The Ising model and the Metropolis algorithm: Ferromagnetic phase transitions FYS3150 Project 4

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Abstract The aim of the project is to determine the Curie temperature T_c of a system which undergoes a phase transition from ferromagnetic to paramagnetic when temperature is raised above a critical temperature, which corresponds to the Curie temperature. We want to solve the problem by modeling it with the Ising model and apply Monte Carlo methods through the Metropolis algorithm.

I. INTRODUCTION

The report comprises a theory and method section followed by a result and discussion section, ending with a conclusion. The programes made in this project is main.cpp, isingmodel.cpp and python scripts for analysis.

II. THEORY AND METHOD

A. The physics problem

1. How does magnetism arise in the first place?

Magnetism arises from two sources; electric current and the magnetic moment induced by the spin of elementary particles, magnetic moment being a measure of the strength of a magnetic source and a quantity which determines the torque a magnet will experience in an external magnetic field. This project investigates the latter source. In the context of the magnetization of materials the electrons' magnetic moments are most important, as the magnetic moments of the nuclei is much weaker. The electrons in a material are normally arranged such that their magnetic moments, orbital and intrinsic, cancel out - they combine into pairs with opposite intrinsic magnetic moments in accordance with the Pauli exclusion principle or combine into filled subshels with no net orbital motion. However, there are solids which may have unpaired electrons or non-filled subshells. Whether the material produce a magnetic field is then determined by the direction of the magnetic moments contributed by these electrons.

2. Ferromagnetic and paramagnetic

Ferromagnetic and paramagnetic materials are substances with such unpaired electrons. In paramagnetic materials the magnetic moments of these tend to point in random directions, cancelling each other out so that the material is not magnetic (in the abcense of an external field). The magnetic moments of unpaired electrons

in ferromagnetic materials, on the other hand, have a tendency to allign parallel to one another, as this leads to lowered energy levels. This effect makes a material magnetic, even in the absence of a magnetic field.

3. Phase transition

If a ferromagnetic material is raised above a certain temperature, it loses its ferromagnetic properties. It becomes paramagnetic as the magnetic moments of impaired electrons align randomly, as a result of the thermal tendency to disorder (increasing entropy) becoming stronger than the tendency towards lower energy (this drag and pull effect makes up the thermodynamical potential Helmholtz free energy, described below).

4. The Curie temperature

The temperature at which a ferromagnetic substance experiences this shift is called its Curie temperature, T_c . It is the critical temperature at which the system undergoes a phase transition from ferromagnetic to paramagnetic. We now have the physics alibi for modeling this phase transition - if done successfully it allows us to determine T_c .

B. The canonical ensemble and the Ising model

In order to derive the thermodynamical properties of interest, we operate with the canonical ensemble. An ensemble is a collection of possible states a system might be in, that is it is the probability distribution for the state of the system. The canonical ensemble describes the possible states available to a system at a fixed temperature T, that is in thermal equilibrium with a heat bath. The system can exchange energy with the heat bath, meaning the available states may differ in total energy. This gives a probability for each microstate i dependent on energy, that is $P_i = \frac{1}{Z}e^{-E_i\beta}$. Here $\beta = \frac{1}{kT}$, k being the Boltzman constant. $e^{-E_i\beta}$ is the Boltzman factor and Z is the partition function, given below.

We now introduce the Ising model which allows us to model the behaviour of the unpaired electrons' magnetic moments in the framework of the canonical ensemble, taking into account the interaction between neighbours. We model the direction of the magnetic moments of the electron spins as discrete variables which can be either +1 or -1. The spins are arranged in a graph and allowed to interact with its neighbours. In this project we study the two dimensional Ising model, where the graph corresponds to a square lattic with L number of spins in each direction.

Given this framework the relevant properties can be computed. We first need the partition function, the sum of the Boltzman factor of every microstate available to the system, given as

$$Z = \sum_{i=1}^{M} e^{-E_i \beta} = \sum_{E} \Omega(E) e^{-E\beta}$$
 (1)

Where $M = 2^N = 2^{L \times L}$ is the number of microstates in our two dimensional case. The first expression is a sum over all microstates. In the second expression we see that microstates with the same energy will have the same Boltzman factor, making it a sum over all energy levels while multiplying the Boltzman factor with the corresponding degeneracy.

The energy of each microstate is given by

$$E_i = -J \sum_{\langle kl \rangle}^N s_k s_l \tag{2}$$

The energy expectation value for the canonical ensamble is given as

$$\langle E \rangle = -\frac{\delta lnZ}{\delta \beta} \langle E \rangle = \sum_{i=1}^{M} E_i P_i = \frac{1}{Z} \sum_{i=1}^{M} E_i e^{-E_i \beta}$$
 (3)

using the probability $P_i = \frac{1}{Z}e^{-E_i\beta}$ for a microstate i. The mean magnetization of the system is

$$\langle \mathcal{M} \rangle = \sum_{i=1}^{M} \mathcal{M}_i P_i = \frac{1}{Z} \sum_{i=1}^{M} \mathcal{M}_i e^{-E_i \beta}$$
 (4)

where the magnetization of one microstate M_i is given as

$$\mathcal{M}_i = \sum_{j=1}^N s_j \tag{5}$$

that is the sum over all the N spins in one configuration.

As stated earlier, the mechanisms driving the spins to parallel or random alignement is the combination of a strive towards lower energy and higher entropy, steered by a system's need to minimize the thermodynamical potential, which for the canonical ensemble is the Helmholtz free energy *F*:

$$F = \langle E \rangle - ST \tag{6}$$

S being the entropy.

Phase transitions may be defined as discontinuity in some order derivative of quantities related to the thermodynamical potential of a given system. A first order phase transition is marked by a discontinuity in the first order derivative, while a second order phase transition is marked by a discontinuity in the second order derivative. While the expectation energy is a first order quantity, other thermodynamical properties such as the heat capacity C_v are of the second order. This quantity is defined

$$C_v = -\frac{\delta \langle E \rangle}{\delta \beta} = \frac{\delta^2 lnZ}{\delta \beta^2} \tag{7}$$

$$=\frac{1}{k_B T^2} \sigma_E^2 \tag{8}$$

Where $\sigma_E^2=\langle E^2\rangle-\langle E\rangle^2$ is the variance, relating to the energy fluctuations in the system.

Finally, magnetic susceptibility is given as

$$\chi = \frac{1}{k_B T} \sigma_{\mathcal{M}}^2 = \frac{1}{k_B T} (\langle \mathcal{M}^2 \rangle - \langle \mathcal{M} \rangle^2)$$
 (9)

1. Finite size

The Ising model has limitations of which the consequences are not neglible. In order for the theory of the canonical ensemble to be true, we need to operate in the thermodynamical limit, that is the number of spins in each direction L goes to infinity. Thus for every computation we conduct with a fixed value of L, we must be aware of the effects of a finite lattice size.

The finite lattice size require us to handle the end points in the lattice. In this project we do so by periodic boundary conditions.

C. Analytical calculation L = 2 lattice

We first look into the properties of the Ising model analytically, for a small L = 2 system, labeling the spins as s_1 , s_2 , s_3 and s_4 in the following manner

$$s_1$$
 s_2 s_3 s_4

Thus the energy of a microstate for this lattice size, using equation 2, becomes

$$E_i = -J(s_1s_2 + s_2s_1 + s_1s_3 + s_3s_1$$
 (10)

$$+s_2s_4+s_4s_2+s_3s_4+s_4s_3$$
 (11)

Calculating the energies for the $M = 2^N = 2^{L \times L} = 16$ microstates we find that there are three possible energies, that is -8J, 0 and 8J. The energy -8J is shared by the two microstates with all spins either up or down, that is $\Omega(-8J) = 2$. The energy 0 is shared by the microstates of three up and one down (four states), one up and three down (four states), and two up and two down states except the ones where equal spins are on the diagonals (four states). Thus $\Omega(0) = 12$. The energy 8J corresponds to the two up and two down states where equal spins are on the diagonal (two states), giv $ing \Omega(8I) = 2.$

Equation 1 now gives the 2×2 partition function

$$\begin{split} Z &= \Omega(-8J)e^{-(-8J)\beta} + \Omega(0)e^{-0.\beta} + \Omega(8J)e^{-8J\beta} \\ &= 2e^{8J\beta} + 12 + 2e^{-8J\beta} \\ &= \frac{1}{4}(\cosh(8J\beta) + 3) \end{split}$$

From equation 3 we find the mean energy

$$\langle E \rangle = -8J \frac{\sinh(8J\beta)}{3 + \cosh(8J\beta)}$$

Equations refeq:Mexp and 5 gives

$$\langle \mathscr{M} \rangle = 0$$

Calculating the expectation value of the absolute magnetization one finds

$$\langle |\mathcal{M}| \rangle = 2 \frac{e^{8J\beta} + 2}{\cosh(8I\beta) + 3}$$

The reason it is of interest to take the absolute value of the magnetization here is the effects of the finite lattice size. The finite lattice size allows for jumps in the magnetization due to a change of sign which we would not see in the thermodynamical limit.

From equation 7 we calculate the specific heat

$$C_v = \frac{64J^2}{kT^2} \left(\frac{\cosh(8J\beta)}{3 + \cosh(8J\beta)} - \frac{\sinh^2(8J\beta)}{(3 + \cosh(8J\beta))^2} \right)$$
 (12)

While equation 9 gives the magnetic suceptibility

$$\chi = \frac{8}{k_B T} \frac{e^{8J\beta} + 1}{3 + \cosh(8J\beta)}$$

The system's need to minimize the thermodynamical potential is what defines its equilibrium, or steady state.

- 1. Establish an initial state with energy E_h by positioning yourself at a random configu-
- Change the initial configuration by flipping e.g., one spin only. Compute the energy of this trial state E_t . Calculate $\Delta E = E_t E_b$. The number of values ΔE is limited to five for the Ising model
- in two dimensions, see the discussion below. 4. If $\Delta E \leq 0$ we accept the new configuration, meaning that the energy is lowered and
- we are hopefully moving towards the energy minimum at a given temperature. Go to step 7. 5. If $\Delta E > 0$, calculate $w = e^{-(\beta \Delta E)}$.
- 6. Compare w with a random number r. If

then accept the new configuration, else we keep the old configuration

- 7. The next step is to update various expectations values.
 8. The steps (2)-(7) are then repeated in order to obtain a sufficently good representation

FIG. 1: The steps of the Metropolis algorithm [1]. It lets us simmulate our system's development towards a steady state. Computing this state for a range of temperatures we can calculate the expectation values of thermodynamical properties at each temperature, allowing us to observe at what temperature a phase transition might occur.

Thus we start to see signs of how temperature affects the ferromangetic property of the material. If the temperature is low, the Blotzman factor indicates that a low energy level will be highly favourable, as this probability dominates. It leads the spins to a parallel allignement. If the temperature rises on the other hand, the proabilities become more equal, making random allignments more probable. It is in the steady state of the system that we are interested in observing the behaviour of the thermodynamic properties we have defined.

D. Markov chains

In order to describe a system which moves towards a steady state or equilibrium given an initial configuration we introduce Markov chains. A Markov process is a random walk with a selected probability for making a move. In our system the random walk is the step from one spin configuration to the next. Markov chains are in fact a discretized description of diffusion and gives a framework for the rules of Brownian motion - the behaviour exhibited by small fractions of any system when exposed to random fluctuations of the medium.

In order to simulate our systems evolution towards a steady state we use Markov chains repeatidly in Monte Carlo simulations. From Markov chains we get two conditions needed for reaching the steady state - detailed balance and ergodicity. This informs our choice of algorithm - the Metropolis algorithm.

E. Metropolis algorithm

In order to simulate our system's evolution towards a steady state using Markov chains we will employ the Metropolis algorithm, which meets the conditions

The steps of the Metropolis algorithm [1] are as shown

in figure 1. We decide the number of repetitions of step two to seven "in order to achieve a sufficiently good representation" by the size of the lattice. By repeating the steps $N=L\times L$ times we allow for a maximum of N possible flips. Thus our model scales according to the lattice size. In our model we view one Monte Carlo cycle as one time step, and so it is natural that more spin flips occur in one time step for a bigger system.

Observe that we flip one spin at the time, if accepted we compute the corresponding change of expectation values. After one complete sweep of the lattice, we measure the new expectation values due to these changes. As stated, there are a limited set of energy changes available, this follows from the outline of the Ising model in two dimensions in the previous section. The possible changes are $\Delta E = 8J$, 4J, 0, -4J, -8J.

The algorithm determines whether a proposed move is implemented based on a transition probability and an acceptance probability. The strength of the algorithm is that the transition probability need not be known. In our case we are relieved from computing the partition function *Z* for each proposed spin flip. In fact the acceptance of a flip need only depend on the change in energy - if the energy becomes smaller we accept with a probability of one, else we compute the Boltzman factor for the energy difference and compare it to a random number.

F. Code

The code I have written implements the Metropolis algorithm as stated above and writes the thermodynamical quantities of interest to an output file specified by the user. This is read by a python script which plots the quantities of interest.

I made two versions of the code. The first is primarily for testing the quality of model and code. The test this code allows for include: does the L=2 case reproduce the analytical results found earlier? At what point is it safe to say that the system has reached equlibrium so that we can measure expectation values (we call this thermalization)? How is it in our interest to look at the absolute value of the magnetization? Should it extend to our definition of the magnetic suceptibility? Does the number of spin flips it accepts seem reasonable? With which probabilities does it actually produce different energy levels at different temperatures? Results of such tests can be seen in the following section.

The second version of the code implements these findings to the end of giving the best possible representation of our system. Some decisions made in this code is to start with an ordered configuration at low temperatures and random at higher temperatures assuming these are the more likely states leading quicker to thermalization. Thermalization is accepted at the point where $\langle E \rangle$ and C_v has changed less than 5% since the previous measurement. the time needed to reach a steady state is longer for temperatures close to the critical temperature

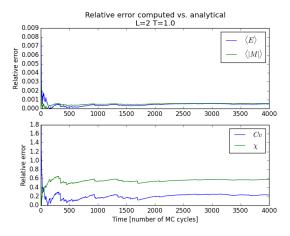


FIG. 2: The relative error of the computed thermodynamical quantities vs the analytical values. The error is small, indicating that the code is functional.

than for temperatures away.

Furthermore, the second version parallelizes the computation of temperatures using MPI. Thus, this is the code used to finally model how the thermodynamical properties develop as we increase temperature. I ran the code for a number of 1 million Monte Carlo cycles using two processors on a laptop with one CPU. Please see the codes for further comments and details.

G. Phase transition and how to find T_c

In order to compute the actual T_C corresponding to the thermodynamical limit from our T_C restricted by the lattice size, we make use of the followin relation

$$T_c(L = \infty) = T_c(L) - aL^{-\frac{1}{\nu}}$$
 (13)

where the exact result gives v = 1 and a is a constant.

III. RESULTS AND DISCUSSION

A. Computations compared with analytical

The relative error is plotted in figure 2. There is a good agreement and stabilization of the error at a 10³ number of Monte Carlo cycles.

B. Thermalization and rate of acceptance

See figures 3, 4, 5 and 6 for the expectation values plotted against time. We first notice that the expectation values take on the expected values, for example the energy

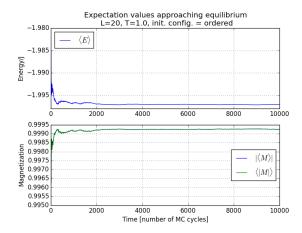


FIG. 3: The graph shows how the magnetization and the energy of the system evolves with time, that is the number of Monte Carlo cycles in the simulation

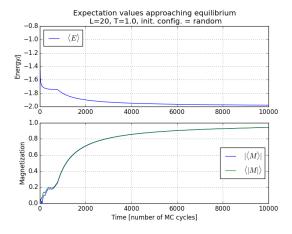


FIG. 4: The graph shows how the magnetization and the energy of the system evolves with time, that is the number of Monte Carlo cycles in the simulation.

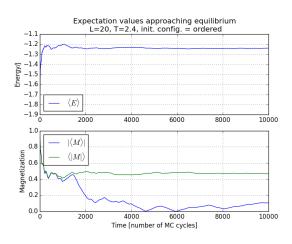


FIG. 5: The graph shows how the magnetization and the energy of the system evolves with time, that is the number of Monte Carlo cycles in the simulation.

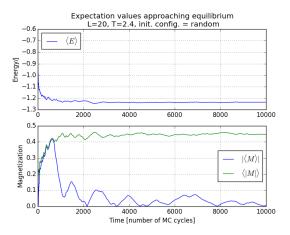


FIG. 6: The graph shows how the magnetization and the energy of the system evolves with time, that is the number of Monte Carlo cycles in the simulation.

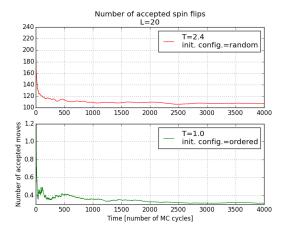


FIG. 7: This graph shows the number of spins in our model flipped over time at two different temperatures.

is lowest energy level, -2 per spin, at the low temperature. Furthermore, we see how the different two different computations of the absolute magnetization differs. We see that

langle|M|

rangle has the smoothest graph, this is our motivation for calculating the suceptibility later on with this value. It is "cheating" in order to avoid some unwanted effects of our finite lattice.

Finally we see why we want to put in the test of thermalization - the system clearly needs to reach the steady state before we start measurements, or else our expectation values will be skewed.

See figure 7 for the plot of accepted moves against time. While few moves are accepted at a low temperature, many more, (about 30%) are accepted at T=2.4. This is as we would expect above and below the critical temperature.

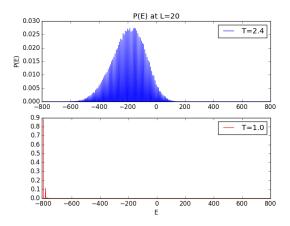


FIG. 8: Exp

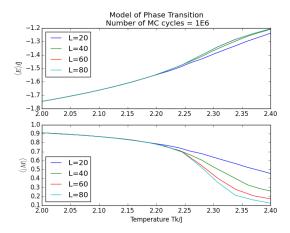


FIG. 9: Behaviour of energy and magnetism during the phase transition

C. P(E)

See figure 8. For the energy probabilites. Again it corresponds to what we know about the ferromagnetic material. The variance of E was computed as 9 for low T and 3283 for high T. As the standard deviation is the squareroot of this, meaning energy values are mostly the same at the low temperature and spread out over a wider range of levels at the higher temperature, these are good results.

D. Phase transition

See figures 9 and 10. We clearly see the signs of the transition - the slope of the energy is steeper, the mag-

netism is about to disappear (but never really does due to our finite lattice size), and the specific heat and suceptibility peak. This all happen around 2.3, which is close to the expected value 2.69. As we saw in the equation relating the finite lattice result to the actual T_c , the two

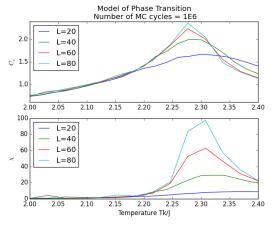


FIG. 10: Behaviour of specific heat and magnetic suceptibility during phase transition

become closer as L increases, and we see this effect in these figures.

IV. CONCLUSION

In this project we used the Ising model and the Metropolis algorithm to model the phase transition from ferromagnetic to paramagnetic material. In this we tried to obtain the Curie temperature. This was successful as the model exhibited a phase transition around 2.3, the exact result being 2.69. For future work one could look into models that differ from the Ising and Metropolis methods - for example methods of flipping several spins at a time, seeing a quicker development of the heat capacity near the critical temperature.

V. APPENDIX

[1] M. Hjort-Jensen Computational Physics: Lecture Notes UiO 2015.