconds**



**Russian roulette: terminate 90% of all contributions with
luminance less than 0.125: 4.8 seconds**



**Russian roulette: terminate 90% of all contributions with
luminance less than 1: 3.6 seconds**

Next time:

- Variance reduction—how do we get the most out of our samples?

